

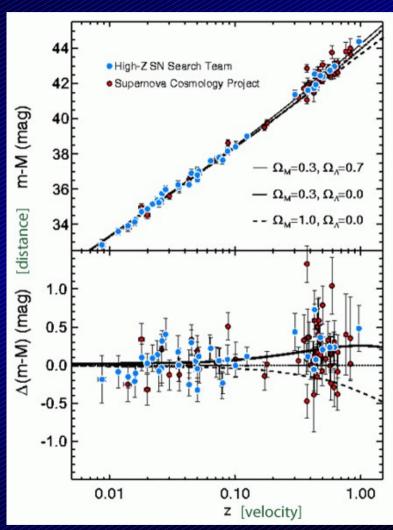


The Story of





Supernova data



A single parameter (Λ) can fit the data perfectly; and it is already part of the theory!



The Story of



One problem:

Why is got that tiny value?

This problem lead to many interesting alternatives (predictions/understandings):

Such as: quintessence, k-essence, inhomogeneous cosmologies, chameleons, quintom, phantom fields, frustrated domain walls, string landscape, LTB, f(R), DGP, cosmological casimir effect, chaplygin gas, holographic dark energy, XCDM, quartessence...



Early universe



We know inflationary models fit the data, but also have certain problems...

Use same philosophy to study alternatives to inflation...

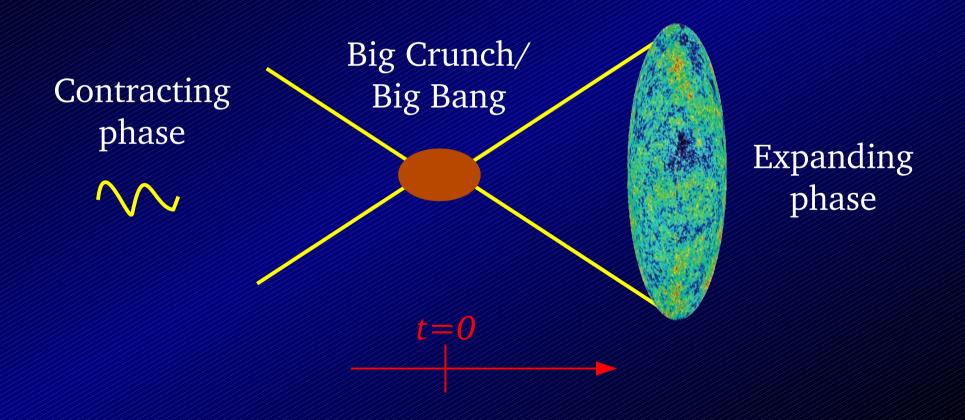
Consider, in particular, bouncing cosmologies



Bouncing models



Key: Pre-Big bang phase where CMB pertns. are created



Two flavours: singular & non-singular



Bouncing models



- Consider singular models
- Assume spectrum of scale invariant perturbations
- Need to understand how these perturbations go through the singularity

What do we know?

- GR breaks down near the singularity
- Quantum gravity should give an answer (e.g. string/M theory)





Cannot solve the problem in general

Study a simple model

$$ds^{2} = a^{2}(t) \left(-dt^{2} + \sum_{i=1}^{3} (dx^{i})^{2} \right)$$

$$a(t) = |t|^{1/2}$$





Higher dimensional motivation

$$ds^{2} = -dt^{2} + t^{2}d\theta^{2} + \sum_{i=1}^{3} (dx^{i})^{2}$$

Milne Universe

KK reduction

$$ds^{2} = |t| \left(-dt^{2} + \sum_{i=1}^{3} (dx^{i})^{2} \right)$$





Higher dimensional motivation



$$ds^{2} = -dt^{2} + t^{2}d\theta^{2} + \sum_{i=1}^{3} (dx^{i})^{2}$$

Milne Universe

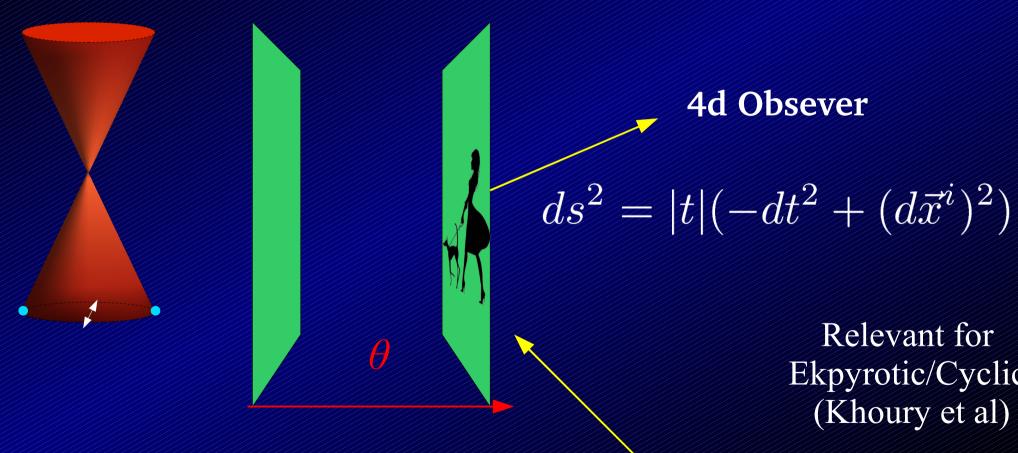
If
$$\theta \in (-\infty, \infty)$$
 flat space in unusual coordinates

If
$$\theta \in [0, 2\pi)$$
 there indeed a conical singularity





Actually, if a \mathbb{Z}_9 symmetry is assumed



 $ds^{2} = -dt^{2} + t^{2}d\theta^{2} + \sum (dx^{i})^{2}$

Ekpyrotic/Cyclic



Fields on Milne



Scalar field in 5d (near singularity)

$$ds^{2} = -dt^{2} + t^{2}d\theta^{2} + \sum (dx^{i})^{2}$$

Get Bessel equation, and in the long-wavelength limit

$$\phi(t) = A + B \ln(t)$$

Diverges at t=0

However, can match across using analytic continuation of Bessel functions (Tolley & Turok, 2006)



String/M-theory analysis



Can we do better with string/M-theory?

Heterotic string effective action

$$S = \int d^{10}x \sqrt{-g}e^{-2\phi} \left(R + 4(\partial\phi)^2 + \mathcal{O}(\alpha') \right)$$

(Gross & Sloan, 1987)

$$\mathcal{O}(\alpha') = \frac{\alpha'}{8} (R_{abcd} R^{abcd} - 4R_{ab} R^{ab} + R^2) + \dots$$

Gauss-Bonnet



String/M-theory analysis



Background

GN & Turok, 2007

$$a^{2}(t) = a_{0}^{2}(t) + \alpha' a_{1}^{2}(t) + \dots$$

$$\phi(t) = \phi_0(t) + \alpha' \phi_1(t) + \dots$$

0th order

$$a_0^2(t) = t$$

$$\phi_0(t) = A \ln(t) + B$$

$$a^{2}(t) = t \left(1 - \frac{\alpha'}{8t^{3}} - \mathcal{O}\left(\frac{\alpha'^{2}}{t^{6}}\right) \right)$$

$$\phi(t) = A \ln(t) + B - \frac{3\alpha'}{16t^3} - \mathcal{O}\left(\frac{\alpha'^2}{t^6}\right)$$

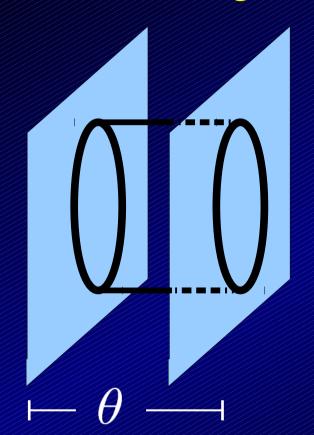
Singularity is reached sooner, so theory breaks down!



String/M-theory analysis



Are we missing something?



 $\frac{KK}{0\text{-modes}}$ 4d

Membranes

Strings (IIA, Het)

(Turok, Perry & Steinhardt, 2002)

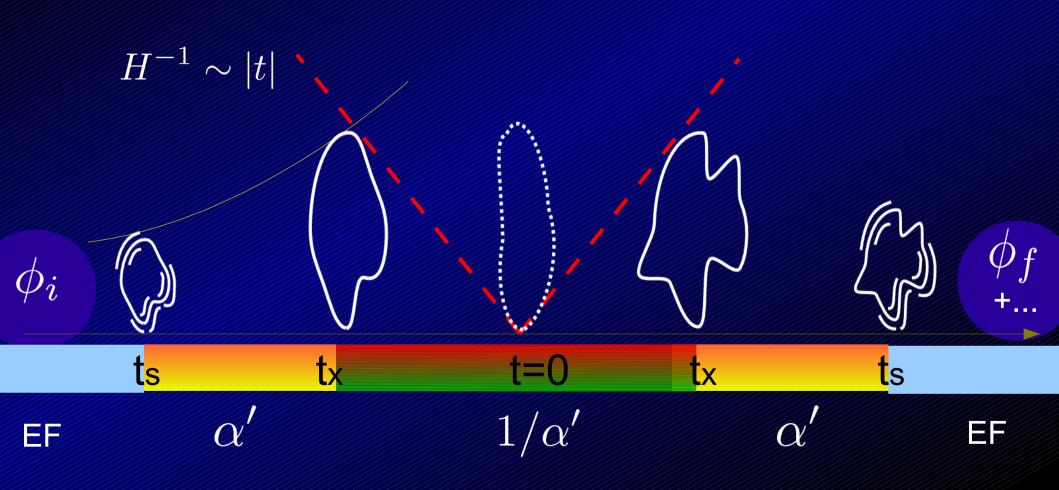
Relevant for cosmological perturbations (Include graviton & dilaton)



Classical evolution



Solutions are regular across the singularity



GN & Turok, 2006

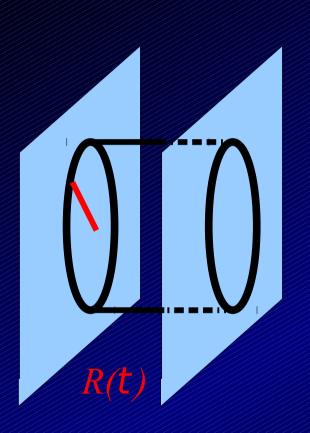


Loop quantisation



In the case of circular loops,

Copeland, GN & Turok, 2010



the Hamiltonian constraint is:

$$H = -\pi_0^2 + \pi_R^2 + t^2 R^2 \simeq 0$$



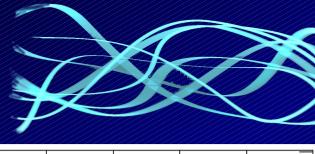
$$\hat{H}\Psi = \left(\partial_t^2 - \partial_R^2 + t^2 R^2\right)\Psi \simeq 0$$

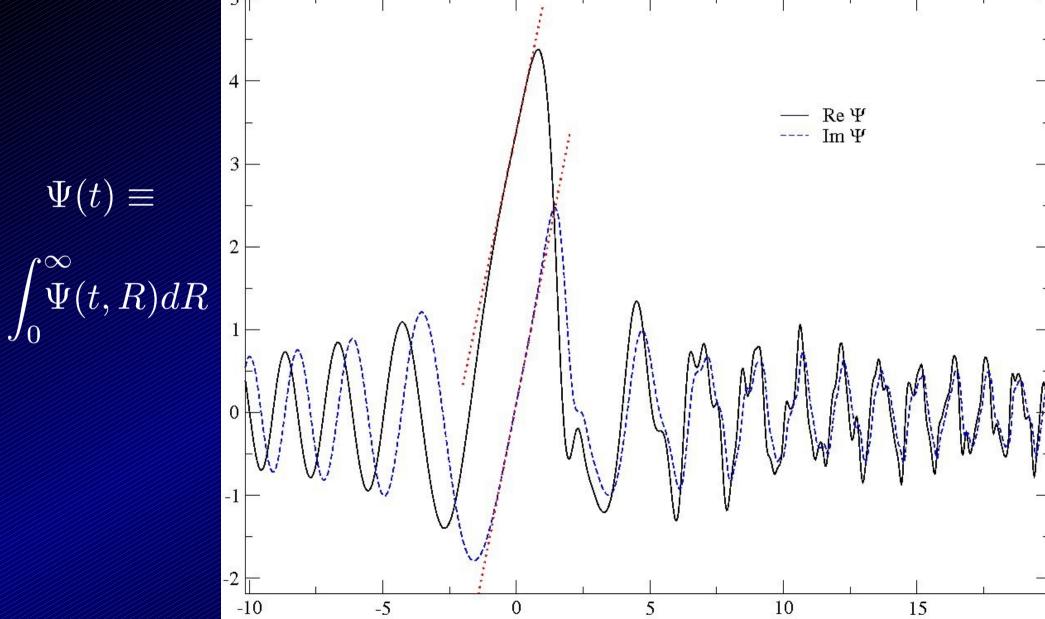
String wave function $\Psi(t,R)$

c.f. Harmonic oscillator



String wave function







Particle production



$$\Psi(t,R) = \sum A_n(t)H_n(x)e^{-x^2/2},$$

$$x \equiv \sqrt{|t|}R$$

Hermite polynomials

for large |t|:
$$A(t) \rightarrow e^{\pm E_n|t|}$$
,

$$E_n \sim \sqrt{(2n+1)}|t|$$

So can send IN vacuum (positive frequency mode) and read the OUT state in terms harmonic oscillator states.

$$|\Psi_n\rangle_{out} = \sum_m \alpha_{nm} |\Psi_m^+\rangle + \beta_{nm} |\Psi_m^-\rangle$$

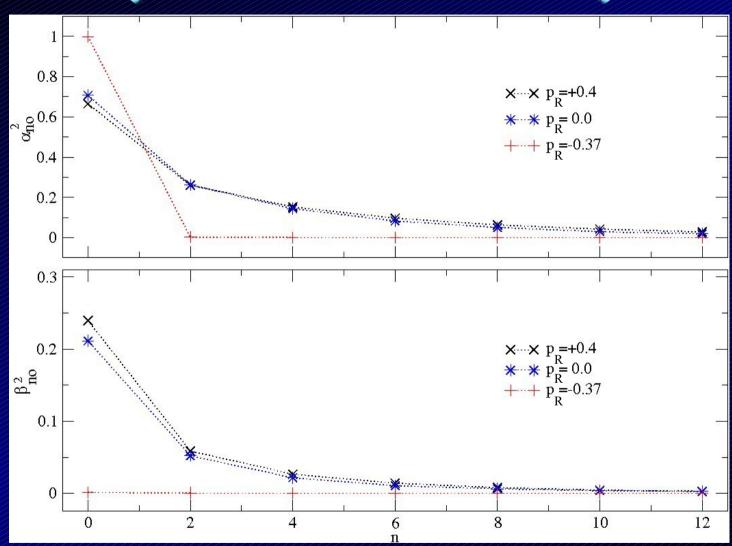
Particle production:

Bogoliubov coeffs.

$$\langle N_n \rangle = \sum |\beta_{nm}|^2$$



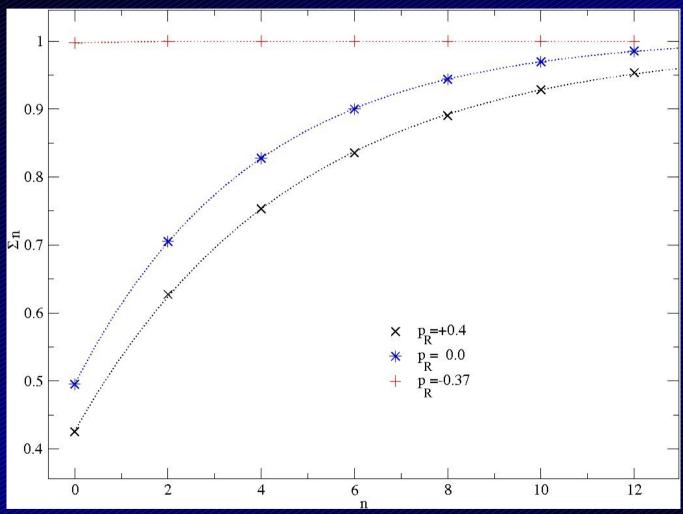
Particle production



Finite Particle Production (decays exponentially with occupation number)







m

Unitarity is preserved if:
$$\sum |\alpha_{nm}|^2 - |\beta_{nm}|^2 = 1$$



conclusions



- Classical string (winding membranes) are regular across a Milne 5d singularity, featuring
 - String travels at the speed of light at t=0
 - Higher oscillation modes resolve the singularity
 - Ultralocal behaviour (string bits)
- Circular loops can be quantised and there is finite particle production
- Unitarity is preserved
- Can be extended to more complicated singularities (Kasner metrics, maybe to FRWL and some BH's)



M-theory model



Winding membranes (strings): Nambu-Goto action

$$S = -\mu_2 \int d^3\sigma \sqrt{-\text{Det}(\partial_A X^M \partial_B X^N g_{MN})}$$

11d Milne

Winding membranes \blacksquare field independent of: $\sigma^3 = y$

$$S = -\int \mu_2 \theta_0 |t| d^2 \sigma \sqrt{-\text{Det}(\partial_\alpha X^\mu \partial_\beta X^\nu \eta_{\mu\nu})}$$

Efective Tension



DyNamics



Two equivalent descriptions

1

- String living on flat spacetime
- Tension: $\mu_1 = \mu_2 \theta_0 |t|$
- Tensionless at t=0
- Like harmonic oscillator with a time-dependent frequency (quantum)

2

- String living on FRWL $g_{\mu\nu}=|t/t_s|\eta_{\mu\nu}$
- Fixed tension: $\mu_1 = \mu_2 \theta_0 |t_s|$
- Speed of light at *t*=0
- Better to study classical behaviour



DyNamics



Bare in mind...

Quantum corrections:

- String interactions are suppressed $(g_s = |t/t_s|^{3/2})$
- α'-corrections are under control GN & Turok.

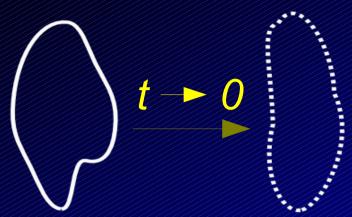
Small perturbations (ripples on the orbifolds) lead to 11d *Kasner* backgrounds:

$$ds^{2} = -dt^{2} + |\theta_{0}t|^{2p_{10}}d\theta^{2} + \sum_{i=0}^{9} |\theta_{0}t|^{2p_{i}}(dx^{i})^{2}$$
$$\sum_{n=1}^{10} p_{n}^{2} = 1 = \sum_{n=1}^{10} p_{n}$$



Classical evolution





String breaks into bits!!

Like ultra-locality

in BKL analysis

Consider Hamiltonian for winding membrane

$$H \sim \int d\sigma \left[\mathcal{A}(\Pi_{\mu}\Pi_{\nu}\eta^{\mu\nu} + \mu_{2}\theta_{0}^{2}t^{2}X^{\prime\;\mu}X^{\prime\;\nu}\eta_{\mu\nu}) + \mathcal{A}^{i}\Pi^{\mu}X^{\prime}_{\mu,i} \right]$$

Arbitrary function

- Interaction term, coupling ~ $1/\alpha'$
- |t|>>1 α' expansion
- |t| << 1 $1/\alpha'$ expansion



String Spectrum



Left/right-mover decomposition:

$$\vec{X}_R = \dot{\vec{X}} - \vec{X}' = \sum_{i=-\infty}^{+\infty} \frac{1}{2} \vec{\alpha}_n e^{in\xi_-}$$

$$\vec{X}_L = \dot{\vec{X}} + \vec{X}' = \sum_{i=-\infty}^{+\infty} \frac{1}{2} \vec{\tilde{\alpha}}_n e^{in\xi_+}$$

$$\xi_{\pm} = \tau \pm \sigma$$

$$\vec{\alpha}_n^{\dagger} = \vec{\alpha}_{-n}$$

$$\vec{\tilde{\alpha}}_n^{\dagger} = \vec{\tilde{\alpha}}_{-n}$$

Constraint
$$\vec{X}_R^2 = 1 = \vec{X}_L^2$$
 ——

curves in S d-1

Kibble & Turok, 1982

String's massless bosonic sector: $\phi, h_{\mu\nu}, B_{\mu\nu}$

spin: 0

2



String Spectrum



Left/right-mover decomposition:

$$\vec{X}_R = \dot{\vec{X}} - \vec{X}' = \sum_{i=-\infty}^{+\infty} \frac{1}{2} \vec{\alpha}_n e^{in\xi_-}$$

$$\xi_{\pm} = \tau \pm \sigma$$

$$ec{X}_L = \dot{ec{X}} + ec{X}' = \sum_{i=-\infty}^{+\infty} rac{1}{2} ec{ ilde{lpha}}_n ec{lpha}_n$$

Consider dilaton

Constraint $ec{X}_R^2 = 1 = ec{X}_L^2$

mode only!

 $\mathsf{S}^{\mathsf{d-1}}$

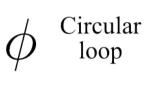
Kibble & Turok, 1982

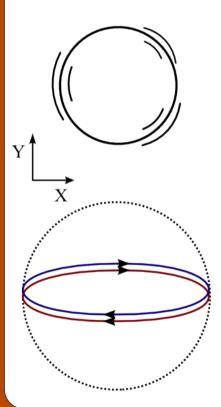
String's massless bosonic sector: ϕ ,

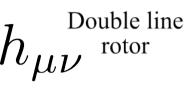
sector: ϕ , $h_{\mu\nu}$, $B_{\mu\nu}$ spin: 0 2 1

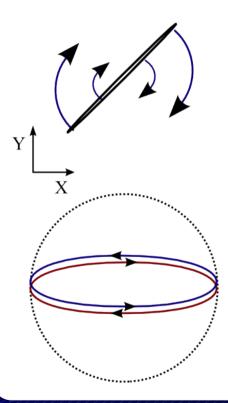
String Spectrum

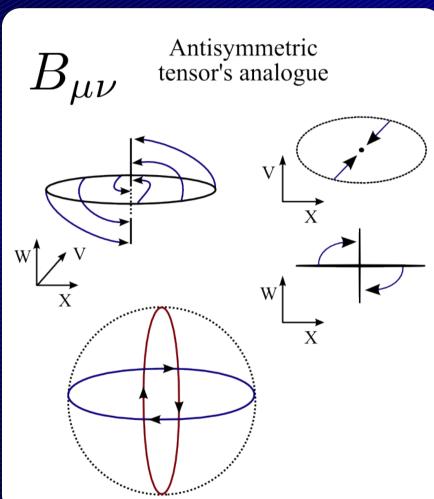














Effective Action



Heterotic string

$$S = \int d^{10}x \sqrt{-g}e^{-2\phi} \left(R + 4(\partial\phi)^2 + \mathcal{O}(\alpha')\right)$$
(Gross & Sloan, 1987)

$$\mathcal{O}(\alpha') = \frac{\alpha'}{8} (R_{abcd} R^{abcd} - 4R_{ab} R^{ab} + R^2) + \text{ terms involving}$$
 the dilaton

Gauss-Bonnet



Background



$$a^{2}(t) = a_{0}^{2}(t) + \alpha' a_{1}^{2}(t) + \dots$$

$$\phi(t) = \phi_0(t) + \alpha' \phi_1(t) + \dots$$

0th order

$$a_0^2(t) = t$$

$$\phi_0(t) = \frac{3}{2}\ln(t)$$

1st, 2nd order

$$a^{2}(t) = t \left(1 - \frac{\alpha'}{8t^{3}} - \mathcal{O}\left(\frac{\alpha'^{2}}{t^{6}}\right) \right)$$

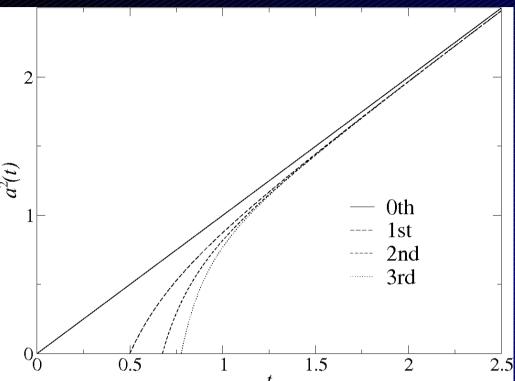
$$\phi(t) = \frac{3}{2}\ln(t) - \frac{3\alpha'}{16t^3} - \mathcal{O}\left(\frac{\alpha'^2}{t^6}\right)$$

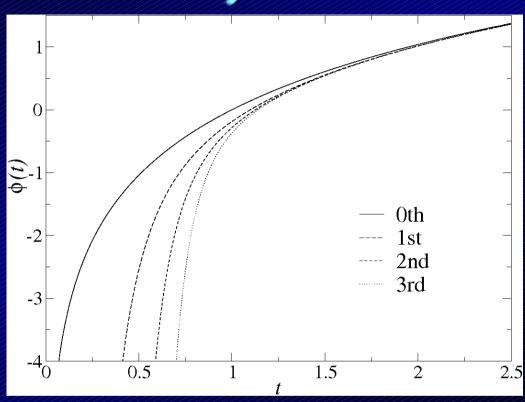
Note that
$$a^2(t) = e^{2\phi/3}$$



Background







- •Singularity is reached sooner
- •Corrections become important around t=tx



Perturbations



Tensor modes

$$ds^{2} = a(t)^{2} \left(-dt^{2} + (\delta_{ij} + h_{ij}) dx^{i} dx^{j} \right)$$

$$h_{ij}(t, \vec{x}) = e^{ik \cdot x} \hat{h}_{ij}(t)$$

$$h_{ij}(t) = h_{ij}^{0}(t) + \alpha' h_{ij}^{1}(t) + \dots$$

0th order

Bessel's equation. Consider

 $k \operatorname{cas} \Theta$

$$h_{ij}^0(t) = A_{ij} + B_{ij} \ln(t)$$

$$h_{ij}(t) = A_{ij} + B_{ij} \ln(t) - \frac{\alpha' B_{ij}}{6t^3} - \mathcal{O}\left(\frac{\alpha'^2 B_{ij}}{t^6}\right)$$



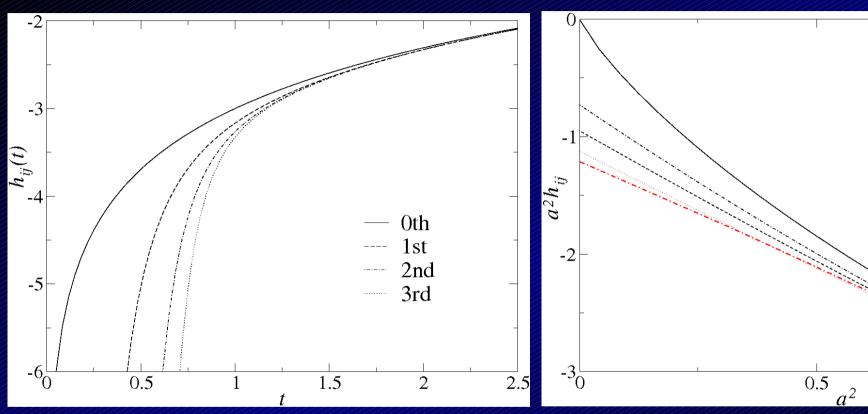
Perturbations



Oth 1st 2nd

3rd

Re-summed



Scalar perturbations + gauge choice = tensor perts.