The destruction and recreation of the X-ray corona in an accreting massive black hole

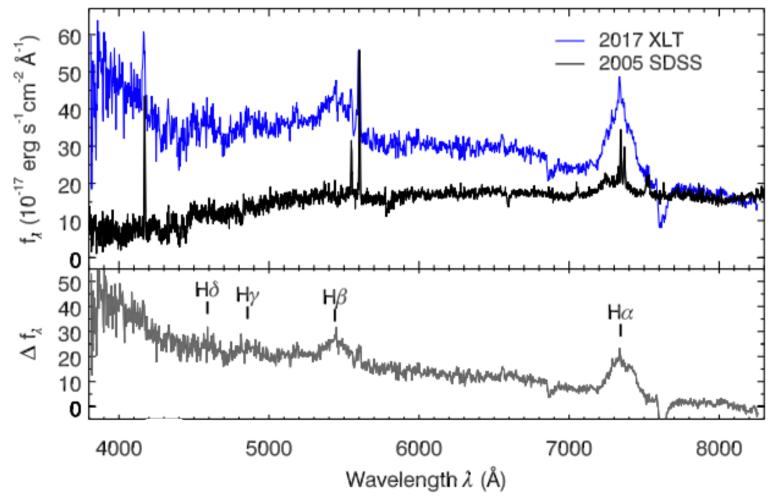
Claudio Ricci

Universidad Diego Portales, Chile Kavli Institute for Astronomy and Astrophysics, China George Mason University, USA

Erin Kara (MIT), Michael Loewenstein (NASA), Benny Trakhtenbrot (Tel Aviv U.), <u>lair Arcavi</u> (Tel Aviv U.), Ron Remillard (MIT), Andrew C. Fabian (Cambridge U.), Keith C. Gendreau (NASA), Zaven Arzoumanian (NASA), Ruancun Li (KIAA), Luis C. Ho (KIAA), Chelsea L. MacLeod (CFA), Ed Cackett (Wayne State U.), Diego Altamirano (Southampton U.), Poshak Gandhi (Southampton U.), Peter Kosec (Cambridge U.), <u>Dheeraj Pasham</u> (MIT), Jack Steiner (MIT), <u>Chi-Ho Chan</u> (Jerusalem U.)

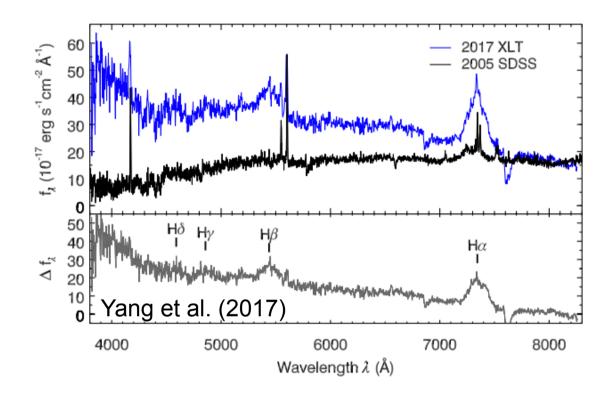
Changing-look AGN

Type 1 <--> Type 2



Yang et al. (2017)

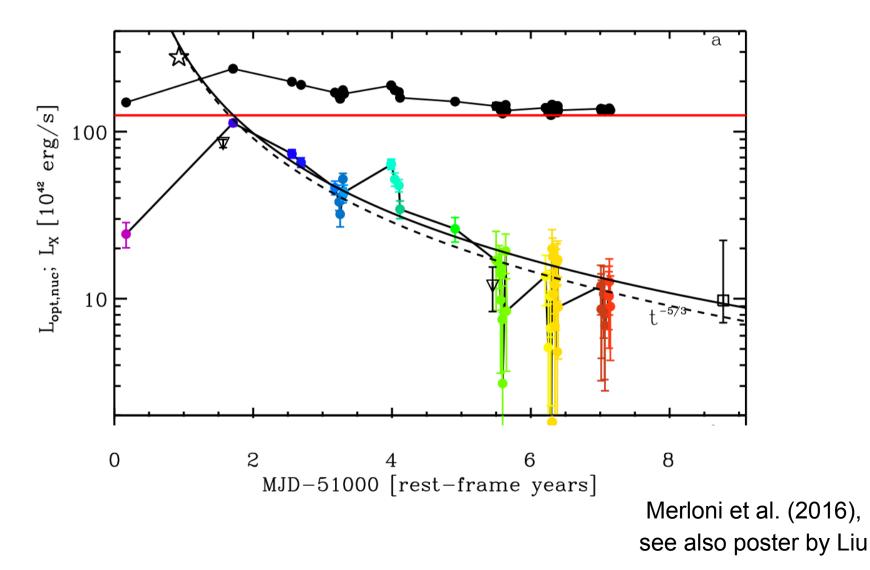
Mechanisms triggering Changing-look events



- State transitions (as in BH binaries; e.g. Noda+18)
- **Disk instabilities** (e.g. Stern+18, Ross+18)
- Tidal disruption events (e.g. Merloni+16)

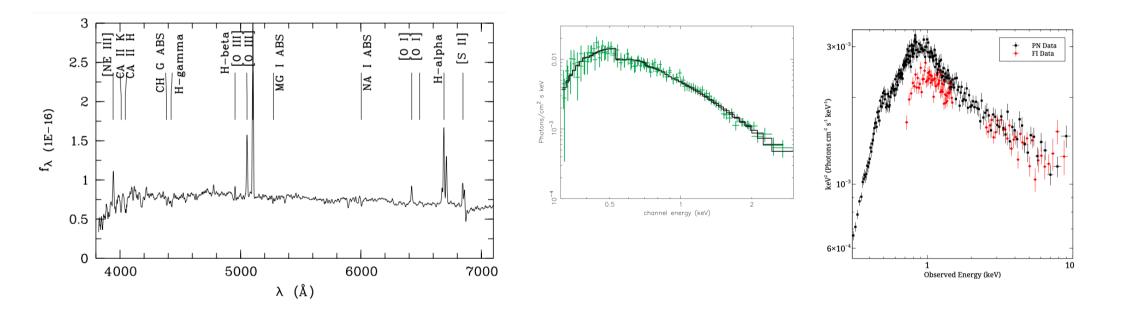
TDEs and changing-look AGN

The changing-look AGN SDSS J0159+0033





Source previously classified as an AGN both in the optical and in the X-rays (Lx~1e43 erg/s)



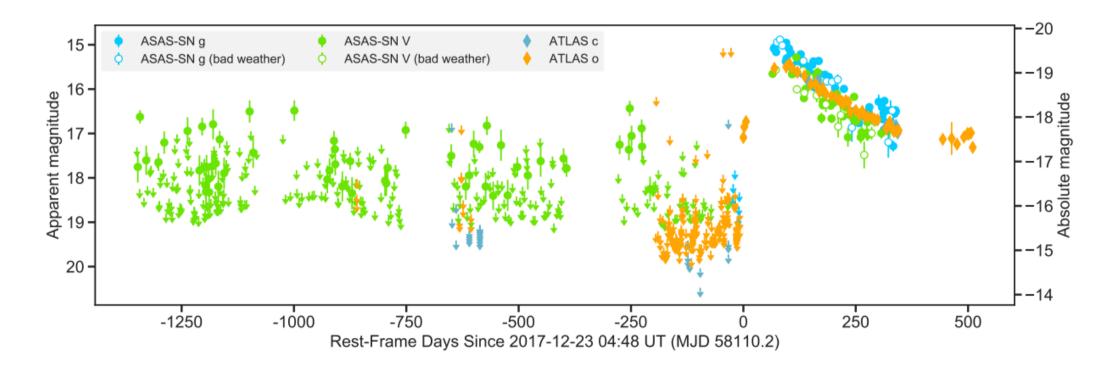
Boller et al. (2003)

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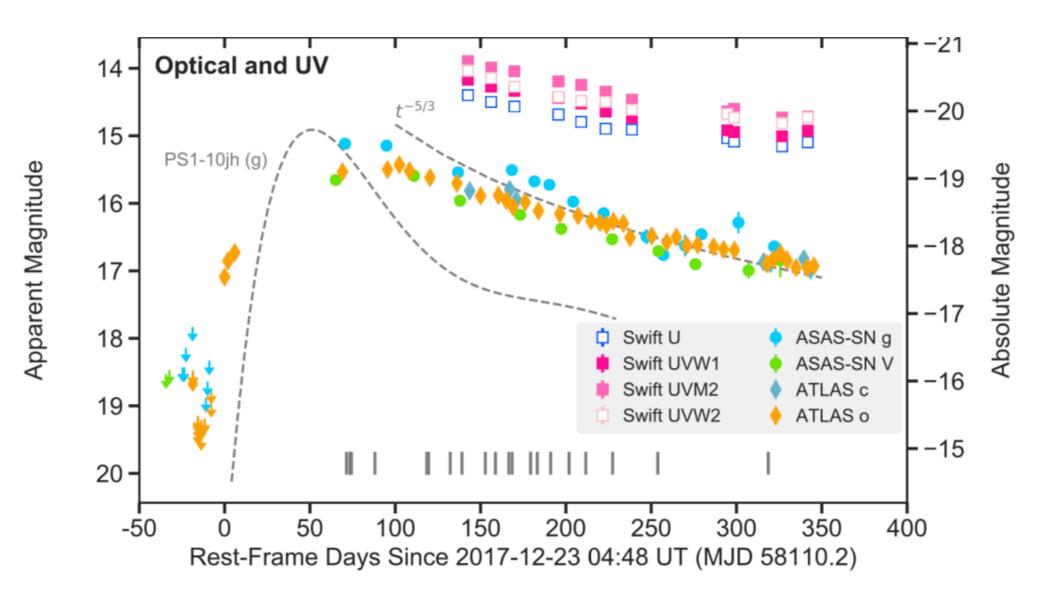
Gallo et al. (2013)



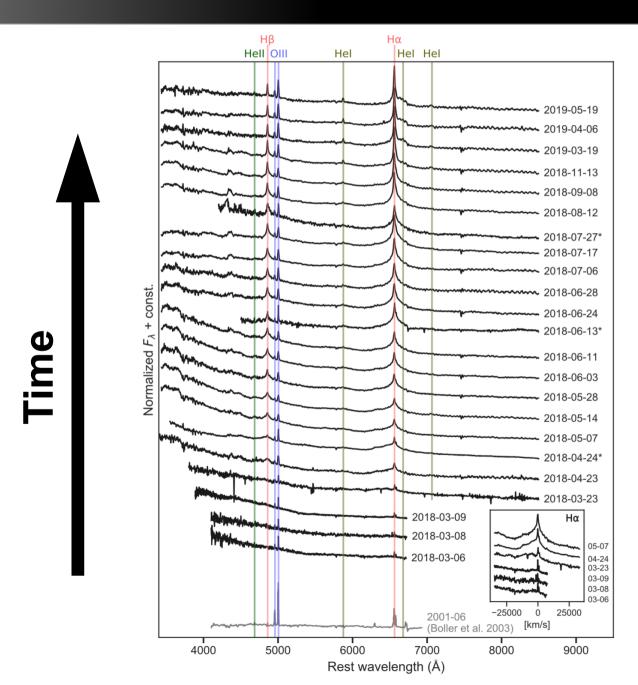
The optical/UV outburst of 1ES 1927+654



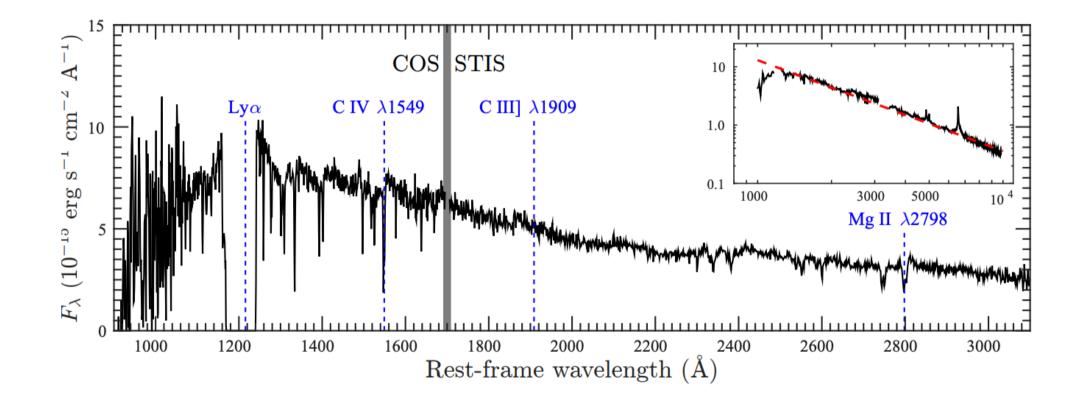
The optical/UV outburst of 1ES 1927+654



The changing-look AGN 1ES 1927+654



The UV spectrum

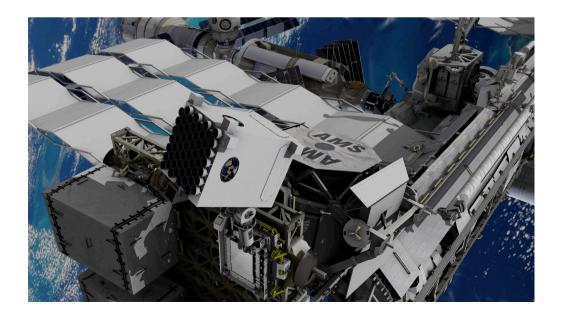


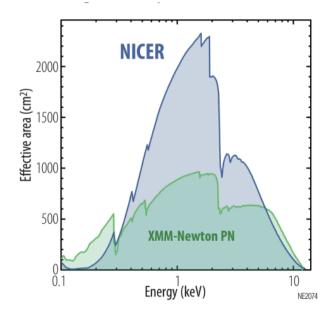
\sim 1.1 Ms (13 full days)

265 NICER (~700 ks)

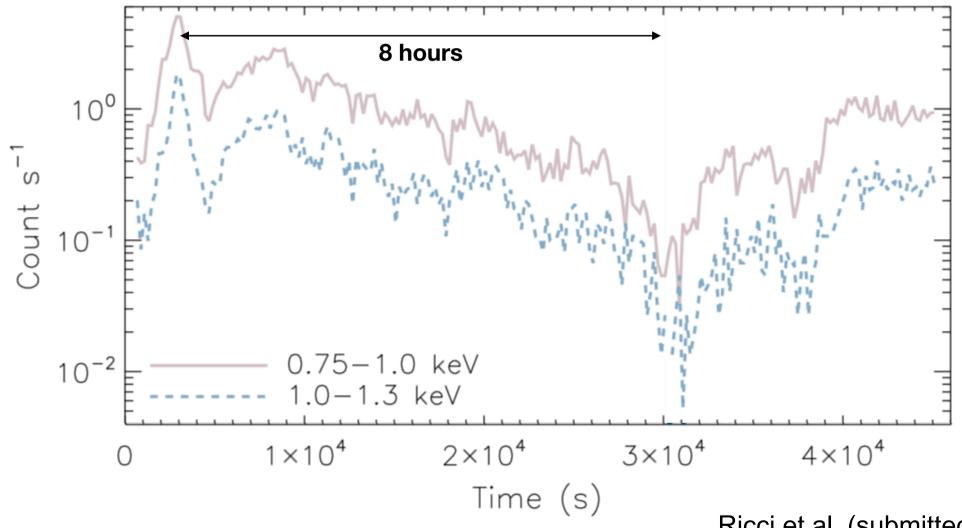
14 Swift (26 ks)

4 simultaneous XMM-Newton/NuSTAR (~400 ks)

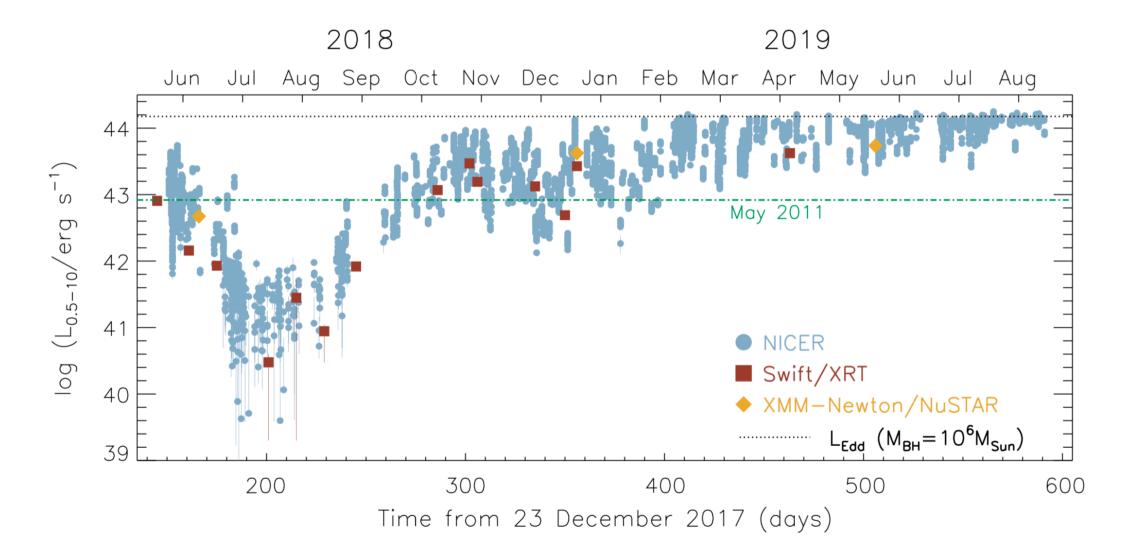




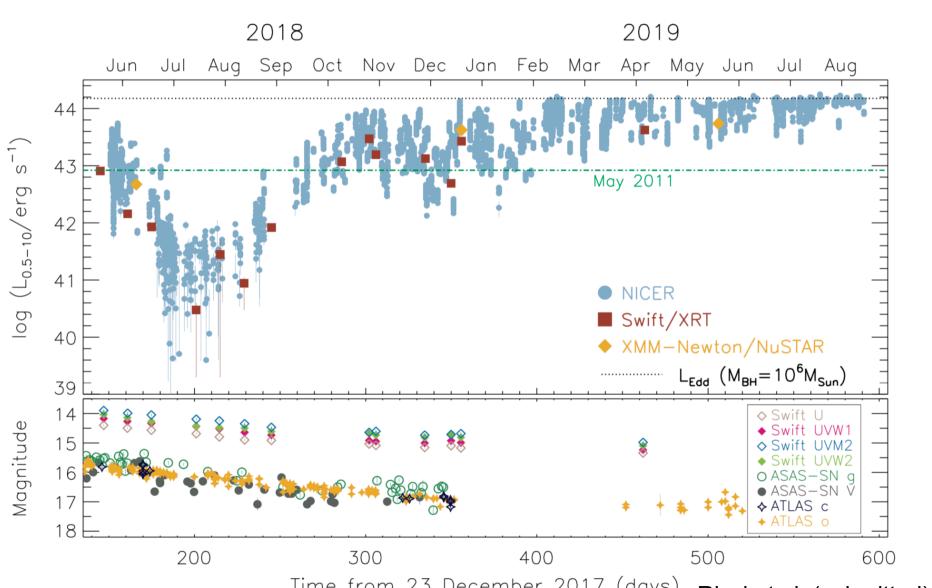
Extreme variability on short timescales..



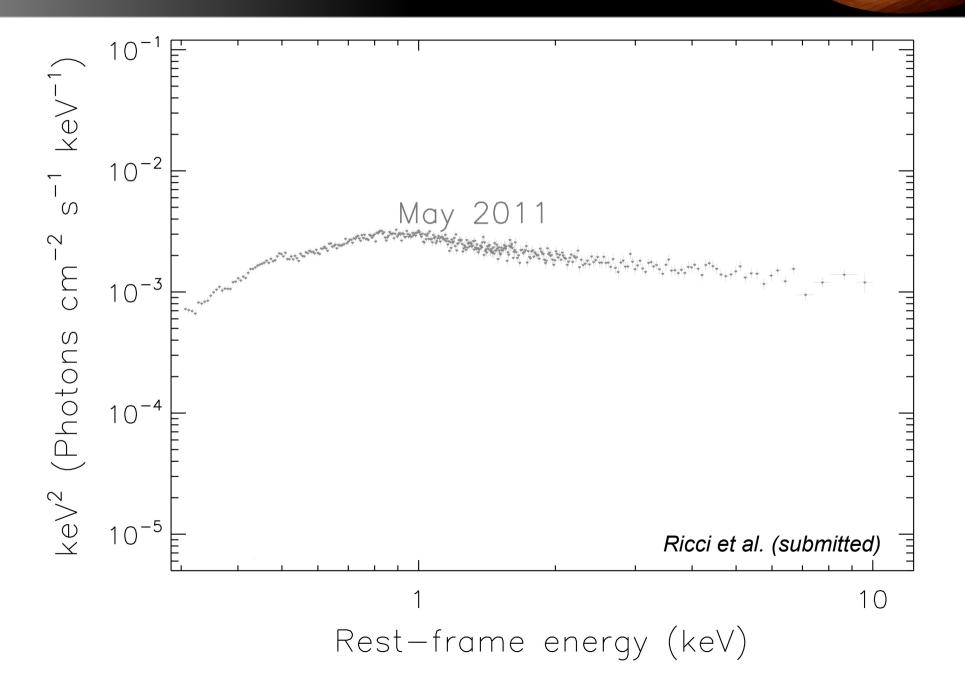
...and on long timescales

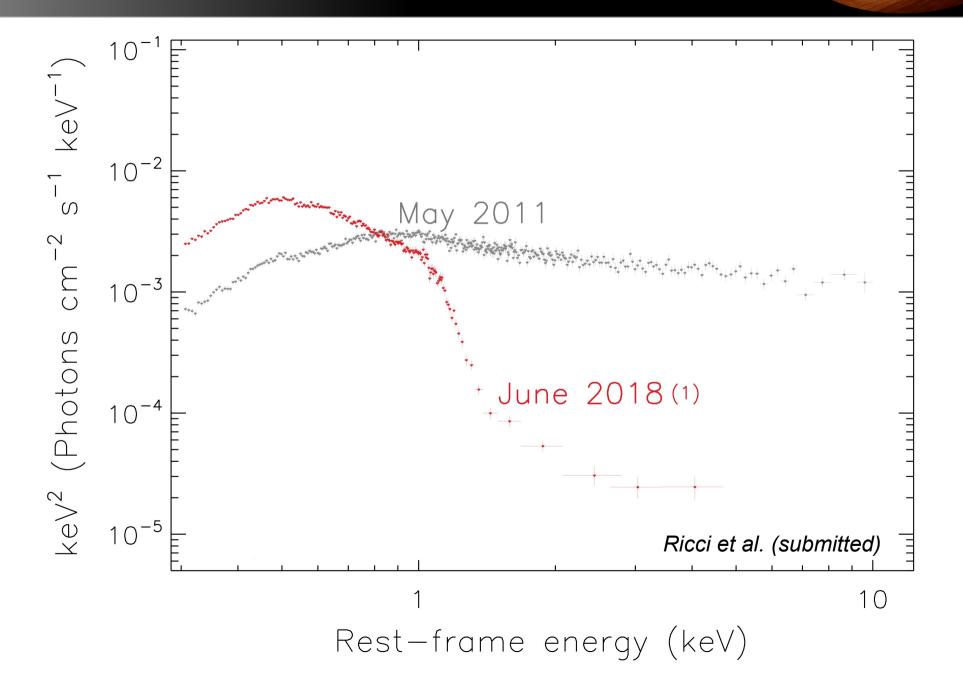


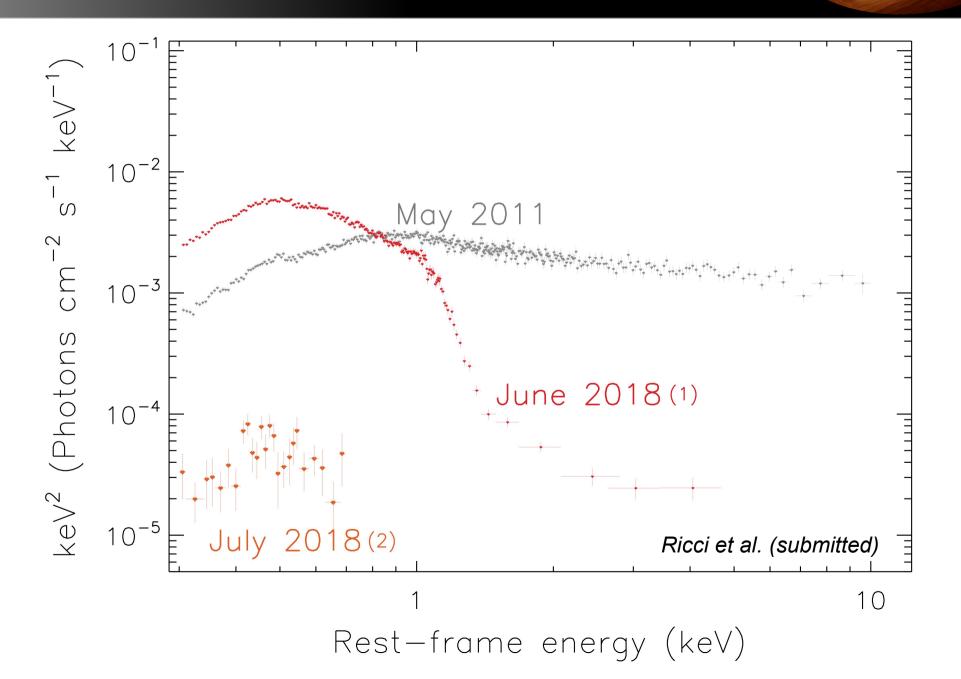
X-ray and UV variability are disconnected

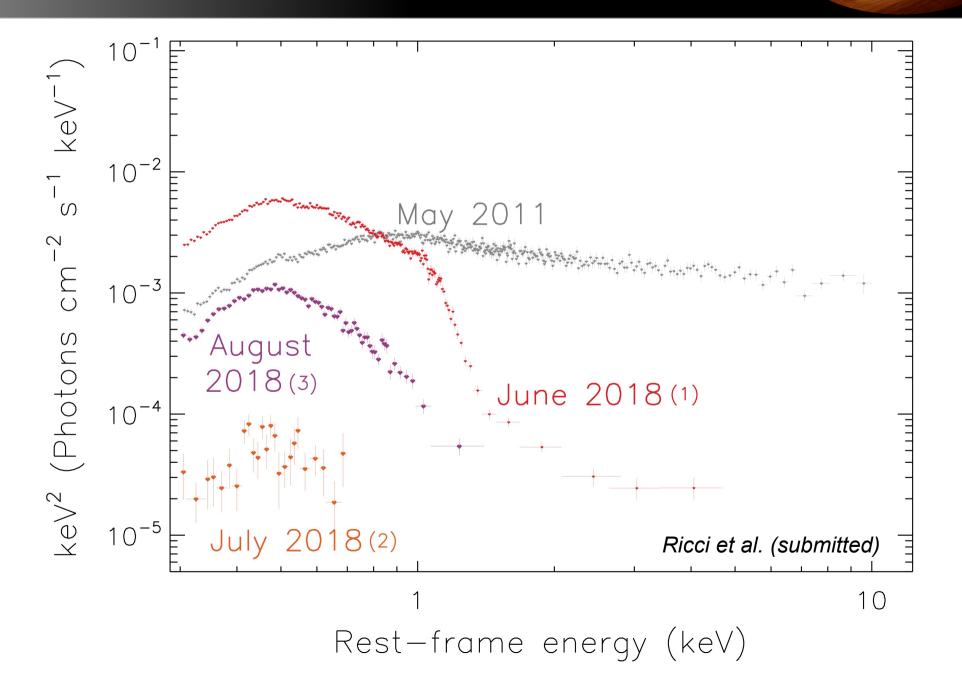


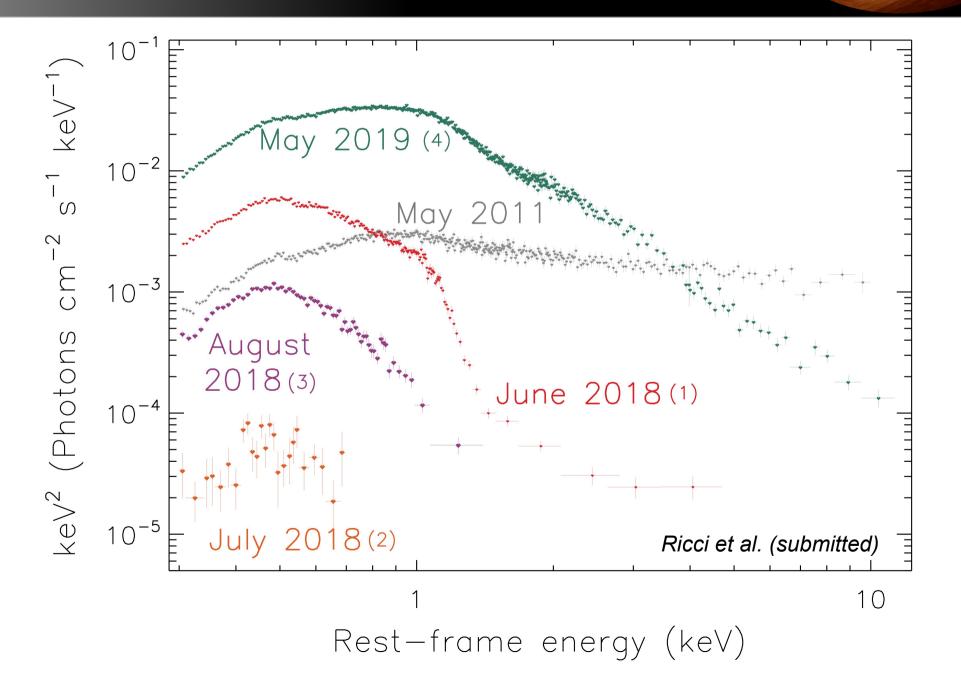
Time from 23 December 2017 (days) Ricci et al. (submitted)

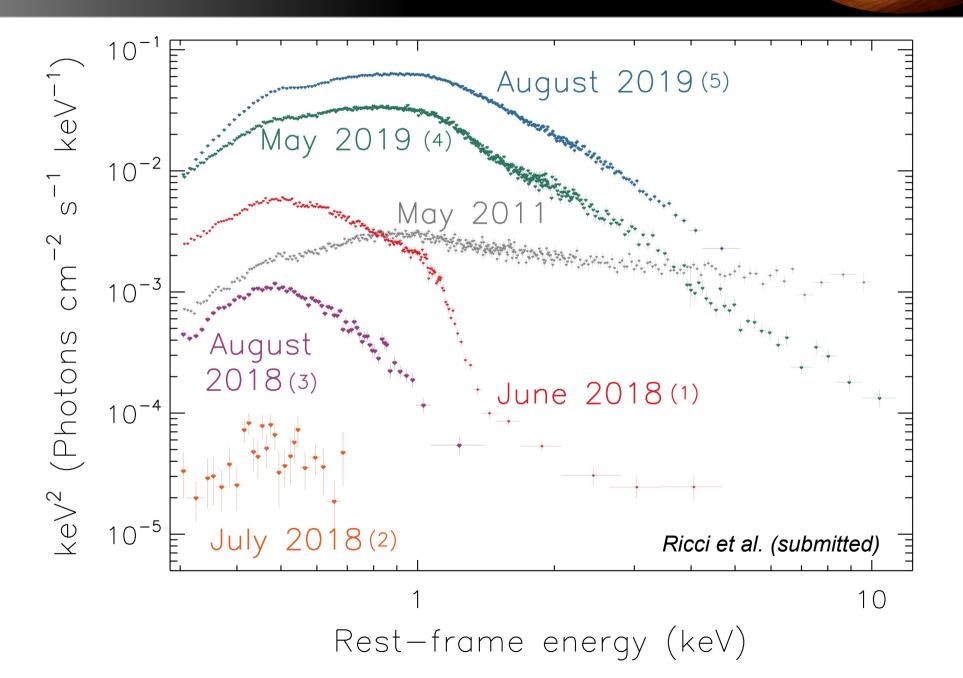




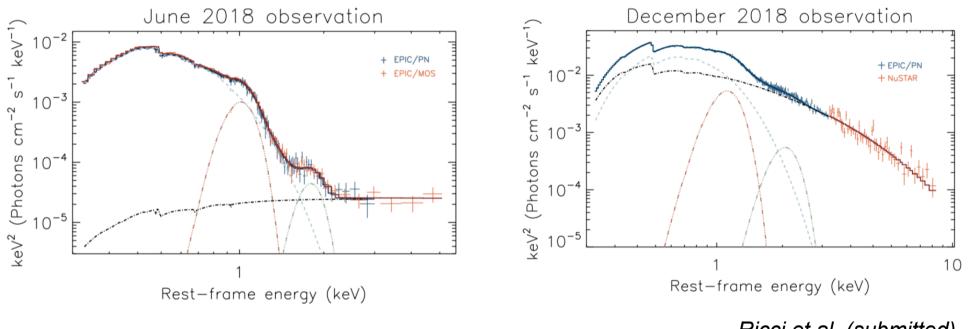








X-ray spectroscopy



Ricci et al. (submitted)

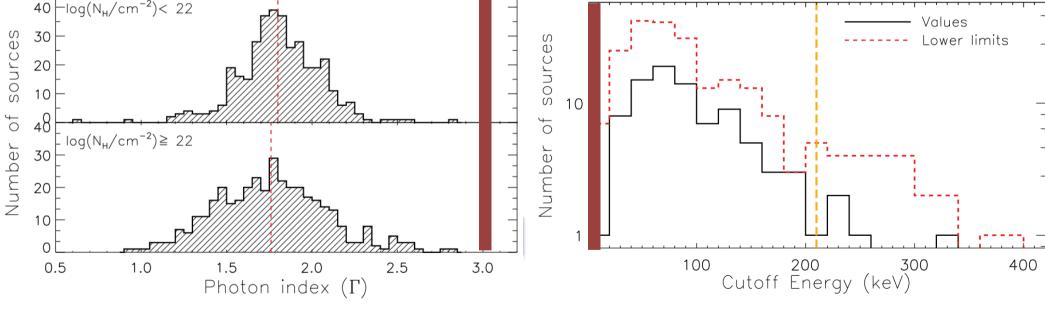
Very soft ($\Gamma \simeq 3$)

Very low energy cutoff! $(E_{\rm C} \simeq 2 - 3 \, \rm keV)$

X-ray spectroscopy



Cutoff energy $(E_{\rm C} \simeq 200 \, {\rm keV})$

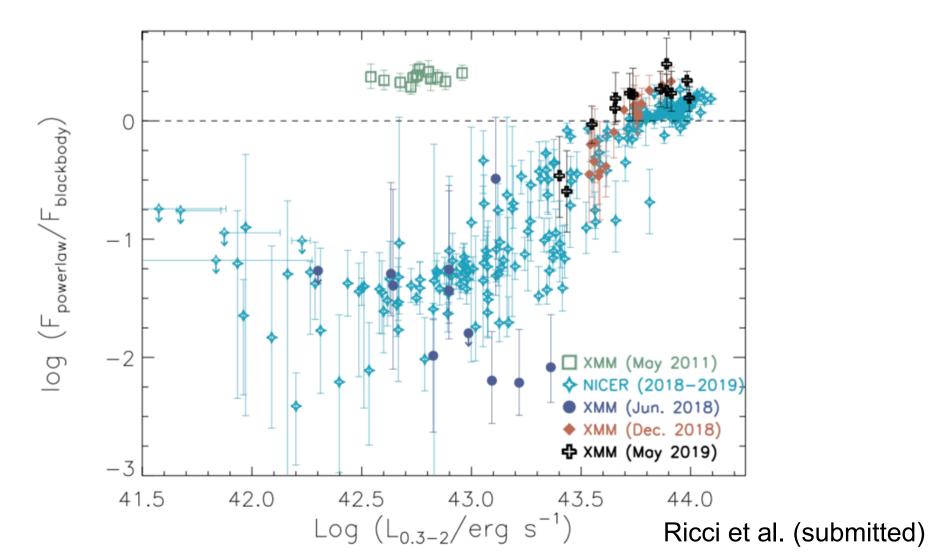


Ricci et al. 2017d

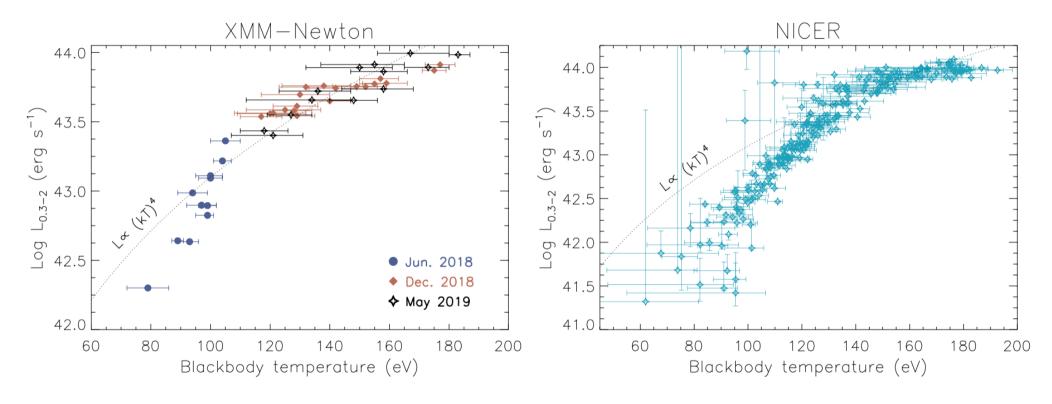
Ricci et al. 2018

X-ray spectral variability

Is the *normal* X-ray continuum reappearing?



X-ray spectral variability



Ricci et al. (submitted)

See also R. Saxton's talk, also Shu+17,18

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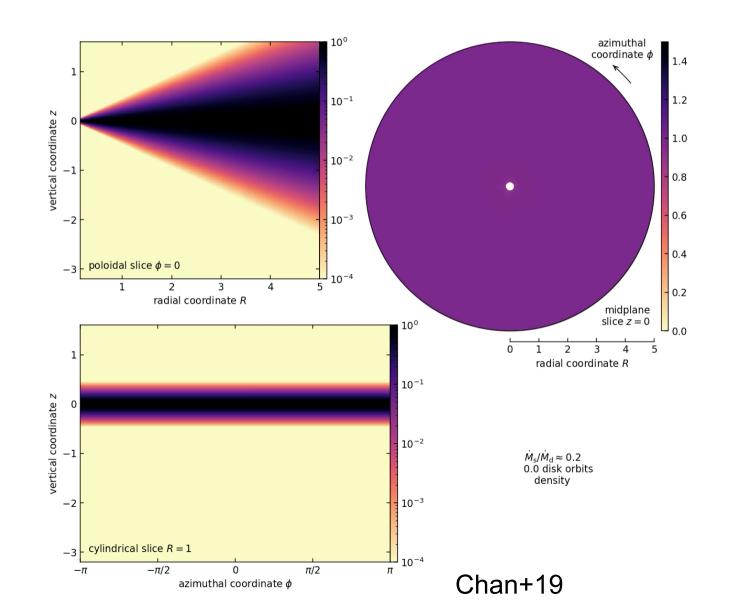
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- 3) After the event the emission was dominated by a blackbody component

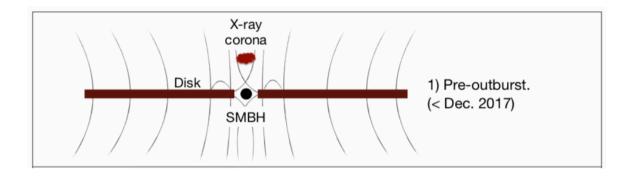
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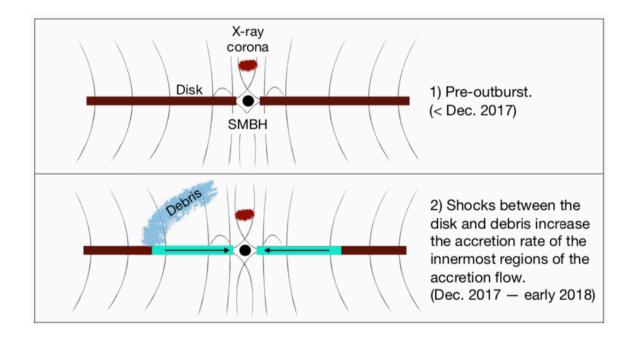
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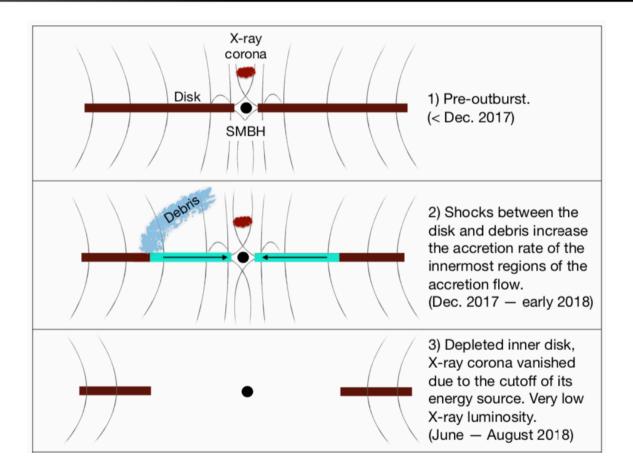
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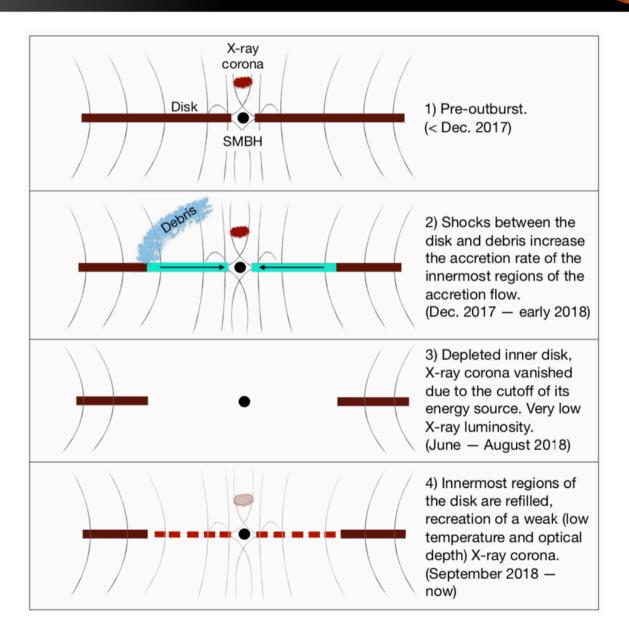
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- 7) No correlation between X-ray and UV variability















- The destruction of the X-ray corona and of the innermost regions of 1ES1927+654 could be caused by the interaction with a star

- Changing-look events in AGN can be associated with dramatic and quick transformations of the innermost regions of accreting SMBHs (and TDEs?)

- Future studies with *eRosita* and the *Einstein probe* might find many more objects such as 1ES 1927+654