Monte Carlo studies of the rotational symmetry breaking in dimensionally reduced super Yang-Mills models

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1 Introduction and the model

Matrix models as a constructive definition of superstring theory

IKKT model (IIB matrix model)

 \Rightarrow Promising candidate for the constructive definition of superstring theory.

N. Ishibashi, H. Kawai, Y. Kitazawa and A. Tsuchiya, hep-th/9612115

$$S = \underbrace{-rac{N}{4} ext{tr} \left[A_{\mu}, A_{
u}
ight]^2}_{=S_B} + \underbrace{rac{N}{2} ext{tr} \, ar{\psi}_lpha(\Gamma_\mu)_{lphaeta} [A_\mu, \psi_eta]}_{=S_F}.$$

- The original IKKT model is ten-dimensional $(\mu, \nu = 1, 2, \dots, D = 10)$. \Rightarrow In the following, we study the simplified six-dimensional model $(\mu, \nu = 1, 2, \dots, D = 6)$.
- SO(6) rotational symmetry and SU(N) gauge symmetry.
- Presence of $\mathcal{N} = 2$ supersymmetry.

•
$$Z = \int dAd\psi d\bar{\psi} e^{-S} = \int dAe^{-S_B} \underbrace{(\det \mathcal{M})}_{=\int d\psi d\bar{\psi} e^{-S_F}} = \int dAe^{-S_0} e^{i\Gamma}.$$

 $Z_0 = \int dAe^{-S_B} |\det \mathcal{M}| = \int dAe^{-S_0}$: phase-quenched model, Γ : complex phase det \mathcal{M} is complex \rightarrow Complex phase is important in SO(6) breakdown.

Results of Gaussian Expansion Method (GEM)

T. Aoyama, J. Nishimura and T. Okubo, arXiv:1007.0883, J. Nishimura, T. Okubo and F. Sugino arXiv:1108.1293 Observable for probing dimensionality : $T_{\mu\nu} = \frac{1}{N} \operatorname{tr} (A_{\mu}A_{\nu})$. $\lambda_n \ (n = 1, \dots, 6)$: eigenvalues of $T_{\mu\nu} \ (\lambda_1 \ge \lambda_2 \ge \dots \ge \lambda_6)$ Extended $d \dim (\lambda_1 = \dots = \lambda_d = R^2)$ and shrunken $(6 - d) \dim (\lambda_{d+1} = \dots = \lambda_6 = r^2)$:

- Universal "compactification" scale: $r^2 \simeq 0.223$, independent of d.
- Constant volume property: The volume is independent of d: $V = R^d \times r^{6-d} = l^6$ The extent of phase-quenched model Z_0 is $l^2 \simeq 0.627$.
- The free energy takes the minimum at d = 3
 - \rightarrow Spontaneous Symmetry Breakdown (SSB) of SO(6) to SO(3).



2 Monte Carlo simulation

Simulation of the phase-quenched model



Large-N limit $\rightarrow \langle \lambda_1 \rangle_0 = \langle \lambda_2 \rangle_0 = \cdots = \langle \lambda_6 \rangle_0 \simeq 0.6.$ $\langle * \rangle_0 = (V.E.V. \text{ of the phase-quenched model } Z_0).$ No SO(6) breakdown in the phase-quenched model. Consistent with the GEM result $\langle \lambda_{1,\dots,6} \rangle_0 = l^2 \simeq 0.627.$

Factorization method

An approach to overcome the overlap problem in Monte Carlo simulation.

K. N. Anagnostopoulos and J. Nishimura, hep-th/0108041,

J. Ambjorn, K. N. Anagnostopoulos, J. Nishimura and J. J. M. Verbaarschot, hep-lat/0208025.

K. N. Anagnostopoulos, T. Azuma and J. Nishimura, arXiv:1009.4504,1108.1534.

Overlap problem: Discrepancy of a distribution function between the phase-quenched model Z_0 and the full model Z.



Factorization property of the distribution function

 $(\tilde{\lambda}_n \stackrel{\text{def}}{=} \lambda_n / \langle \lambda_n \rangle_0$: deviation from $1 \Rightarrow$ effect of the phase)

$$egin{aligned} &
ho_n(x) \ \stackrel{ ext{def}}{=} \left< \delta(x- ilde{\lambda}_n)
ight>^{ ext{reweighting}} rac{\langle \delta(x- ilde{\lambda}_n) e^{i\Gamma}
angle_0}{\langle e^{i\Gamma}
angle_0} \ &= rac{1}{\langle e^{i\Gamma}
angle_0} imes \langle \delta(x- ilde{\lambda}_n)
angle_0} imes rac{\langle \delta(x- ilde{\lambda}_n) e^{i\Gamma}
angle_0}{\langle \delta(x- ilde{\lambda}_n)
angle_0} \ &= rac{1}{C}
ho_n^{(0)}(x) w_n(x) \ &rac{\langle \delta(x- ilde{\lambda}_n)
angle_0}{=w_n(x)} \end{aligned}$$

where

$$C = \langle e^{i\Gamma}
angle_0 = \langle \cos \Gamma
angle_0, ~~
ho_n^{(0)}(x) = \langle \delta(x - ilde{\lambda}_n)
angle_0, ~~w_n(x) = \langle e^{i\Gamma}
angle_{n,x} = \langle \cos \Gamma
angle_{n,x}, \ \langle *
angle_{n,x} = [ext{V.E.V. for the partition function} ~~Z_{n,x} = \int dA e^{-S_0} \delta(x - ilde{\lambda}_n)].$$

Simulation of partition function $Z_{n,x} \Rightarrow x$ is trapped at $\tilde{\lambda}_n$.

The system visits the configurations important for full partition function Z. Resolution of overlap problem.

Monte Carlo evaluation of $\langle \tilde{\lambda}_n \rangle$

 $w_n(x) > 0 \Rightarrow \langle \tilde{\lambda}_n \rangle$ is the minimum of the free energy $\mathcal{F}_n(x) = -\frac{1}{N^2} \log \rho_n(x)$: We solve $\mathcal{F}'_n(x) = 0$, namely $\frac{1}{N^2} f_n^{(0)}(x) = -\frac{d}{dx} \left\{ \frac{1}{N^2} \log w_n(x) \right\}$, where $f_n^{(0)}(x) \stackrel{\text{def}}{=} \frac{\partial}{\partial x} \log \rho_n^{(0)}(x)$. Scaling properties in the large-N limit:

$$rac{f_n^{(0)}(x)}{N^2} = \left\{egin{array}{ccc} \mathrm{O}(N^0) \;\; x < 1 & \ 1 \ \mathrm{O}\left(rac{1}{N}
ight) \;\; x > 1 & \ rac{1}{N^2}\log w_n(x) = \mathrm{O}(N^0) \end{array}
ight.$$

• Existence of the hard-core potential in $\frac{1}{N^2}f_n^{(0)}(x)$ at large N.

• $\frac{1}{N^2}\log w_n(x)$ becomes almost constant at large x.



Results

Around $x \simeq 1$, $\frac{1}{N} f_n^{(0)}(x)$ scales at large N: Ansatz around $x \simeq 1 \rightarrow \frac{x}{N} f_n^{(0)}(x) \simeq g(x) = \{a_{1,n}(x-1) + a_{2,n}(x-1)^2 + a_{3,n}(x-1)^3\}.$ At small $x, \, rac{1}{N^2} f_n^{(0)}(x) - rac{g(x)}{Nx} \propto rac{7-n}{2x}.$ $w_n(x) ext{ behaves as } \Phi_n(x) = \lim_{n o +\infty} rac{1}{N^2} \log w_n(x) \propto \left\{egin{array}{c} x^{7-n} & (x < 1) \ \cosh x & (x > 1) \end{array}
ight.$ $\begin{array}{c} N=24 \\ N=32 \\ -\Phi'(x) \end{array} \\ 0 \\ 0 \\ N=12 \end{array}$ f_{n=4}⁽⁰⁾ (x)/N² - g(x)/(Nx) 0.8 0.6 0.4 0.2 0 -0.014 . 6 0.4 0.8 2 4 10 0 0.2 0.6 1 0 8 Х Х Result for n = 4: The solution of $\frac{1}{N^2} f_n^{(0)}(x) = -\frac{d}{dx} \left\{ \frac{1}{N^2} \log w_n(x) \right\}$ is $x_s = 0.31(1)$. (c.f.) GEM result for shrunken directions: $\langle \tilde{\lambda}_n \rangle = \frac{\langle \lambda_n \rangle}{\langle \lambda_n \rangle_0} = \frac{r^2}{l^2} \simeq \frac{0.223}{0.627} = 0.355 \cdots$

Comparison of the free energy \rightarrow evaluate $\Delta = -\mathcal{F}_3(x_s) + \mathcal{F}_4(x_s)$.

 $\Delta < 0 \Rightarrow n = 3$ direction does not shrink \rightarrow SO(3) spacetime is preferred to SO(2).

$$\Delta = \frac{\log \rho_3(x_s) - \log \rho_4(x_s)}{N^2} = \frac{\log w_3(x_s) - \log w_4(x_s)}{N^2} - \underbrace{\int_{SO(2)}^{SO(3)} \frac{1}{N^2} \frac{d}{dx} \log \rho^{(0)}}_{\rightarrow 0(\text{as } N \rightarrow +\infty)}$$

$$rac{1}{N^2}\log w_n(x) ext{ for } n=3,4,5,6, \ N=24:$$
 Very subtle, but $rac{1}{N^2}\log w_{n=4}(x_s)$ is the highest at $x_s\simeq 0.3.$

Constant volume property

Geometric mean
$$L^2 = \left(\prod_{i=1}^6 \langle \lambda_i \rangle_{n,x} \right)^{rac{1}{6}}$$
 for $n=3,4,5,6$:



 $L^2 \simeq 0.627$ at x = 1 (which corresponds to setting λ_n at the value $\langle \lambda_n \rangle_0$). Consistent with the GEM result $V = R^d \times r^{6-d} \simeq l^6$, where $l^2 \simeq 0.627$.

3 Conclusion

Monte Carlo simulation of 6d IKKT model \Rightarrow spontaneous breakdown of SO(6) symmetry.

- Numerical evidence for symmetry breakdown $SO(6) \rightarrow SO(3)$.
- Constant volume property.

Future works

- Simulation of larger $N \Rightarrow$ study the finite-N effect.
- Ultimately, we would like to study the 10d IKKT model.