

Time-domain metric reconstruction for self-force applications

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Structure of talk

- Why another method for self-force calculations?
- Formulation of method in Kerr
- Implementation for circular orbits in Schwarzschild
- Going forward to Kerr

Why another method for self-force calculations?

Method for obtaining $h_{\alpha\beta}$	frequency domain	time domain
direct: $E_{\mu\nu}(h_{\alpha\beta}) = T_{\alpha\beta}$	✓	✓
metric reconstruction: $\text{Teuk}(\Psi) = \mathcal{T}, \quad \Psi \Rightarrow h_{\alpha\beta}$	✓	this talk

Reconstruction method is computationally cheaper, but so far only formulated and implemented via frequency-domain decomposition.

Reasons to want to have a time-domain method for metric reconstruction:

- Highly eccentric or unbound (incl. high-energy) orbits
- Orbital evolution under the self-force
- Test of f-domain calculations

Review of metric reconstruction

vacuum spacetime

- Let $h_{\alpha\beta}$ be a vacuum perturbation of the Kerr metric, with ψ_4, ψ_0 associated Weyl scalars.
- Find a “Hertz potential” Φ that satisfies

$$\boxed{\text{Teuk}_s \Phi = 0} \quad \text{and} \quad \boxed{D_s^4 \Phi = \psi_{-s} \quad \text{or} \quad \mathcal{L}_s^4 \Phi = \psi_s}.$$

- Then the original perturbation can be reconstructed via

$$h_{\alpha\beta} = \text{Re} \left(e_{\mathbf{a}(\alpha} e_{\mathbf{b}\beta)} \mathcal{D}^{\mathbf{ab}} \Phi \right) + h_{\alpha\beta}^{\text{gauge}} + \delta M_{\alpha\beta} + \delta J_{\alpha\beta}.$$

Review of metric reconstruction

vacuum spacetime

Two variants:

- *Ingoing Radiation Gauge (IRG)*

$$\boxed{\text{Teuk}_{-2}\Phi = 0} \quad \text{and} \quad \boxed{D^4\Phi = \psi_0 \quad \text{or} \quad \mathcal{L}^4\Phi = \psi_4}.$$

$$h_{\alpha\beta}^{\text{rec}} = \text{Re} \left(e_{\mathbf{a}(\alpha} e_{\mathbf{b}\beta)} D^{\mathbf{ab}} \Phi \right), \quad \ell^\alpha h_{\alpha\beta}^{\text{rec}} = 0.$$

- *Ongoing Radiation Gauge (ORG)*

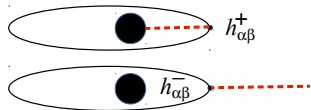
$$\boxed{\text{Teuk}_{+2}\Phi = 0} \quad \text{and} \quad \boxed{\tilde{D}^4\Phi = \psi_4 \quad \text{or} \quad \tilde{\mathcal{L}}^4\Phi = \psi_0}.$$

$$h_{\alpha\beta}^{\text{rec}} = \text{Re} \left(e_{\mathbf{a}(\alpha} e_{\mathbf{b}\beta)} \tilde{D}^{\mathbf{ab}} \Phi \right), \quad n^\alpha h_{\alpha\beta}^{\text{rec}} = 0.$$

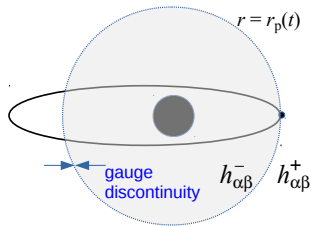
Review of metric reconstruction

point-particle source

- In presence of matter sources, the procedure fails to yield a valid solution even in vacuum away from sources [LB & Ori 2001; Price & Whiting 2007]
- Reconstruction for bound orbits, with string-like gauge singularities [Ori 2003]
- “No-string” reconstruction with gauge discontinuity on a sphere [Keidl, Friedman et al 2007-12; vdMeent 2015–]
- Self-force from a reconstructed metric [Pound, Merlin & LB (2014)]
- Determination of $\delta M_{\alpha\beta} + \delta J_{\alpha\beta}$ [Merlin et al 2016; vdMeent 2017]



Ori's “half-string” solutions



Friedman's “no-string” solution

New method: metric reconstruction in the t domain

Basic idea: Obtain Φ^\pm by solving the appropriate Teukolsky equation in 1+1D, with suitable boundary conditions at infinity and on the horizon, and with suitable junction conditions along the particle's worldline. From it obtain the "no-string" perturbation via

$$h_{\alpha\beta}^\pm = \text{Re} \left(e_{\mathbf{a}(\alpha} e_{\mathbf{b}\beta)} \mathcal{D}^{\mathbf{ab}} \Phi^\pm \right) + \delta M_{\alpha\beta}^\pm + \delta J_{\alpha\beta}^\pm$$

or an ℓ -by- ℓ application thereof.

Need 3 things:

- 1 1+1D decomposition of the Teukolsky equation
- 2 Boundary conditions for 1+1D solutions
- 3 Junction conditions along the particle's worldline

1+1D decomposition of the Teukolsky equation

Decompose into spin-weighted *spherical* harmonics (even in Kerr):

$$\Phi_s^\pm = (r\Delta^s)^{-1} \sum_{\ell=2}^{\infty} \sum_{m=-\ell}^{\ell} \phi_{s\ell m}^\pm(t, r) {}_sY_{\ell m}(\theta, \tilde{\varphi}).$$

Substitution into Teukolsky's master equation (in vacuum) gives

$$\sum_{\ell m} {}_sY_{\ell m}(\theta, \tilde{\varphi}) \left[\tilde{\square} \phi_{s\ell m}^\pm - a^2 \sin^2 \theta (\phi_{s\ell m}^\pm)_{,tt} + 2ias \cos \theta (\phi_{s\ell m}^\pm)_{,t} \right] = 0.$$

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After reexpanding in ${}_sY_{\ell m}$ we get, for each ℓ ,

$$\phi_{,uv}^\ell + U(r)\phi_{,u}^\ell + V(r)\phi_{,v}^\ell + W(r)\phi^\ell + K(r) \left[a^2 C_0^\ell \phi_{,tt}^\ell + \mathcal{I}(\phi^{\ell\pm 1}, \phi^{\ell\pm 2}) \right] = 0,$$

with the ℓ -mode coupling term

$$\mathcal{I} = -a^2 \left(C_{++}^\ell \phi_{,tt}^{\ell+2} + C_+^\ell \phi_{,tt}^{\ell+1} + C_-^\ell \phi_{,tt}^{\ell-1} + C_{--}^\ell \phi_{,tt}^{\ell-2} \right) + 2ias \left(c_+^\ell \phi_{,t}^{\ell+1} + c_-^\ell \phi_{,t}^{\ell-1} \right).$$

Boundary conditions for the fields $\phi_{slm}^{\pm}(v, u)$

behavior at $r \gg M$

- time-dependent modes:

$$\text{physical: } \phi^+ \sim e^{-i\omega u} \quad \Leftrightarrow \text{bounded,}$$

$$\text{nonphysical: } \phi^+ \sim r^{2s} e^{-i\omega v} \quad \Leftrightarrow \text{blows up for ORG.}$$

- static modes:

$$\text{physical: } \phi^+ \sim r^{-\ell+s} \quad \Leftrightarrow \text{bounded,}$$

$$\text{nonphysical: } \phi^+ \sim r^{\ell+s+1} \quad \Leftrightarrow \text{blows up.}$$

In an ORG reconstruction, all nonphysical modes blow up at infinity

Boundary conditions for the fields $\phi_{slm}^{\pm}(v, u)$

behavior near the event horizon

- time-dependent modes:

$$\text{physical: } \phi^{-} \sim e^{-i\omega v}, \quad \hookrightarrow \text{bounded,}$$

$$\text{nonphysical: } \phi^{-} \sim \Delta^s e^{-i\omega u} e^{-2im\Omega_H r_*} \hookrightarrow \text{blows up for IRG.}$$

- static modes:

$$\text{physical: } \phi^{-} \sim \text{const} \quad \hookrightarrow \text{bounded,}$$

$$\text{nonphysical: } \phi^{-} \sim \Delta^s \quad \hookrightarrow \text{blows up for IRG.}$$

In an IRG reconstruction, all nonphysical modes blow up on the horizon

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In an IRG reconstruction, all nonphysical modes blow up on the horizon

This motivates a “mixed-gauge” approach where $h_{\alpha\beta}^{+}$ is reconstructed in ORG while $h_{\alpha\beta}^{-}$ is reconstructed in IRG. (We won't implement it here.)

Junction conditions on the particle's worldline

- Radial inversion relations decompose into ${}_s Y_{\ell m}$ modes without any mode coupling, even in Kerr! E.g., for IRG:

$$8r\Delta^2\mathcal{D}_l^4\left(\Delta^2\bar{\phi}_{\ell m}^\pm/r\right) = (-1)^m\psi_{0,\ell,-m}^\pm,$$

where

$$\mathcal{D}_l := \Delta^{-1}\left((r^2 + a^2)\partial_v - ima\right).$$

- Angular relations don't have this nice feature, except for $a = 0$ where they become very simple (essentially $\phi_{,t}^\pm \sim \psi_0^\pm$). See [Lousto & Whiting 2002](#).

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- The jumps across the worldline in ϕ and its v derivatives thus obey

$$\sum_{n=0}^4 F_n(v) [\partial_v^n \phi_{\ell m}] = [\bar{\psi}_{0,\ell,-m}],$$

where $[\bar{\psi}_{0,\ell,-m}]$ can be inferred directly from the source of the Teukolsky equation.

Junction conditions on the particle's worldline

- $[\partial_v^{n \geq 2} \phi]$ can be expressed in terms of $[\phi]$ and $[\phi, v]$ and their τ derivatives along the orbit, via a repeated application of the vacuum Teukolsky equation.
- We obtain a coupled set of ODEs for $[\phi]$ and $[\phi, v]$:

$$\sum_{n=0}^3 \left(a_n(\tau) \frac{d^n [\phi]}{d\tau^n} + b_n(\tau) \frac{d^n [\phi, v]}{d\tau^n} \right) + \mathcal{I} \text{ terms} = [\bar{\psi}_{2, \ell, -m}]$$
$$\sum_{n=0}^4 \left(c_n(\tau) \frac{d^n [\phi]}{d\tau^n} + d_n(\tau) \frac{d^n [\phi, v]}{d\tau^n} \right) + \mathcal{I} \text{ terms} = [(\bar{\psi}_{2, \ell, -m}), v]$$

- These should be solved with suitable initial or periodicity conditions.
- Once $[\phi]$ and $[\phi, v]$ are known, the jumps in all higher (u , v or mixed) derivatives are easily obtained using $[\dot{\phi}] = \dot{u}_p [\phi, u] + \dot{v}_p [\phi, v]$ and the Teukolsky equation.

Implementation for circular orbits in Schwarzschild

1. Field equation and junction conditions

- Specializing to Schwarzschild and IRG:

$$\phi_{,uv}^{\ell} + U(r)\phi_{,u}^{\ell} + V(r)\phi_{,v}^{\ell} + W(r)\phi^{\ell} = 0,$$

with

$$U(r) = \frac{2M}{r^2}, \quad V(r) = -\frac{2f}{r}, \quad W(r) = \frac{f}{4} \left(\frac{(\ell+2)(\ell-1)}{r^2} - \frac{2M}{r^3} \right).$$

- Specializing further to circular orbits, jump equations become **algebraic**:

$$\begin{aligned} a_{\Sigma}[\phi] + b_{\Sigma}[\phi, v] &= [\bar{\psi}_{2,\ell,-m}], \\ c_{\Sigma}[\phi] + d_{\Sigma}[\phi, v] &= [\partial_v \bar{\psi}_{2,\ell,-m}], \end{aligned}$$

so can be solved analytically.

Implementation for circular orbits in Schwarzschild

1. Field equation and junction conditions

- Even in this simple case, the jumps are complicated...

$$\begin{aligned} [\phi_{\ell m}^{\text{IRG}}] &= \tilde{\Delta}^{-1} (d_{\Sigma} [\bar{\psi}_{2,\ell,-m}] - b_{\Sigma} [\partial_{\nu} \bar{\psi}_{2,\ell,-m}]), \\ [\partial_{\nu} \phi_{\ell m}^{\text{IRG}}] &= \tilde{\Delta}^{-1} (a_{\Sigma} [\partial_{\nu} \bar{\psi}_{2,\ell,-m}] - c_{\Sigma} [\bar{\psi}_{2,\ell,-m}]), \end{aligned}$$

where

$$\tilde{\Delta} = \frac{1}{4} f_0^4 r_0^8 [\lambda^2 (\lambda + 2)^2 + (12mM\Omega)^2],$$

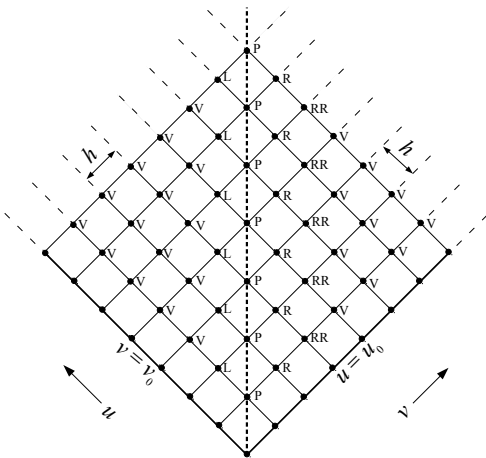
and, e.g., $c_{\Sigma} = \sum_{n=0}^4 (-im\Omega)^n c_n$ with

$$\begin{aligned} c_0 &= r_0^3 f_0^2 (1-y) \lambda (\lambda + 2), \\ c_1 &= f_0 r_0^4 [\lambda (\lambda + 5) - 2(\lambda^2 + 2\lambda - 6)y - 2(4\lambda + 17)y^2 + 12y^3], \\ c_2 &= 2r_0^5 [3\lambda + (15 - 7\lambda)y + 2(\lambda - 23)y^2 + 24y^3], \\ c_3 &= 2f_0 r_0^6 (\lambda + 22y), \\ c_4 &= 16Mr_0^6. \end{aligned}$$

Here $\lambda := (\ell + 2)(\ell - 1)$, $f_0 := 1 - 2M/r_0$, $y := M/r_0$.

Implementation for circular orbits in Schwarzschild

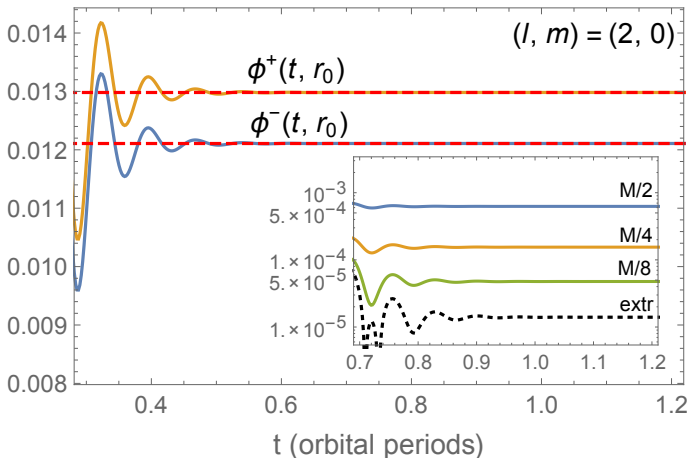
2. Numerical method



- Teukolsky equation discretized on a 1+1 mesh in uv coords
- Finite-difference scheme is 2nd-order convergent [local error is $O(h^4)$]
- FD scheme at near-particle points incorporates jump information
- Start with zero initial data on $u = u_0$ and $v = v_0$; wait for junk radiation to dissipate.

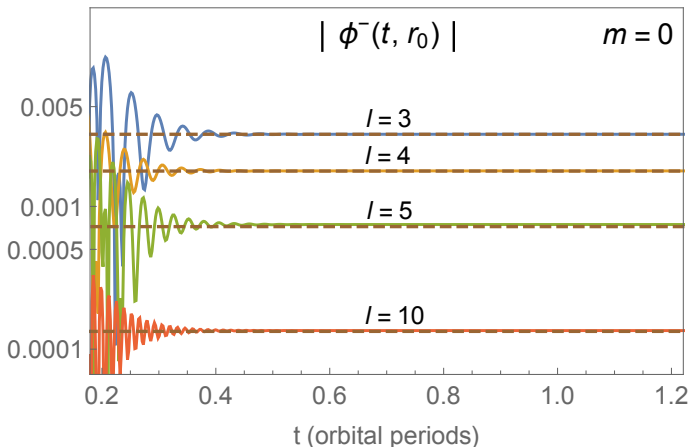
Implementation for circular orbits in Schwarzschild

3. Results and tests



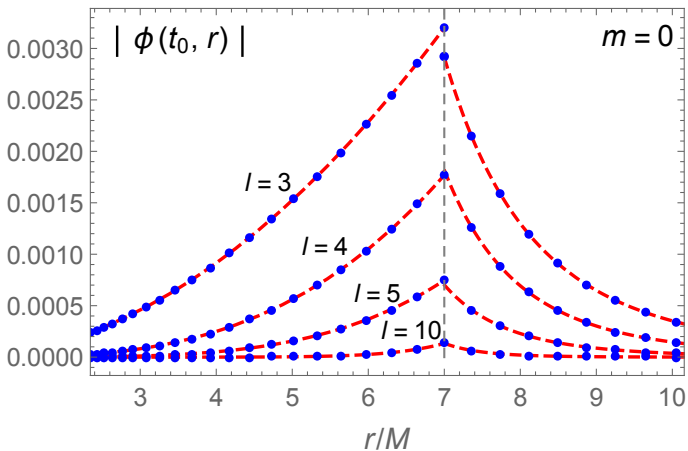
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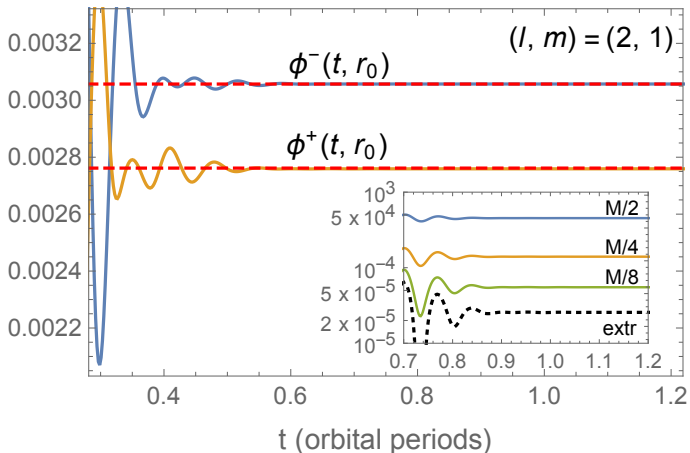
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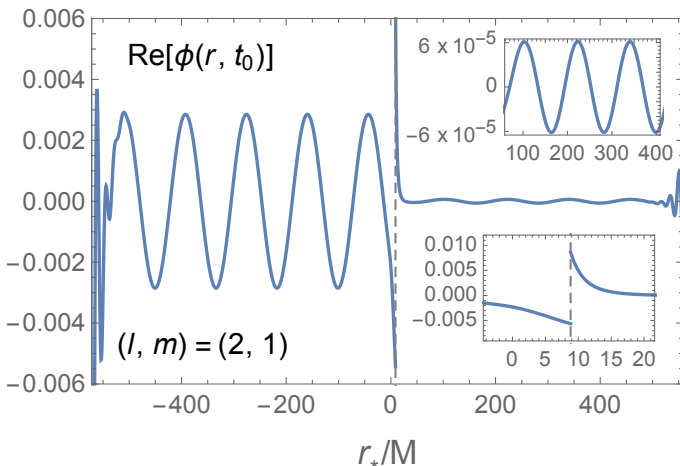
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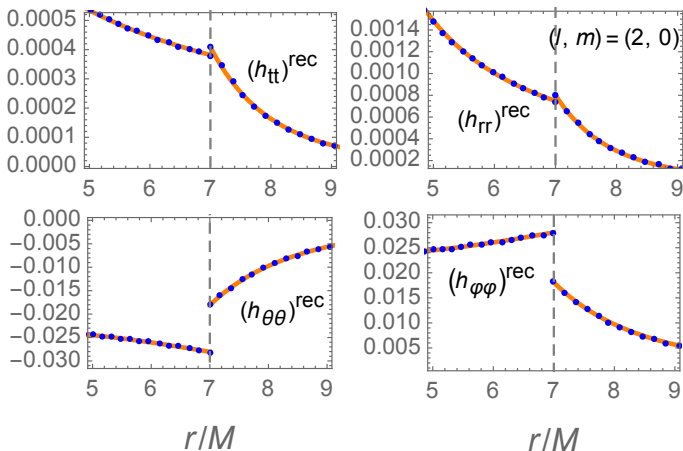
Implementation for circular orbits in Schwarzschild

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Implementation for circular orbits in Schwarzschild

3. Results and tests



Going forward (1/4)

Extension to eccentric or unbound orbits

Main challenge is to solve 4th-order ODEs that determine the jumps of ϕ across the orbit.

- For periodic orbits: can decompose into harmonics of orbital frequencies
- For evolving bound orbits: can impose approximate periodic initial conditions at initial point
- For scatter orbits: can impose initial conditions at infinity; need to find a way to eliminate nonphysical unstable solutions.

Going forward (2/4)

Extension to Kerr

Need to incorporate mode coupling in both field & jump equations.

- Preliminary results show that's easily done in vacuum.
- In particle case, may need to deal with coupling at the level of the mode-sum regularization formula

Going forward (3/4)

Exploring the mixed-gauge approach

- Our IRG implementation seems to “automatically” select the physical (retarded) solution. Need to understand why and whether that’s always guaranteed.
- A mixed-gauge approach would have the advantage that a globally bounded solution is automatically the physical one.
- To set up a the mixed-gauge calculation, need (1) a formulation of the self-force from a mixed-gauge perturbation (easy), and (2) a method for determining the jump condition (hard!)

Going forward (4/4)

Improving the numerical method

Method could be directly implemented on existing platforms (e.g., Diener's or Thornburg's), to include

- Mesh refinement
- Hyperboloidal slicing
- Compactification
- Parallelization

Method is also amenable to finite-element discretization a la Canizares & Sopena.