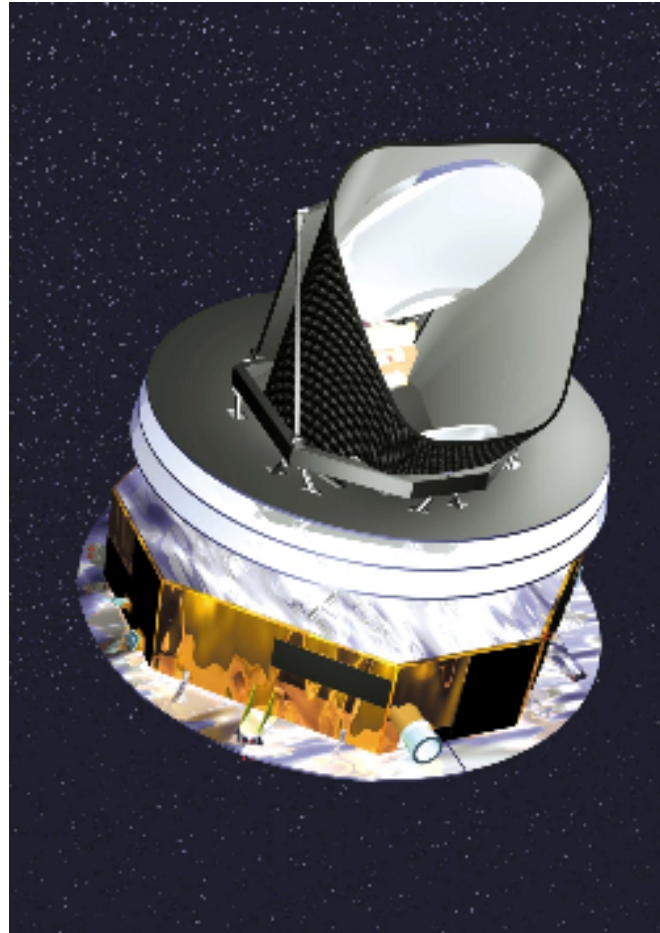


PLANCK



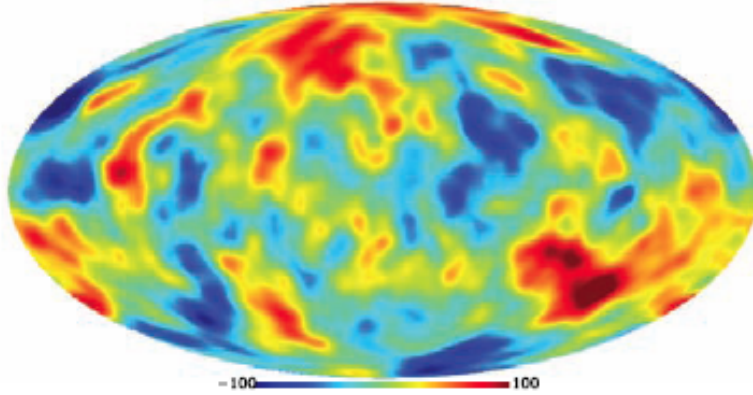
launched on
14 May, 2009

full-sky survey started on 27 Aug, 2009

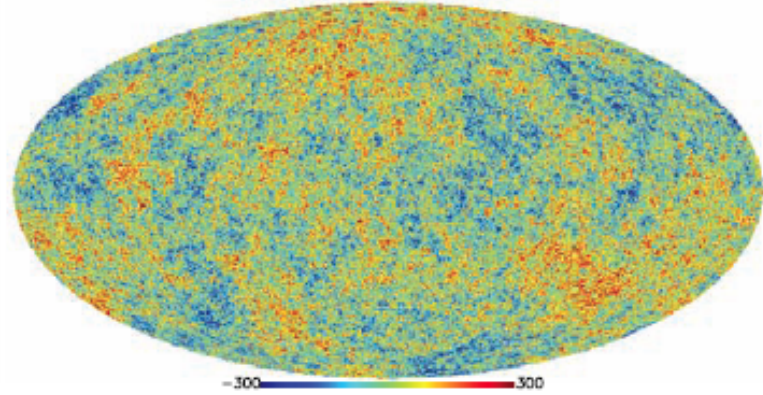
Angular resolution

CMB Fluctuations (μK)

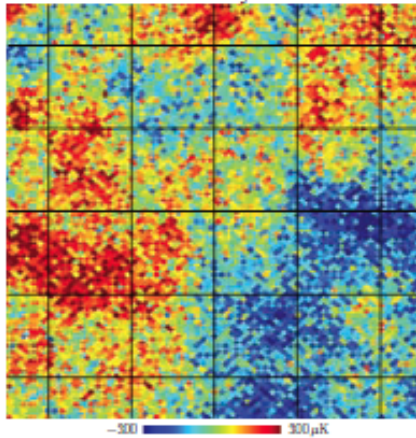
COBE-DMR resolution



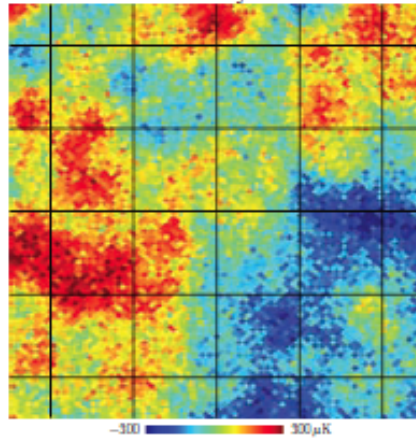
Planck resolution



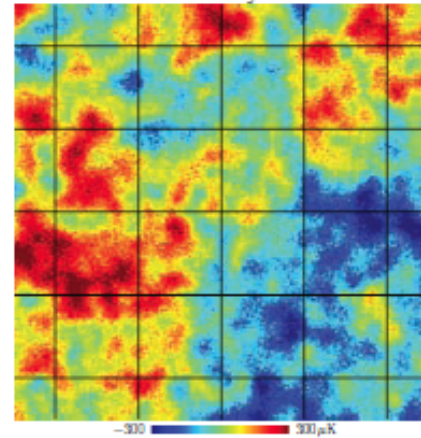
WMAP 2 years



WMAP 8 years



Planck 1 year

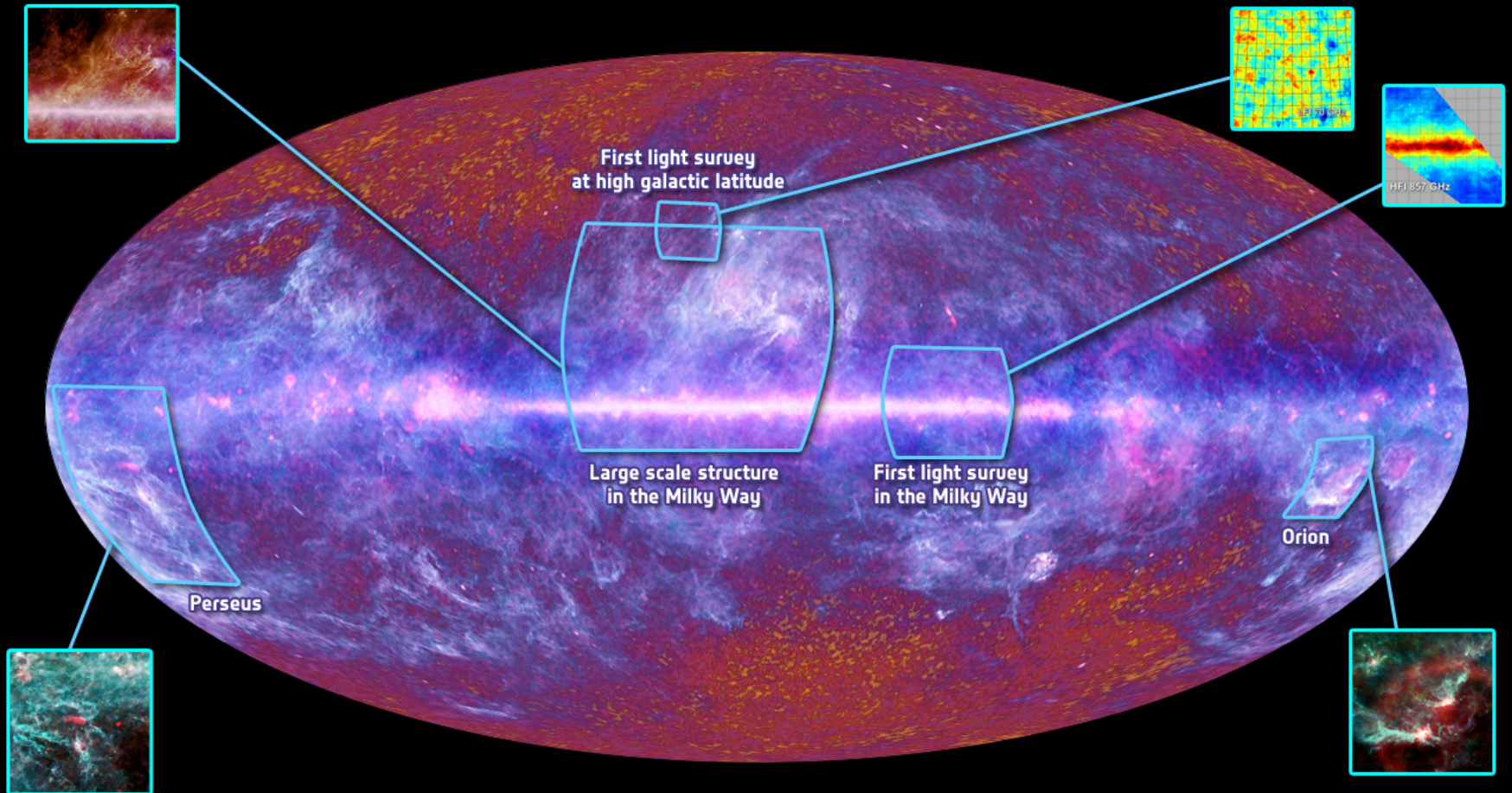


WMAP 2years

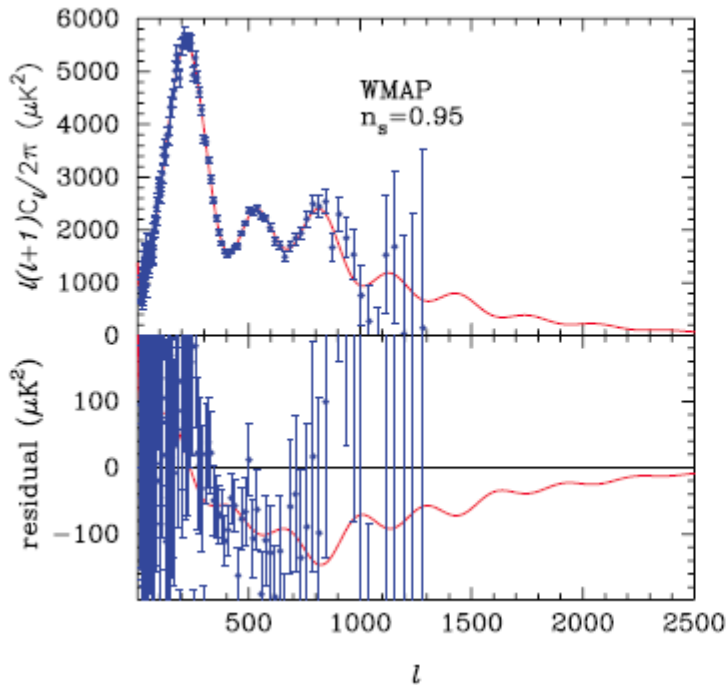
WMAP 8years

Planck 1year

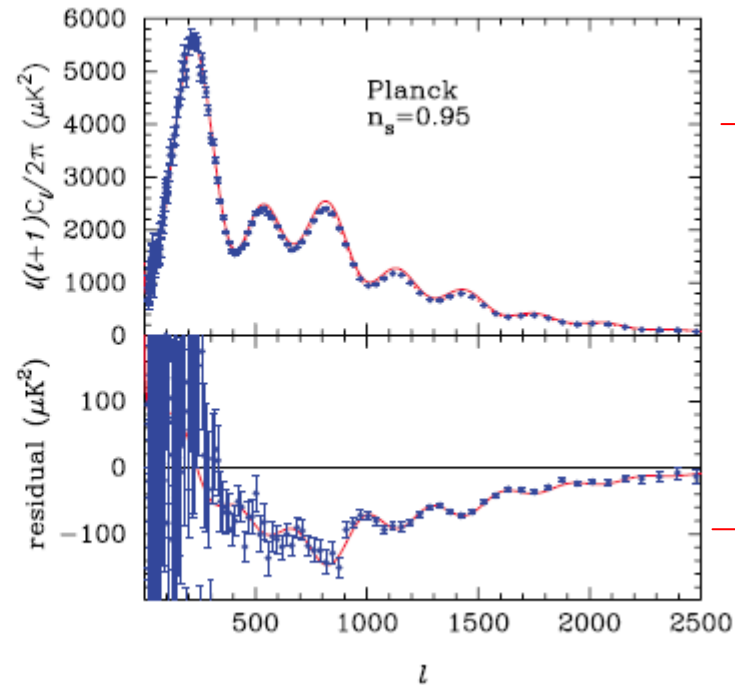
Planck 1yr Map



Scalar spectral index: $P_R(k) \propto k^{n_s}$



WMAP 4 years



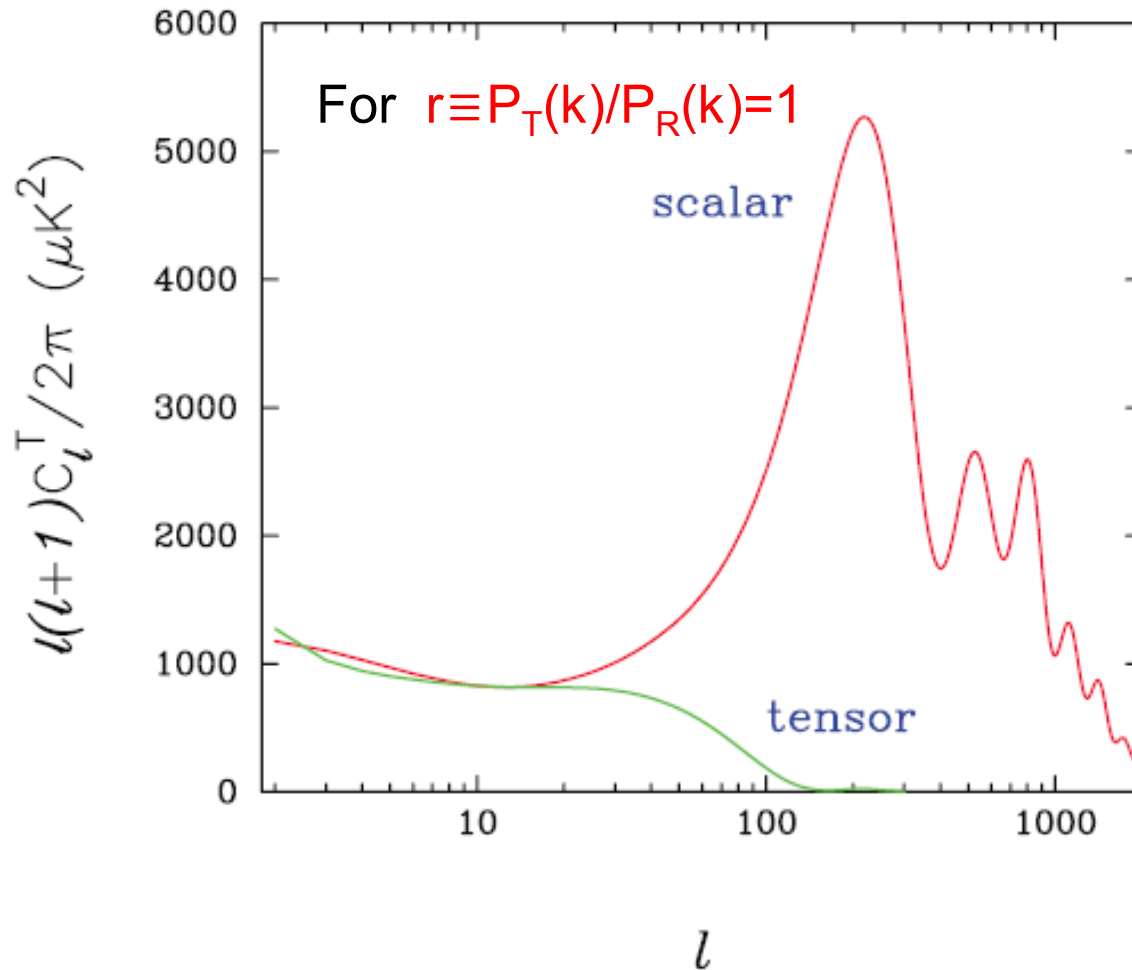
PLANCK 1 year

— $n_s = 1$

— $n_s = 0.95$

determination of spectral index to within 1 % accuracy

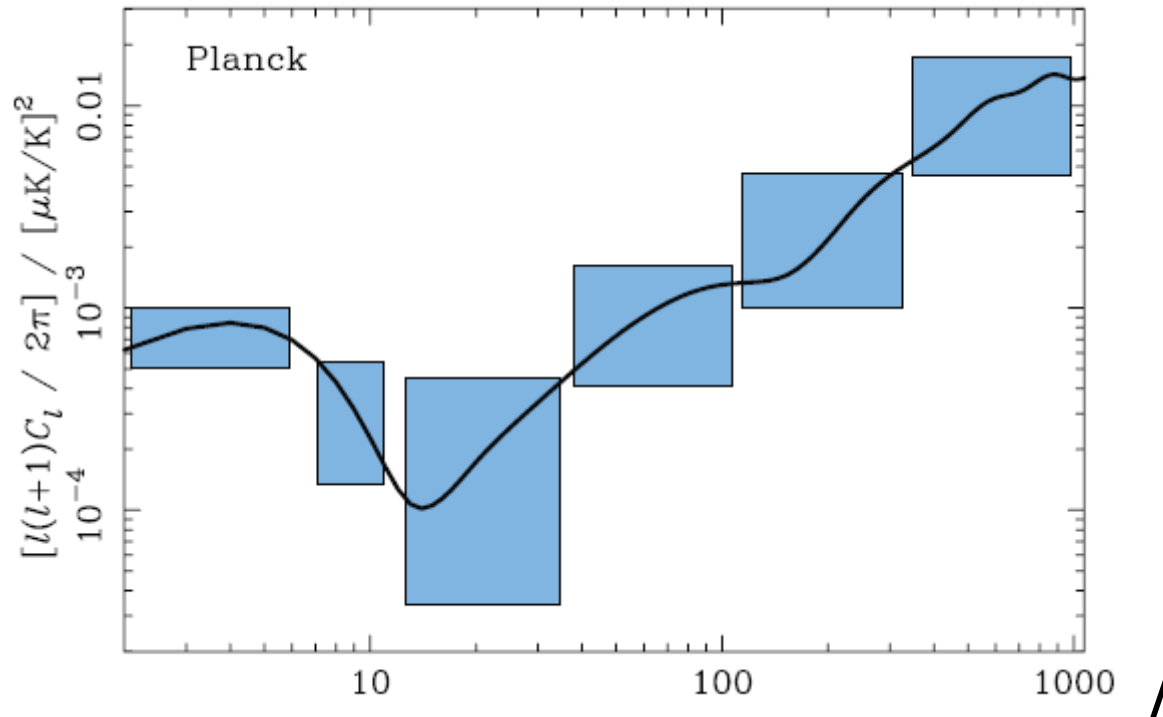
Tensor (gravitational wave) modes



In a “standard” (chaotic inflation) model, $r \sim 0.1$

B-mode (odd parity) polarization

B-mode is a unique signature of tensor modes



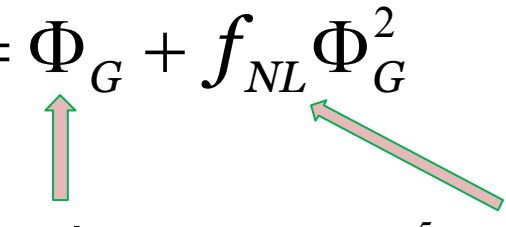
— $r = P_T(k) / P_S(k) = 0.1$

Planck sensitivity: $r > 0.05$ may be rejected at $>80\%$ CL

Detection of $r > 0.05$ **proves** inflation at $V^{1/4} \approx 10^{16} \text{GeV}$

Non-Gaussianity

Curvature perturbation (gravitational potential):

$$\Phi = \Phi_G + f_{NL} \Phi_G^2$$


Gaussian perturbation: $\Phi_G \sim 10^{-5}$

current WMAP bound:

$$-10 < f_{NL} < 74 \quad (95\% \text{ CL})$$

Planck's sensitivity: $|f_{NL}| \gtrsim 5$

Standard single-field slow-roll inflation gives $|f_{NL}| \ll 1$

Detection of non-zero f_{NL} implies non-conventional inflation

Tests of Fundamental Theory

extra-dims, string landscape, ...

Cosmology in 21st Century

High Precision Cosmology

- gravitational waves from Inflation
- non-Gaussian perturbations
- extra dimensions / string cosmology
- origin of dark energy
- ...

fundamental laws of nature may be revealed.
(final theory?)

Cosmological perturbation theory
will play a major role