

Frequency domain calculations of eccentric orbit EMRIs on Schwarzschild

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Motivation and Scope

- Calculate gravitational self-force for EMRIs
 - SCO (mass = μ) orbiting MBH (M)
 - BH perturbation theory, expansion in $\epsilon = \mu/M$
 - Successive field and motion corrections
- Waveforms for use in LISA detection and parameter estimation

Cumulative phase $\Phi = \kappa_1 \frac{1}{\epsilon} + \kappa_2 \epsilon^0 + \kappa_3 \epsilon + \dots$,
 $\Phi \sim 10^6$, $\delta\Phi \lesssim 0.1$ \rightarrow need for second order

But transient resonances $\rightarrow \epsilon^{-1/2}$ terms for generic orbits on Kerr
Flanagan and Hinderer (2010), arXiv:1009.4923v2 [gr-qc]

- Present scope: Eccentric orbits on Schwarzschild
 - Accurate 1st-order RWZ perturbations (done)
 - Metric in RW gauge (done)
Hopper and Evans (2010), PRD 82:084010; gr-qc:1006.4907

Gauge generators Ξ_μ and metric in Lorenz gauge (in progress)
1st-order self-force correction with high accuracy (coming)

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Accuracy Requirements

Example: Assume $\epsilon = \mu/M = 10^{-6}$, $\Delta\Phi/\Phi \simeq 10^{-7}$

| | | | |
|-----------------|------------------|------------------------|-------------------------|
| (metric errors) | $\mathcal{O}(1)$ | $\mathcal{O}(10^{-6})$ | $\mathcal{O}(10^{-12})$ |
| | | | |

Metric : $g_{\mu\nu} + {}_1 p_{\mu\nu} + {}_2 p_{\mu\nu} + \dots$

| | | | |
|--|--|--|--|
| | | | |
|--|--|--|--|

Self – force : ${}_1 f_\mu + {}_2 f_\mu + \dots$

| | | | |
|--|--|--|--|
| | | | |
|--|--|--|--|

(self – force errors) $\mathcal{O}(1)$ $\mathcal{O}(10^{-6})$ $\mathcal{O}(10^{-12})$

Analytic and Numerical Approach

High-level summary

- (1) Use RWZ formalism; find $\Psi_{\ell m}^{\text{odd}} = \Psi_{\ell m}^{\text{CPM}}(t, r)$, $\Psi_{\ell m}^{\text{even}} = \Psi_{\ell m}^{\text{ZM}}(t, r)$
- (2) Obtain metric in RW gauge
- (3) Obtain gauge generators to go from RW to Lorenz gauge
- (4) Obtain metric in Lorenz gauge
- (5) (Next) use to obtain the self-force

... but how can we do with requisite accuracy

Analytic and Numerical Approach

- Frequency domain (FD) solution for master functions (**Step 1**)
 - Standard Green function method for $R_{\ell mn}(r)$
 - Find normalized homogeneous solutions $R_{\ell mn}^{\pm}$ (normalized)
- Time domain (TD) solution for master functions and metric (**Step 2**)
 - Sum $R_{\ell mn}^{\pm}$, return to TD with $\Psi_{\ell m}^{\pm}(t, r)$
 - Extend homogeneous solutions (EHS) to $r = r_p(t)$ to find $\Psi_{\ell m}(t, r)$
 - Avoid Gibbs phenomenon
- Barack, Ori, Sago (2008), PRD 78:084021
- Reconstruct TD metric $p_{\mu\nu}^{\text{RW}}$ in RW gauge
- FD solution for gauge generator amplitudes $\xi_{(i)}^{\ell mn}$ (**Step 3**)
 - Solve separated equations
- Sago, Nakano, and Sasaki (2003), PRD 67:104017
- TD solution for Ξ_μ and metric $p_{\mu\nu}^{\text{L}}$ in Lorenz gauge (**Step 4**)
 - Using EHS and (new) EPS methods

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Master Functions and Regge-Wheeler Metric

- Schwarzschild geometry:

$$ds^2 = -f dt^2 + \frac{1}{f} dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2, \quad f \equiv 1 - \frac{2M}{r}.$$

- $\Psi_{\ell m}^{\text{even}}$, $\Psi_{\ell m}^{\text{odd}}$ functions of $p_{\mu\nu}^{\ell m}$ and derivatives in RW gauge
- Both satisfy wave equations

$$\left[-\frac{\partial^2}{\partial t^2} + \frac{\partial^2}{\partial r_*^2} - V_\ell(r) \right] \Psi_{\ell m}(t, r) = S_{\ell m}(t, r)$$

- Singular sources

$$S_{\ell m}(t, r) = \tilde{G}_{\ell m}(t) \delta[r - r_p(t)] + \tilde{F}_{\ell m}(t) \delta'[r - r_p(t)]$$

- $p_{\mu\nu}^{\ell m}$ recovered from $\Psi_{\ell m}^{\text{even/odd}}$, their derivatives, and delta sources

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Master Functions

- $\Psi_{\ell m}^{\text{even/odd}}$ are linear combinations of $h_{\mu\nu}^{\ell m}$ and its derivatives in Regge-Wheeler gauge
- Even-parity master function is the Zerilli-Moncrief function

$$\Psi_{\text{even}}(t, r) \equiv \frac{2r}{\ell(\ell+1)} \left[K + \frac{1}{\Lambda} (f^2 h_{rr} - rf \partial_r K) \right],$$

- Odd-parity master function is the Cunningham-Price-Moncrief function

$$\Psi_{\text{odd}}(t, r) \equiv \frac{r}{\lambda} \left[\partial_r h_t - \partial_t h_r - \frac{2}{r} h_t \right].$$

Frequency Domain Formalism

Fourier coefficients for field and source in eccentric orbit:

$$R_{\ell mn}(r) = \frac{1}{T_r} \int_0^{T_r} dt \Psi_{\ell m}(t, r) e^{i\omega_{mn} t}$$

$$Z_{\ell mn}(r) = \frac{1}{T_r} \int_0^{T_r} dt S_{\ell m}(t, r) e^{i\omega_{mn} t}.$$

An ODE for each ℓ, m, n :

$$\frac{d^2 R_{\ell mn}(r)}{dr_*^2} + [\omega_{mn}^2 - V_\ell(r)] R_{\ell mn}(r) = Z_{\ell mn}(r), \quad \omega_{mn} \equiv m\Omega_\varphi + n\Omega_r.$$

Ω_r : Radial libration frequency, Ω_φ : Average angular rate of advance

$$\Psi_{\ell m}(t, r) = \sum_{-\infty}^{+\infty} R_{\ell mn}(r) e^{-i\omega_{mn} t}, \quad S_{\ell m}(t, r) = \sum_{-\infty}^{+\infty} Z_{\ell mn}(r) e^{-i\omega_{mn} t}$$

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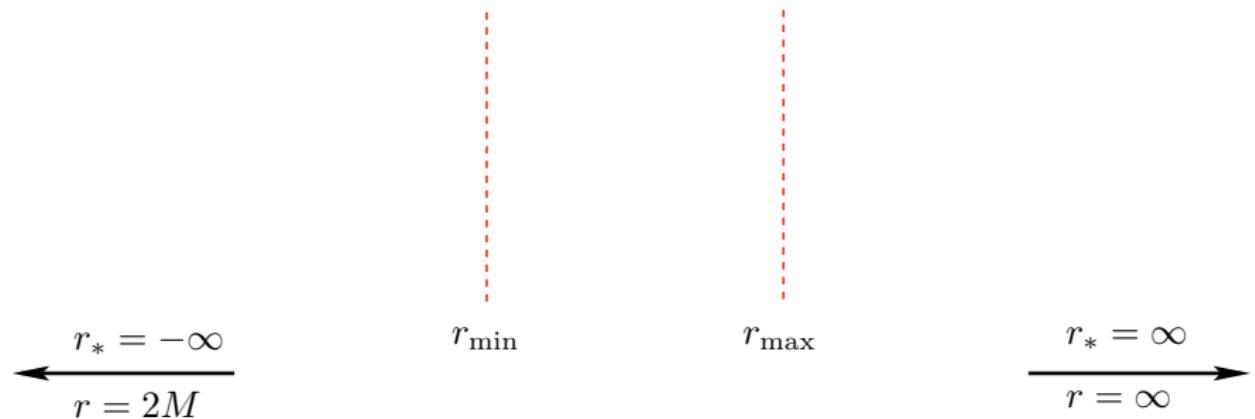
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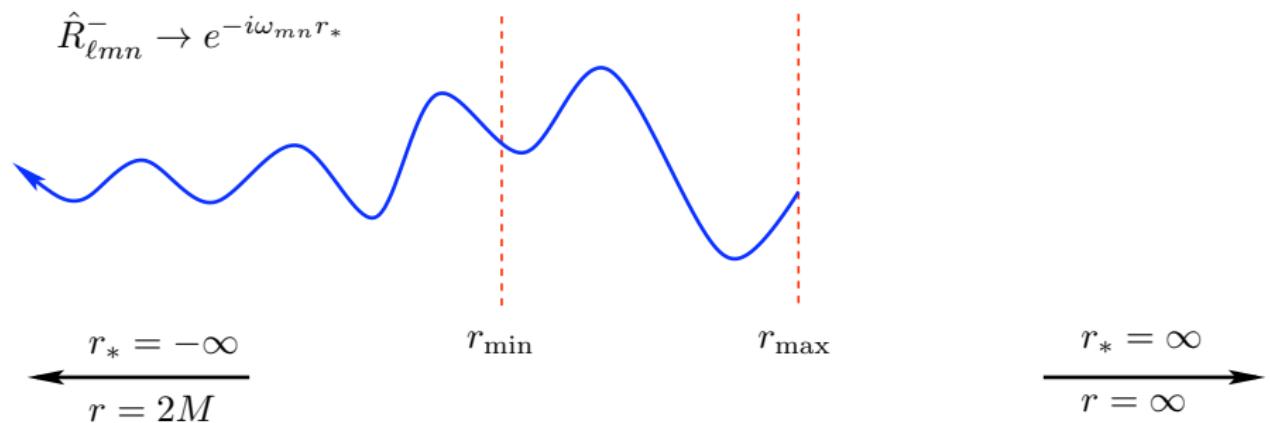
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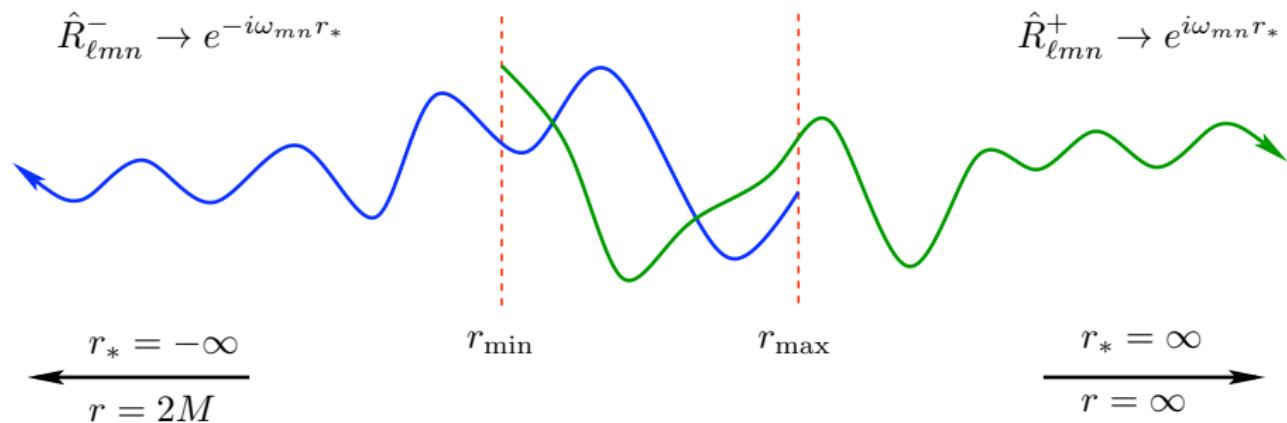
Standard Frequency Domain Method



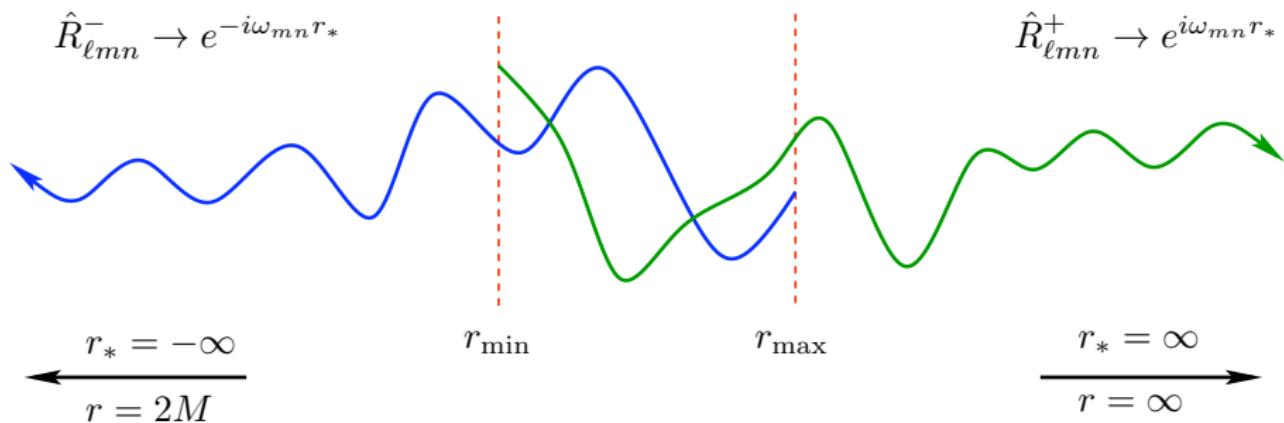
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$$R_{\ell mn}^{\text{std}}(r) = \hat{R}_{\ell mn}^+(r) \int_{r_{\min}}^r \frac{\hat{R}_{\ell mn}^-(r') Z_{\ell mn}(r')}{W_{\ell mn}} dr' + \hat{R}_{\ell mn}^-(r) \int_r^{r_{\max}} \frac{\hat{R}_{\ell mn}^+(r') Z_{\ell mn}(r')}{W_{\ell mn}} dr'$$

Standard Frequency Domain Method

- Normalized homogeneous solutions outside r_{\min} and r_{\max} :

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allow computation of emitted radiation

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- A partial sum up of $R_{\ell mn}^{\text{std}}$ cannot represent a discontinuity

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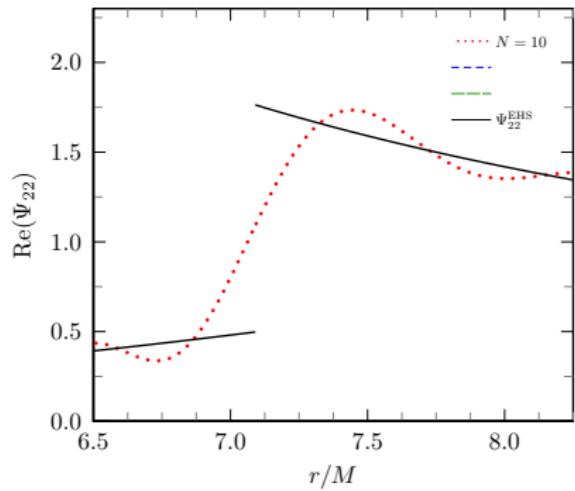
allow computation of emitted radiation

- A partial sum up of $R_{\ell mn}^{\text{std}}$ cannot represent a discontinuity
- **Gibbs phenomenon appears in TD.**

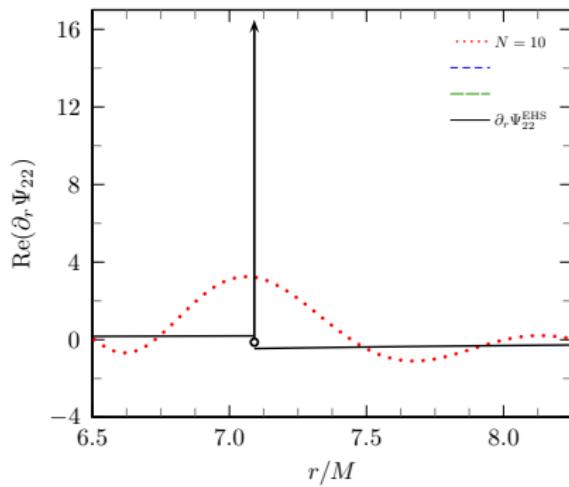
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Standard Frequency Domain Method

$$\Psi_{22}(t_0, r)$$



$$\partial_r \Psi_{22}(t_0, r)$$



$$p = 7.50478$$

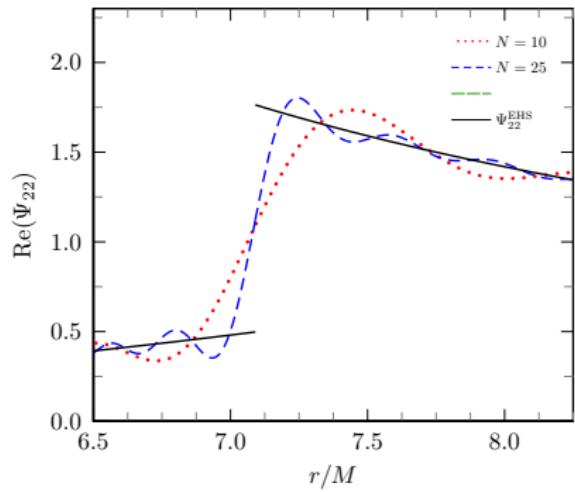
$$e = 0.188917$$

$$t_0 = 51.78M$$

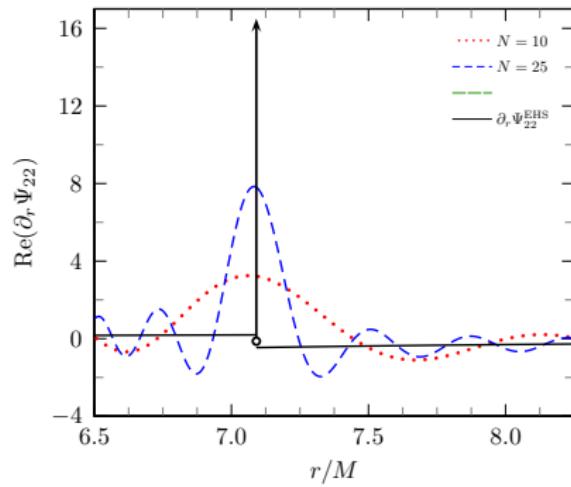
$$\Psi_{\ell m}(t_0, r) = \sum_{n=-10}^{10} R_{\ell m n}^{\text{std}}(r) e^{-i\omega_{mn} t_0}$$

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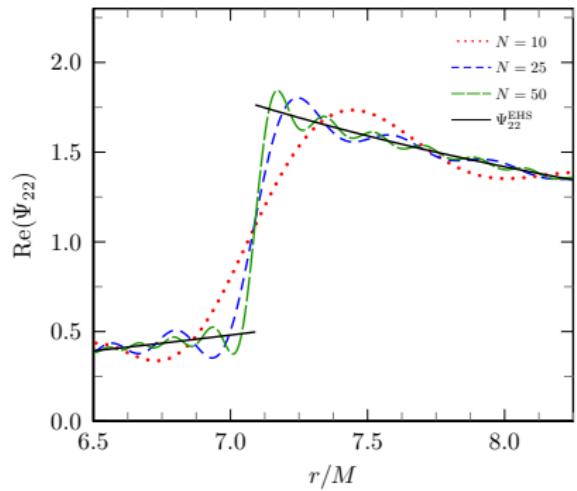
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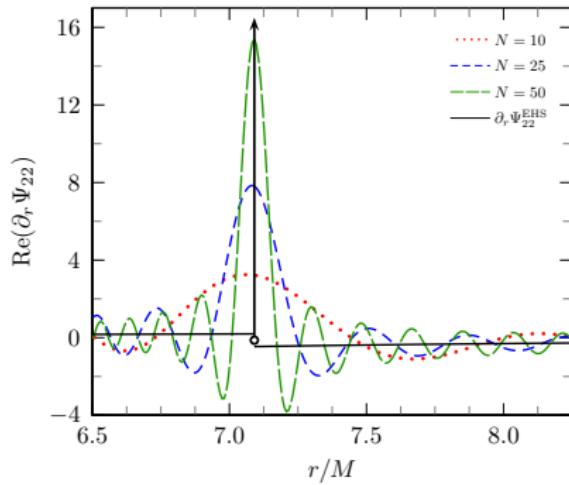
$$\Psi_{\ell m}(t_0, r) = \sum_{n=-25}^{25} R_{\ell m n}^{\text{std}}(r) e^{-i\omega_{mn} t_0}$$

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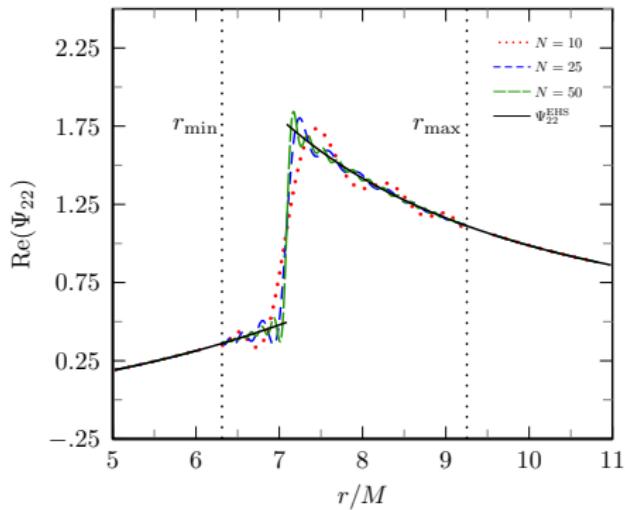
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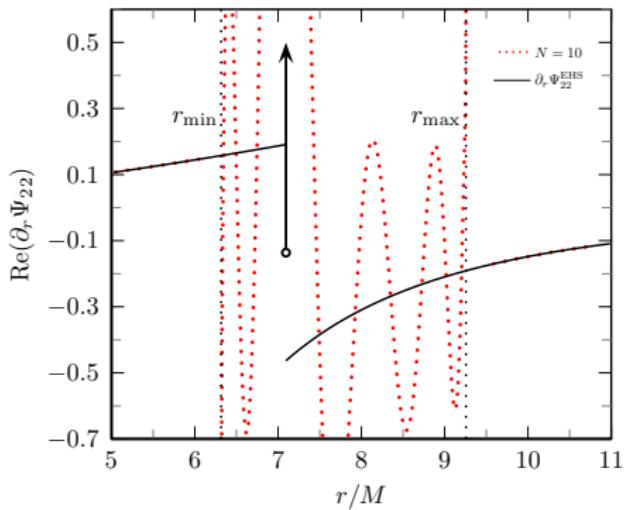
$$\Psi_{\ell m}(t_0, r) = \sum_{n=-50}^{50} R_{\ell m n}^{\text{std}}(r) e^{-i\omega_{mn} t_0}$$

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Method of Extended Homogeneous Solutions

- Barack, Ori, and Sago (2008) demonstrated method for scalar charge on Schwarzschild and for monopole moment
- We applied EHS to “all” radiative gravitational modes (RWZ)
- Define *Time-Domain Extended Homogeneous Solutions*:

$$\Psi_{\ell m}^{\pm}(t, r) = \sum_n R_{\ell mn}^{\pm} e^{-i\omega_{mn} t} = \sum_n C_{\ell mn}^{\pm} \hat{R}_{\ell mn}^{\pm} e^{-i\omega_{mn} t}$$

- Valid in entire range: $2M < r < \infty$
- Argument: Solution to TD RWZ equations is given by

$$\Psi_{\ell m}(t, r) = \Psi_{\ell m}^{+}(t, r) \theta[r - r_p(t)] + \Psi_{\ell m}^{-}(t, r) \theta[r_p(t) - r]$$

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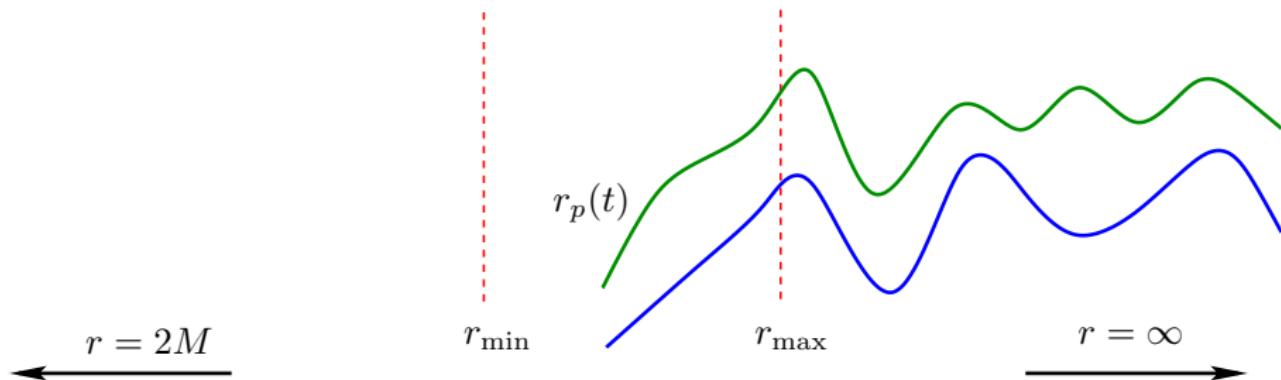
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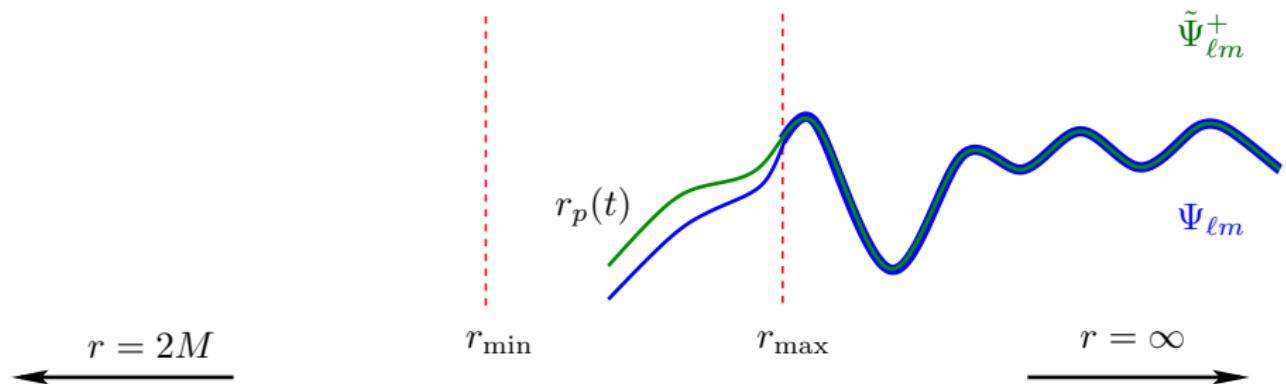
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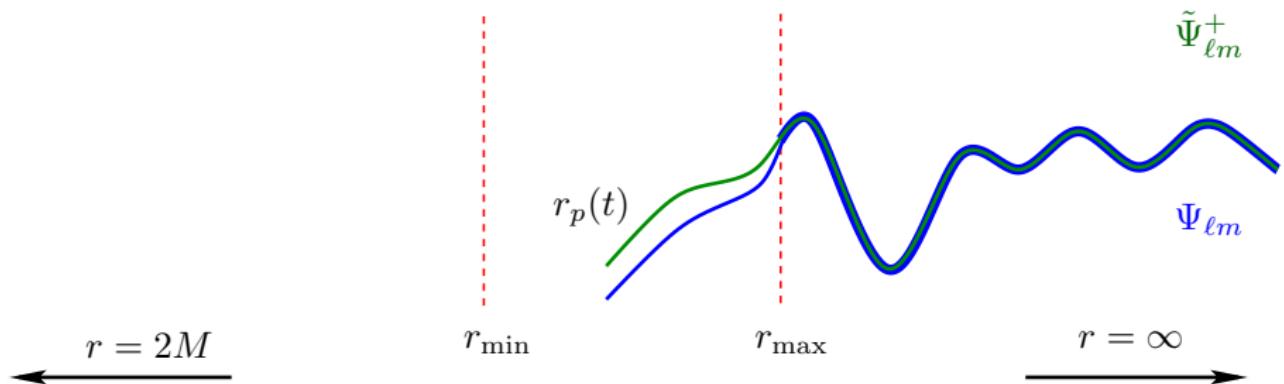
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- They coincide throughout $r > r_{\max}(t)$ since there $R^{\text{std}} = R^+$
- While $R^{\text{std}} \neq R^+$ for $r_p(t) < r < r_{\max}$, Ψ^{std} and Ψ^+ satisfy the same source-free wave equation
- Therefore, they must coincide for all of $r > r_p(t)$.
- The same is true for $\Psi_{\ell m}$ and $\tilde{\Psi}_{\ell m}^-$ throughout $r < r_p(t)$.

Method of Extended Homogeneous Solutions



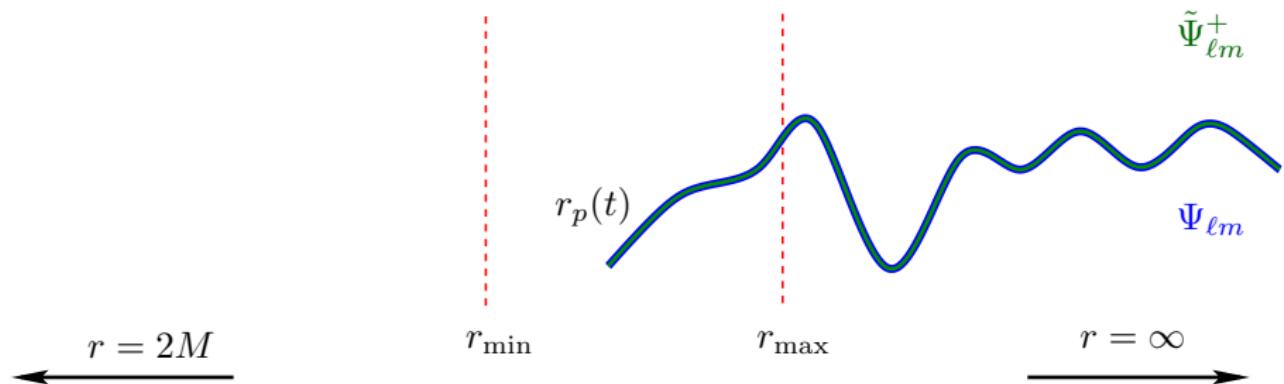
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- While $R^{\text{std}} \neq R^+$ for $r_p(t) < r < r_{\max}$, $\Psi_{\ell m}^{\text{std}}$ and $\Psi_{\ell m}^+$ satisfy the same source-free wave equation
- Therefore, they must coincide for all of $r > r_p(t)$.
- The same is true for $\Psi_{\ell m}$ and $\tilde{\Psi}_{\ell m}^-$ throughout $r < r_p(t)$.

Method of Extended Homogeneous Solutions



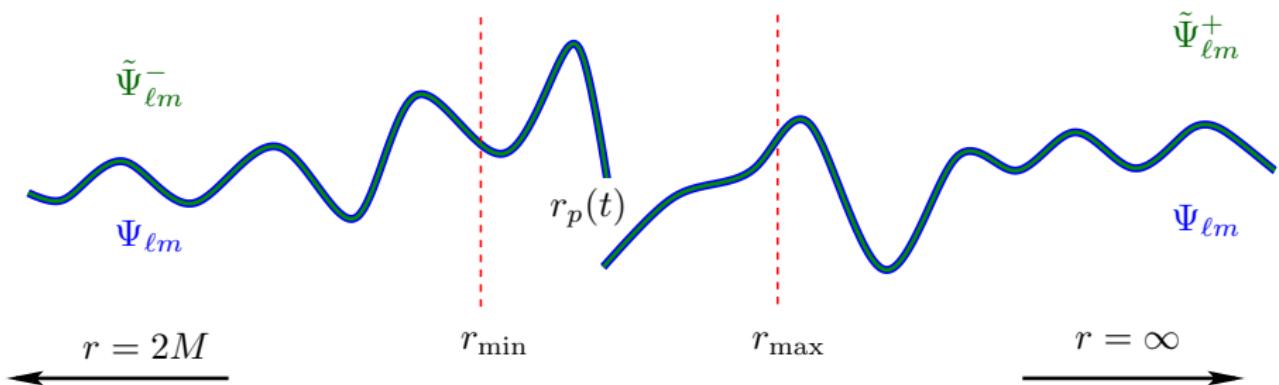
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Method of Extended Homogeneous Solutions



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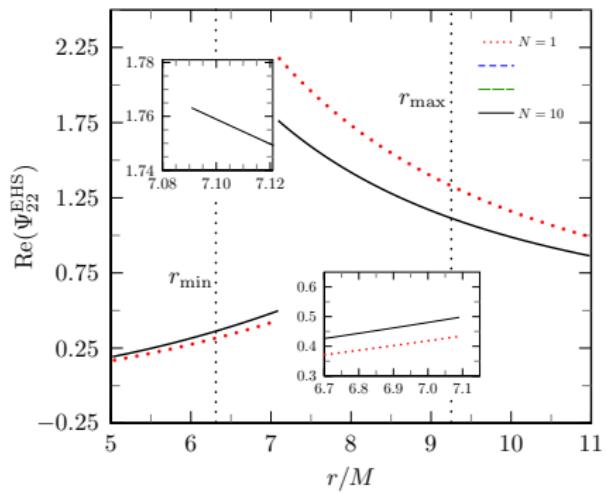
Method of Extended Homogeneous Solutions



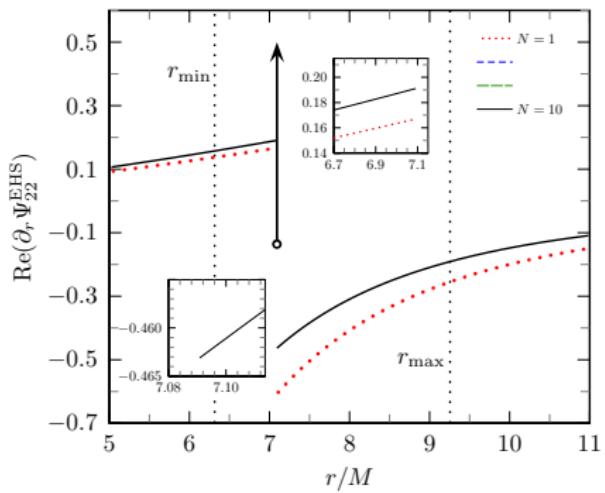
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Master Functions via Extended Homogeneous Solns

$$\Psi_{22}(51.78M, r)$$



$$\partial_r \Psi_{22}(51.78M, r)$$



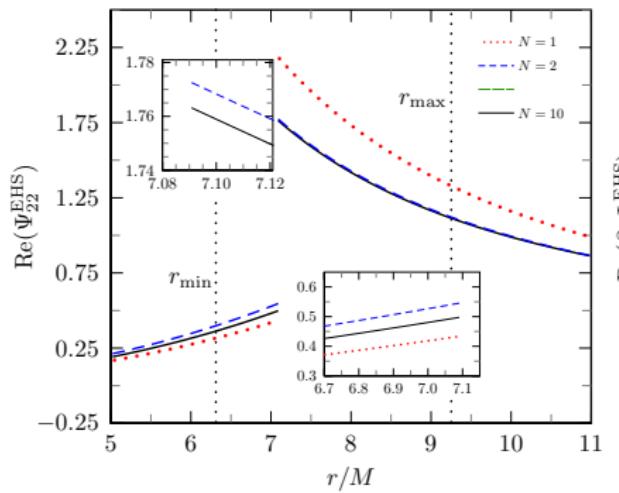
$$p = 7.50478$$

$$e = 0.188917$$

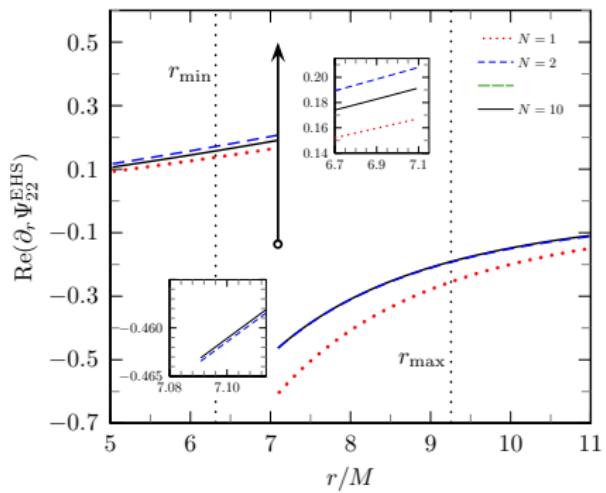
$$\Psi_{\ell m}^{\pm}(t_p, r) = \sum_{n=-1}^1 R_{\ell mn}^{\pm} e^{-i\omega_{mn} t}$$

Master Functions via Extended Homogeneous Solns

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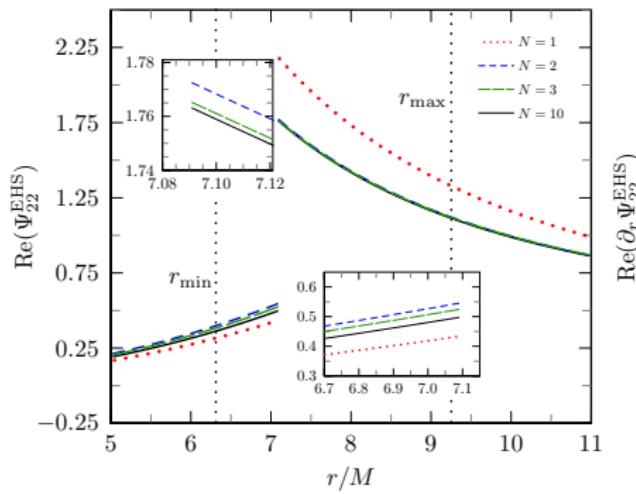
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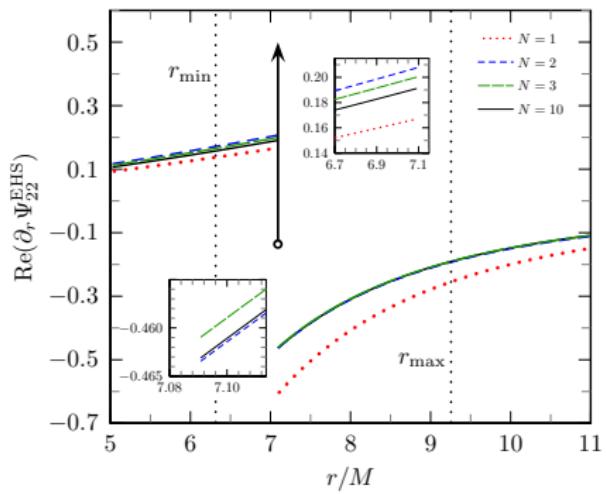
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Master Functions via Extended Homogeneous Solns

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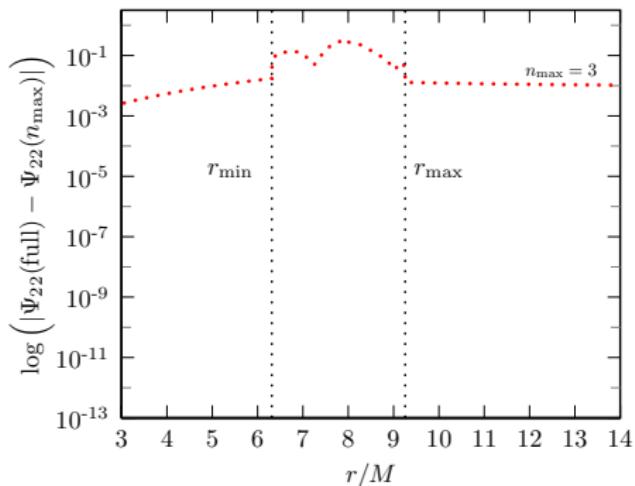
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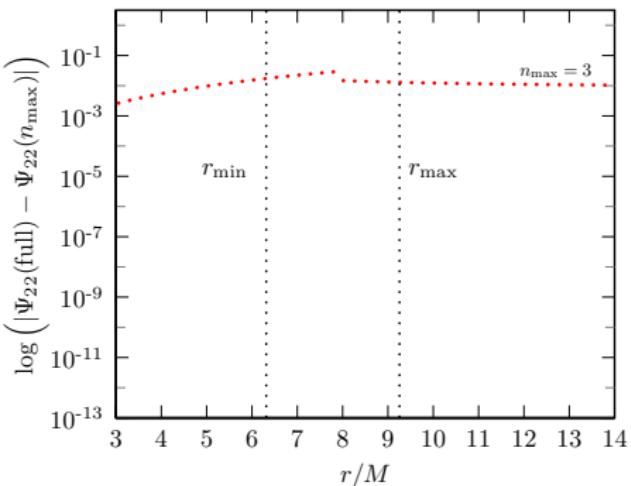
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Fourier Convergence of the Master Functions

Standard Method



Extended Homogeneous Solutions



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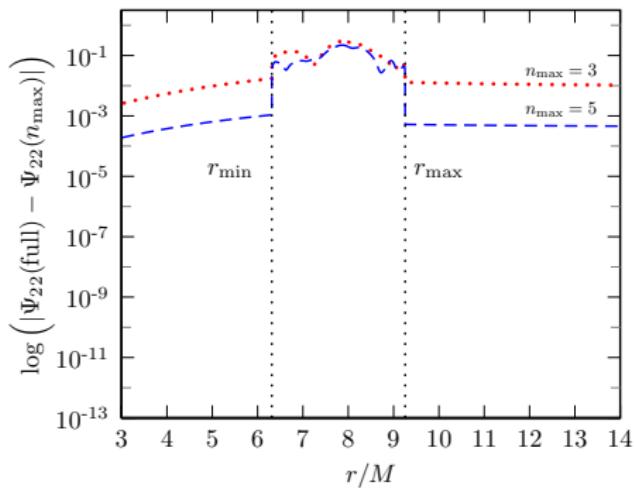
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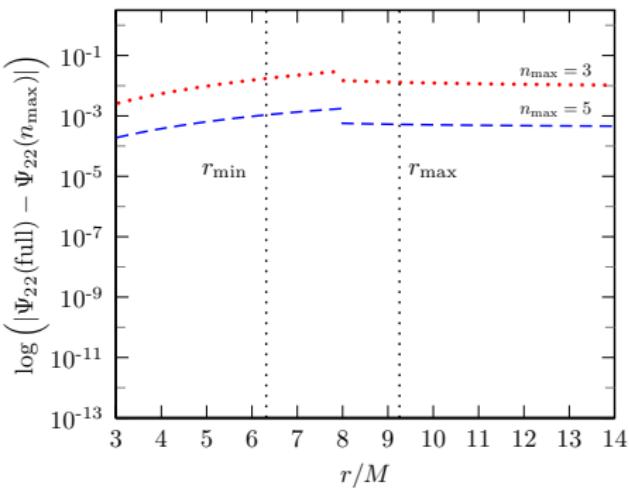
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Fourier Convergence of the Master Functions

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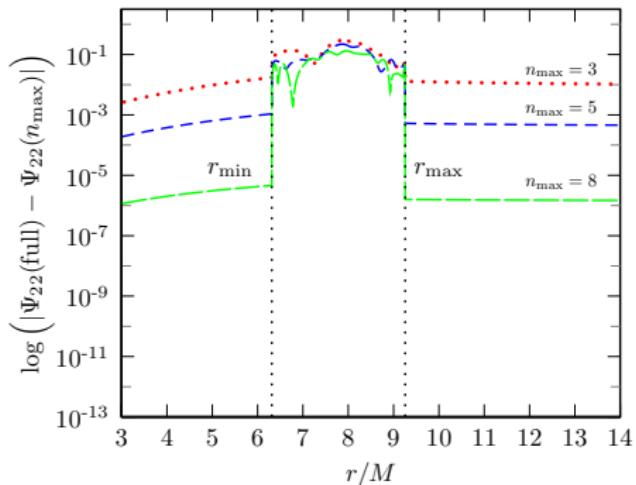
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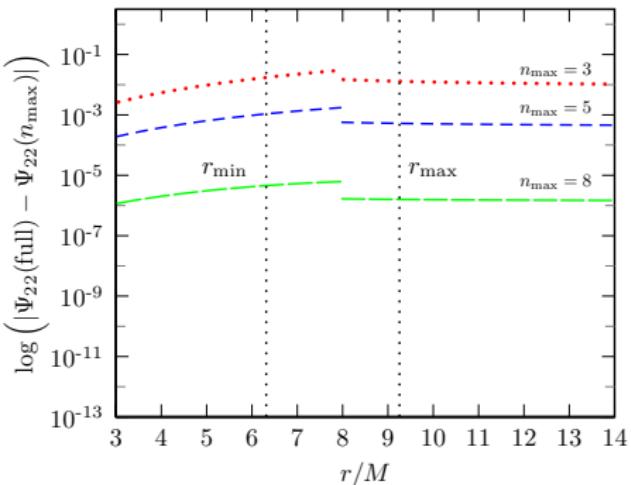
$$\Psi_{\ell m}(t_p, r) = \sum_{n=-5}^5 R_{\ell mn}(r) e^{-i\omega_{mn} t}$$

Fourier Convergence of the Master Functions

Standard Method



Extended Homogeneous Solutions



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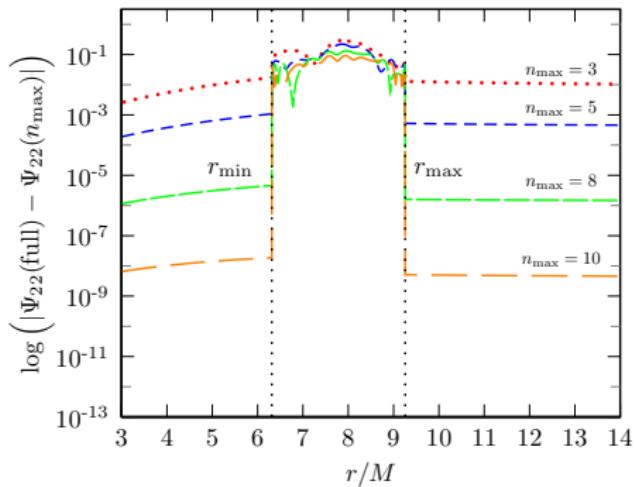
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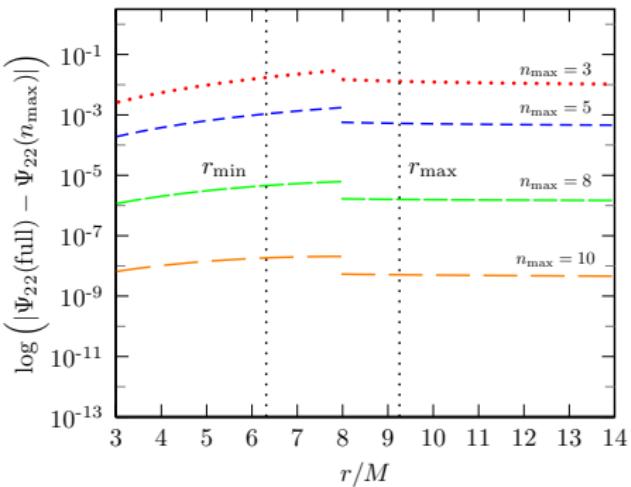
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Fourier Convergence of the Master Functions

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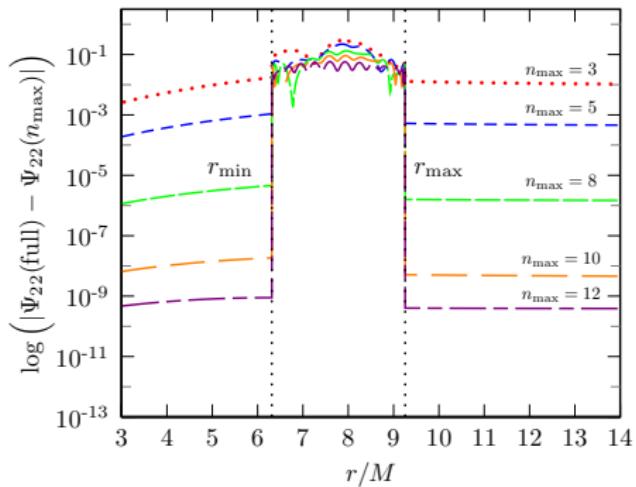
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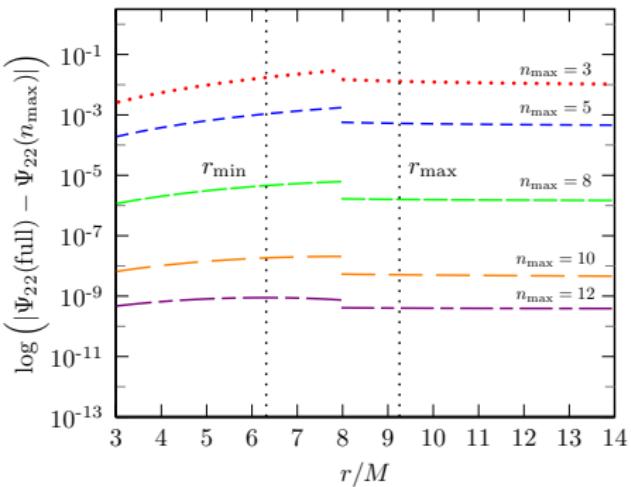
$$\Psi_{\ell m}(t_p, r) = \sum_{n=-10}^{10} R_{\ell mn}(r) e^{-i\omega_{mn} t}$$

Fourier Convergence of the Master Functions

Standard Method



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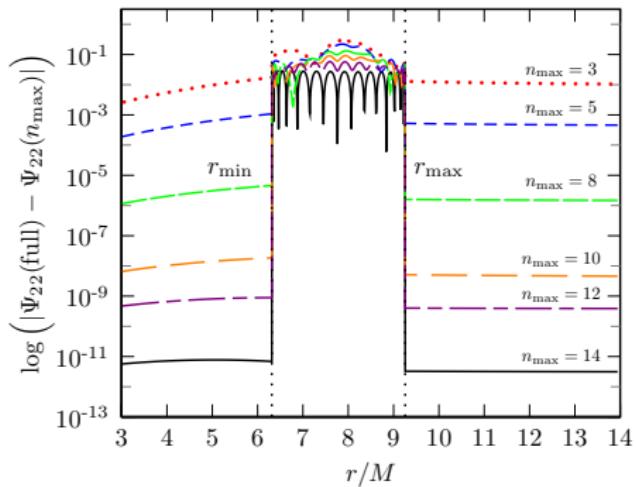
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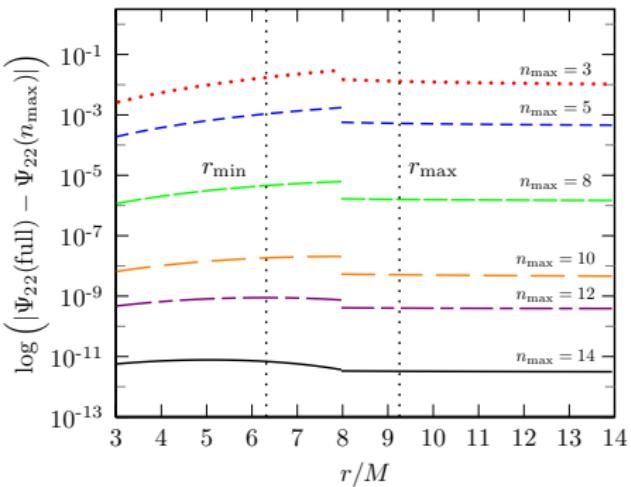
$$\Psi_{\ell m}(t_p, r) = \sum_{n=-12}^{12} R_{\ell mn}(r) e^{-i\omega_{mn} t}$$

Fourier Convergence of the Master Functions

Standard Method



Extended Homogeneous Solutions



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$$\Psi_{\ell m}^{\pm}(t_p, r) = \sum_{n=-14}^{14} R_{\ell mn}^{\pm}(r) e^{-i\omega_{mn} t}$$

Metric Perturbation Reconstruction

- Even-parity sector:

$$K(t, r) = f \partial_r \Psi_{\text{even}} + A \Psi_{\text{even}} - \frac{r^2 f^2}{(\lambda + 1)\Lambda} Q^{tt},$$

$$h_{rr}(t, r) = \frac{\Lambda}{f^2} \left[\frac{\lambda + 1}{r} \Psi_{\text{even}} - K \right] + \frac{r}{f} \partial_r K,$$

$$h_{tr}(t, r) = r \partial_t \partial_r \Psi_{\text{even}} + r B \partial_t \Psi_{\text{even}} - \frac{r^2}{\lambda + 1} \left[Q^{tr} + \frac{rf}{\Lambda} \partial_t Q^{tt} \right],$$

$$h_{tt}(t, r) = f^2 h_{rr} + f Q^\sharp,$$

$$\Lambda(r) \equiv \lambda + \frac{3M}{r}, \quad \lambda \equiv \frac{(\ell + 2)(\ell - 1)}{2}.$$

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- Q 's are the even-parity source terms $\propto \delta[r - r_p(t)]$

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Metric Perturbation Reconstruction

- Odd-parity sector:

$$h_t(t, r) = \frac{f}{2} \partial_r (r \Psi_{\text{odd}}) - \frac{r^2 f}{2\lambda} P^t,$$
$$h_r(t, r) = \frac{r}{2f} \partial_t \Psi_{\text{odd}} + \frac{r^2}{2\lambda f} P^r,$$

- P 's are odd-parity source terms $\propto \delta[r - r_p(t)]$
- Again, assume $\Psi = \Psi^+ \theta(z) + \Psi^- \theta(-z)$, $z \equiv r - r_p(t)$
- Both of the odd-parity amplitudes are C^{-1} , with no singular terms:

$$h_t^\pm(t, r) = \frac{f}{2} \partial_r (r \Psi^\pm), \quad h_t^S(t) = 0,$$
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Metric Perturbation Reconstruction

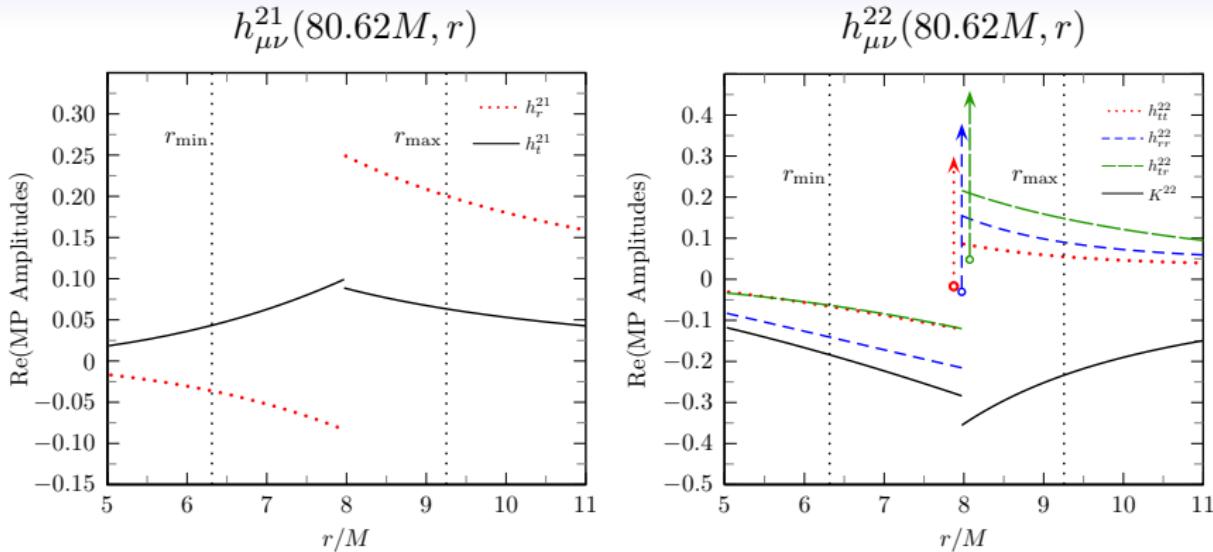
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Metric Perturbation Reconstruction



$h_{\mu\nu}^{\ell m}$ recovered from: $\Psi_{\ell m}$, $\partial_\mu \Psi_{\ell m}$ and $T_{\mu\nu}$

- Even parity metric amplitudes h_{tt} , h_{tr} , and h_{rr} are point singular
Time dependent singular amplitudes are computable
- Metric amplitudes K (even) and h_r (odd) and h_t (odd) are C^{-1}

Transforming to Lorenz gauge

- Gauge transformation from Regge-Wheeler (RW) to Lorenz (L)

$$x_{\text{RW}}^\mu \rightarrow x_{\text{L}}^\mu = x_{\text{RW}}^\mu + \Xi^\mu, \quad |\Xi^\mu| \sim |p_{\mu\nu}| \ll 1$$

- Metric perturbation transforms as

$$p_{\mu\nu}^{\text{RW}} \rightarrow p_{\mu\nu}^{\text{L}} = p_{\mu\nu}^{\text{RW}} - \Xi_{\mu|\nu} - \Xi_{\nu|\mu}$$

Stroke $|$ is covariant differentiation with respect to $g_{\mu\nu}$

- Demand $p_{\mu\nu}^{\text{L}}$ satisfy the Lorenz gauge condition, $\bar{p}_{\mu\nu}^{\text{L}}{}^{\nu} = 0$
- Therefore

$$\Xi_{\mu|\nu}{}^{\nu} = \bar{p}_{\mu\nu}^{\text{RW}}{}^{\nu} = p_{\mu\nu}^{\text{RW}}{}^{\nu} - \frac{1}{2}g^{\alpha\beta}p_{\alpha\beta|\mu}^{\text{RW}}$$

Equations for the Gauge Generator Amplitudes

- Define spherical harmonic amplitudes

$$\Xi_t = \xi_t(t, r) Y^{\ell m}$$

$$\Xi_r = \xi_r(t, r) Y^{\ell m}$$

$$\Xi_A = \xi_{(e)}(t, r) Y_A^{\ell m} + \xi_{(o)}(t, r) X_A^{\ell m}$$

- One, separate odd-parity equation

$$\left[-\frac{\partial^2}{\partial t^2} + \frac{\partial^2}{\partial r_*^2} - f(r) \frac{\ell(\ell+1)}{r^2} \right] \xi_{(o)}(t, r) = 2f\Psi_{RW} + fp(t)\delta[r - r_p(t)]$$

- Three, coupled even-parity equations

$$\square \xi_t + M_t(\xi_t, \xi_r) = F_t(\Psi_{ZM}) + \text{singular term}$$

$$\square \xi_r + M_r(\xi_r, \xi_t, \xi_{(e)}) = F_r(\Psi_{ZM}) + \text{singular term}$$

$$\square \xi_{(e)} + M_{(e)}(\xi_{(e)}, \xi_r) = F_{(e)}(\Psi_{ZM}) + \text{singular term}$$

Alternative: Sago, Nakano, and Sasaki

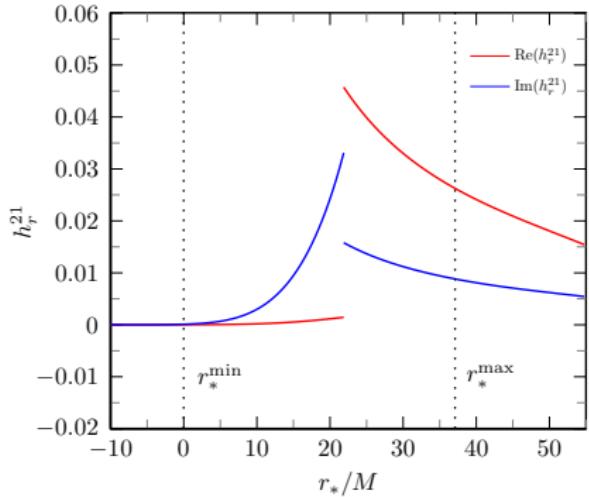
- Formalism devised by Sago, Nakano, and Sasaki (2002)
- The odd-parity part same as before
- The even-parity splits further into scalar part and divergence-free vector part

$$\Xi_{\text{even}}^{\mu} = \Xi_{(s)}{}^{\mu} + \Xi_{(v)}^{\mu}$$

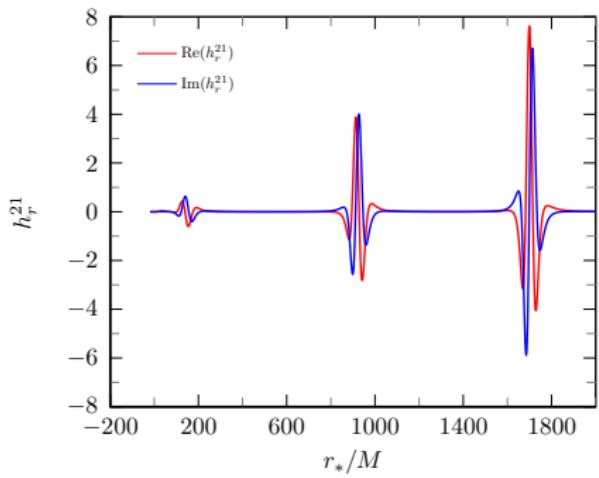
- $r\Xi_{(s)}$ satisfies 4th order wave equation
- Antisymmetric gradient of $\Xi_{(v)}$ projects to two scalars ϕ_0 and ϕ_2
 ϕ_0 and ϕ_2 enter into $s = -1$ and $s = +1$ Teukolsky equations
- A last equation sourced by ϕ_0 and ϕ_2 determines ξ_t , ξ_r , and $\xi_{(e)}$
- Direct even-parity gauge solution: 3 coupled 2nd order equations
- SNS even-parity gauge solution: 1 separate 4th order, 3 separate 2nd order equations

Sneak Peek: $h_r^{\ell m}$ in Regge-Wheeler gauge

$h_r^{21}(t_o, r_*)$ locally



$h_r^{21}(t_o, r_*)$ asymptotically



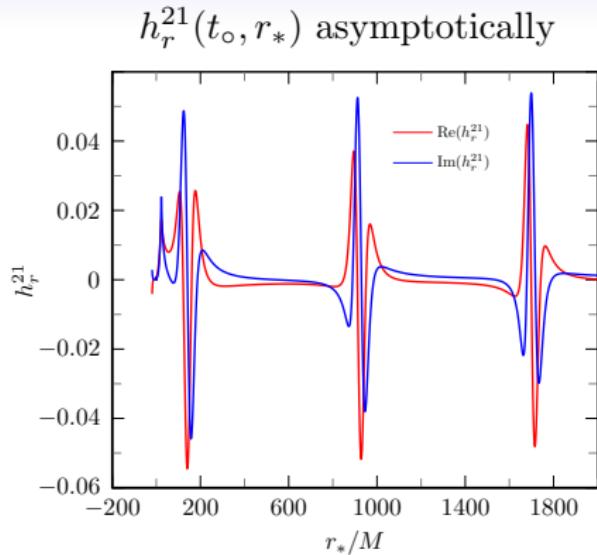
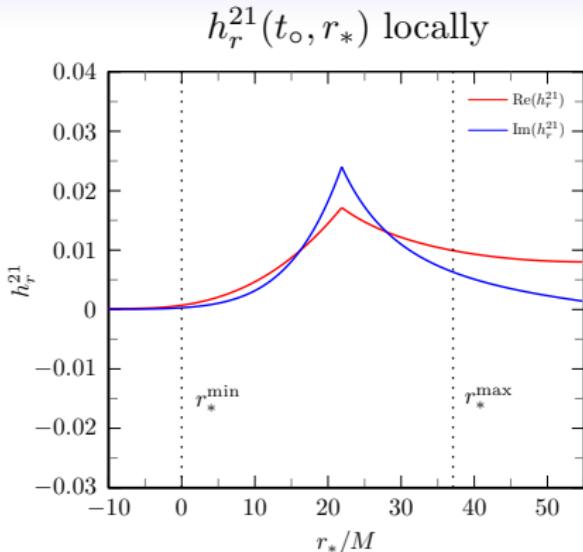
$$p = 8.75455$$

$$e = 0.764124$$

$$t_o = 143.45M$$

$$-50 \leq n \leq 50$$

Sneak Peek $h_r^{\ell m}$ in Lorenz gauge



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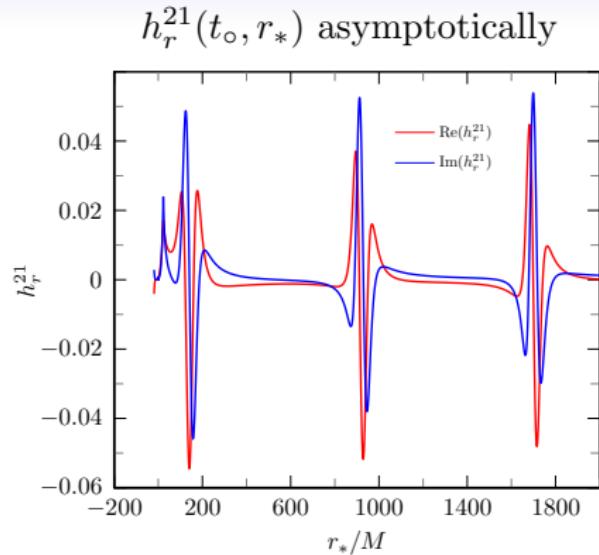
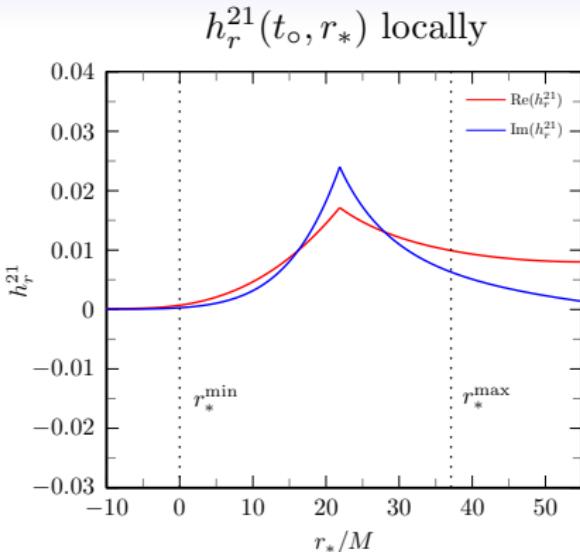
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- Now C^0 at the particle
- Asymptotically \sim wave

Sneak Peek h_r^{lm} in Lorenz gauge



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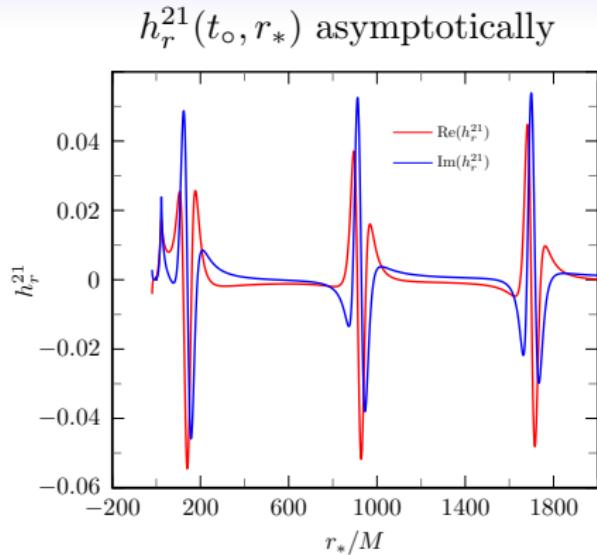
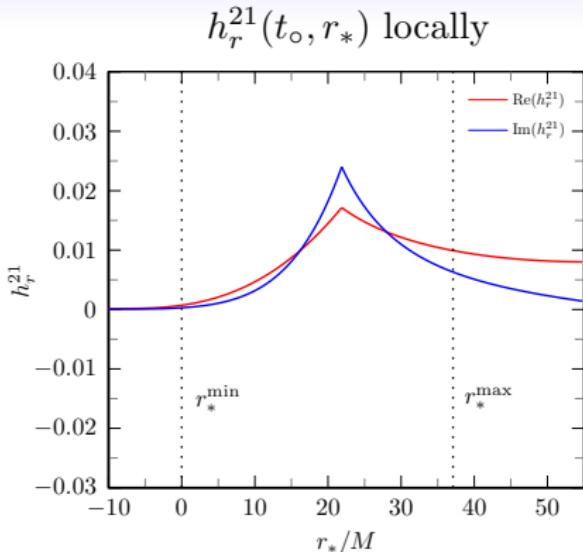
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Conclusions

- Highly accurate (10^{-12}) $\Psi_{\ell m}^{\text{even/odd}}$ for eccentric orbits
 Use FD engine
- EHS method obtains master functions and metric (RW) accurate at $r = r_p(t)$
- TD jump conditions provide checks and convergence criteria
- Reasonable runtimes $\lesssim 6$ hours for $0 \leq e \lesssim 0.8$
- Frequency domain fully competitive with time domain for most e .
- Avoids transients
- Gauge transformation to Lorenz in progress (Seth Hopper)
- Using Sago, Nakano, and Sasaki instead of direct, coupled 1+3 gauge equations
- Some new solution techniques coming out of latter effort

Jump Conditions in the Time Domain

- Now we equate the respective δ and δ' coefficients from the two sides, leaving

$$[\![\Psi]\!]_p(t) = \frac{\mathcal{E}^2}{f_p^2 U_p^2} \tilde{F}(t)$$

$$\begin{aligned} [\![\partial_r \Psi]\!]_p(t) = & \frac{\mathcal{E}^2}{f_p^2 U_p^2} \left[\tilde{G}(t) + \frac{1}{U_p^2 r_p^2} \left(3M - \frac{\mathcal{L}^2}{r_p} + \frac{5M\mathcal{L}^2}{r_p^2} \right) \tilde{F}(t) \right. \\ & \left. - 2\dot{r}_p \frac{d}{dt} ([\![\Psi]\!]_p) \right]. \end{aligned}$$

- $U^2(r, \mathcal{L}^2) \equiv f(1 + \mathcal{L}^2/r^2)$
- \mathcal{E} and \mathcal{L} are the specific energy and angular momentum
- These expressions provide a powerful check on the validity of our method.
- They also provide useful stopping conditions, telling us when we've computed enough frequency modes.

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- These expressions provide a powerful check on the validity of our method.
- They also provide useful stopping conditions, telling us when we've computed enough frequency modes.

Jump Conditions in the Time Domain

- Now we equate the respective δ and δ' coefficients from the two sides, leaving

$$[\![\Psi]\!]_p(t) = \frac{\mathcal{E}^2}{f_p^2 U_p^2} \tilde{F}(t)$$

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Master Functions

2 degrees of freedom corresponding to 2 wave polarizations

$\Psi_{\ell m}^{\text{even/odd}}$ are linear combinations of $h_{\mu\nu}^{\ell m}$ and its derivatives in Regge-Wheeler gauge:

$$\Psi_{\text{even}}(t, r) \equiv \frac{2r}{\ell(\ell+1)} \left[K + \frac{1}{\Lambda} (f^2 h_{rr} - rf \partial_r K) \right],$$

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Both satisfy wave equations

$$\left[-\frac{\partial^2}{\partial t^2} + \frac{\partial^2}{\partial r_*^2} - V_\ell(r) \right] \Psi_{\ell m}(t, r) = S_{\ell m}(t, r), \quad r_* : \text{tortoise coord}$$

Singular sources: $S_{\ell m}(t, r) = \tilde{G}_{\ell m}(t) \delta[r - r_p(t)] + \tilde{F}_{\ell m}(t) \delta'[r - r_p(t)]$

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Metric Perturbation Reconstruction

- Even-parity sector:

$$K(t, r) = f \partial_r \Psi_{\text{even}} + A \Psi_{\text{even}} - \frac{r^2 f^2}{(\lambda + 1)\Lambda} Q^{tt},$$

$$h_{rr}(t, r) = \frac{\Lambda}{f^2} \left[\frac{\lambda + 1}{r} \Psi_{\text{even}} - K \right] + \frac{r}{f} \partial_r K,$$

$$h_{tr}(t, r) = r \partial_t \partial_r \Psi_{\text{even}} + r B \partial_t \Psi_{\text{even}} - \frac{r^2}{\lambda + 1} \left[Q^{tr} + \frac{rf}{\Lambda} \partial_t Q^{tt} \right],$$

$$h_{tt}(t, r) = f^2 h_{rr} + f Q^\sharp,$$

$$\Lambda(r) \equiv \lambda + \frac{3M}{r}, \quad \lambda \equiv \frac{(\ell + 2)(\ell - 1)}{2}.$$

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- Q 's are the even-parity source terms $\propto \delta[r - r_p(t)]$

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Metric Perturbation Reconstruction

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- P 's are odd-parity source terms $\propto \delta[r - r_p(t)]$
- Again, assume $\Psi = \Psi^+ \theta(z) + \Psi^- \theta(-z)$, $z \equiv r - r_p(t)$
- Both of the odd-parity amplitudes are C^{-1} , with no singular terms:

$$h_t^\pm(t, r) = \frac{f}{2} \partial_r (r \Psi^\pm), \quad h_t^S(t) = 0,$$
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Black Hole Perturbation Theory

| Field Equations | Equations of Motion |
|---|--|
| \mathcal{O} | $\frac{Du^\mu}{d\tau} = 0$ |
| $\square \bar{h}_{\alpha\beta} + 2R^\gamma{}_\alpha{}^\delta{}_\beta \bar{h}_{\gamma\delta} = -16\pi T_{\alpha\beta}$ | $\frac{Du^\mu}{d\tau} = -\frac{1}{2} (g^{\mu\nu} + u^\mu u^\nu)$ |
| $\bar{h}^{\alpha\beta}{}_{ \beta} = 0$ | $\times (2h_{\nu\lambda;\rho} - h_{\lambda\rho;\nu}) u^\lambda u^\rho$ |
| $g_{\mu\nu} = g_{\mu\nu} + {}_1 h_{\mu\nu} \dots$ | MiSaTaQuWa Equations |
| $\square_2 \bar{h} = (1 + {}_1 \bar{h}) T[z + \delta z] + (\nabla {}_1 \bar{h})^2$ | |

(Schematic)

Black Hole Perturbation Theory

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|---|--|
| \mathcal{O} $G_{\mu\nu}(g_{\alpha\beta}) = 0$ |  $\frac{Du^\mu}{d\tau} = 0$ |
| $\square \bar{h}_{\alpha\beta} + 2R^\gamma{}_\alpha{}^\delta{}_\beta \bar{h}_{\gamma\delta} = -16\pi T_{\alpha\beta}$ $\bar{h}^{\alpha\beta}{}_{ \beta} = 0$ | $\frac{Du^\mu}{d\tau} = -\frac{1}{2} (g^{\mu\nu} + u^\mu u^\nu)$ $\times (2h_{\nu\lambda;\rho} - h_{\lambda\rho;\nu}) u^\lambda u^\rho$ |
| $g_{\mu\nu} = g_{\mu\nu} + {}_1 h_{\mu\nu} \dots$ | MiSaTaQuWa Equations |
| $\square_2 \bar{h} = (1 + {}_1 \bar{h}) T[z + \delta z] + (\nabla {}_1 \bar{h})^2$ | |

(Schematic)

Black Hole Perturbation Theory

| | Field Equations | Equations of Motion |
|----------|---|---|
| 0 | $G_{\mu\nu}(g_{\alpha\beta}) = 0$ | $\frac{Du^\mu}{d\tau} = 0$ |
| 1 | $\square \bar{h}_{\alpha\beta} + 2R^\gamma{}_\alpha{}^\delta{}_\beta \bar{h}_{\gamma\delta} = -16\pi T_{\alpha\beta}$ $\bar{h}^{\alpha\beta}{}_{ \beta} = 0$ | $\frac{Du^\mu}{d\tau} = -\frac{1}{2} (g^{\mu\nu} + u^\mu u^\nu) \times (2h_{\nu\lambda;\rho} - h_{\lambda\rho;\nu}) u^\lambda u^\rho$ |
| | $g_{\mu\nu} = g_{\mu\nu} + {}_1 h_{\mu\nu} \dots$ | MiSaTaQuWa Equations |
| | $\square_2 \bar{h} = (1 + {}_1 \bar{h}) T[z + \delta z] + (\nabla {}_1 \bar{h})^2$ | |

(Schematic)

Black Hole Perturbation Theory

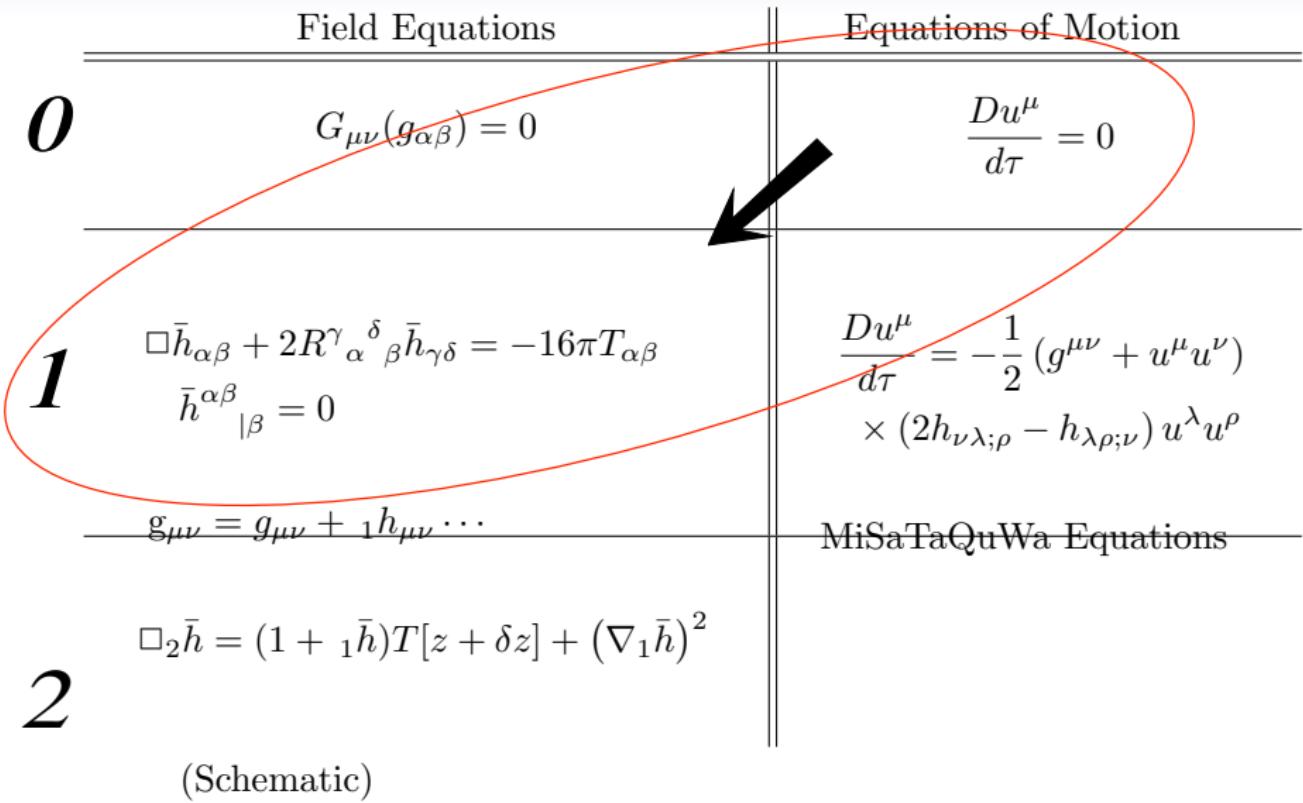
| | Field Equations | Equations of Motion |
|----------|---|---|
| 0 | $G_{\mu\nu}(g_{\alpha\beta}) = 0$ | $\frac{Du^\mu}{d\tau} = 0$ |
| 1 | $\square \bar{h}_{\alpha\beta} + 2R^\gamma{}_\alpha{}^\delta{}_\beta \bar{h}_{\gamma\delta} = -16\pi T_{\alpha\beta}$ $\bar{h}^{\alpha\beta}{}_{ \beta} = 0$ |  $\begin{aligned} \frac{Du^\mu}{d\tau} = & -\frac{1}{2} (g^{\mu\nu} + u^\mu u^\nu) \\ & \times (2h_{\nu\lambda;\rho} - h_{\lambda\rho;\nu}) u^\lambda u^\rho \end{aligned}$ |
| | $\underline{g}_{\mu\nu} = g_{\mu\nu} + {}_1h_{\mu\nu} \dots$ | MiSaTaQuWa Equations |
| | $\square_2 \bar{h} = (1 + {}_1\bar{h})T[z + \delta z] + (\nabla {}_1\bar{h})^2$ | |
| | (Schematic) | |

Black Hole Perturbation Theory

| | Field Equations | Equations of Motion |
|----------|---|---|
| 0 | $G_{\mu\nu}(g_{\alpha\beta}) = 0$ | $\frac{Du^\mu}{d\tau} = 0$ |
| 1 | $\square \bar{h}_{\alpha\beta} + 2R^\gamma{}_\alpha{}^\delta{}_\beta \bar{h}_{\gamma\delta} = -16\pi T_{\alpha\beta}$ $\bar{h}^{\alpha\beta}{}_{ \beta} = 0$ | $\frac{Du^\mu}{d\tau} = -\frac{1}{2} (g^{\mu\nu} + u^\mu u^\nu) \times (2h_{\nu\lambda;\rho} - h_{\lambda\rho;\nu}) u^\lambda u^\rho$ |
| 2 | $g_{\mu\nu} = g_{\mu\nu} + {}_1 h_{\mu\nu} \dots$ |  MiSaTaQuWa Equations |
| | $\square_2 \bar{h} = (1 + {}_1 \bar{h}) T[z + \delta z] + (\nabla {}_1 \bar{h})^2$ | |

(Schematic)

Black Hole Perturbation Theory



Analytic and Numerical Method

| SH & CRE (2010) | Barack and Sago (2010) |
|--|---|
| Frequency domain | Time domain |
| RWZ formalism | Lorenz gauge |
| Master functions: $\Psi_{\ell m}^{\text{odd}}$, $\Psi_{\ell m}^{\text{even}}$ | Metric perturbations: $h_{\mu\nu}^{\ell m}$ |
| ODEs but multiple harmonics | Coupled PDEs |
| Exact bndy conditions, periodic | IV problem, transients |
| Smeared source, libration region | Moving delta source |
| ODE accuracy, TD convergence | 4th-order PDE scheme |
| Gibbs problem (use EHS) | Time dep. jump conditions |
| Reconstructed metric (RW gauge) | Metric in Lorenz gauge |
| Gauge transform and self-force | Self-force |
| Kerr/higher order? | Kerr/higher order? |