

# Precision measurements of the $^{24}\text{Mg}(\alpha, p\gamma)^{27}\text{Al}$ cross section and $^{27}\text{Al}(p, \alpha\gamma)^{24}\text{Mg}$ cross sections

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Type I X-ray bursts are important astrophysical events that occur in binary star systems where a H and He are accreted onto the surface of a neutron star. To understand the production of nuclei in these explosive environments and related to this, their energy generation and light curves, there have been a number of model calculations that span the parameter space given nuclear physics uncertainties [1]. A number of reactions were identified as having an impact on the X-ray burst energy generation, one of which is the  $^{24}\text{Mg}(\alpha, p)^{27}\text{Al}$  reaction. The cross sections for this reaction used in the model calculations were based on the inverse reaction  $^{27}\text{Al}(p, \alpha)^{24}\text{Mg}$ . The data available for the direct reaction is scarce, especially the cross sections to excited states of  $^{27}\text{Al}$ . In order to more precisely determine the role of the  $^{24}\text{Mg}(\alpha, p)^{27}\text{Al}$  reaction, we performed an experiment to measure both the  $^{24}\text{Mg}(\alpha, p\gamma)^{27}\text{Al}$  and  $^{27}\text{Al}(p, \alpha\gamma)^{24}\text{Mg}$  reactions at the University of Notre Dame using 5U Van de Graaff accelerator and HAGRID, an array of LaBr<sub>3</sub> gamma-ray detectors to detect secondary  $\gamma$  rays from the excited states of  $^{24}\text{Mg}$  and  $^{27}\text{Al}$ . An R-matrix analysis of the differential cross sections will be presented and the influence of these cross sections on previously calculated resonance strengths will be discussed.

[1] Parikh et al., *Astrophysical Journal Suppl.* 178, 110 (2008).

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