

# Metric reconstruction via Hertz potential in the time domain

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- 1 Background and Motivation
- 2 Our new method: Metric reconstruction in the time domain
- 3 Cases
  - Static particle in flat space
  - Circular orbit in Schwarzschild
  - Generic orbit in Schwarzschild
  - Kerr
- 4 Implementation
- 5 Outlook

# Background and Motivation

# Motivation: Long Term evolution of a particle on Kerr background

	Solve For $h^{Lorentz}$	Solve Teukolsky equation
Frequency domain	(Akcaý etal, Osburn etal) <ul style="list-style-type: none"><li>• Hard to evolve orbit</li><li>• Hard to get high eccentricity</li><li>• Degenerate modes</li></ul>	(Shah etal ,van de Meent) <ul style="list-style-type: none"><li>• Hard to evolve orbit</li><li>• Hard to get high eccentricity</li></ul>
Time domain	(Barack, Sago, Dolan) <ul style="list-style-type: none"><li>• Computationally expansive</li><li>• Some unstable modes</li></ul>	<b>OUR NEW METHOD:</b> <ul style="list-style-type: none"><li>• Any orbit</li><li>• No "bad modes"</li><li>• Evolution</li></ul>

# Background I : Metric reconstruction in vacuum

The CCK metric reconstruction procedure involves three steps:

- Solving the Teukolsky equation for  $\psi_s$

$$T_s \psi_s = 0$$

- Finding Hertz potential that satisfies both the Teukolsky equation  $T_{-s} \Psi^* = 0$  and a differential equation with  $\psi_s$  as a source

$$D_s^4 \Psi = \psi_s$$

- Operating on the Hertz potential with another differential operator to obtain metric perturbation in a (traceless) radiation gauge

$$h_{\alpha\beta} = H_{\alpha\beta} \Psi$$

# Background I : Metric reconstruction in vacuum

The Kinnersley tetrad in Boyer-Linquist coordinates is given by  $e_a^\alpha = (\ell^\alpha, n^\alpha, m^\alpha, \bar{m}^\alpha)$ . We can use ORG or IRG gauges defined as

$$IRG : l^\alpha g_{\alpha\beta} = 0 \quad ORG : n^\alpha g_{\alpha\beta} = 0$$

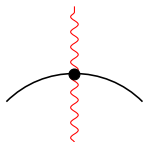
Reconstruction of the metric via Hertz potential is done via

$$\begin{aligned} H_{\alpha\beta}^{ORG} = & \\ & - \varrho^{-4} \left\{ n_\alpha n_\beta (\bar{\delta} - 2\alpha) (\bar{\delta} - 4\alpha) + \bar{m}_\alpha \bar{m}_\beta (\Delta + 5\mu - 2\gamma) (\Delta + \mu - 4\gamma) \right. \\ & \left. - n_{(\alpha} \bar{m}_{\beta)} \left[ (\bar{\delta} - 2\alpha) (\Delta + \mu - 4\gamma) + (\Delta + 4\mu - 4\gamma) (\bar{\delta} - 4\alpha) \right] \right\} \\ & \pm \text{c.c.} \end{aligned}$$

similar for IRG

# Background II : Metric reconstruction with point particle

- Gauges: full-string, half-string, no-string<sup>1</sup>



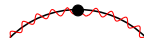
Singular along  $\ell^\alpha$



Regular at  $\ell^{\alpha+}$



Regular at  $\ell^{\alpha-}$



Discontinuous across a surface

- No-string reconstruction in f-domain done by Friedman, Shah et al, Merlin , van de Meent
- Here we will discuss no-string gauge in time domain

<sup>1</sup>Courtesy of C.Merlin

## Background III : Self-force from reconstructed metric

- In either of the two half string gauges via

$$F_{\alpha}^{\pm} = \sum_l (F_{\alpha}^{l\pm} - A_{\alpha}^{\pm} L - B_{\alpha} - C_{\alpha}/L) + D_{\alpha}^{\pm}$$

- radial limit  $\pm$
  - A,B,C as in Lorenz gauge
  - $D_{\alpha}^{\pm}$  non zero in general (depends on  $u^{\alpha}$  extension)
- In no-string gauge

$$F_{\alpha} = \sum_l (\bar{F}_{\alpha}^l - B_{\alpha}) \quad \bar{F}^l = \frac{1}{2}(F^+ + F^-)$$



# Our new method: Metric reconstruction in the time domain

Evolve Hertz Potential directly in the time domain

## Need:

- Evolution equation (Teukolsky equation) ✓
- Boundary conditions
- Junction condition at the particle

# $\Psi$ Boundary conditions

$\psi_4$  and  $\psi_0$  satisfy the vacuum Teukolsky equation for spin-weights  $\mp 2$  when  $\Psi^{ORG}$  or  $\Psi^{IRG}$  satisfy the spin-weight  $\pm 2$  Teukolsky equations.

The asymptotic behaviour will be  
(Ingoing)

$$\Psi^{IRG} \propto e^{-i\omega v} / r^5 \quad r \rightarrow \infty, \quad \boxed{\Psi^{IRG} \propto e^{-i\omega v} \Delta^{-s} \quad r \rightarrow 2M}$$

(Outgoing)

$$\boxed{\Psi^{IRG} \propto e^{-i\omega u} / r \quad r \rightarrow \infty}, \quad \Psi^{IRG} \propto e^{-i\omega u} \quad r \rightarrow 2M$$

Similar for  $\Psi^{ORG}$ .

## Regularity condition at the Horizon

$\Delta^s \Psi$  is smooth with  $\Delta = (r - r_-)(r - r_+)$

# Junction condition for $\Psi$

Obtained from inversion relation and field equation in vacuum

$$D^4\Psi = \psi_0, \quad T\Psi = 0$$

given the junction conditions for  $\psi_0$

- 1 On either sides of the world line

$$(\ell^\mu \partial_\mu)^4 \Psi^\pm = \psi_0^\pm$$

- 2 Take Right-Left difference (denoted by  $[\dots]$ )
- 3 use  $T\Psi^\pm = 0$  to write LHS in terms of  $[\Psi]$  and  $[\Psi_\nu]$
- 4 Solve for  $[\Psi]$  and  $[\Psi_\nu]$  in terms of  $[\psi_0]$  and  $[\psi_{0,\nu}]$  (known)

# Cases

# Case1: Static particle in flat space

- The inversion relation reduces to

$$\frac{d^4}{dr^4}\Psi^* = \psi_0$$

- By taking derivatives of  $T\Psi^* = 0$  repeatedly express  $\frac{d^4}{dr^4}\Psi^*$  and  $\frac{d^5}{dr^5}\Psi^*$  in terms of  $\Psi^*$  and  $\Psi_r^*$ .
- Take 'Right-Left' difference to obtain

$$\begin{aligned} -\frac{16(l^2 + l + 15)}{r^3}[\partial_r\Psi^*] + \frac{(l^2 + l + 60)(l + 3)(l - 2)}{r^4}[\Psi^*] &= [\psi_0] \\ \frac{1800 + (l^2 + l)(l^2 + l + 198)}{r^4}[\partial_r\Psi^*] - \frac{20(l^2 + l + 24)(l^2 + l - 6)}{r^5}[\Psi^*] &= [\partial_r\psi_0] \end{aligned}$$

This can be inverted algebraically to derive  $[\Psi^*]$  and  $[\Psi_r^*]$

## Case2: Schwarzschild for circular orbits

- In this case the inversion relations can be written as

$$[\Psi_{vvv}] = [\psi_0]$$

- Mixed derivative jumps can be derived via TE in vacuum

$$[\Psi_{vu}] = -(A_p^u[\Psi_u] + A_p^v[\Psi_v] + V_p[\Psi])$$

- Taking  $\tau$  derivatives we derive the set

$$[\dot{\Psi}] = \dot{u}[\Psi_u] + \dot{v}[\Psi_v] = 0$$

$$[\dot{\Psi}_v] = \dot{u}[\Psi_{vu}] + \dot{v}[\Psi_{vv}] = 0$$

$$[\ddot{\Psi}] = 2\dot{v}\dot{u}[\Psi_{vu}] + \dot{u}^2[\Psi_{uu}] + \dot{v}^2[\Psi_{vv}] = 0$$

This can be done till the 5th derivative, all the jumps of the form  $[\Psi_{v\dots vu\dots u}]$  will be written in terms of  $[\Psi_v]$ ,  $[\Psi]$  so we can solve as the previous case

## Case3: Schwarzschild for generic orbits

- Same inversion relation as previous case
- Second derivatives terms can be derived by solving

$$[\dot{\Psi}] = \dot{u}[\Psi_u] + \dot{v}[\Psi_v]$$

$$[\dot{\Psi}_v] = \dot{u}[\Psi_{vu}] + \dot{v}[\Psi_{vv}]$$

$$[\ddot{\Psi}] = \ddot{v}[\Psi_v] + \ddot{u}[\Psi_u] + 2\dot{v}\dot{u}[\Psi_{vu}] + \dot{u}^2[\Psi_{uu}] + \dot{v}^2[\Psi_{vv}]$$

Same procedure to the next order looking at  $\frac{d}{d\tau}[\Psi_{vv}]$  and  $\frac{d^3}{d\tau^3}[\Psi]$ .



## Case3: Schwarzschild for generic orbits

We solve iteratively till the 5th order where  $[\Psi_{vvvv}]$  and  $[\Psi_{vvvvv}]$  will be written in terms of  $([\Psi], \frac{d}{d\tau}[\Psi], \dots, \frac{d^4}{d\tau^4}[\Psi])$  and  $([\Psi_v], \dots, \frac{d^4}{d\tau^4}[\Psi_v])$  so we obtain the system

$$\sum_{n=0}^3 a_n \frac{d^n}{d\tau^n}[\Psi] + \sum_{n=0}^3 b_n \frac{d^n}{d\tau^n}[\Psi_v] = [\psi_0]$$
$$\sum_{n=0}^4 \tilde{a}_n \frac{d^n}{d\tau^n}[\Psi] + \sum_{n=0}^4 \tilde{b}_n \frac{d^n}{d\tau^n}[\Psi_v] = [\psi_{0,v}]$$

where the RHS is known. This is a system of coupled ODEs of 3rd and 4th order to be solved numerically.

## Case4: Kerr Case

- Use the decomposition in spin-weighted spherical harmonics

$$\Psi = \sum_{lm} Y_{slm} \Psi_{lm}$$

The 1+1D Teukolsky equation has mixed l-modes terms

$$[\Psi'_{uv}] + A^u[\Psi'_u] + A^v[\Psi'_v] + V[\Psi'] + \sum_{\substack{q=l-2 \\ l \neq n}}^{l+2} \mathcal{I}_{l,q}[\Psi^q] = 0$$

- Truncate the number of l modes and write the system in a band diagonal matrix form

$$\begin{bmatrix} \tilde{T}_2 & \mathcal{I}_{2,3} & \mathcal{I}_{2,4} & 0 \\ \mathcal{I}_{3,2} & \tilde{T}_3 & \mathcal{I}_{3,4} & \cdots \\ \mathcal{I}_{4,2} & \mathcal{I}_{4,3} & \cdots & \cdots \\ 0 & \cdots & \cdots & \cdots \end{bmatrix} \begin{pmatrix} \vdots \\ \Psi' \\ \vdots \end{pmatrix} = 0$$

This equation will be used to reduce the order as previously and write  $\{[\Psi'], [\Psi'_v]\}$  in terms of their  $\tau$  derivatives

## Case4: Kerr Case

- The inversion relation reduces to

$$[(\partial_v + \frac{ima}{\Delta})^4 \Psi'] = \sum_{n=0}^4 c_n [\partial_v^n \Psi'] = [\psi'_0]$$

- Each term can be written in terms of  $\{[\Psi'], [\Psi'_v]\}$  and its  $\tau$  derivatives. As before these are coupled ODEs

$$\sum_{n=0}^4 A_n \frac{d^n}{d\tau^n} [\vec{\Psi}] + \sum_{n=0}^3 B_n \frac{d^n}{d\tau^n} [\vec{\Psi}_v] = [\vec{\psi}_0]$$

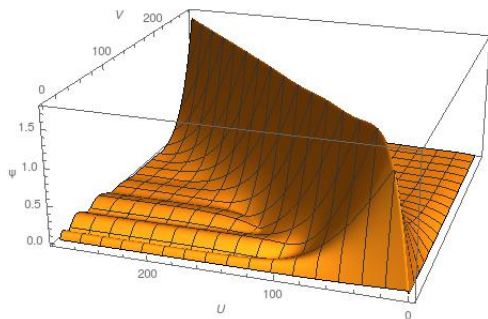
$$\sum_{n=0}^4 \tilde{A}_n \frac{d^n}{d\tau^n} [\vec{\Psi}] + \sum_{n=0}^3 \tilde{B}_n \frac{d^n}{d\tau^n} [\vec{\Psi}_v] = [\vec{\psi}_{0,v}]$$

with  $A, B, \tilde{A}, \tilde{B}$  band diagonal matrices. The system can be solved numerically.

# Implementation

## Numerics

- 1 Use  $(u,v)$  coordinates grid
- 2 Set initial condition to zero on  $(u,v)$  axis
- 3 Evolve via finite difference scheme
- 4 Discard junk radiation and record late time behaviour



Time evolution in  $(u, v)$  for circular orbit in Schwarzschild for  $\psi_0$

# Outlook

# Conclusion and Outlook

- We propose to solve Teukolsky equation for  $\Psi$  in time-domain via finite difference scheme using the derived jumps
- Numerical solutions can be used to reconstruct metric perturbation in a no-string radiation gauge and derive the self force based on Pound et al (2013)
- We are currently implementing numerical code that uses this method

ありがとうございます。

(Thank you)