

# Towards a 2nd Law for Lovelock Theory

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**“Towards a second law for Lovelock theories”**

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# Introduction and Motivation

- Black Holes are "Theoretical Laboratories" for understanding quantum gravity.
- String theory  $\Rightarrow$  Despite being successful in some aspects  $\Rightarrow$  unresolved issues remaining.
- It is, therefore in principle useful to use "*general principles*" to constrain admissible low energy theories.
- The second law of thermodynamics is one such principle that we can test on low energy solutions of gravity.

# Introduction and Motivation

- In Einstein's theory of gravity BH's are thermodynamic objects :
  - They have energy, temperature and entropy.
  - The entropy is given by the area of BH horizon.
  - The entropy satisfies a 1st law of thermodynamics.
  - In dynamical situations the horizon area always increases  $\Rightarrow$  The 2nd law
- The proof of 1st and 2nd law in general relativity requires classical EOM's + certain conditions on the matter stress tensor (Null Energy Conditions).

# Introduction and Motivation

- The Einstein-Hilbert action can not be the complete story : The low energy limit of any UV complete quantum theory of gravity generate higher derivative corrections to the leading two derivative action :

$$I = \int d^d x \sqrt{-g} [R + \mathcal{L}_{matter} + \mathcal{L}_{HD}]$$

- Precise form of  $\mathcal{L}_{HD}$  depends on particular nature of UV completion.
- Additional corrections in the action  $\Rightarrow$  modifications in the EOM's.
- However, BH solutions will continue to exist.
- However, since the EOM's have changed, the earlier proofs of 1st and 2nd law (for two derivative theories of gravity) will no longer be valid.

# Introduction and Motivation

- Wald argued that for stationary black hole solutions of higher derivative gravity theories, the entropy is a Noether charge associated with time translations along the horizon generating Killing field.
  - Wald entropy was constructed to explicitly satisfy the first law of thermodynamics, which being an equilibrium statement, can be understood in the stationary solution.
- The 2nd law involves dynamics and thus is a statement beyond equilibrium.
- However, we do not have a general proof that "Wald Entropy" satisfies 2nd law for any higher derivative theories of gravity.
- An open question since then has been whether there is a notion of second law of black hole mechanics in higher derivative gravity.

# Introduction and Motivation

- The precise statement of a 2nd law :

If perturbing one equilibrium configuration (say  $eq_1$ ) one can reach another equilibrium configuration (say  $eq_2$ ) in course of time evolution, then  
the total entropy of  $eq_2 \geq$  the total entropy of  $eq_1$

- How does this translate to gravity :
  - Equilibrium configuration  $\Rightarrow$  Metric with Killing horizon
  - Equilibrium entropy  $\Rightarrow$  Wald entropy on Killing horizon
  - Suppose,  $metric_1$  and  $metric_2$  both have Killing horizons and dynamical perturbations around  $metric_1$  eventually evolves to  $metric_2$ .
  - Second law :  $Wald\ entropy_{metric_2} \geq Wald\ entropy_{metric_1}$
- We will actually consider a local version of second law, a stronger statement.

# Introduction and Motivation

- Our eventual target:

To investigate the 2nd law for any dynamical BH solution in  
arbitrary higher derivative theories of gravity  
or  
To find a concrete counter example

- What we have attempted in this work is an initial small step towards this final goal.
- We have examined whether 2nd law holds for dynamical black holes in Lovelock theories of gravity, one special case of higher derivative theories of gravity.
- More precisely, we have extended the Wald entropy (a notion of equilibrium entropy) to the dynamical setting satisfying a local version of 2nd law for Lovelock theories.



# Strategy of the Proof

- Naively, 2nd law is a non-local statement, comparing only the initial and final configuration :  $S_{Wald}^{final} - S_{Wald}^{initial} \geq 0$

- To prove this we explicitly construct a function of the system variables, the "entropy function":  $S_{total} = S_{Wald} + S_{cor}$

- The conditions on the entropy function  $S_{total}$  will be
  - it reduces to the familiar notion of entropy  $S_{Wald}$ , whenever the geometry has a Killing horizon

$$(1) S_{total} \stackrel{equilibrium}{=} S_{Wald}$$

- and it is monotonically increasing under time evolution as long as the time evolution is governed by the higher derivative EOM

$$(2) \partial_\nu S_{total} \geq 0 \Rightarrow \text{for all } \nu$$

# Strategy of the Proof

- Existence of such a  $S_{\text{total}}$  is enough to prove the 2nd law

$$(1) S_{\text{total}}^{\text{equilibrium}} = S_{\text{Wald}} \quad \text{and} \quad (2) \partial_v S_{\text{total}} \geq 0$$
$$\Rightarrow S_{\text{Wald}}^{\text{final}} - S_{\text{Wald}}^{\text{initial}} = \int_{\text{initial}}^{\text{final}} \partial_v S_{\text{total}} dv \geq 0$$

- The steps of the proof

- Define entropy density :  $S_{\text{total}} = \int_{\Sigma_v} d^{d-2}x \sqrt{h} \rho_{\text{total}}$
- Define  $\Theta$  :  $\partial_v S_{\text{total}} = \int_{\Sigma_v} d^{d-2}x \sqrt{h} \Theta$
- We aim to show  $\partial_v \Theta \leq 0$
- We assume :  $v \rightarrow \infty, \Theta \rightarrow 0$ .
- This implies  $\Theta \geq 0$  for every finite  $v$ .
- Finally we obtain that :  $\partial_v S_{\text{total}} \geq 0$

# Strategy and Final result

- Note that our method is a bit indirect : in the sense
  - We don't directly try to prove  $\partial_v S_{\text{total}} \geq 0$ ,
  - Rather a stronger and local version :  $\Theta \geq 0$
  - Also there is no claim for uniqueness.

## Our Result :

- For Lovelock theory we have been able to construct an  $S_{\text{total}}$  that works in any time evolution, maintaining spherical symmetry
- Outside spherical symmetry, our construction works provided a very particular total derivative term is always non-negative in the course of evolution.

# Basic Set Up

- Higher derivative terms in the action come with dimensionfull coefficients  $\Rightarrow \alpha_2 \sim \ell_s^2 \Rightarrow$  length dimension = 2

$$I = \int d^d x \sqrt{-g} \left[ R + \alpha_2 \left( R^2 - 4R_{\mu\nu}R^{\mu\nu} + R_{\mu\nu\alpha\beta}R^{\mu\nu\alpha\beta} \right) \right]$$

- Our initial equilibrium configuration = Stationary BH solution in this theory.
- Perturbing away from equilibrium are parametrized by
  - (a) the amplitude of the departure from equilibrium :  $\alpha$ ,
  - (b) the characteristic frequency :  $\omega$
- An entropy function should, at the very least, carry information about these three parameters :  $\alpha_2, \alpha, \omega$

# Basic Set Up : Assumptions

- The following summarizes what is known to date:
  - (a) Wald entropy can be constructed for arbitrary  $\alpha_2$  with  $\alpha = 0$  (Iyer-Wald),
  - (b) Entropy functions are constructed for  $f(R)$  theories for finite range of  $\alpha_2$  but arbitrary  $\alpha, \omega$ . (Jacobson-Kang-Myers)
  - (c) Some developments are there for  $f(\text{Lovelock})$  and general four derivative theories but for  $\alpha \ll 1$ . (Padmanavan  $\dots$ , Wall-Sarkar  $\dots$ )
- We aim to construct an entropy function in higher derivative gravity, perturbatively in the couplings  $\alpha_2 \ll 1$ , and the frequency  $\omega l_s \ll 1$ , but valid for arbitrary amplitudes  $\alpha$
- We work perturbatively in the higher derivative terms,
  - (a) The corrections to Einstein-Hilbert theory are treated in a gradient expansion,
  - (b) the effective small dimensionless parameter :  $\omega l_s \sim \sqrt{\alpha_2} \partial_\nu$
  - (c) no assumption about the amplitude.

# Basic Set Up : Assumptions

- At horizon higher derivative corrections are suppressed compared to the Einstein-Hilbert term  
⇒ Around horizon, curvature-scales are large compared to fluctuations of the BH horizon  $\sim$  determined by  $\alpha_2 \sim \ell_s$
- A classical description of gravity is valid, i.e., no loop correction etc. enter our discussion
- Despite the 2-derivative theory dominates on the horizon scale, why do we need to modify the entropy for 2nd law ?  
⇒ Because :

Although the leading area contribution is large, it is possible under evolution, the area variation is anomalously small and overwhelmed by the higher derivative  $\mathcal{O}(\omega \ell_s)$  contributions, spoiling the monotonicity of the entropy.

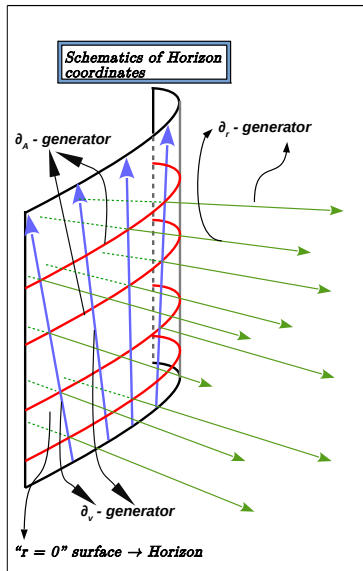
# Basic Set Up : Choice of Coordinates

- Around horizon we can choose a coordinate system  $\{x^\mu\}$  such that the effect of higher derivative corrections are handled in a derivative expansion  $\Rightarrow \omega l_s \sim \sqrt{\alpha_2} \partial_\mu \ll 1$
- The geometry must have a horizon  $\Rightarrow$  a null hypersurface.
- On the horizon  $\mathcal{H}$ :
  - (a) define coordinate  $v \Rightarrow$  affine parameter along the null generators  $\partial_v$ ,
  - (b) along constant  $v$  slices,  $\Sigma_v$  : define the spatial coordinates  $x_A$  .  
 $\Rightarrow \{v, x^A\}$  are coordinates on  $\mathcal{H}$  and  $\{x^A\}$  are coordinates on  $\Sigma_v$
- Away from  $\mathcal{H}$  : define coordinate  $r \Rightarrow$  affinely parametrized along null geodesics  $\partial_r$  piercing through the horizon at angle

$$(\partial_v, \partial_r) \Big|_{\mathcal{H}} = 1, \quad (\partial_r, \partial_A) \Big|_{\mathcal{H}} = 0$$

- Choose the origin of the  $r$  coordinate such that horizon is at  $r = 0$ .

# Basic Set Up : Choice of Coordinates





## Basic Set Up : Choice of Metric

- In our coordinates the metric will take the following form

$$ds^2 = 2dvdr - f(r, v, x^A)dv^2 + 2k_A(r, v, x^A)dvdx^A + h_{AB}(r, v, x^A)dx^A dx^B$$

such that  $f(r, v, x^A)\Big|_{\mathcal{H}} = k_A(r, v, x^A)\Big|_{\mathcal{H}} = \partial_r k_A(r, v, x^A)\Big|_{\mathcal{H}} = 0$

- Note that any metric with a horizon could be expressed in the above form
- Our construction will be in terms of explicit derivatives of  $f(r, v, x^A)$ ,  $k_A(r, v, x^A)$  and  $h_{AB}(r, v, x^A)$
- Following notation would be useful

$$\mathcal{K}_{AB} = \frac{1}{2}\partial_v h_{AB}\Big|_{\mathcal{H}}, \quad \bar{\mathcal{K}}_{AB} = \frac{1}{2}\partial_r h_{AB}\Big|_{\mathcal{H}}$$

- $h_{AB}(r = 0, v, x^A)$  is the induced metric on  $\mathcal{H}$

# Gauss-Bonnet Theory

- We consider Gauss-Bonnet theory :

$$I = \int d^d x \sqrt{-g} \left[ R + \alpha_2 \left( R^2 - 4R_{\mu\nu} R^{\mu\nu} + R_{\mu\nu\alpha\beta} R^{\mu\nu\alpha\beta} \right) \right]$$

- We will add correction to  $S_{\text{Wald}}$ , and  $\rho$  and  $\Theta$  can be written as

$$\begin{aligned} S_{\text{total}} &= S_{\text{Wald}} + S_{\text{cor}} = \int_{\Sigma_V} d^{d-2}x \sqrt{h} [\rho_{\text{eq}} + \rho_{\text{cor}}] \\ \partial_V S_{\text{total}} &= \int_{\Sigma_V} d^{d-2}x \sqrt{h} [\Theta_{\text{eq}} + \Theta_{\text{cor}}] \\ \Theta_{\text{eq}} &= \frac{1}{\sqrt{h}} \partial_V (\sqrt{h} \rho_{\text{eq}}), \quad \Theta_{\text{cor}} = \frac{1}{\sqrt{h}} \partial_V (\sqrt{h} \rho_{\text{cor}}) \end{aligned}$$

- Finally we want to show :  $\partial_V [\Theta_{\text{eq}} + \Theta_{\text{cor}}] \leq 0$
- Condition on the correction term :  $S_{\text{cor}} \stackrel{\text{equilibrium}}{=} 0$

# Gauss-Bonnet Theory : Final Result

- We consider Gauss-Bonnet theory with Wald entropy

$$I = \int d^d x \sqrt{-g} \left[ R + \alpha_2 \left( R^2 - 4R_{\mu\nu} R^{\mu\nu} + R_{\mu\nu\alpha\beta} R^{\mu\nu\alpha\beta} \right) \right]$$

$$S_{\text{Wald}} = \int_{\Sigma_V} d^{d-2}x \sqrt{h} \underbrace{(1 + 2\alpha_2 \mathcal{R})}_{(=\rho_{\text{eq}})}$$

- We add the correction

$$S_{\text{cor}} = \int_{\Sigma_V} d^{d-2}x \sqrt{h} \rho_{\text{cor}}, \quad \rho_{\text{cor}} = \alpha_2^2 \sum_{n=0}^{\infty} \kappa_n \alpha_2^n \partial_V^n (h_{2(0)})_B^A \partial_V^n (h_{2(0)})_A^B$$

where  $(h_{2(0)})_B^A = \delta_{BB_1 B_2}^{AA_1 A_2} \mathcal{K}_{A_1}^{B_1} \overline{\mathcal{K}}_{A_2}^{B_2}$

- We showed that

$$\partial_V [\Theta_{\text{eq}} + \Theta_{\text{cor}}] \leq 0, \quad \Rightarrow \quad A_n = 2\kappa_n - \frac{\kappa_{n-1}^2}{A_{n-2}}$$

for  $n = -2, -1, 0, \dots$ , and  $\kappa_{-2} = -1/2$ ,  $\kappa_0 = -1$ ,  $\kappa_{-1} = -2$

# Gauss-Bonnet Theory : Details of proof

- We consider Gauss-Bonnet theory with Wald entropy

$$I = \int d^d x \sqrt{-g} \left[ R + \alpha_2 \left( R^2 - 4R_{\mu\nu}R^{\mu\nu} + R_{\mu\nu\alpha\beta}R^{\mu\nu\alpha\beta} \right) \right]$$

$$S_{\text{Wald}} = \int_{\Sigma_\nu} d^{d-2}x \sqrt{h} \underbrace{(1 + 2\alpha_2 \mathcal{R})}_{(=\rho_{eq})}, \text{ and } \Theta_{eq} = \frac{1}{\sqrt{h}} \partial_\nu (\sqrt{h} \rho_{eq})$$

- Next We compute

$$\partial_\nu \Theta_{eq} = \text{Term 1} + \text{Term 2} + \text{Term 3} + \text{Term 4} + \text{Term 5}$$

$$\text{Term 1} = -T_{\nu\nu}, \quad \text{Term 2} = -\mathcal{K}_{AB}\mathcal{K}^{AB}, \quad \text{Term 3} = \alpha^2 \mathcal{K}_B^A \mathcal{K}_{B'}^{A'} \mathcal{M}_{AA'}^{BB'}$$

$$\text{Term 4} = \alpha^2 \mathcal{K}_B^A \partial_\nu \left[ \delta_{AB_1 B_2}^{BA_1 A_2} \mathcal{K}_{A_1}^{B_1} \bar{\mathcal{K}}_{A_2}^{B_2} \right], \quad \text{Term 5} = \nabla_A \mathcal{Y}^A$$

- Note that  $\mathcal{M}_{AA'}^{BB'}$  is some specific four indexed and two derivative tensor but no  $\partial_\nu$ .

# Gauss-Bonnet Theory : Details of proof

$$\partial_\nu \Theta_{eq} = \text{Term 1} + \text{Term 2} + \text{Term 3} + \text{Term 4} + \text{Term 5}$$

$$\text{Term 1} = -T_{\nu\nu}, \quad \text{Term 2} = -\mathcal{K}_{AB}\mathcal{K}^{AB}, \quad \text{Term 3} = \alpha^2 \mathcal{K}_B^A \mathcal{K}_{B'}^{A'} \mathcal{M}_{AA'}^{BB'}$$

$$\text{Term 4} = \alpha^2 \mathcal{K}_B^A \partial_\nu \left[ \delta_{AB_1 B_2}^{BA_1 A_2} \mathcal{K}_{A_1}^{B_1} \overline{\mathcal{K}}_{A_2}^{B_2} \right], \quad \text{Term 5} = \nabla_A \mathcal{Y}^A$$

- For Einstein theory ( $\alpha = 0$ ) 2-nd law is valid for wald entropy,

$$\partial_\nu \Theta_{eq} = -T_{\nu\nu} - \mathcal{K}_{AB}\mathcal{K}^{AB} \Rightarrow \partial_\nu \Theta_{eq} \leq 0$$

- Term 1 : Null Energy condition, and Term 2  $\leq 0$  .
- Term 3  $\ll$  Term 2  $\Rightarrow$  Term 2 + Term 3 =  $\mathcal{K}_B^A \mathcal{K}_{B'}^{A'} \left[ \delta_A^B \delta_{A'}^{B'} + \alpha^2 \mathcal{M}_{AA'}^{BB'} \right]$
- Term 4 is naively small compared to Term 2 : But not always true

$$\text{Term 2} + \text{Term 4} = \mathcal{K}_B^A \left[ \mathcal{K}_A^B + \alpha^2 \partial_\nu \left[ \delta_{AB_1 B_2}^{BA_1 A_2} \mathcal{K}_{A_1}^{B_1} \overline{\mathcal{K}}_{A_2}^{B_2} \right] \right]$$

$$\text{Term 4} \gtrsim \text{Term 2} \Rightarrow \partial_\nu \Theta_{eq} \not\leq 0$$

Need to add corrections to Wald entropy to handle it.

# Gauss-Bonnet Theory : Details of proof

- We decide the correction term

$$\begin{aligned} S_{\text{total}} &= S_{\text{Wald}} + S_{\text{cor}} = \int_{\Sigma_v} d^{d-2}x \sqrt{h} [\rho_{\text{eq}} + \rho_{\text{cor}}] \\ \partial_v S_{\text{total}} &= \int_{\Sigma_v} d^{d-2}x \sqrt{h} [\Theta_{\text{eq}} + \Theta_{\text{cor}}] \\ \partial_v \Theta_{\text{eq}} &= -T_{vv} - \mathcal{K}_{AB} \mathcal{K}^{AB} + \alpha^2 \mathcal{K}_B^A \partial_v H_A^B + \nabla_A \mathcal{Y}^A + \text{Negligible Terms} \\ H_A^B &= \delta_{AB_1 B_2}^{BA_1 A_2} \mathcal{K}_{A_1}^{B_1} \bar{\mathcal{K}}_{A_2}^{B_2} \end{aligned}$$

- We add  $S_{\text{cor}}$  to adjust  $\partial_v \Theta_{\text{cor}}$  such that

$$\begin{aligned} \partial_v \Theta_{\text{cor}} &= \alpha^4 \gamma \partial_v H_A^B \partial_v H_B^A \\ \partial_v [\Theta_{\text{eq}} + \Theta_{\text{cor}}] &= -T_{vv} - \left[ \mathcal{K}_B^A - \frac{\alpha^2}{2} \partial_v H_B^A \right]^2 - \alpha^4 \left( \gamma - \frac{1}{4} \right) \partial_v H_A^B \partial_v H_B^A \\ \Rightarrow \partial_v [\Theta_{\text{eq}} + \Theta_{\text{cor}}] &\leq 0 \quad \text{if } \gamma \leq 1/4 \end{aligned}$$

- We need to impose :  $\nabla_A \mathcal{Y}^A = \nabla_A \nabla_A \mathcal{Z}^{AB} = 0$ , which is true for spherically symmetric evolution.

# Extension to Lovelock Theories

- We extended the analysis beyond Gauss-Bonnet to Lovelock theories

$$I \equiv \int \sqrt{-g} \left[ R + \alpha_m \ell_S^{2m-2} \mathcal{L}_m + \mathcal{L}_{matter} \right]$$

$$\mathcal{L}_m = \delta_{\rho_1 \sigma_1 \dots \rho_m \sigma_m}^{\mu_1 \nu_1 \dots \mu_m \nu_m} R^{\rho_1 \mu_1 \sigma_1 \nu_1} \dots R^{\rho_m \mu_m \sigma_m \nu_m}$$

$\delta_{\rho_1 \sigma_1 \dots \rho_m \sigma_m}^{\mu_1 \nu_1 \dots \mu_m \nu_m}$  = determinant of  $(n \times n)$  matrix whose  $(ij)$ -th element is  $\delta_{\nu_j}^{\mu_i}$

- The correction we need to add

$$S_{\text{total}} = S_{\text{Wald}} + S_{\text{cor}} = \int_{\Sigma_V} d^{d-2}x \sqrt{h} [\rho_{\text{eq}} + \rho_{\text{cor}}]$$

$$\rho_{\text{eq}} = \left. \frac{\delta \mathcal{L}_{\text{grav}}}{\delta R^{\nu}{}_{\nu}{}^r{}_r} \right|_{R \rightarrow \mathcal{R}}, \quad \mathcal{L}_{\text{grav}} = R + \alpha_m \ell_S^{2m-2} \mathcal{L}_m$$

$$\rho_{\text{cor}} = \sum_{n=0}^{\infty} \kappa_n \left[ \ell_S^n \partial_V^n \left( \frac{1}{2} \frac{\delta^2 \mathcal{L}_{\text{grav}}}{\delta R^A{}_{A_1}{}_{C_1}{}_{\nu}{}_{\delta} R^{\nu}{}_{B_1}{}_{D_1}{}_{B}} \right) \Big|_{R \rightarrow \mathcal{R}} \mathcal{K}_{A_1}^{C_1} \bar{\mathcal{K}}_{B_1}^{D_1} \right]^2$$

- The replacement rule :  $R \rightarrow \mathcal{R} \Rightarrow$  replace all the curvature tensors of the spacetime with those intrinsic to  $\Sigma_V$

# Extension to Lovelock Theories

$$S_{\text{total}} = S_{\text{Wald}} + S_{\text{cor}} = \int_{\Sigma_v} d^{d-2}x \sqrt{h} [\rho_{\text{eq}} + \rho_{\text{cor}}]$$

$$\rho_{\text{eq}} = \left. \frac{\delta \mathcal{L}_{\text{grav}}}{\delta R^{\nu}{}_{\nu}{}^r{}_{r}} \right|_{R \rightarrow \mathcal{R}}, \quad \mathcal{L}_{\text{grav}} = R + \alpha_m \ell_s^{2m-2} \mathcal{L}_m$$

$$\rho_{\text{cor}} = \sum_{n=0}^{\infty} \kappa_n \left[ \ell_s^n \partial_v^n \left( \frac{1}{2} \left. \frac{\delta^2 \mathcal{L}_{\text{grav}}}{\delta R^A{}_{A_1}{}_{C_1}{}_{\nu}{}_{\delta} R^{\nu}{}_{B_1}{}_{D_1}{}_{B}} \right|_{R \rightarrow \mathcal{R}} \mathcal{K}_{A_1}^{C_1} \overline{\mathcal{K}}_{B_1}^{D_1} \right) \right]^2$$

- $\partial_v S \geq 0$  in any time evolution maintaining spherical symmetry provided the  $\kappa_n$ 's satisfy the following recursive inequality

$$A_n = 2 \kappa_n - \frac{\kappa_{n-1}^2}{A_{n-2}} \leq 0, \quad \text{for } n = -2, -1, 0, \dots$$

$$\text{initial condition : } \kappa_{-2} = -\frac{1}{2}, \quad \kappa_{-1} = -2.$$

- The Obstruction term in the form of a total derivative is still there.



# Conclusions

- For Lovelock theory we have been able to construct an  $S_{\text{total}}$  satisfying *2nd* law, surely for spherical symmetry.
- We need to understand the implications of the obstruction term.
- Our construction is not unique, for example
  - Instead of  $\partial_\nu \Theta \leq 0$  we prove  $\partial_\nu (Z \Theta) \leq 0$  for some  $Z \geq 0$ , and it would do the job.  
In fact this is how 2nd law is proved for  $f(R)$  theories.
- Field redefinitions and foliation dependence.
- The method is indirect : It is possible that  $\partial_\nu S_{\text{total}}$  not monotonically decreasing. To obtain some constraints on the structure of the higher derivative corrections we need a direct method  $\partial_\nu S_{\text{total}} < 0$
- Possible connection with Holographic Entanglement entropy  $\Rightarrow$  Myers, Dong, Camps analysis.

**Thank You For Attention**