

# Completion of metric reconstruction for a particle orbiting a Kerr black hole

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# Outline

- Strategy for computing the gravitational SF from reconstructed MP. Adam Pound, Cesar Merlin and Leor Barack. [[Phys. Rev. D 89, 024009](#)]
- Numerical Implementation in Schwarzschild. Cesar Merlin and Abhay G. Shah [[Phys. Rev. D 91, 024005](#)]
- The completion of the solution. Cesar Merlin, Leor Barack, Amos Ori, Adam Pound and Maarten van de Meent.

# Review

- The original formulation of the SF was given in the Lorenz-gauge [Poisson *et al.* (2011)]. For Kerr the field equations in the LG are not separable.
- The treatment of black-hole perturbations for Kerr is much simpler in a radiation gauge, where it is possible to reconstruct the perturbations from the perturbed Weyl scalars. [Chrzanowski-Cohen-Kegeles (1975), Wald (1978), Ori (2008), Keidl *et al.* (2010)]
- Gauge invariant quantities have been successfully computed using this idea. [Shah *et al.* (2011-12)]

# Completion problem

- The solution in the RG needs to be completed by including certain completion piece. [Wald (1973)]

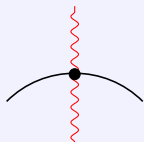
$$h_{\alpha\beta} = h_{\alpha\beta}^{(\text{rec})} + h_{\alpha\beta}^{(\text{comp})}.$$

- In Schwarzschild the completion is the  $\ell = 0, 1$  modes of the Einstein Field Equations. [Detweiler-Poisson (2004), Barack-Lousto (2005), Price (2007)]
- To measure the mass and AM for a reconstructed perturbation one can evaluate some complicated surface-integrals. [Abbot-Desner (1982), Dolan-Barack (2011)]

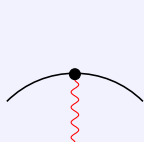
- Shah *et al.* obtained the completion imposing boundary-conditions at the event horizon and at infinity, by assuming that  $h_{\alpha\beta}^{(\text{rec})}$  does not contain any mass and AM. [Shah *et al.* (2011-12)]
- A numerical method to determine an homogeneous piece of the ‘Hertz potential’ for a ring configuration was recently proposed. [Sano-Tagoshi (2014-15)]
- We want to obtain  $h_{\alpha\beta}^{(\text{comp})}$  rigorously.

# Singular structure of the radiation gauges

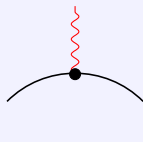
Pathological singularity in the radiation-gauge



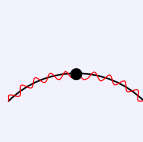
Singular along  $\ell^\alpha$



Regular at  $\ell^{\alpha+}$



Regular at  $\ell^{\alpha-}$



Discontinuous across a surface through the particle

Half-string solutions

$$h_{\tau A}^\pm = \mp \frac{2m x_A}{s(s \pm z)}$$

$$h_{zA}^\pm = \pm \frac{2m x_A}{s(s \pm z)}$$

Full-string solution

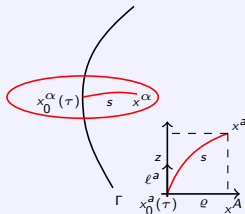
$$h_{\tau A} = \frac{2m z x_A}{s \varrho^2}$$

$$h_{zA} = -\frac{2m z x_A}{s \varrho^2}$$

No-string solution

$$h_{\tau A} = h_{\tau A}^+ \theta(z) + h_{\tau A}^- \theta(-z)$$

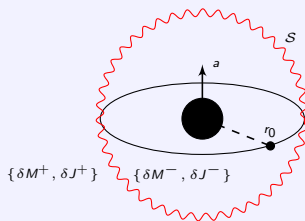
$$h_{zA} = h_{zA}^+ \theta(z) + h_{zA}^- \theta(-z)$$



# Completion of the solution

Wald showed (1973) that the only things we can add to the metric reconstruction are:

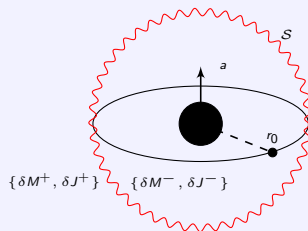
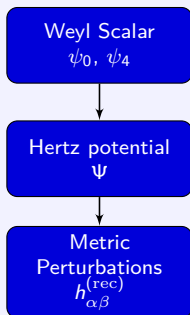
- Mass and angular momentum perturbations ( $\delta M$  and  $\delta J$ ).
- Perturbations to other algebraically special solutions (C-metric and Kerr-NUT).
- Gauge perturbations.



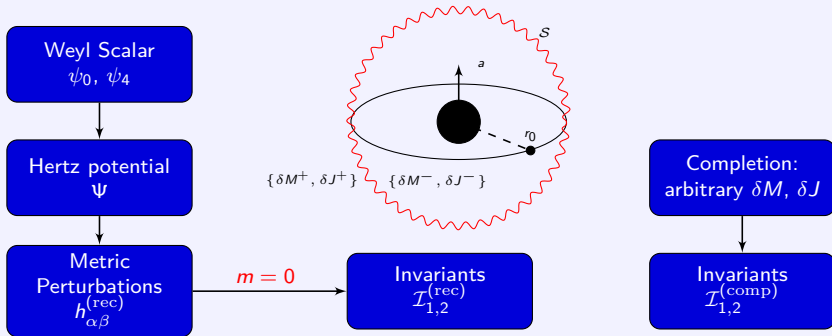




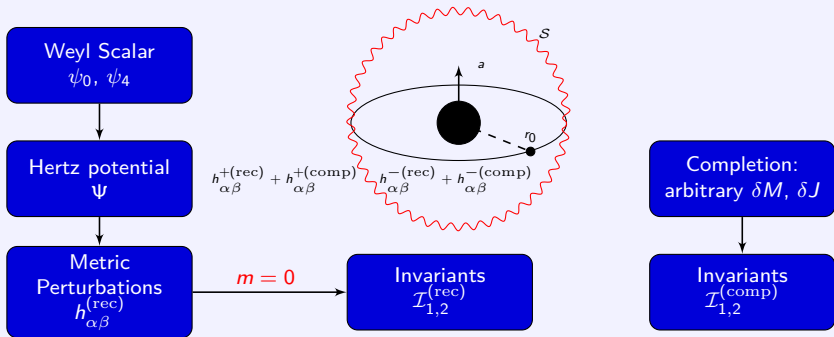
# Implementation



# Implementation



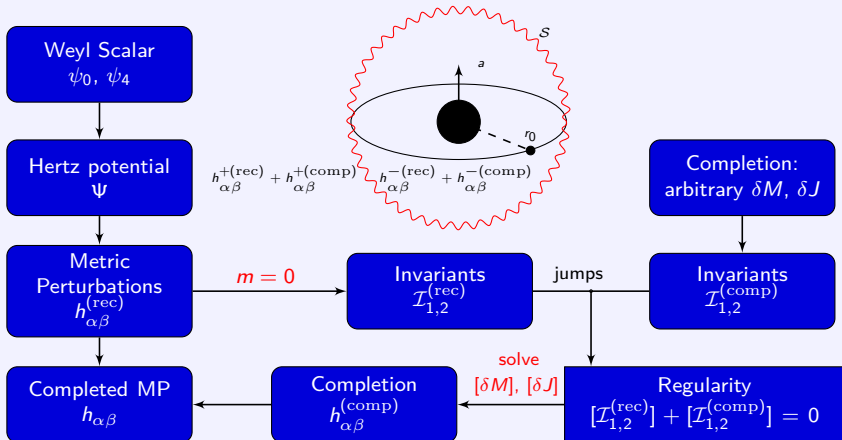
# Implementation



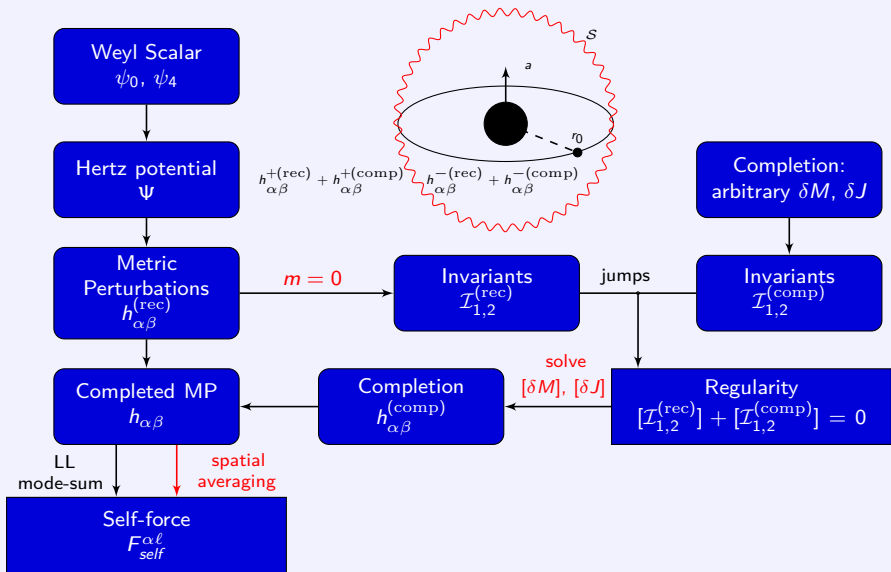
In Schwarzschild we move to a gauge in which  $\text{Re}(\delta\tilde{\psi}_2) = 0$ ,  
 in which  $\mathcal{I}_1 \equiv \tilde{h}_{rr} = h_{rr} - 2\tilde{\xi}_{r,r} + 2\Gamma_{rr}^r \tilde{\xi}_r$ , and  $\mathcal{I}_2 \equiv \text{Im}(\delta\psi_2)$ ,  
 with  $\tilde{\xi}^r = \frac{r^4}{3M} \text{Re}(\delta\psi_2)$ .



# Implementation



# Implementation



# Completion for circular orbits

Our method for circular orbits:

- Obtain the  $m = 0$  sector of the reconstructed part of the MP, using the CCK procedure mode by mode. This is done on each of the two sides of the sphere  $\mathcal{S}$  with  $r = r_0$ .
- Construct two auxiliary invariants  $\mathcal{I}_{1,2}^{(\text{rec})}(r, \theta)$  on each side of  $\mathcal{S}$ , and obtain the jump across  $r = r_0$  on each mode of the invariants.
- Sum all the  $\ell$  modes of the jumps (analytically).
- Construct the completion pieces in Boyer-Lindquist coordinates:

$$h_{\alpha\beta}^{(\text{comp})\pm} = \left( \left. \frac{\partial g_{\alpha\beta}^{\text{Kerr}}(M, J)}{\partial M} \right|_{J=\text{const.}} \right) \delta M^{\pm} + \left( \left. \frac{\partial g_{\alpha\beta}^{\text{Kerr}}(M, J)}{\partial J} \right|_{M=\text{const.}} \right) \delta J^{\pm}.$$

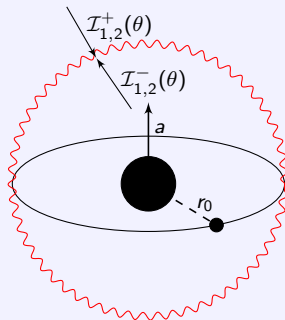
- Impose continuity for  $[\mathcal{I}_{1,2}^{(\text{rec})}](\theta) + [\mathcal{I}_{1,2}^{(\text{comp})}](\theta)$  across  $\mathcal{S}$  off-the-particle.
- Solve for the missing amplitudes  $[\delta M]$  and  $[\delta J]$ .

# Example: Circular equatorial orbits in Schwarzschild

The jump of the  $m = 0$  part of the reconstructed solution for the two invariants in Schwarzschild is

$$[\mathcal{I}_1^{(\text{rec})}](\theta) = \sum_{\ell=2}^{\infty} \left[ \frac{8m\mathcal{E}\pi}{3Mf_0^2} Y_{\ell}(\theta) \bar{Y}_{\ell}(\theta_0) + \frac{4m\mathcal{E}\pi(r_0 - M)}{3Mr_0 f_0^3} Y_{\ell}(\theta) \bar{Y}_{\ell}''(\theta_0) \right],$$

$$[\mathcal{I}_2^{(\text{rec})}](\theta) = \sum_{\ell=2}^{\infty} \frac{4\pi m\mathcal{L}}{r_0^4} Y_{\ell}(\theta) \bar{Y}_{\ell}'(\theta_0),$$





which we sum analytically as distributions by adding and subtracting the missing  $\ell = 0, 1$  pieces:

$$[\mathcal{I}_1^{(\text{rec})}](\theta) = \frac{8m\mathcal{E}\pi}{3Mf_0^2}\delta(\theta - \theta_0) - \frac{4m\mathcal{E}\pi(r_0 - M)}{3Mr_0f_0^3}\delta'(\theta - \theta_0) - \frac{2m\mathcal{E}}{3Mf_0^2},$$

$$[\mathcal{I}_2^{(\text{rec})}](\theta) = -\frac{4\pi m\mathcal{L}}{r_0^4}\delta'(\theta - \theta_0) + \frac{3m\mathcal{L}}{r_0^4}\cos\theta.$$

From the completion pieces we obtain a piece of the invariants:

$$[\mathcal{I}_1^{(\text{comp})}] = \frac{2[\delta M]}{3Mf_0^2}, \quad \text{and} \quad [\mathcal{I}_2^{(\text{comp})}](\theta) = -\frac{3[\delta J]\cos\theta}{r_0^4}.$$

We now impose the regularity at  $\theta \neq \pi/2$  to obtain

$$[\mathcal{I}_1](\theta \neq \pi/2) = -\frac{2m\mathcal{E}}{3Mf_0^2} + \frac{2[\delta M]}{3Mf_0^2} = 0,$$

$$[\mathcal{I}_2](\theta \neq \pi/2) = \frac{3m\mathcal{L}}{r_0^4}\cos\theta - \frac{3[\delta J]}{r_0^4}\cos\theta = 0.$$

We obtain  $[\delta M] \equiv m\mathcal{E}$ , and  $[\delta J] \equiv m\mathcal{L}$ . The jump of the mass and angular momentum perturbations in the invariants agrees with the specific energy and angular momentum of the particle.

# Second example: Circular equatorial orbits in Kerr

The reconstructed piece (with  $m = 0$  at  $\theta \neq \pi/2$ ) gives

$$\begin{aligned} \left[ \mathcal{I}_1^{(\text{rec})} \right] (\theta) = & -\frac{2i\mu\Sigma_0}{3Mr_0^3\Delta_0^2} \left\{ 2a^6 M\Omega^2 + ar_0^3(4Mr_0 - 6M^2 - 3r_0^2)\Omega - 3a^3 r_0(r_0^2\dot{f}_0 - 2M^2)\Omega - 4a^5 M\Omega \right. \\ & - a^4 \left[ 3M^2 r_0\Omega^2 + 2r_0^3\Omega^2 + M(r_0^2\Omega^2 - 2) \right] + a^2 r_0 \left[ 3M^2(r_0^2\Omega^2 - 1) \right. \\ & \left. \left. + r_0^2(5 + r_0^2\Omega^2) + Mr_0(2r_0^2\Omega^2 - 5) \right] + r_0^3 \left[ 3M^2 + r_0^2 + r_0^4\Omega^2 - 3Mr_0(1 + r_0^2\Omega^2) \right] \right\}, \\ \left[ \mathcal{I}_2^{(\text{rec})} \right] (\theta) = & \frac{i\mu\Sigma_0}{3aMr_0^3} \left\{ 3 \left[ 2a^4 M\Omega + 2r_0^5\Omega + 2a^2 r_0(r_0^2\dot{f}_0 - 3M^2)\Omega - a^5 M\Omega^2 + a^3(r_0^3\Omega^2 - M + 3M^2 r_0\Omega^2 + Mr_0^2\Omega^2) \right. \right. \\ & \left. \left. - ar_0(r_0^4\Omega^2 - 3M^2 - 3Mr_0 + 3r_0^2 + 3Mr_0^3\Omega^2) \right] + a \left[ a^4 M\Omega^2 - 2a^3 M\Omega + 2aMr_0(2r_0 - 3M)\Omega \right. \right. \\ & \left. \left. + a^2(M + 3M^2 r_0\Omega^2 - Mr_0^2\Omega^2 - r_0^3\Omega^2) + r_0 \left( 3M^2 + r_0^2 + r_0^4\Omega^2 - 3Mr_0(1 + r_0^2\Omega^2) \right) \right] \cos(2\theta) \right\} \csc^2 \theta. \end{aligned}$$

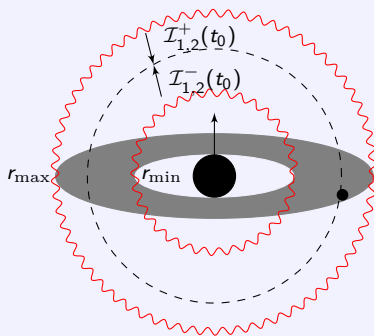
The completion part of the invariants in the Kerr case is just

$$\begin{aligned} \left[ \mathcal{I}_1^{(\text{comp})} \right] (\theta) &= \frac{\Sigma}{3M\Delta^2} \left\{ (5a^2 + r^2)[\delta M] - 3a[\delta J] \right\}, \\ \left[ \mathcal{I}_2^{(\text{comp})} \right] (\theta) &= -\frac{\Sigma}{6aM} \left\{ (a \cos(2\theta) - 9a)[\delta M] + 6[\delta J] \right\} \csc^2 \theta. \end{aligned}$$

With the regularity conditions we determine that the amplitudes are  $[\delta M] = m\mathcal{E}$ , and  $[\delta J] = m\mathcal{L}$ .

# Completion for eccentric orbits

- Obtain the jump on the reconstructed part of the two auxiliary invariants  $[\mathcal{I}_{1,2}](r_0(t_0), \theta)$  for a momentary ring in the libration region  $r_{\min} < r_0(t_0) < r_{\max}$ .
- Sum all the  $\ell$  modes for the partial ring.
- The completion pieces are constructed in the same way as in the circular case.
- Impose continuity for  $[\mathcal{I}_{1,2}^{(\text{rec})}](t_0, \theta) + [\mathcal{I}_{1,2}^{(\text{comp})}](t_0, \theta)$ , at each  $t_0$ .
- Solve for the jump of missing amplitudes  $[\delta M](t_0)$  and  $[\delta J](t_0)$  of the completion pieces.
- Integrate over  $t_0$  covering an orbital period.



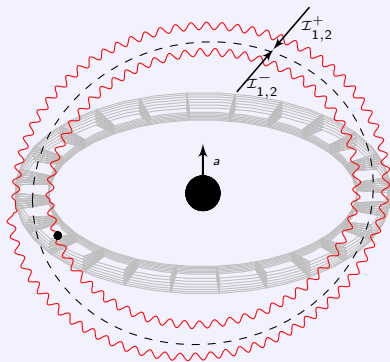
# Results

After analytically performing the integrals we find that the missing amplitudes  $[\delta\tilde{M}]$  and  $[\delta\tilde{J}]$  correspond to the **specific energy  $\mathcal{E}$**  and **angular momentum  $\mathcal{L}$  of the particle**, just like in the circular-orbit case:

$$[\delta\tilde{M}] = m\mathcal{E}, \quad \text{and} \quad [\delta\tilde{J}] = m\mathcal{L}.$$

# More general orbits in Kerr

- At each point in the region  $r_{\min} \leq r_0 \leq r_{\max}$ ,  $\theta_{\min} \leq \theta_0 \leq \theta_{\max}$  we find the jump on the gauge invariants corresponding to a infinitesimal ring.
- We impose continuity (at the level of the Green's function) for each ring and integrate throughout the source region.



# Summary and future work

- We have identified three types of RG gauges. We have two methods to calculate the SF starting from RG MP, using new versions of the mode-sum method.
- To include the completion piece, we derived new gauge invariant quantities. By imposing continuity off the particle we extract the amplitude of the jump in the mass and angular momentum perturbations.
- We have extended our analysis to obtain the completion for equatorial-eccentric orbits. This requires to evaluate a complicated integral in the libration region. Inclined orbits will require extra considerations.
- Our completion piece is required to calculate the GSF for eccentric orbits in Kerr [van de Meent and Shah, (2015)].