

Numerical study of Q-ball formation in gravity mediation

Takashi Hiramatsu

*Yukawa Institute for Theoretical Physics (YITP)
Kyoto University*

COLLABORATION WITH

Masahiro Kawasaki (ICRR)
Fuminobu Takahashi (IPMU)
Masahide Yamaguchi (TITech)

TH, Kawasaki, Takahashi, JCAP 06(2010)008 [arXiv:1003.1779]

TH, Takahashi, Yamaguchi, in preparation

- Affleck-Dine field parametrising “flat directions”

Affleck, Dine, NPB (1985)

- In MSSM, there are a large number of fields like quarks, leptons, gauges, Higgs, and their superpartners.

$$H_u = \begin{pmatrix} 0 \\ \Phi \end{pmatrix} \quad L = \begin{pmatrix} \Phi \\ 0 \end{pmatrix} \quad F = D = 0 \quad \therefore V(\Phi) = 0$$

- Global U(1) symmetry conserves baryon/lepton number

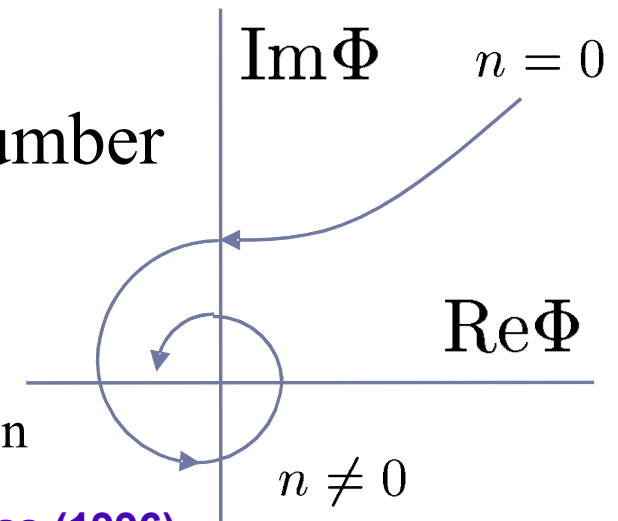
$$n_{B,L} = i\beta(\Phi^* \dot{\Phi} - \Phi \dot{\Phi}^*)$$

- Dynamical generation of baryon/lepton number

- Soft $SUSY$ terms and non-renormalisable terms

lift the potential, driving Φ toward the origin

- A-term like $\Phi^n + \Phi^{*n}$ kicks Φ to angular direction

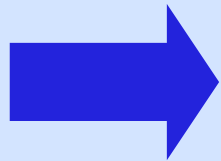


Dine, Randall, Thomas (1996)

Scalar field with global U(1) charge : $\mathcal{L} = |\partial_\mu \Phi|^2 - V(\Phi)$

If $\frac{V(\Phi)}{|\Phi|^2}$ has a minimum for $\Phi \neq 0$

Coleman, NPB (1985)



*non-topological soliton with a given charge, **Q-ball**, exists*

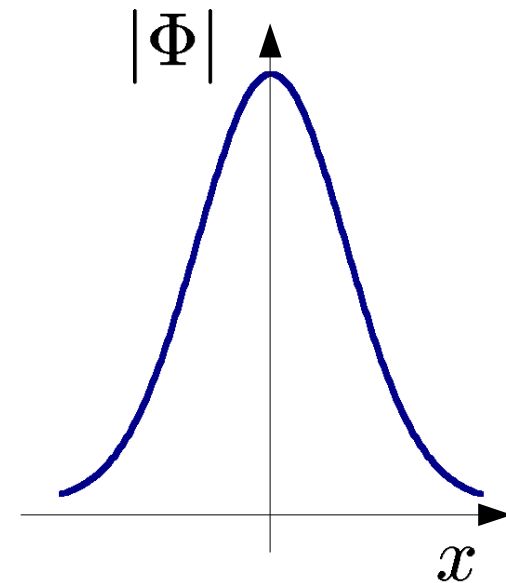
In cosmological context,

- ✓ dark matter candidate
- ✓ baryon/lepton number inside Q-balls protected from spharelon process



decay rate, evaporation rate, etc....

*crucially depends on **charge***



Field equation and potential of Affleck-Dine field in gravity mediation

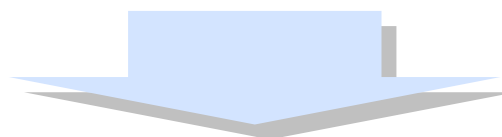
$$\ddot{\Phi} + 3H\dot{\Phi} - \frac{1}{a^2}\nabla^2\Phi = -V'(\Phi)$$

$$V(\Phi) = m^2|\Phi|^2 \left[1 + K \log \left(\frac{|\Phi|^2}{M_*^2} \right) \right] - cH^2|\Phi|^2 + (\text{N.R.})$$

1-loop correction from gauginos

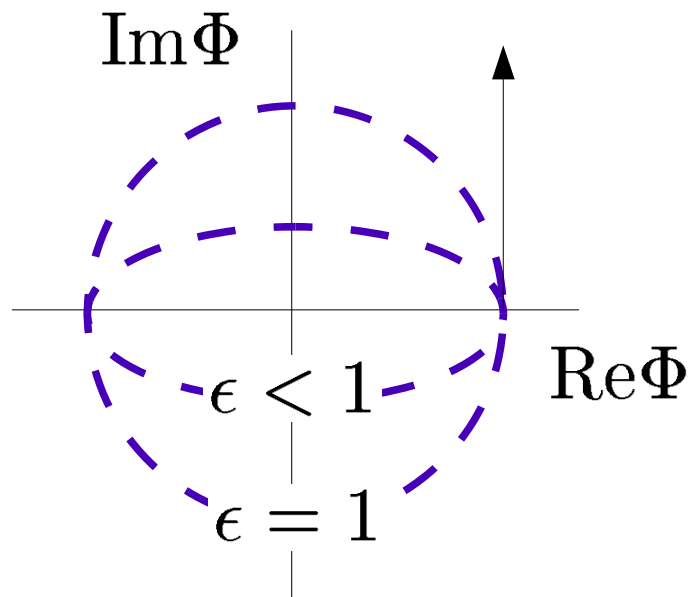
$$K = -0.1 \sim -0.01$$

Enqvist, McDonald, PLB(1998)



$V(\Phi)/|\Phi|^2$ has a minimum at $\Phi \neq 0$

Hence this system has Q-ball solution



Initial condition

(situation after starting to rotate in the phase space)

$$\Phi_{in} = M_*$$

$$\dot{\Phi}_{in} = imM_*\epsilon$$

adding small fluctuations as seed of Q-balls

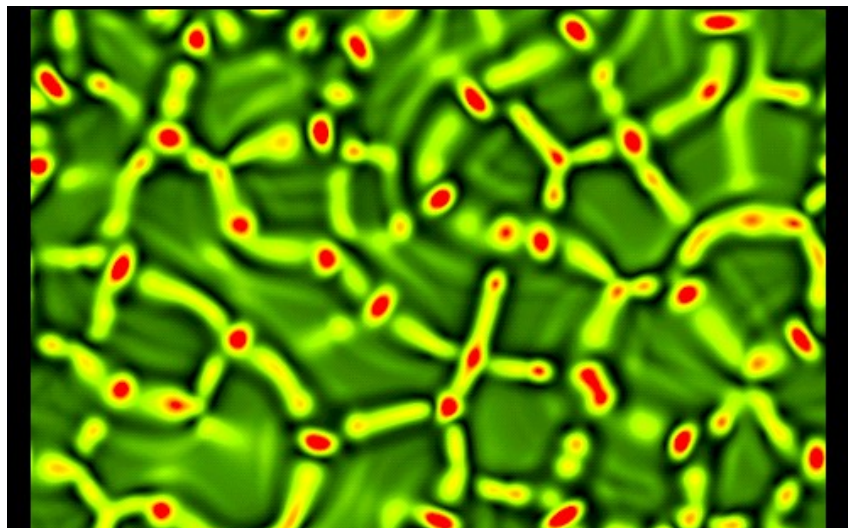
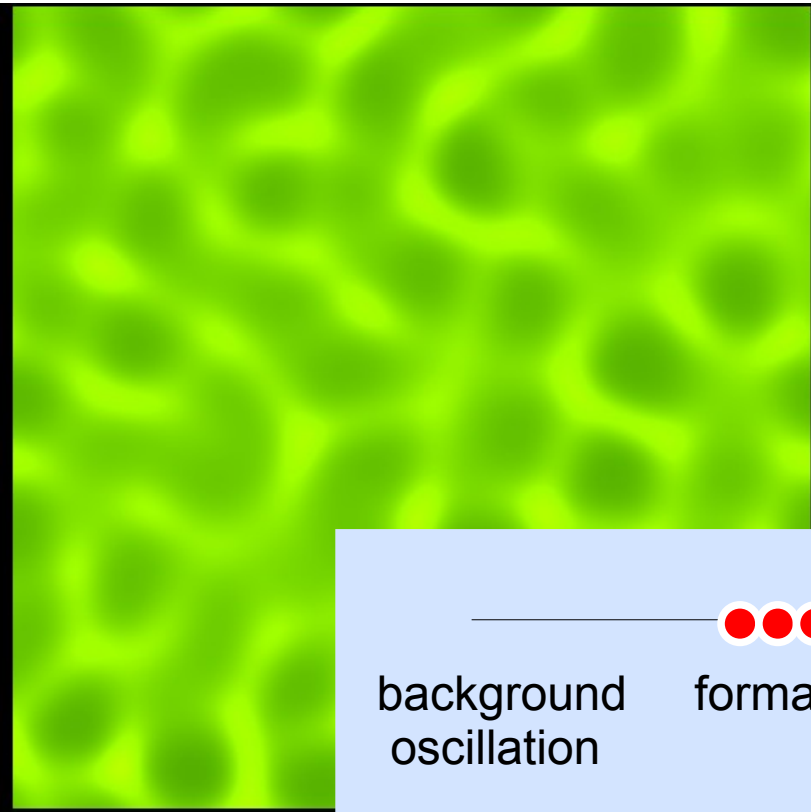
$$\left| \frac{\delta\Phi}{\Phi} \right|_{in} = O(10^{-7})$$

Kasuya, Kawasaki, PRD (2000)

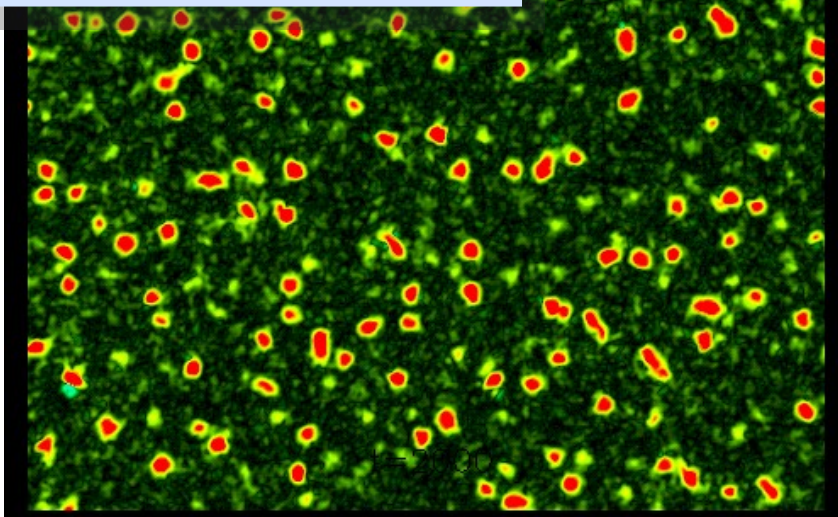
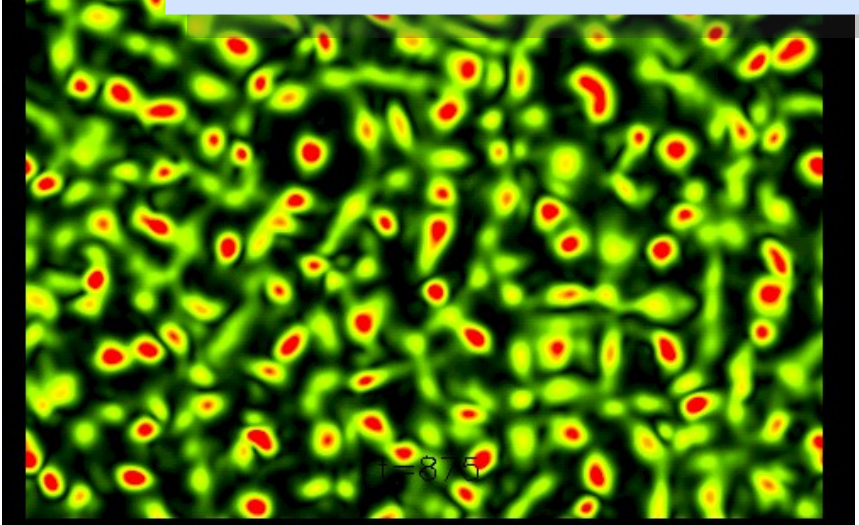
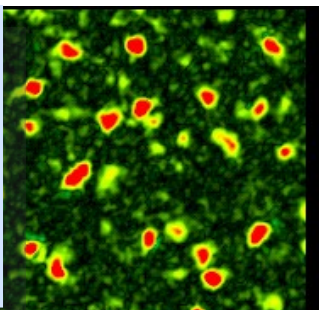
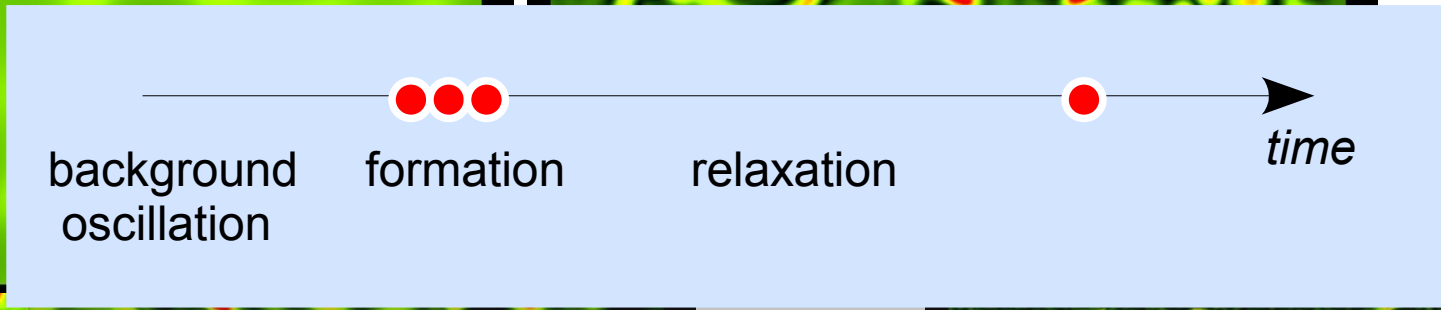
6th-order symplectic integrator by Yoshida (time)+ finite difference (space)
(supported by Aphrodite code)

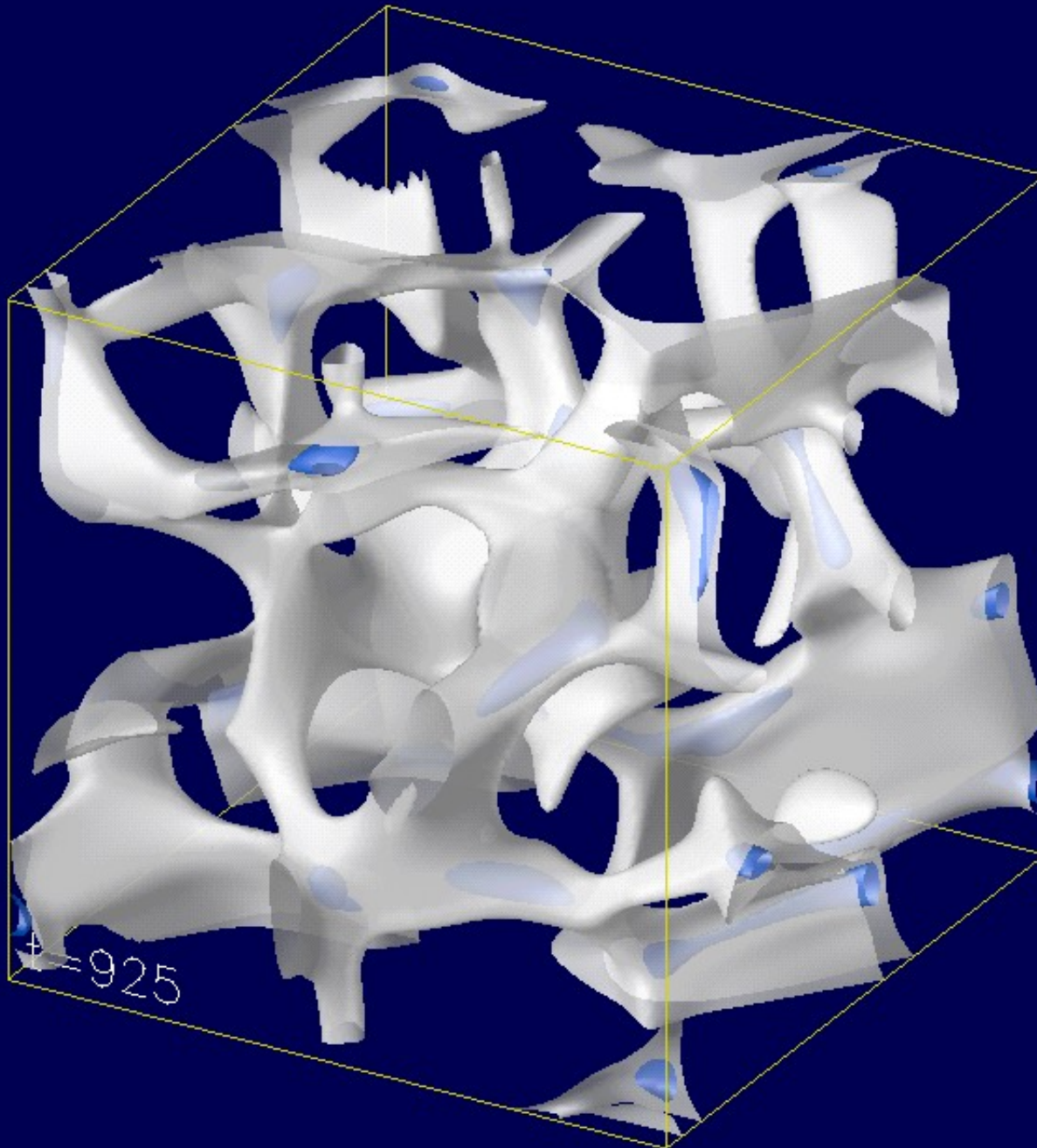
Regarding a region where $|q(t, x)| > q_c$ as a Q-ball with $q_c = q(t_{form})/5$

Result : 2D

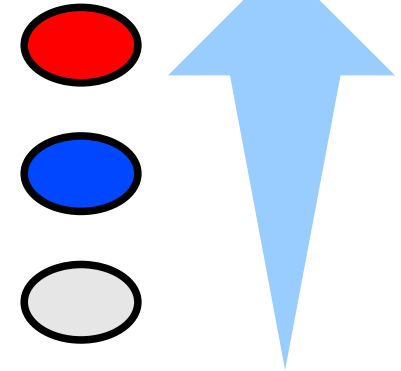


$$\epsilon = 1$$
$$N = 512^2$$





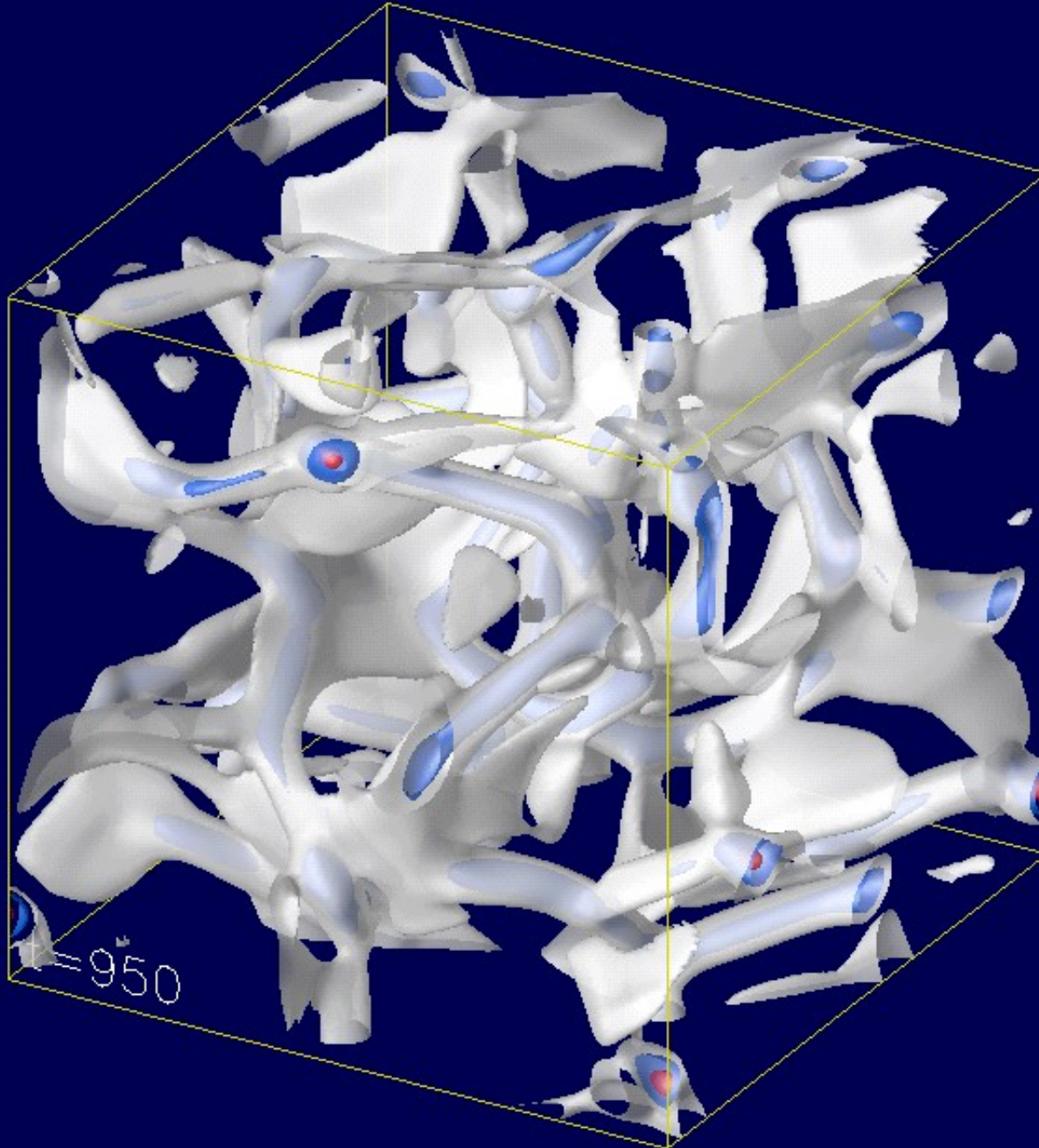
$$N = 128^3$$



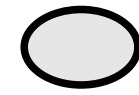
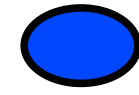
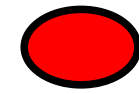
charge density

Filaments

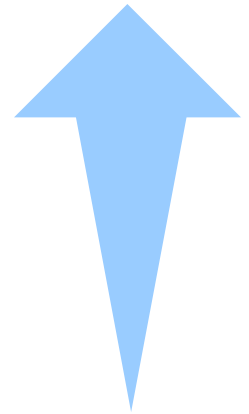
Enqvist, et al. , PRD(2001)
Multamaki, Vilja, PLB(2002)



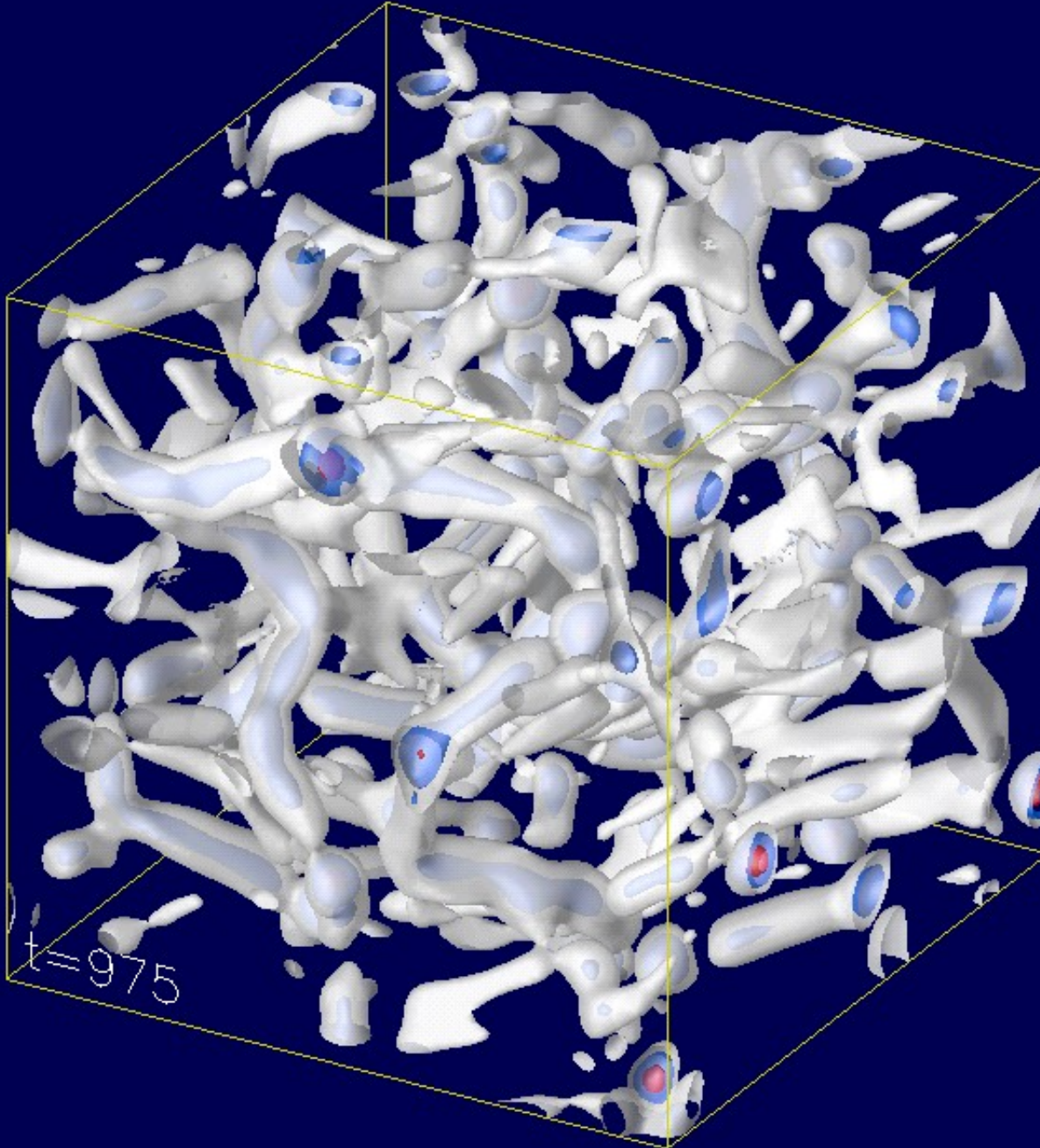
$$N = 128^3$$



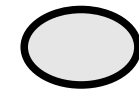
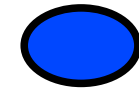
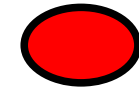
charge density



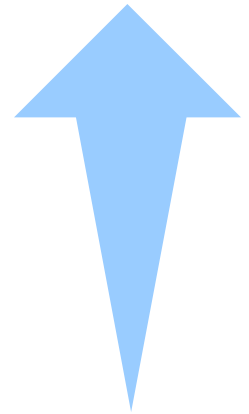
Large Q-balls
at
intersections



$$N = 128^3$$

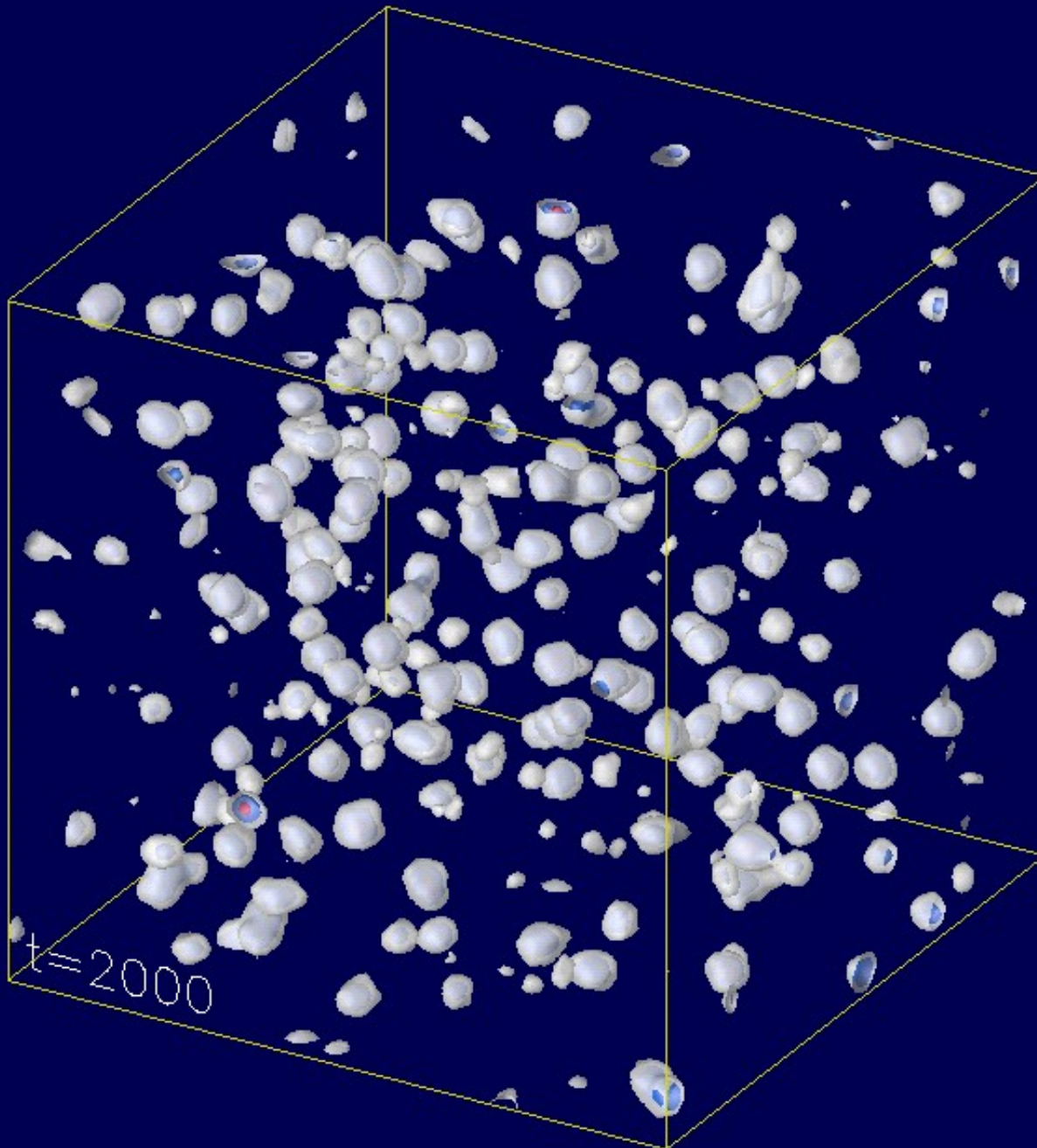


charge density

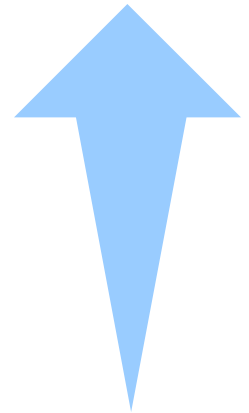
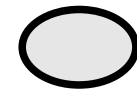
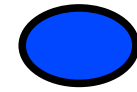
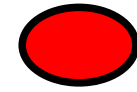


Torn to
small pieces

Result : 3D

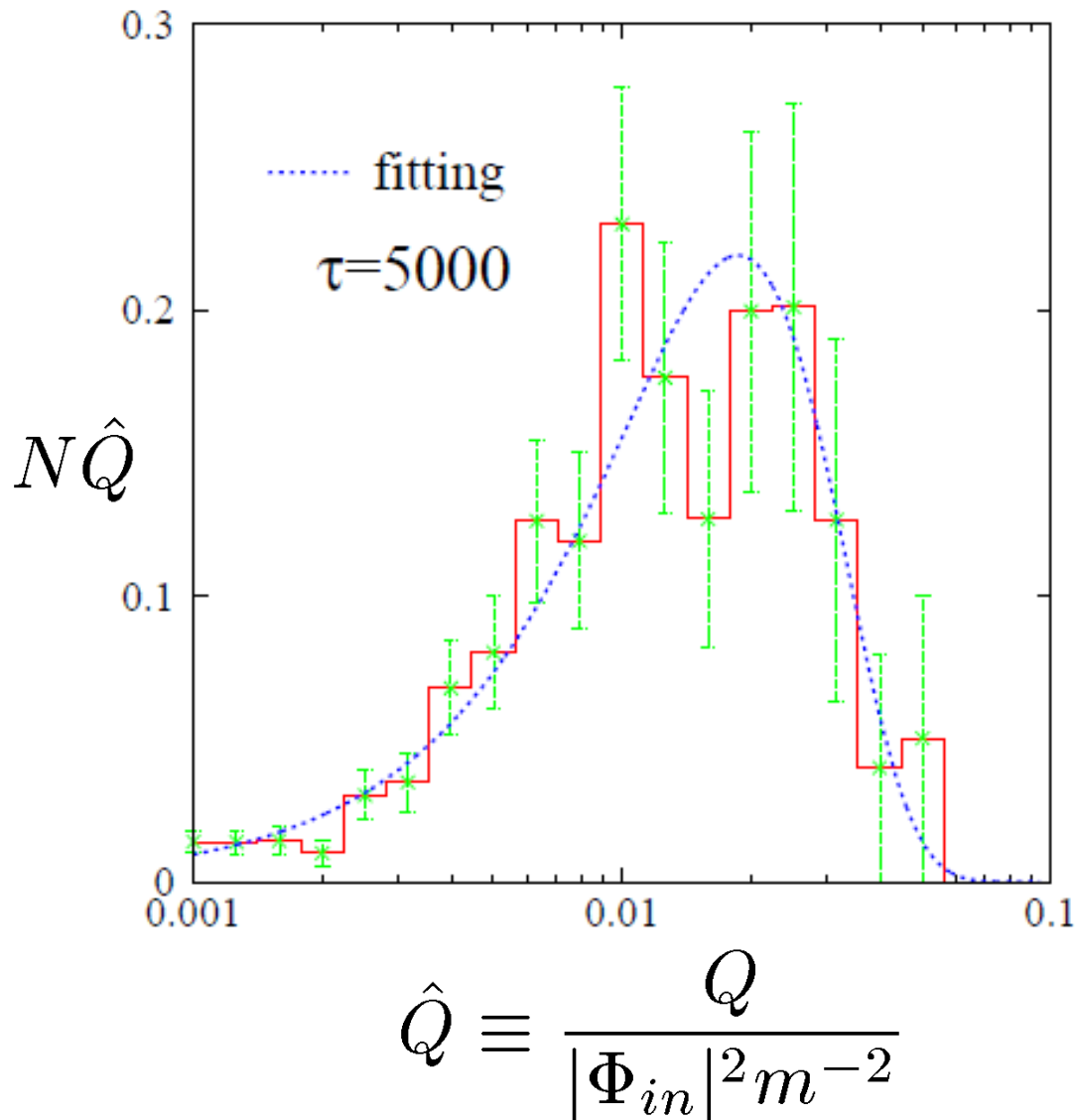


$$N = 128^3$$



charge density

Relaxation



$$f_{NQ} = aQ^b e^{-cQ^2}$$

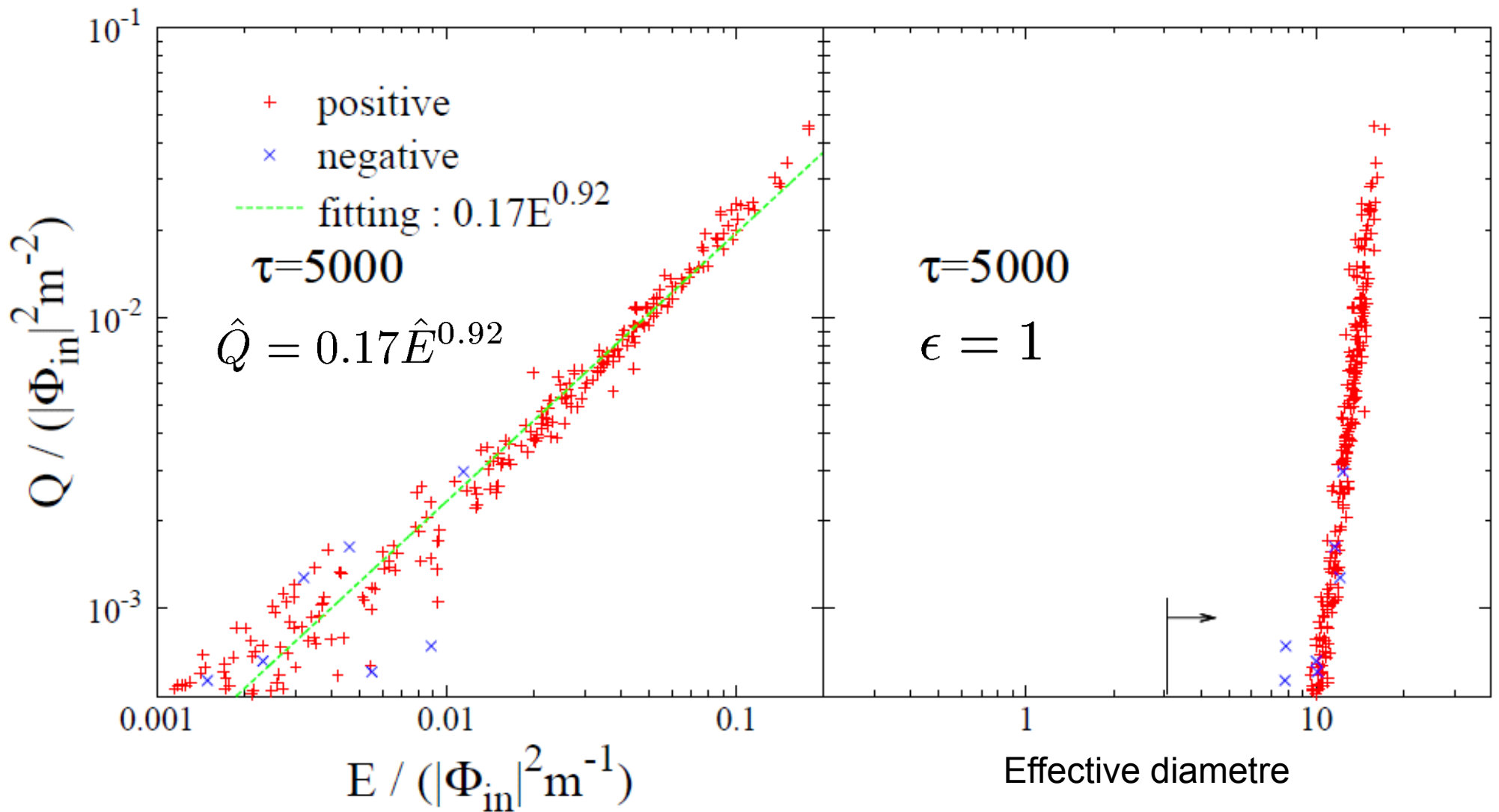
$$a = 71.2, b = 1.29, c = 1.86 \times 10^3$$

$$Q_{\text{peak}} = 1.9 \times 10^{-2} |\Phi_{in}|^2 m^{-2}$$

~60% larger than existing result :

$$Q_{\text{max}}^{\text{KK}} = 1.2 \times 10^{-2} |\Phi_{in}|^2 m^{-2}$$

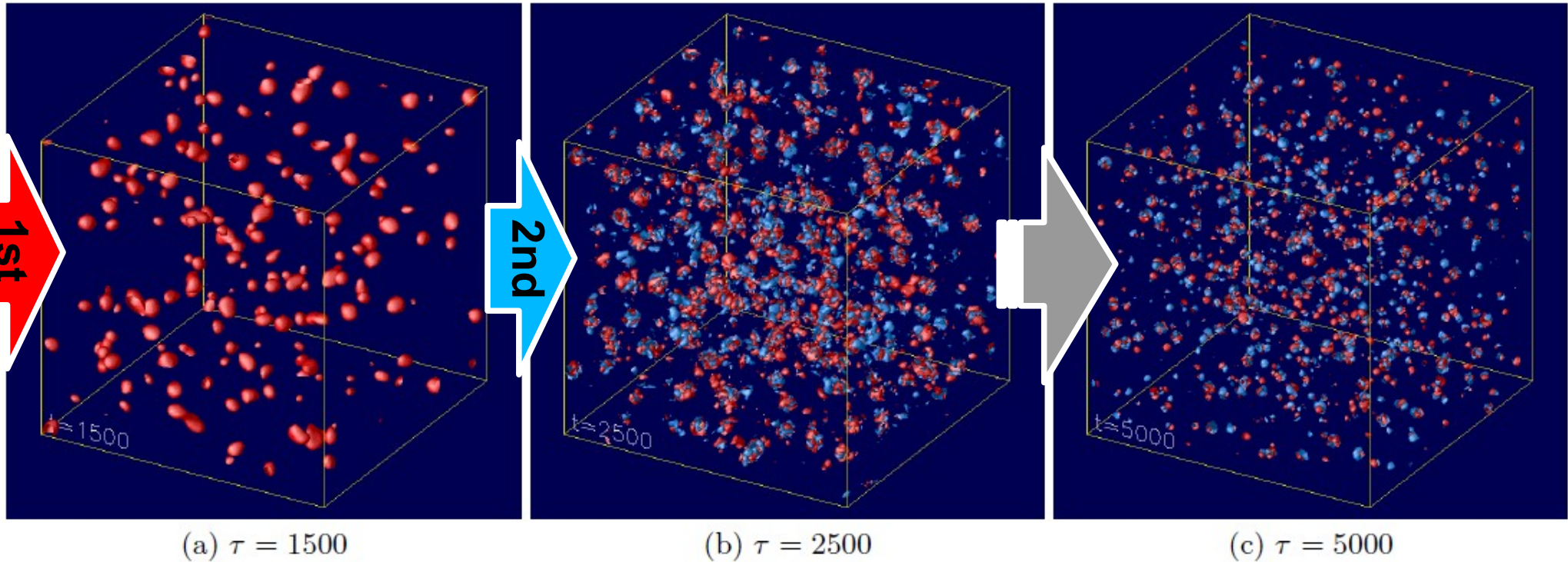
Kasuya, Kawasaki, PRD (2000)



c.f gravity mediation : $Q \propto E, \quad R \sim |K|^{-1/2} m^{-1}$

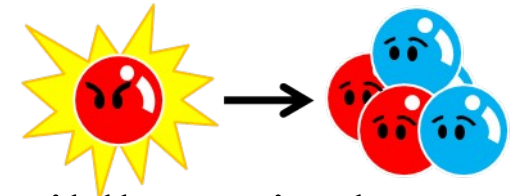
Result : 2nd stage formation in $\epsilon = 0.01$ case

Recall : $\Phi_{in} = M_*$ $\dot{\Phi}_{in} = imM_*\epsilon$



▶ 1st generation Q-ball : **POSITIVE**, EXCITED

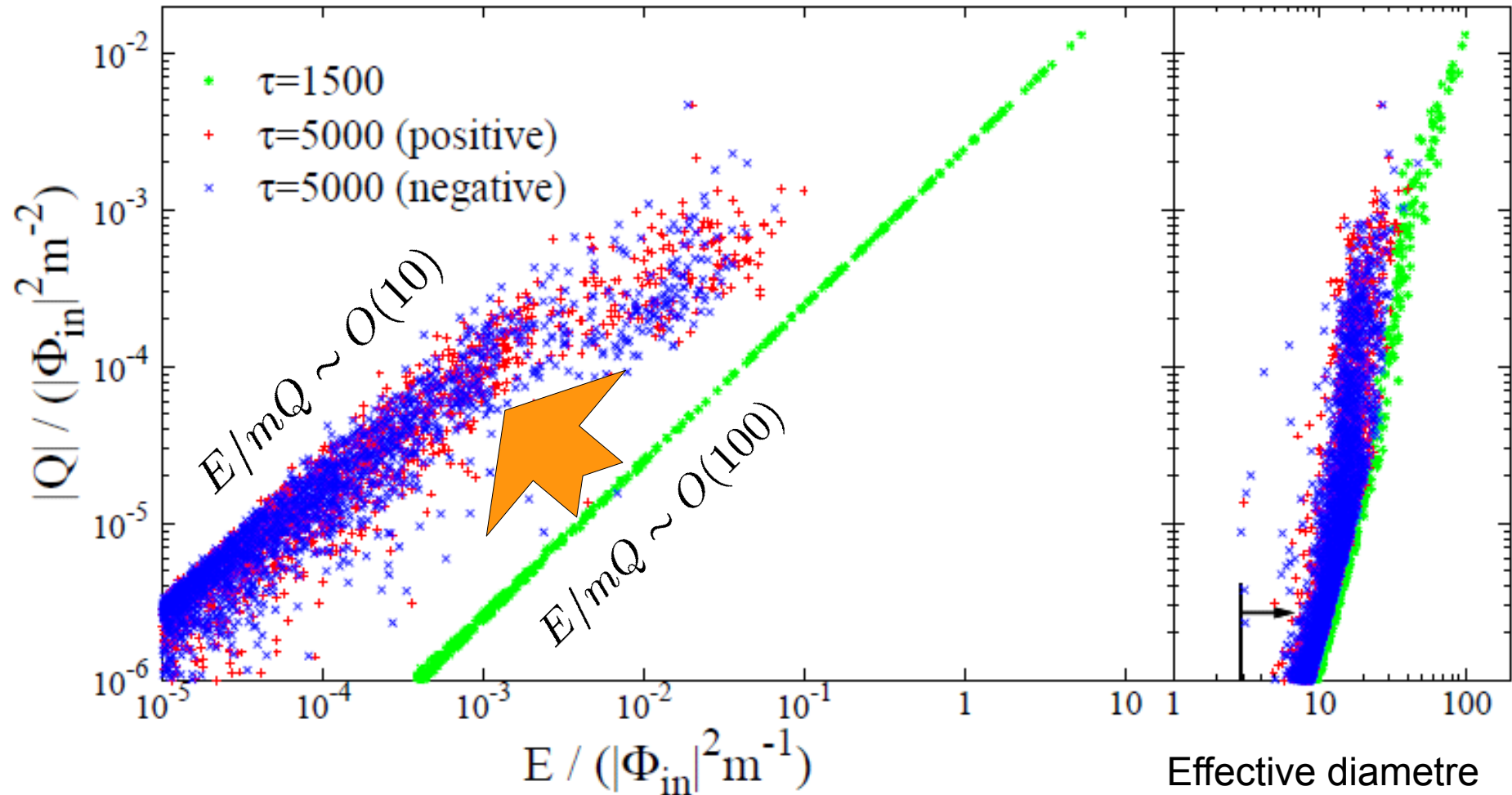
▶ 2nd generation Q-ball : **POSITIVE=NEGATIVE**, mildly excited



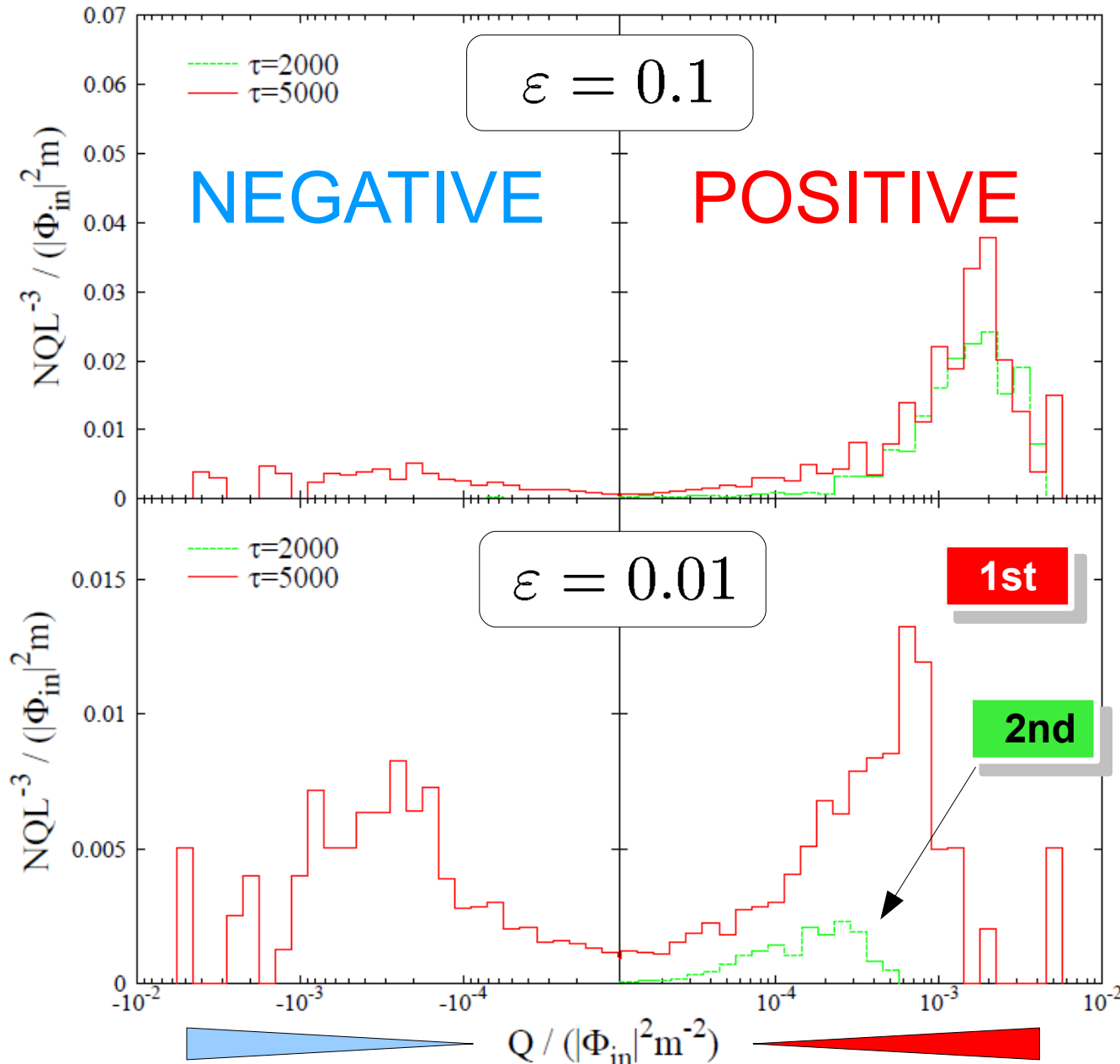
Excited Q-balls release their **excessive energy** , producing **negative** Q-balls

Result : relations

$$\left. \begin{aligned} \Phi_{in} &= M_* \\ \dot{\Phi}_{in} &= imM_*\epsilon \end{aligned} \right\} \frac{E}{mQ} \sim \frac{|\Phi|^2}{|\Phi\dot{\Phi}|} \sim \frac{1}{\epsilon} \ll 1 \longrightarrow \frac{E}{mQ} \Big|_{fin} \sim O(10)$$



Result : charge distributions



► For small ϵ , # of +/- Q-balls eventually become the same.

► Peak charge of 1st-gen Q-balls scales as

$$Q \sim |\Phi \dot{\Phi}| \propto \epsilon$$

► The scaling becomes no longer valid for 2nd-gen Q-balls

cf. this scaling is broken down also in gauge mediation.

Kasuya, Kawasaki, PRD (2001)

▶ Q-balls could be a promising source of GWs

- First numerical simulation : Kusenko et al.

Kusenko, Mazumdar, PRL (2008)

Kusenko, Mazumdar, Multamaki, PRD (2009)

- Analytical estimation with 'thermal-log term' contributions

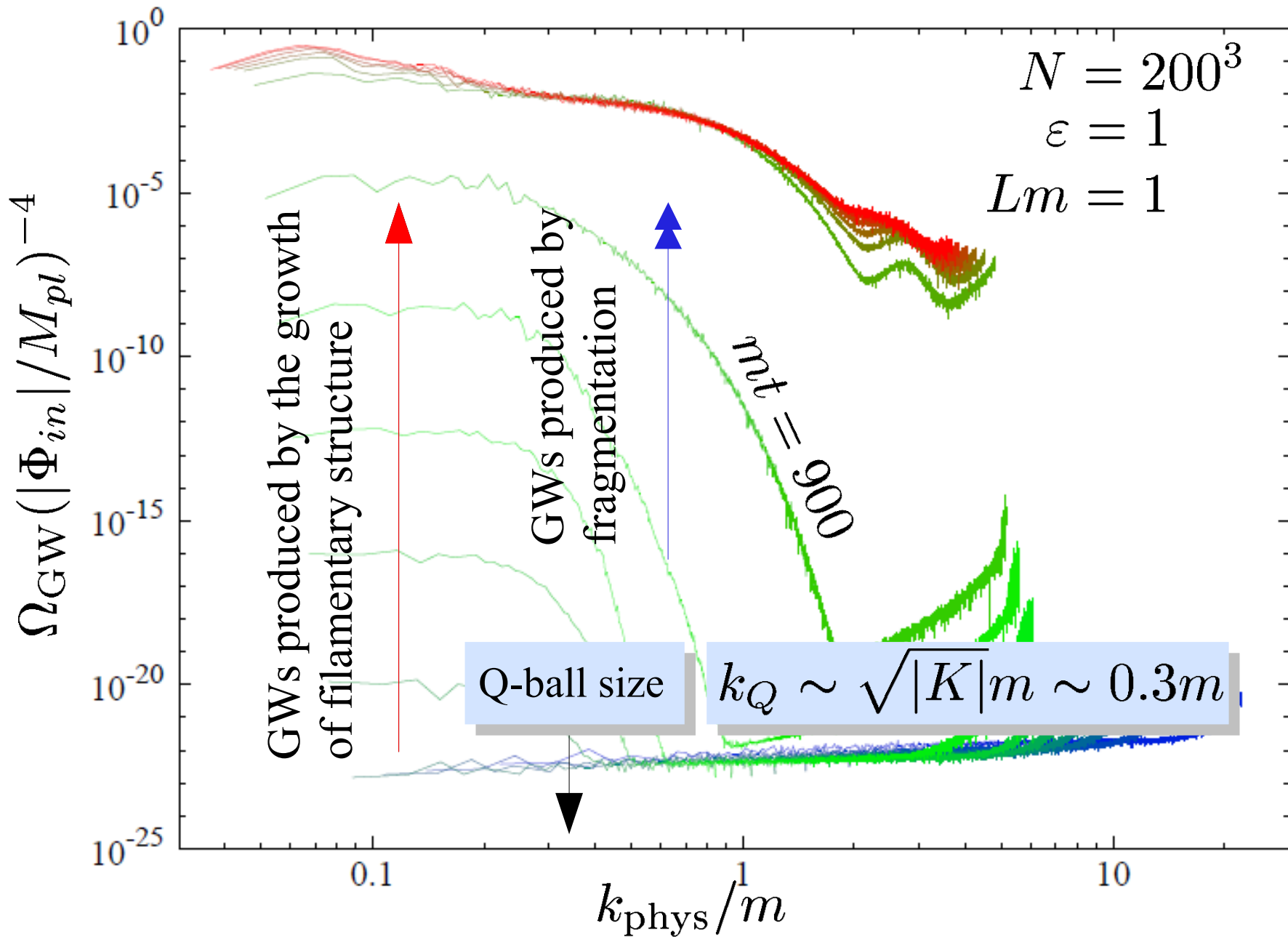
Chiba, Kamada, Yamaguchi, PRD (2010)

▶ Gravitational wave energy and spectrum

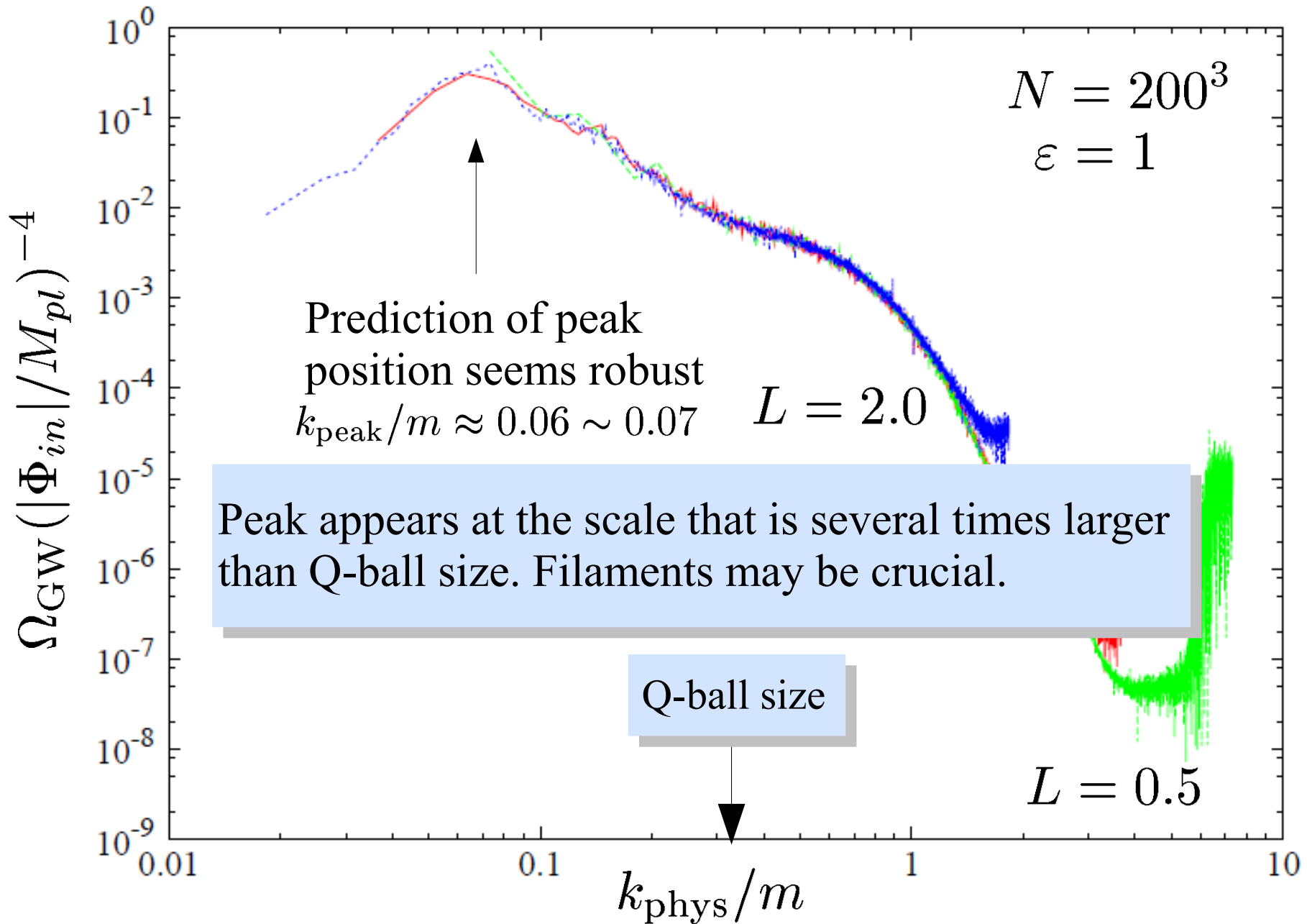
$$\rho_{\text{GW}} = \frac{1}{32\pi G} \langle \dot{h}_{ij} \dot{h}_{ij} \rangle \quad \Omega_{\text{GW}} = \frac{1}{\rho_c} \frac{d\rho_{\text{GW}}}{d \log k}$$

- ▶ The basic structure of the spectrum is reflected by the existence of the filamentary structure before Q-ball formation epoch.

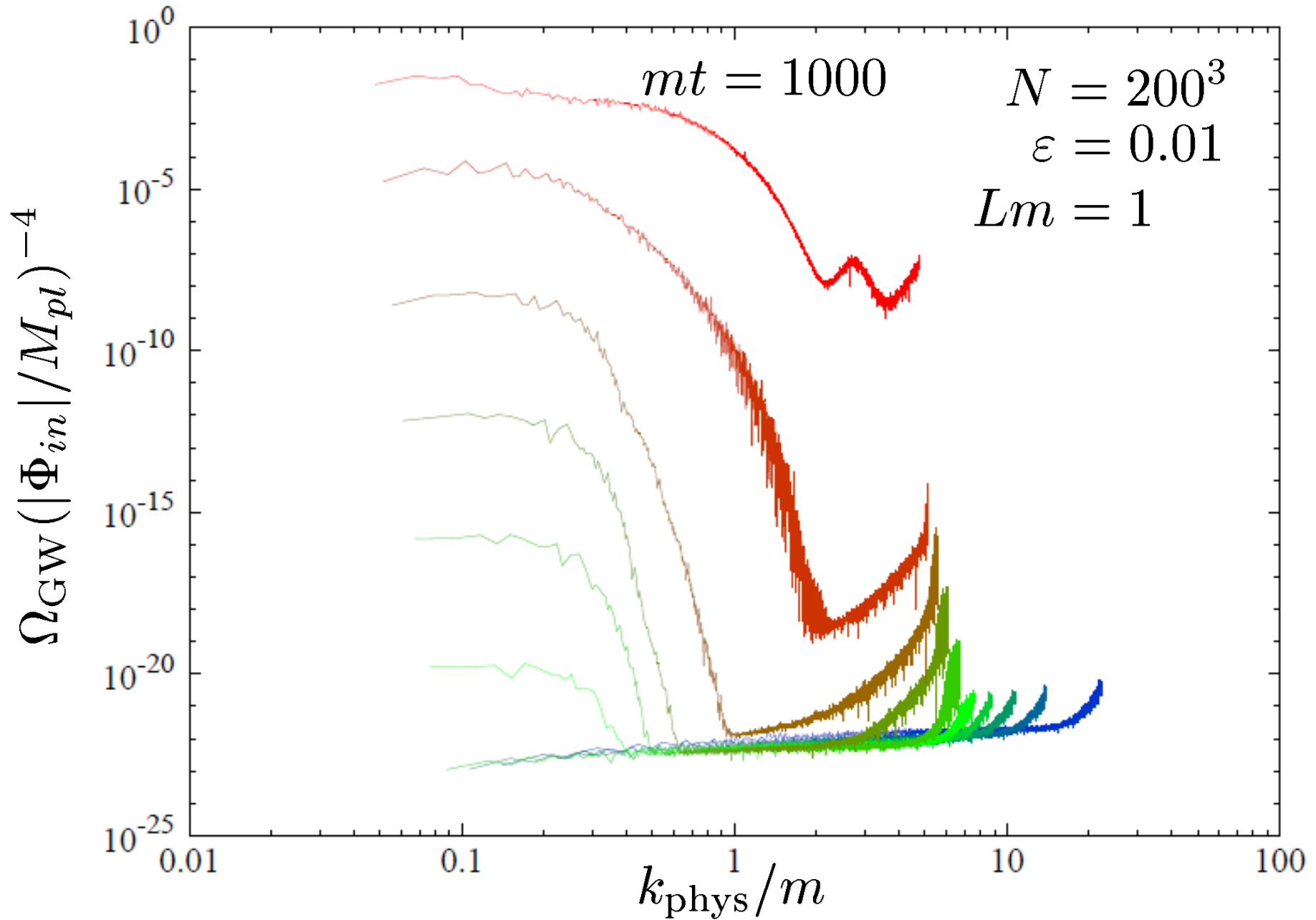
Energy spectrum at Q-ball formation epoch



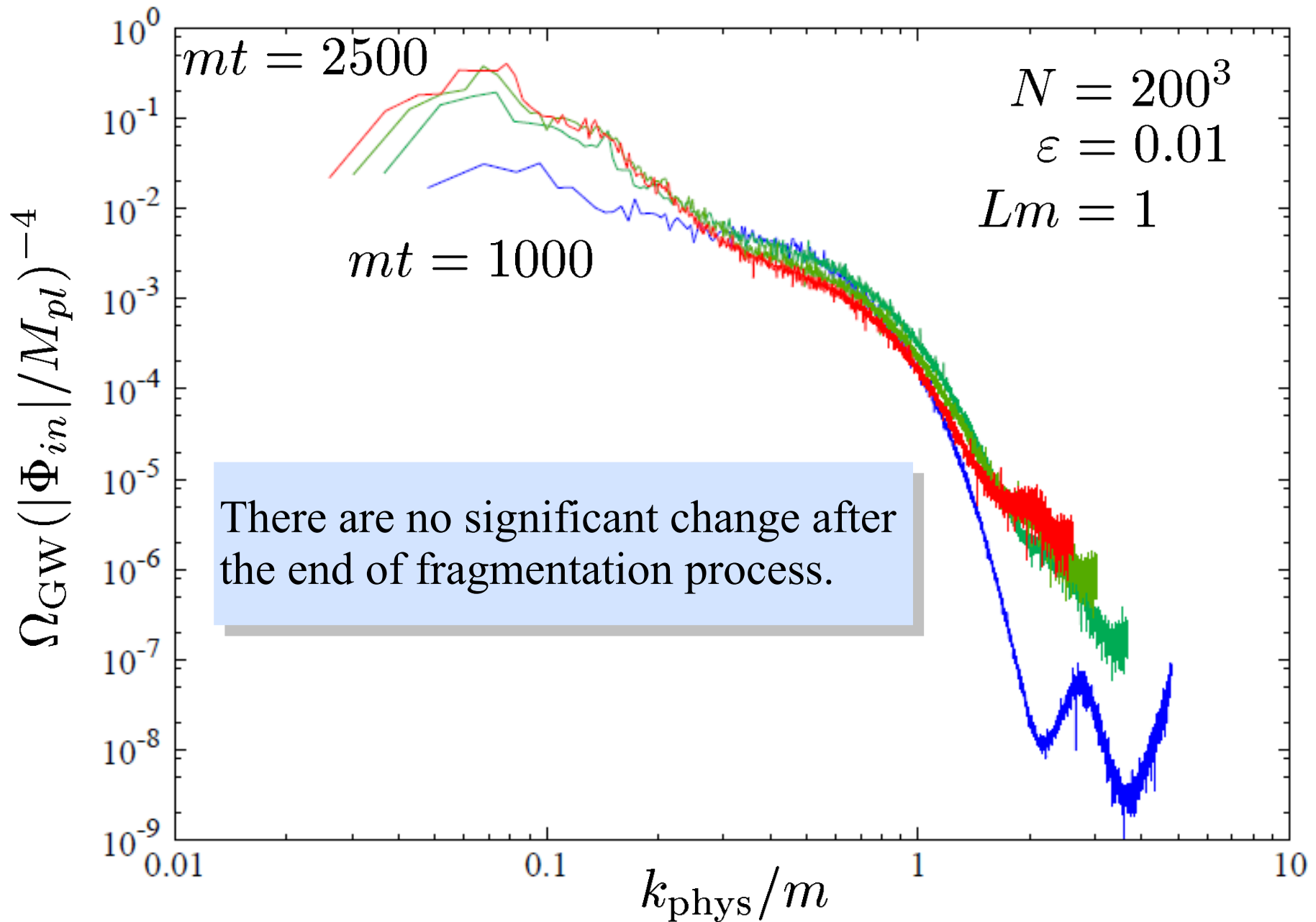
Box size effect



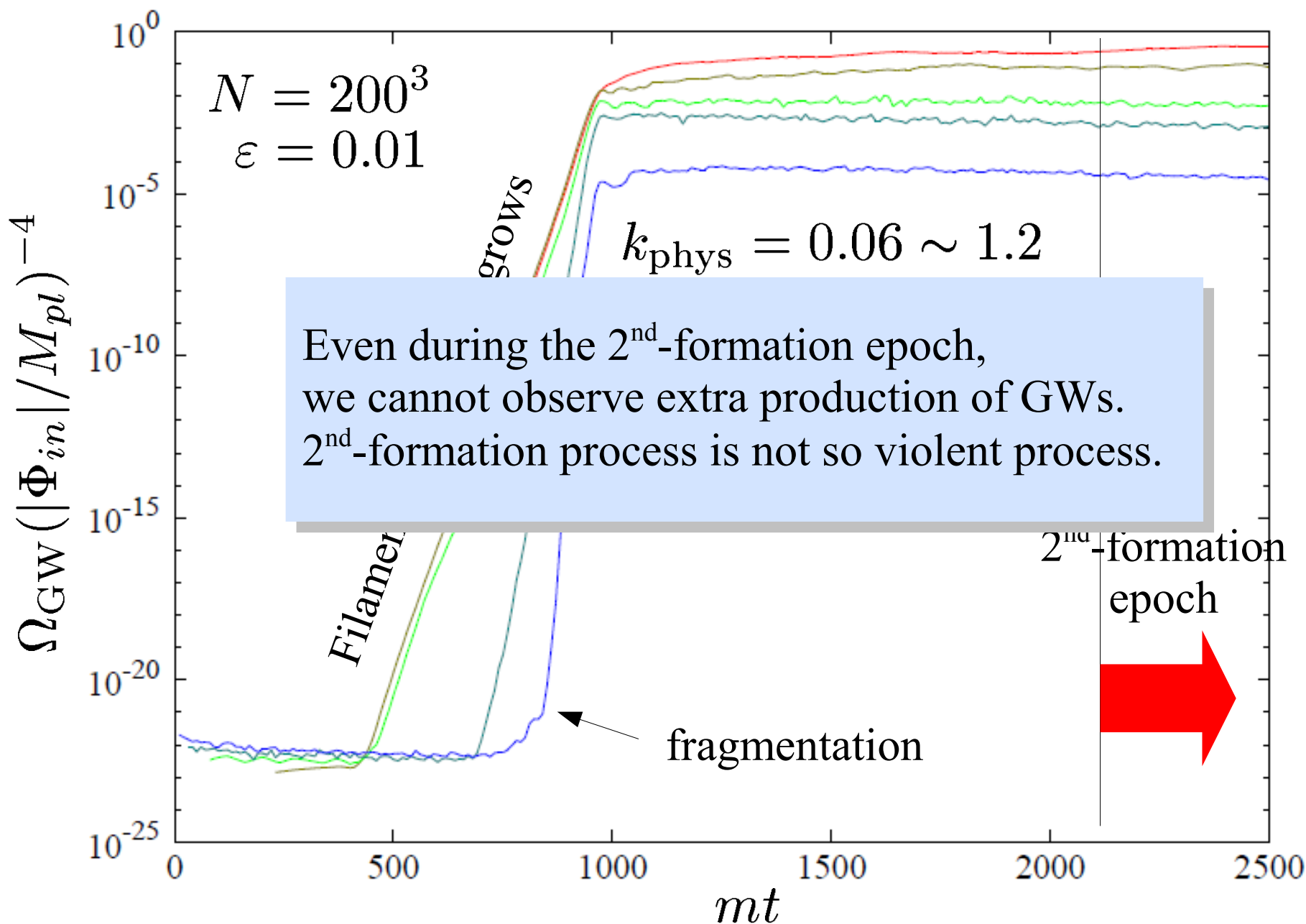
$\varepsilon = 0.01$: early time $mt \leq 1000$



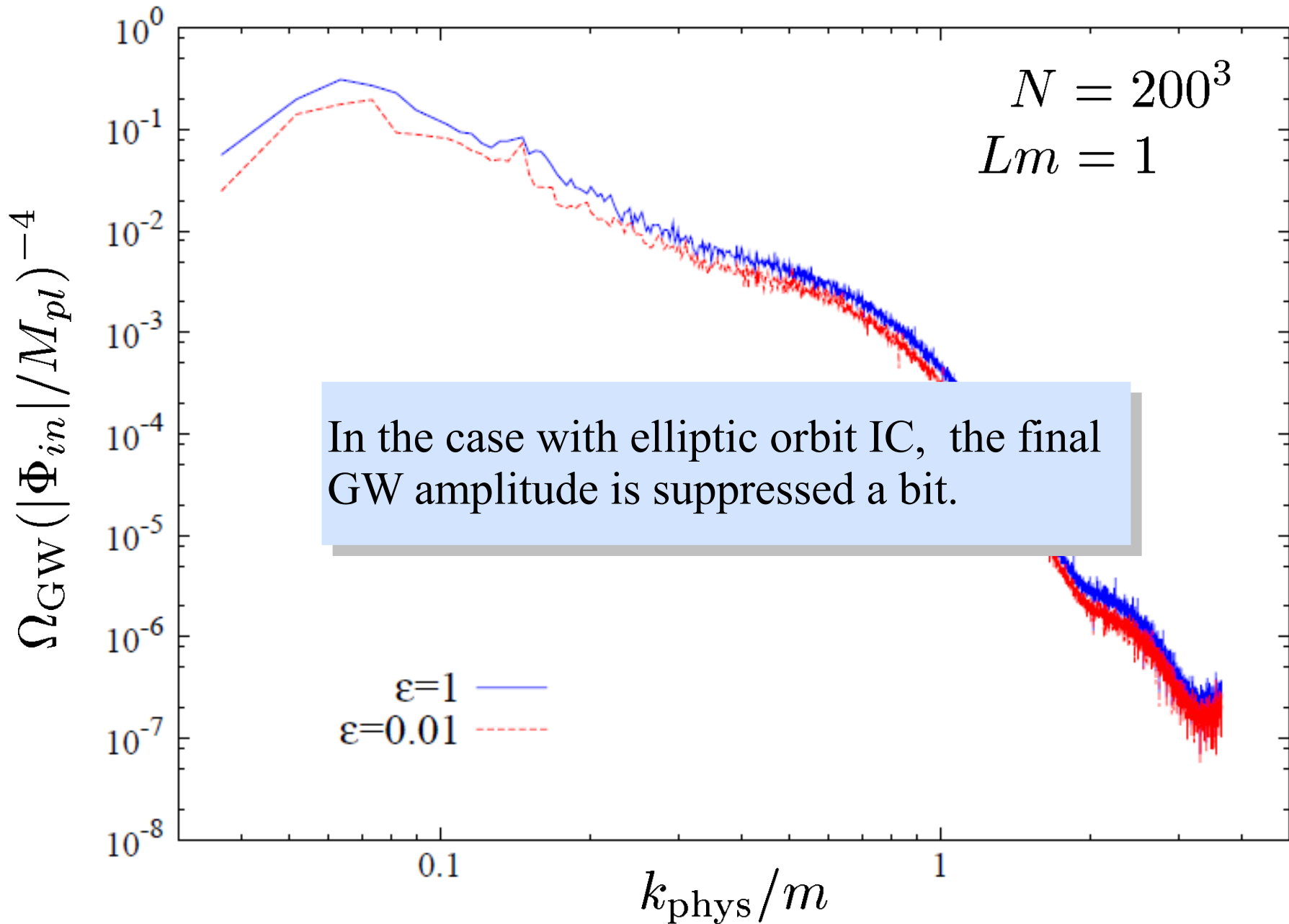
$\varepsilon = 0.01$: late time $1000 < mt$



Time evolution of amplitude for $\varepsilon = 0.01$



Difference between $\varepsilon = 1$ and $\varepsilon = 0.01$



- ▶ 3D simulations for Q-ball formation in gravity mediation.
- ▶ Charge distribution of Q-balls **TH, Kawasaki, Takahashi, JCAP 06(2010)008**
 - 'circular' case : the peak charge is slightly larger than existing results
 - 'elliptic' case : eventually the same numbers of +/- Q-balls appear, and peak charge, scaling , ...
- ▶ Power spectrum of GWs from Q-balls **TH, Takahashi, Yamaguchi, in preparation**
 - early epoch : large scale GWs may be associated with filamentary structure. But it remains unclear now.
 - formation(fragmentation) epoch : small scale significantly grows.
 - relaxation epoch : no more grows even during 2nd-formation process, though it's crucial for the final shape of charge distribution.
- ▶ Filamentary structure plays a crucial role for both charge distributions and the peak amplitude of GWs.