## Evaluation of the neutron capture reaction on <sup>79</sup>Se via a surrogate reaction of $d(^{79}Se, p)$ reaction at OEDO

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<sup>79</sup>Se placed on the path of s-process nucleo-synthesis is important as a stellar thermometer because of the direct  $\beta$  decay channel from the isomeric state in <sup>79</sup>Se [1]. The neutron capture reaction cross section on the nucleus must be evaluated to determine the reaction path more directly. However, it has never been attempted due to difficulty to fabricate a radioactive target of <sup>79</sup>Se.

For the  $(n,\gamma)$  reaction on <sup>79</sup>Se at the astrophysical energy, the compound reaction channel is dominant. Up to date, as an indirect method to determine the  $(n,\gamma)$  cross section of compound reaction channel, the surrogate technique in which the corresponding compound state is populated via an alternative transfer reaction has been employed with measuring the  $\gamma$  rays from the compound state to determine the  $\gamma$  emission channel [2]. Although many studies to determine the transition path from the unbound state have been carried out, the final result was model-dependent.

Instead of measuring  $\gamma$  rays, we have measured the (d, p) reaction on <sup>79</sup>Se in inverse kinematics with the OEDO beam line which can provide the low-energy RI beam with a small beam spot size at RIBF [3]. The recoiled particles were analyzed with the SSD-CsI(Tl) telescope array TiNA, which determined the excitation energy of the compound state. The residual nuclei were analyzed via the SHARAQ spectrometer. The  $\gamma$  emission channel was deduced by comparing the amount of <sup>80</sup>Se with all the reaction products of (d, p) reaction because  $\gamma$  emission means that the residual nucleus didn't change either proton nor neutron number. With the experimental  $\gamma$  emission probability, the neutron capture reaction on <sup>79</sup>Se were evaluated as a function of the neutron energy. In this talk, the detailed experimental setup and result will be discussed.

- [1] F. Kappler, H. Beer, and K. Wisshak, Rep. of Prog. in Phys.52, ('89) 945-1013
- [2] A. Ratkiewicz, et al., PRL 122, 052502 (2019) and references therein.
- [3] S. Michimasa et al., PTEP accepted (2019).