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Complex Langevin analysis of the
spontaneous rotational symmetry
breaking in the Euclidean type IIB matrix
model

(arXiv:1712.07562, arXiv:2001.XXXXX)

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with Konstantinos N. Anagnostopoulos (NTUA), Yuta Ito (KEK),
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Difficulties in simulating complex partition functions.

$$Z = \int dA \exp(-S_0 + i\Gamma), \quad Z_0 = \int dA e^{-S_0}$$

Sign problem:

The reweighting $\langle \mathcal{O} \rangle = \frac{\langle \mathcal{O} e^{i\Gamma} \rangle_0}{\langle e^{i\Gamma} \rangle_0}$ requires configs. $\exp[\mathcal{O}(N^2)]$

$\langle^* \rangle_0 =$ (V.E.V. for the phase-quenched partition function Z_0)

Various methods to address the sign problem:

(**Complex Langevin Method (CLM)**, factorization method, Lefschetz-thimble method...)

In the following, we discuss **CLM**.

2. Euclidean type IIB matrix model

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type IIB matrix model model (a.k.a. **IKKT model**)

⇒ Promising candidate for nonperturbative string theory

[N. Ishibashi, H. Kawai, Y. Kitazawa and A. Tsuchiya, hep-th/9612115]

$$Z = \int dA d\psi e^{-(S_b + S_f)}$$
$$S_b = -\frac{N}{4} \text{tr}[A_\mu, A_\nu]^2, \quad S_f = N \text{tr} \bar{\psi}_\alpha (\Gamma^\mu)_{\alpha\beta} [A_\mu, \psi_\beta]$$

Euclidean case after Wick rotation $A_0 \rightarrow iA_D, \Gamma^0 \rightarrow -i\Gamma_D$.

⇒ Path integral is finite without cutoff.

[W. Krauth, H. Nicolai and M. Staudacher, hep-th/9803117, P. Austing and J.F. Wheeler, hep-th/0103059]

• $A_\mu, \Psi_\alpha \Rightarrow N \times N$ Hermitian traceless matrices.

$$\mu = 1, 2, \dots, D, \quad \alpha, \beta = \begin{cases} 1, 2, 3, 4 & (D = 6) \\ 1, 2, \dots, 16 & (D = 10) \end{cases}$$

• Originally defined in **D=10** (ψ : Majorana-Weyl)

We consider the **simplified D=6 case as well**

(ψ : Weyl, not Majorana $d\psi \rightarrow d\psi d\bar{\psi}$)

2. Euclidean type IIB matrix model

Result of Gaussian Expansion Method (GEM)

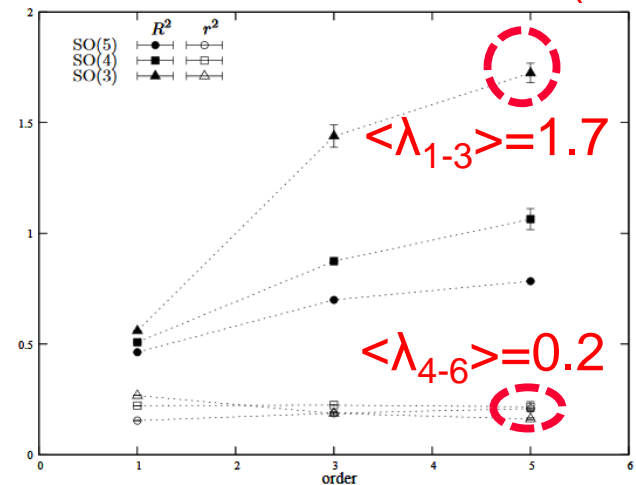
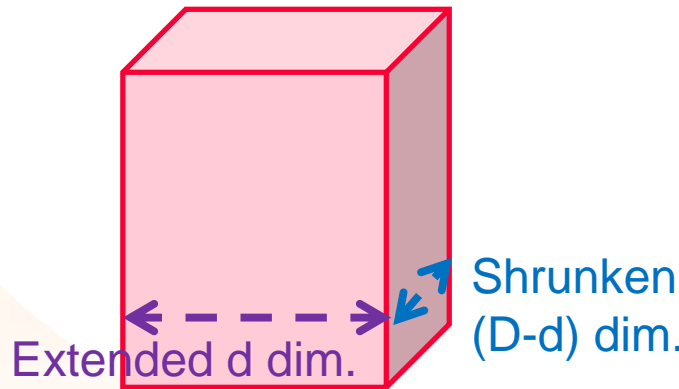
[T.Aoyama, J.Nishimura, and T.Okubo, arXiv:1007.0883, J.Nishimura, T.Okubo and F.Sugino, arXiv:1108.1293]

SSB **SO(6)** → **SO(3)** (In D=10, too, SO(10) → SO(3))
 Dynamical compactification to 3-dim spacetime.

$\lambda_n (\lambda_1 \geq \dots \geq \lambda_D)$: eigenvalues of $T_{\mu\nu} = \frac{1}{N} \text{tr}(A_\mu A_\nu)$

$$\rho_\mu = \frac{\langle \lambda_\mu \rangle}{\sum_{\nu=1}^D \langle \lambda_\nu \rangle} = \begin{cases} 0.30 & (\mu = 1, 2, 3) \\ 0.035 & (\mu = 4, 5, 6) \end{cases}$$

(D = 6) [arXiv:1007.0883](https://arxiv.org/abs/1007.0883) (D=6)



2. Euclidean type IIB matrix model

$$Z = \int dA d e^{-S_b} \underbrace{\left(\int d\psi e^{-S_f} \right)}_{= \det/Pf \mathcal{M} = |\det/Pf \mathcal{M}| e^{i\Gamma}} = \int dA \underbrace{e^{-S}}_{e^{-\{S_b - \log(\det/Pf \mathcal{M})\}}}$$

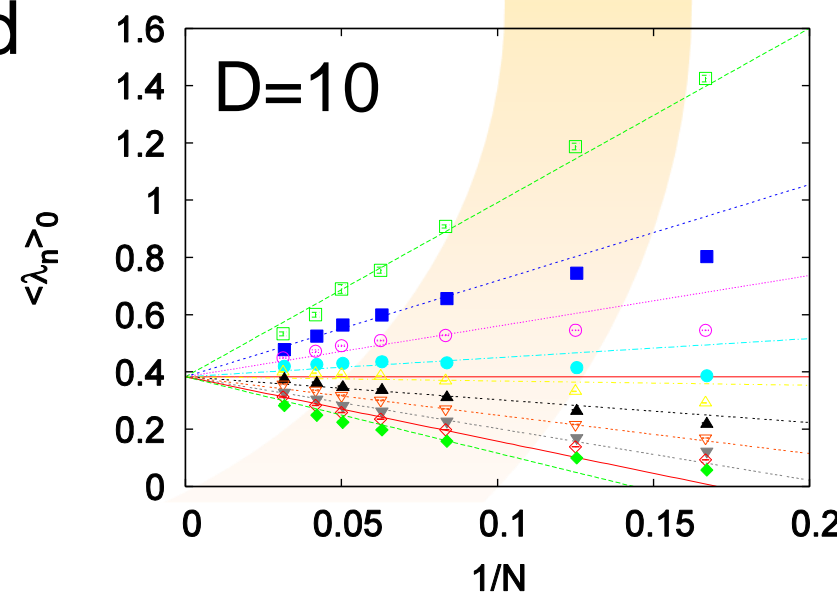
- Integrating out ψ yields **det \mathcal{M}** in **D=6** (Pf \mathcal{M} in D=10)
- det/Pf \mathcal{M} 's **complex phase** contributes to the **Spontaneous Symmetry Breaking (SSB)** of SO(D).

No SSB with the phase-quenched partition function.

[J. Ambjorn, K.N. Anagnostopoulos, W. Bietenholz, T. Hotta and J. Nishimura, hep-th/0003208,0005147, K.N. Anagnostopoulos, T. Azuma, J.Nishimura arXiv:1306.6135, 1509.05079]

$$Z_0 = \int dA e^{-S_0} = \int dA e^{-S_b} |\det/Pf \mathcal{M}|$$

$\langle^* \rangle_0 = \text{V.E.V. for } Z_0$



Complex Langevin Method (CLM)

⇒ Solve the complex version of the Langevin equation.

[Parisi, Phys.Lett. 131B (1983) 393, Klauder, Phys.Rev. A29 (1984) 2036]

$$\frac{d(A_\mu)_{ij}}{dt} = - \frac{\partial S}{\partial (A_\mu)_{ji}} + \eta_{\mu,ij}(t)$$

drift term

$$\frac{\partial S}{\partial (A_\mu)_{ji}} = \frac{\partial S_b}{\partial (A_\mu)_{ji}} - c_d \text{Tr} \left(\frac{\partial \mathcal{M}}{\partial (A_\mu)_{ji}} \mathcal{M}^{-1} \right) \quad c_d = \begin{cases} 1 & (D=6 \rightarrow \det \mathcal{M}) \\ \frac{1}{2} & (D=10 \rightarrow \text{Pf} \mathcal{M}) \end{cases}$$

• A_μ : Hermitian → general complex traceless matrices.

• η_μ : Hermitian-matrix white noise obeying the probability distribution

$$\exp \left(-\frac{1}{4} \int \text{tr} \eta^2(t) dt \right)$$

3. Complex Langevin Method

CLM does not work when it encounters these problems:

(1) Excursion problem: A_μ is too far from Hermitian
 \Rightarrow **Gauge Cooling** minimizes the **Hermitian norm**

$$\mathcal{N} = \frac{-1}{DN} \sum_{\mu=1}^D \text{tr}[(A_\mu - (A_\mu)^\dagger)^2] \quad [\text{K. Nagata, J. Nishimura and S. Shimasaki, arXiv:1604.07717}]$$

A_μ : **Hermitian** \rightarrow **general complex** traceless matrices.
 \Rightarrow We make use of this **extra symmetry**:

After each step of discretized Langevin equation,

$$A_\mu \rightarrow g A_\mu g^{-1}, \quad g = e^{\alpha H}, \quad H = \frac{-1}{N} \sum_{\mu=1}^D [A_\mu, A_\mu^\dagger]$$

α : real parameter, such that \mathcal{N} is minimized.

3. Complex Langevin Method

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(2) Singular drift problem:

The drift term $dS/d(A_\mu)_{ji}$ diverges due to \mathcal{M} 's **near-zero** eigenvalues.

We trust CLM when the distribution $p(u)$ of the **drift norm**

$$u = \sqrt{\frac{1}{DN^3} \sum_{\mu=1}^D \sum_{i,j=1}^N \left| \frac{\partial S}{\partial (A_\mu)_{ji}} \right|^2} \quad \text{falls exponentially as } p(u) \propto e^{-au}.$$

[K. Nagata, J. Nishimura and S. Shimasaki, arXiv:1606.07627]

Look at the **drift term** \Rightarrow Get the drift of CLM!!

Mass deformation [Y. Ito and J. Nishimura, arXiv:1609.04501]

• SO(D) symmetry breaking term $\Delta S_b = \frac{1}{2} N \varepsilon \sum_{\mu=1}^D m_{\mu} \text{tr}(A_{\mu})^2$

Order parameters for SSB of SO(D): $\lambda_{\mu} = \text{Re} \left\{ \frac{1}{N} \text{tr}(A_{\mu})^2 \right\}$

• Fermionic mass term:

$$\Delta S_f = N m_f \text{tr}(\bar{\Psi}_{\alpha} \gamma_{\alpha\beta} \Psi_{\beta}), \quad \gamma = \begin{cases} \Gamma_6 & (D=6) \\ i\Gamma_8 \Gamma_9^{\dagger} \Gamma_{10} & (D=10) \end{cases}$$

Avoids the singular eigenvalue distribution of \mathcal{M} .

This breaks $\text{SO}(6) \rightarrow \text{SO}(5)$ ($\text{SO}(10) \rightarrow \text{SO}(7)$)

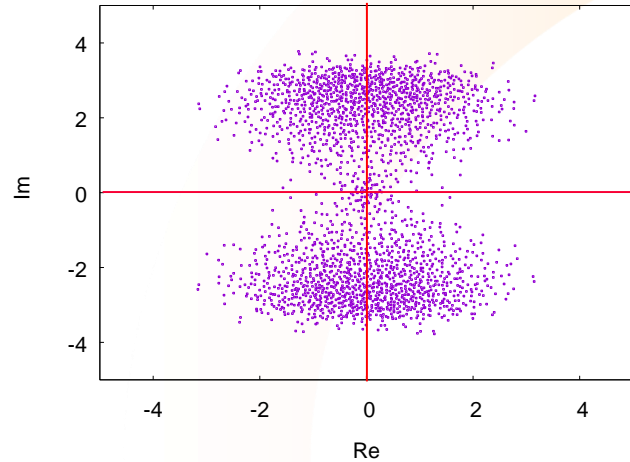
We study the SSB of the remaining symmetry.

Extrapolation (i) $N \rightarrow \infty \Rightarrow$ (ii) $\varepsilon \rightarrow 0 \Rightarrow$ (iii) $m_f \rightarrow 0$.

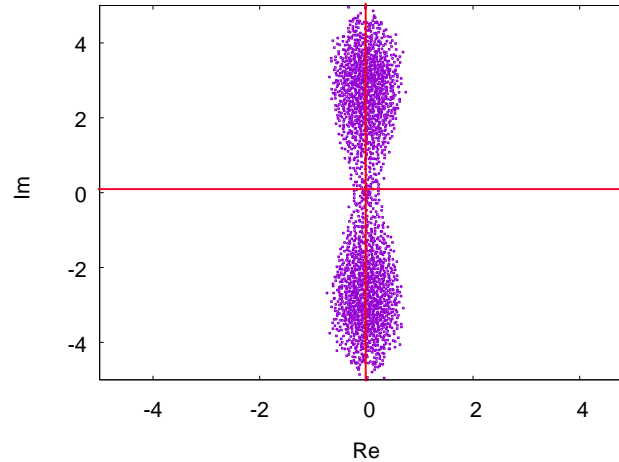
4. Result for D=6

The effect of adding these mass terms

$(\epsilon, m_f) = (0.00, 0.00)$



$(\epsilon, m_f) = (0.25, 0.00)$

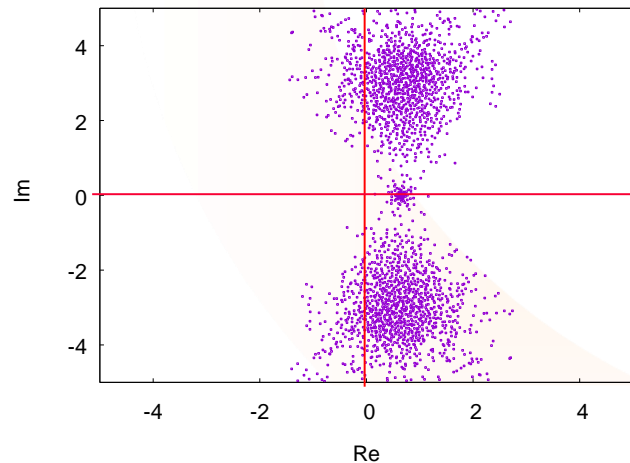


Scattering plots of the eigenvalues of the $4(N^2-1) \times 4(N^2-1)$ matrix \mathcal{M} for $D=6$, $N=24$.

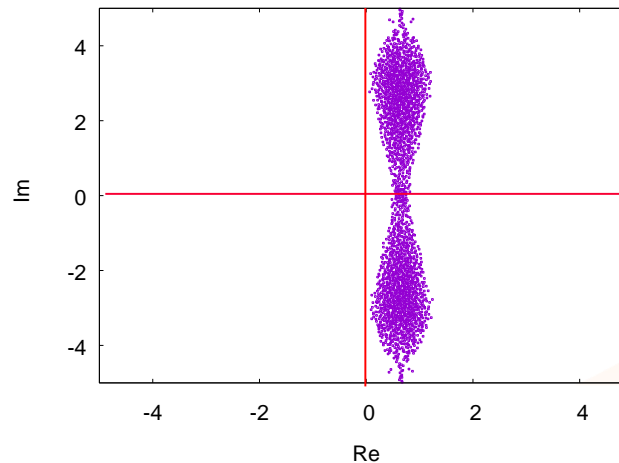
ΔS_b narrows the eigenvalue distribution.

ΔS_f shifts the eigenvalues, to evade the origin.

$(\epsilon, m_f) = (0.00, 0.65)$



$(\epsilon, m_f) = (0.25, 0.65)$



4. Result for D=6

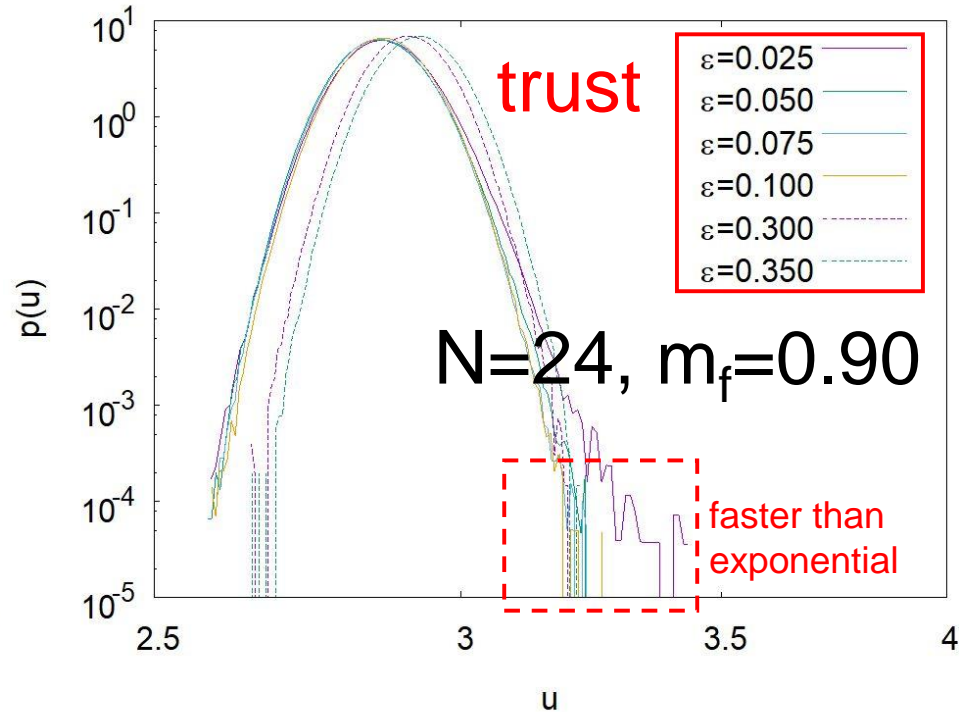
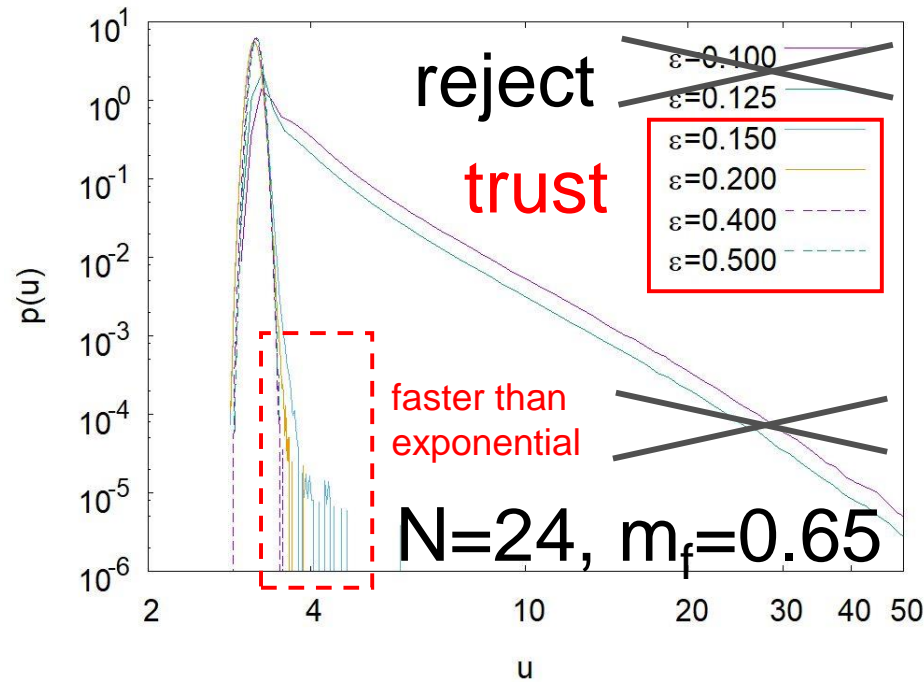
$$\Delta S_b = \frac{1}{2} N \varepsilon \sum_{\mu=1}^D m_{\mu} \text{tr}(A_{\mu})^2$$

$$\Delta S_f = N m_f \text{tr}(\bar{\Psi}_{\alpha} (\Gamma_D)_{\alpha\beta} \Psi_{\beta}) \quad (D = 6)$$

$$m_{\mu} = (0.5, 0.5, 1, 2, 4, 8)$$

$$u = \sqrt{\frac{1}{DN^3} \sum_{\mu=1}^D \sum_{i,j=1}^N \left| \frac{\partial S}{\partial (A_{\mu})_{ji}} \right|^2}$$

's distribution $p(u)$ (log-log)



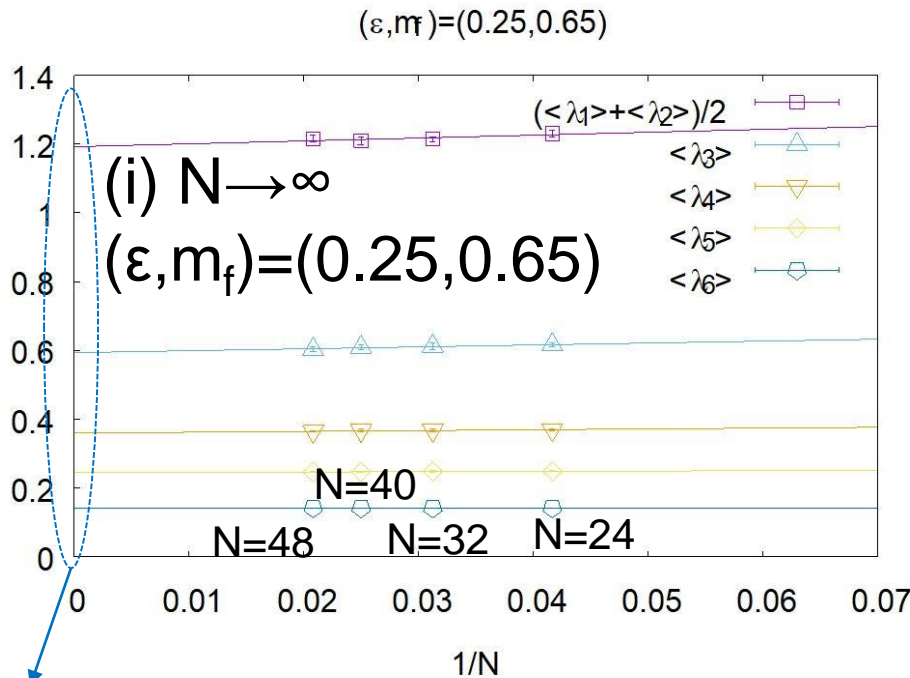
4. Result for D=6

$$\Delta S_b = \frac{1}{2} N \varepsilon \sum_{\mu=1}^D m_{\mu} \text{tr}(A_{\mu})^2$$

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$$m_{\mu} = (0.5, 0.5, 1, 2, 4, 8)$$

(i) $N \rightarrow \infty$ limit for **fixed** (ε, m_f)



$(\varepsilon, m_f) \rightarrow (0, 0)$ extrapolation
for **finite N**
 \Rightarrow We cannot observe
SSB of $SO(D)$.

$\langle \lambda_{\mu} \rangle_{\varepsilon, m_f}$ at large N

4. Result for D=6

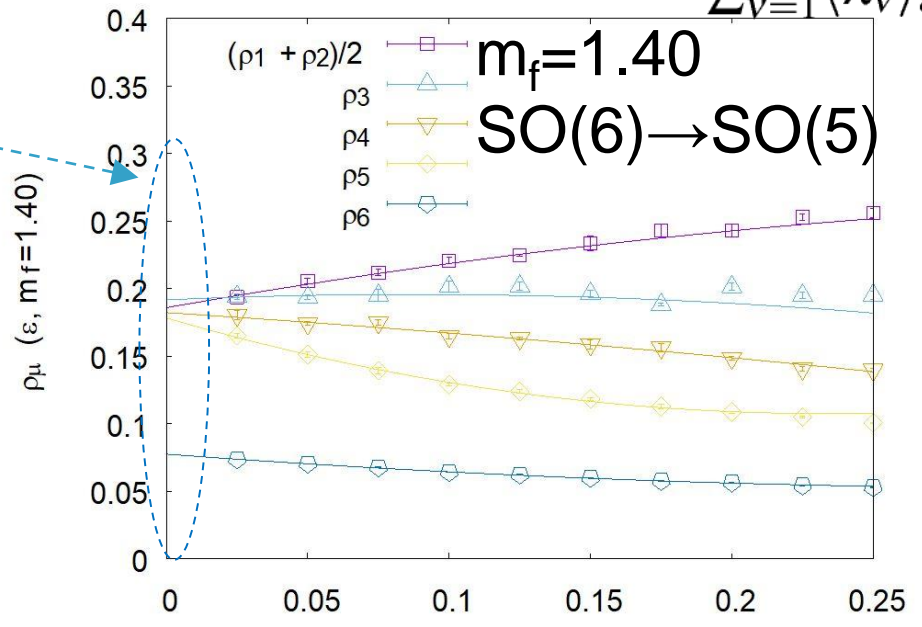
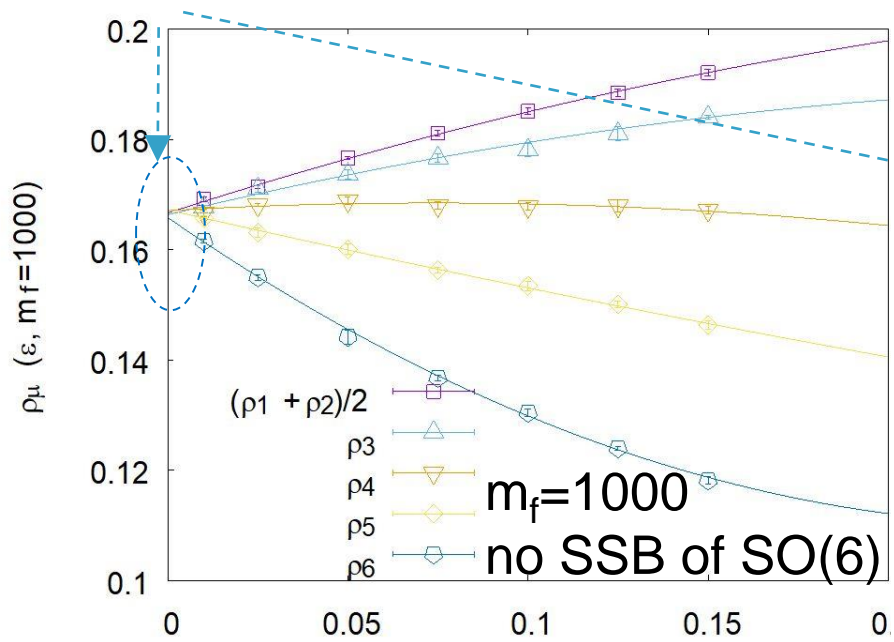
$$\Delta S_b = \frac{1}{2} N \varepsilon \sum_{\mu=1}^D m_\mu \text{tr}(A_\mu)^2$$

$$\Delta S_f = N m_f \text{tr}(\bar{\Psi}_\alpha (\Gamma_D)_{\alpha\beta} \Psi_\beta) \quad (D = 6)$$

$$m_\mu = (0.5, 0.5, 1, 2, 4, 8)$$

$$\rho_\mu(\varepsilon, m_f) = \frac{\langle \lambda_\mu \rangle_{\varepsilon, m_f}}{\sum_{\nu=1}^D \langle \lambda_\nu \rangle_{\varepsilon, m_f}}$$

(ii) $\varepsilon \rightarrow 0$ after $N \rightarrow \infty$



- $m_f \rightarrow \infty$: Ψ decouples from A_μ and reduces to the bosonic IKKT.
- The bosonic IKKT S_b does not break $SO(D)$.

[T. Hotta, J. Nishimura and A. Tsuchiya, hep-th/9811220]

- The SSB of $SO(D)$ is not an artifact of $\varepsilon \rightarrow 0$ but **a physical effect**.

4. Result for D=6

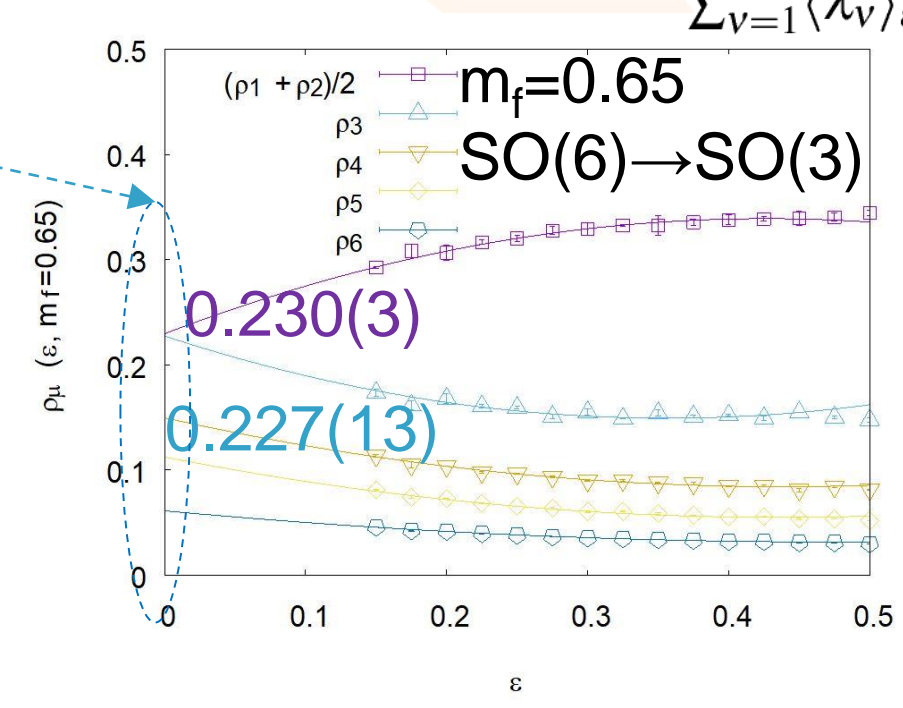
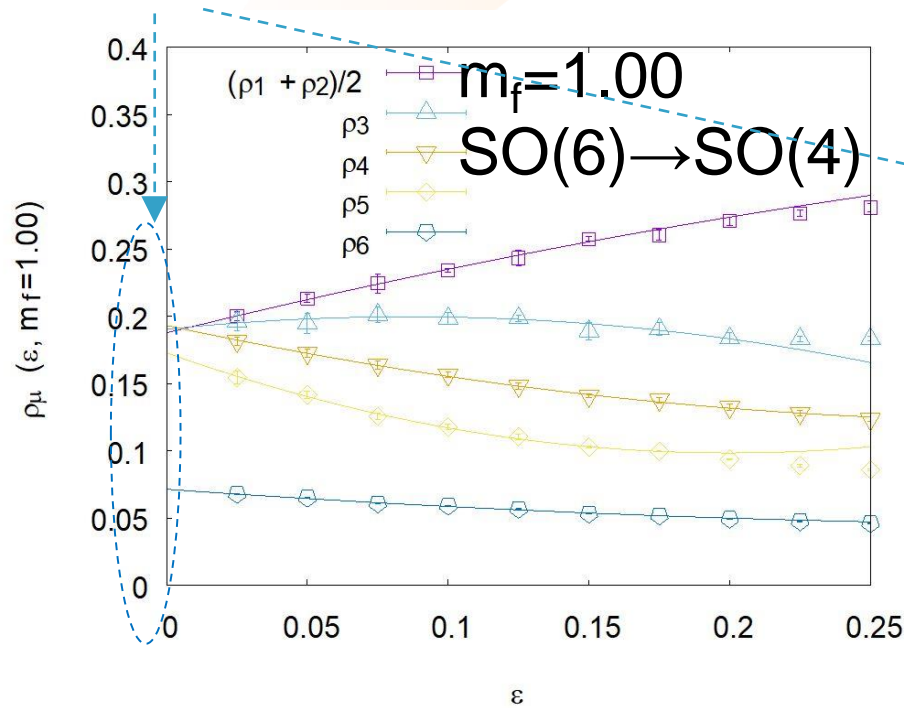
$$\Delta S_b = \frac{1}{2} N \varepsilon \sum_{\mu=1}^D m_{\mu} \text{tr}(A_{\mu})^2$$

$$\Delta S_f = N m_f \text{tr}(\bar{\psi}_{\alpha} (\Gamma_D)_{\alpha\beta} \psi_{\beta}) \quad (D = 6)$$

$$m_{\mu} = (0.5, 0.5, 1, 2, 4, 8)$$

$$\rho_{\mu}(\varepsilon, m_f) = \frac{\langle \lambda_{\mu} \rangle_{\varepsilon, m_f}}{\sum_{\nu=1}^D \langle \lambda_{\nu} \rangle_{\varepsilon, m_f}}$$

(ii) $\varepsilon \rightarrow 0$ after $N \rightarrow \infty$



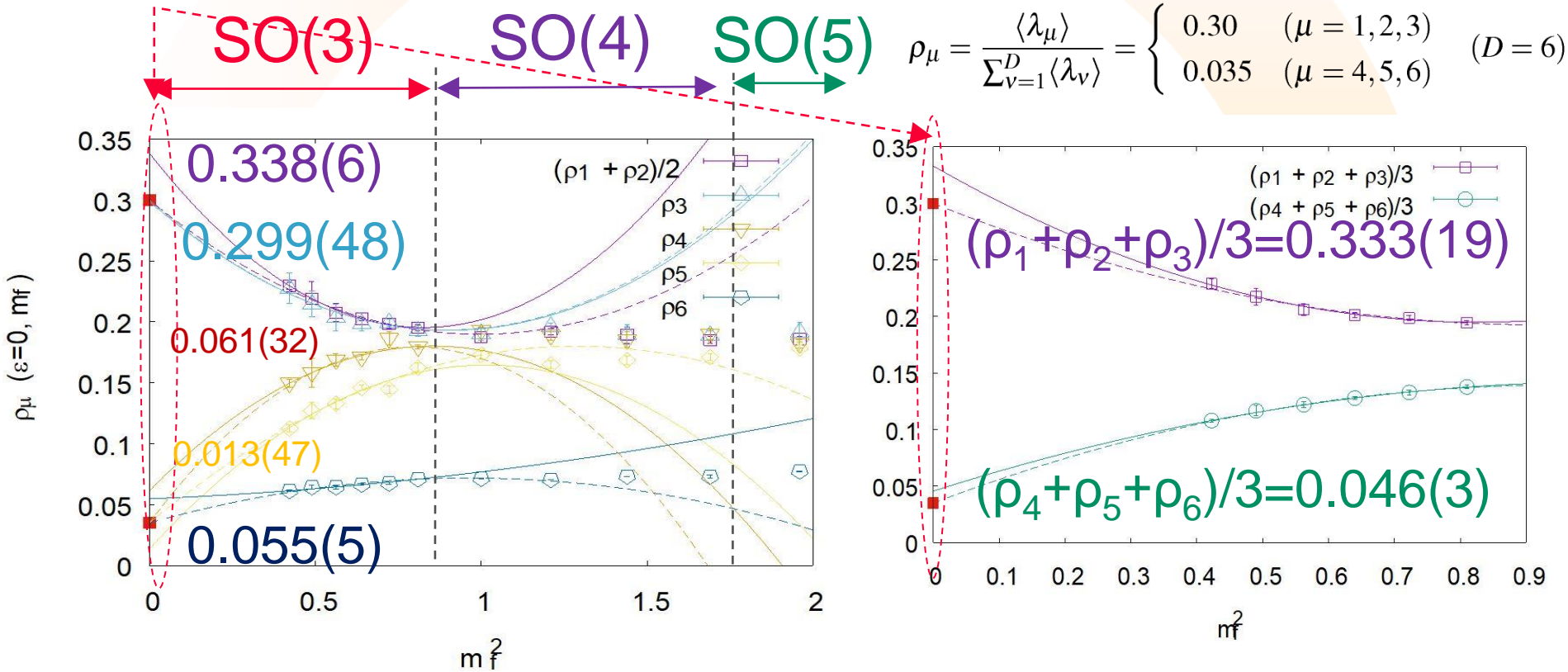
4. Result for D=6

$$\Delta S_b = \frac{1}{2} N \varepsilon \sum_{\mu=1}^D m_\mu \text{tr}(A_\mu)^2$$

$$\Delta S_f = N m_f \text{tr}(\bar{\Psi}_\alpha (\Gamma_D)_{\alpha\beta} \Psi_\beta) \quad (D = 6)$$

$$m_\mu = (0.5, 0.5, 1, 2, 4, 8)$$

(iii) $m_f \rightarrow 0$ after $\varepsilon \rightarrow 0$



(dotted line: $m_f \rightarrow 0$ limit fixed to GEM results)

SSB SO(6) \rightarrow at most SO(3)

Consistent with GEM.

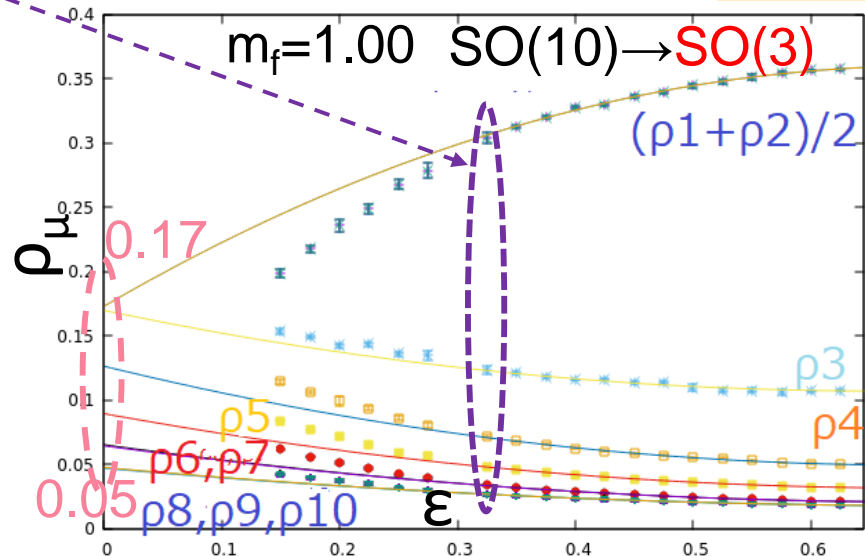
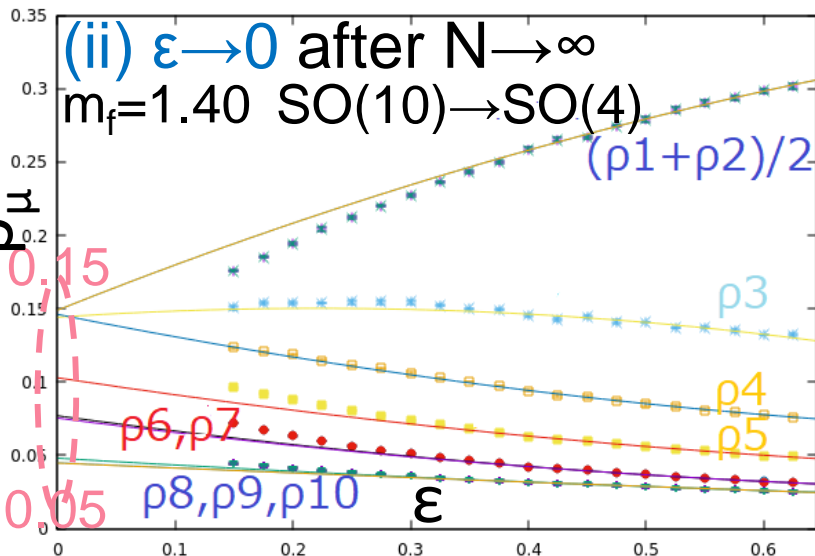
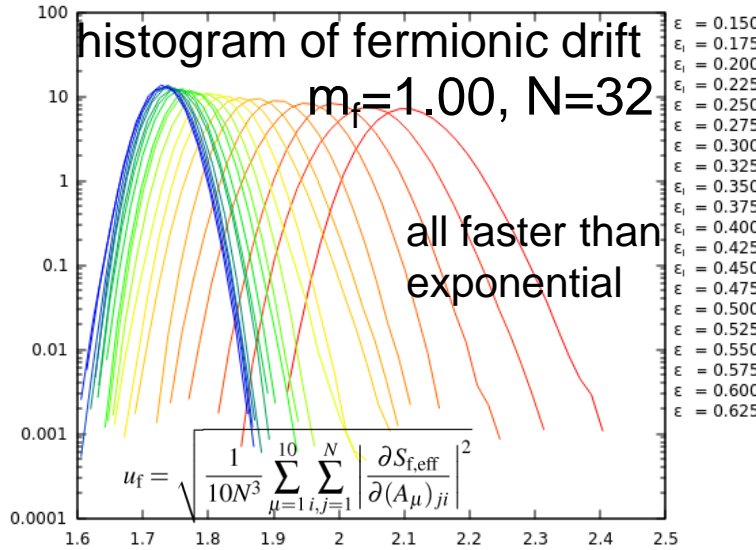
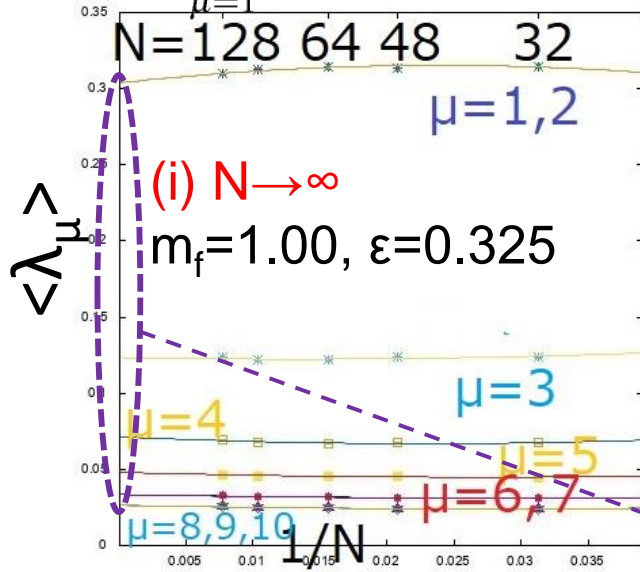
5. Result for D=10

$$m_\mu = (0.5, 0.5, 1, 2, 4, 8, 8, 8, 8, 8)$$

$$\Delta S_b = N \frac{\epsilon}{2} \sum_{\mu=1}^{10} m_\mu \text{tr}(A_\mu)^2$$

$$\Delta S_f = N m_f \text{tr} \left(\bar{\psi}_\alpha (i \Gamma_8 \Gamma_9^\dagger \Gamma_{10})_{\alpha\beta} \psi_\beta \right)$$

$$\rho_\mu(\epsilon, m_f) = \frac{\langle \lambda_\mu \rangle_{\epsilon, m_f}}{\sum_{\nu=1}^{10} \langle \lambda_\nu \rangle_{\epsilon, m_f}}$$



5. Result for D=10

$$m_\mu = (0.5, 0.5, 1, 2, 4, 8, 8, 8, 8, 8)$$

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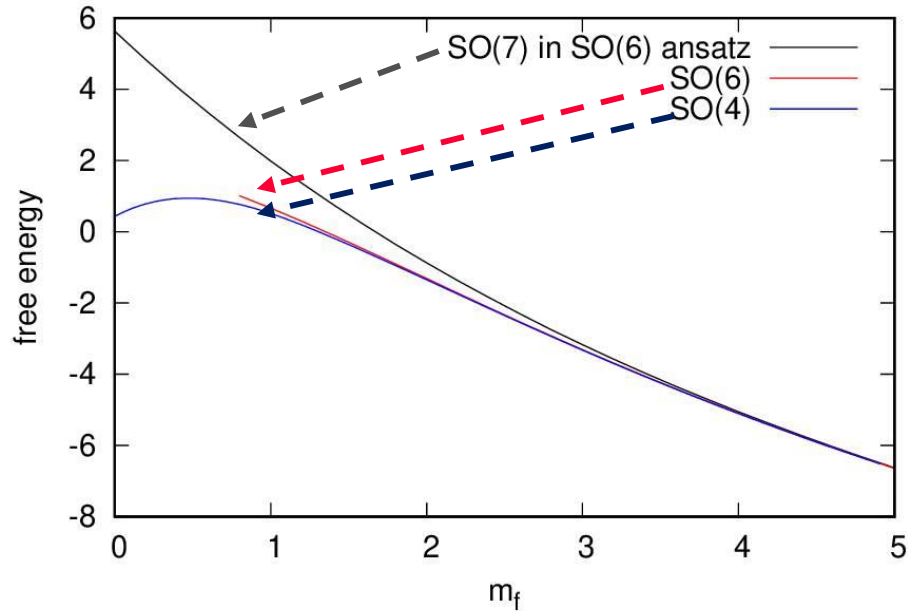
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$$\Delta S_f = Nm_f \text{tr} \left(\bar{\psi}_\alpha (i\Gamma_8 \Gamma_9^\dagger \Gamma_{10})_{\alpha\beta} \psi_\beta \right)$$

(no $\Delta S_b = N \frac{\epsilon}{2} \sum_{\mu=1}^{10} m_\mu \text{tr}(A_\mu)^2$ term)

GEM result for $m_f > 0$ at 3 loop solutions of SO(4) and SO(6) ansatz

$$\rho_\mu(\epsilon, m_f) = \frac{\langle \lambda_\mu \rangle_{\epsilon, m_f}}{\sum_{\nu=1}^{10} \langle \lambda_\nu \rangle_{\epsilon, m_f}}$$



Favors lower symmetry at smaller m_f .

Dynamical compactification of the spacetime in the Euclidean type IIB matrix model.

"Complex Langevin Method" \Rightarrow trend of $SO(D) \rightarrow SO(3)$.

Future works

Application of CLM to other cases

Lorentzian version of the type IIB matrix model

generalization to Gross-Witten-Wadia model

$$S_g = N(\text{atr}U + \text{btr}U^\dagger) \quad [\text{P. Basu, K. Jaswin and A. Joseph, arXiv:1802.10381}]$$

supersymmetric quantum mechanics

[A. Joseph and A. Kumar, arXiv:1908.04153]

Example [G. Aarts, arXiv:1512.05145]

$S(x) = \frac{1}{2} \underbrace{(a+ib)}_{=\sigma} x^2$, ($a, b \in \mathbf{R}$, $a > 0$) $S(x)$ is complex for **real** x .
Complexify to **$z=x+iy$** .

$$S(z) = \frac{1}{2} \sigma z^2 = \frac{1}{2} (a+ib) \overbrace{(x+iy)^2}^{=z^2} = \frac{a(x^2 - y^2)}{2} + ibxy, \quad \frac{\partial S}{\partial z} = \sigma z = (a+ib)(x+iy)$$

Complex Langevin equation for this action

$$\dot{x}(t) = -\operatorname{Re} \left(\frac{\partial S}{\partial z} \right) + \eta(t) = (-ax + by) + \eta(t) \quad \dot{y}(t) = -\operatorname{Im} \left(\frac{\partial S}{\partial z} \right) = (-ay - bx)$$

The **real** white noise satisfies

$$\langle \eta(t_1) \eta(t_2) \rangle = 2\delta(t_1 - t_2) \quad \langle \dots \rangle = \frac{\int \mathcal{D}\eta \dots \exp \left(-\frac{1}{4} \int \eta^2(t) dt \right)}{\int \mathcal{D}\eta \exp \left(-\frac{1}{4} \int \eta^2(t) dt \right)}$$

Solution of the Langevin equation

$$x(t) = e^{-at} \underbrace{[x(0) \cos bt + y(0) \sin bt]}_{=A(t)} + \int_0^t \eta(s) e^{-a(t-s)} \cos[b(t-s)] ds$$

$$y(t) = e^{-at} [y(0) \cos bt - x(0) \sin bt] - \int_0^t \eta(s) e^{-a(t-s)} \sin[b(t-s)] ds$$

$$\begin{aligned} \langle x^2 \rangle &= \lim_{t \rightarrow +\infty} \langle x^2(t) \rangle = \lim_{t \rightarrow +\infty} \left\{ \underbrace{e^{-2at} A(t)^2}_{\rightarrow 0} + 2e^{-at} A(t) \int_0^t \underbrace{\langle \eta(s) \rangle}_{=0} e^{-a(t-s)} \cos[b(t-s)] ds \right. \\ &\quad \left. + \int_0^t \int_0^t \underbrace{\langle \eta(s) \eta(s') \rangle}_{=2\delta(s-s')} e^{-a(2t-s-s')} \cos[b(t-s)] \cos[b(t-s')] ds ds' \right\} \\ &= \lim_{t \rightarrow +\infty} \left\{ 2 \int_0^t e^{-2a(t-s)} \cos^2[b(t-s)] \right\} ds = \frac{2a^2 + b^2}{2a(a^2 + b^2)} \end{aligned}$$

Similarly, $\langle y^2 \rangle = \frac{b^2}{2a(a^2 + b^2)}$, $\langle xy \rangle = \frac{-b}{2(a^2 + b^2)}$

This replicates $\langle z^2 \rangle = \langle x^2 \rangle - \langle y^2 \rangle + 2i\langle xy \rangle = \frac{a - ib}{a^2 + b^2} = \frac{1}{\sigma}$

Fokker-Planck equation

$$\frac{\partial P}{\partial t} = L^\top P \quad \text{where} \quad L^\top = \frac{\partial}{\partial x} \left\{ \underbrace{\operatorname{Re} \left(\frac{\partial S}{\partial z} \right) + \frac{\partial}{\partial x}}_{=ax-by} \right\} + \frac{\partial}{\partial y} \left\{ \underbrace{\operatorname{Im} \left(\frac{\partial S}{\partial z} \right)}_{=ay+bx} \right\}$$

Ansatz for its static solution:

$$P(x, y) = N \exp(-\alpha x^2 - \beta y^2 - 2\gamma xy) = N \exp \left(-\beta \left(y + \frac{\gamma x}{\beta} \right)^2 - \underbrace{\left(\alpha - \frac{\gamma^2}{\beta} \right)}_{=a(a^2+b^2)/(2a^2+b^2)} x^2 \right)$$

$$0 = \partial_t P = L^\top P = \left[\underbrace{(2a - 2\alpha)}_{=0 \rightarrow a=\alpha} + x^2 \underbrace{(4\alpha^2 - 2a\alpha - 2b\gamma)}_{=0 \rightarrow \gamma=a^2/b} + y^2 \underbrace{(4\gamma^2 + 2b\gamma - 2a\beta)}_{=0 \rightarrow \beta=a(1+2a^2/b^2)} + xy \underbrace{(4(2\alpha - a)\gamma + 2b(\alpha - \beta))}_{=0} \right] P$$

Using $\frac{\int_{-\infty}^{+\infty} t^2 e^{-At^2} dt}{\int_{-\infty}^{+\infty} e^{-At^2} dt} = \frac{1}{2A}$ ($A > 0$) we have

$$\langle x^2 \rangle = \frac{\iint x^2 P(x, y) dx dy}{\iint P(x, y) dx dy} = \frac{1}{2} \cdot \frac{a(a^2 + b^2)}{2a^2 + b^2} = \frac{2a^2 + b^2}{2a(a^2 + b^2)}$$