Stability of spherical symmetric spacetimes in Modified Gravity

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with Teruaki Suyama, Takahiro Tanaka [PRD (2011)]



Introduction

Mistery of Dark Energy

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- Complex problem

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- Motivated different approaches

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- New gravity and new matter: Extended Brans-Dicke,
 Galileon

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- Interesting phenomenology [Review by ADF, Tsujikawa '10]

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- Two extra new scalar fields: F, ξ

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Gauss-Bonnet gravity

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- Instability $c_s^2 < 0$ on FLRW before DE [ADF, Tsujikawa '10]

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- In other backgrounds, both propagate

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- Write down the action expanded at 2nd order
- Or linearize the equations of motion
- Look for behavior of perturbations

• Study vacuum spher. symm. $(z = \cos \theta)$

$$ds^{2} = -Adt^{2} + dr^{2}/B + \frac{r^{2}dz^{2}}{1 - z^{2}} + r^{2}(1 - z^{2})d\phi^{2}$$

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- Expand action at second order

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$$\delta g_{03}=(1-z^2)h_0\partial_z Y_{l0}$$
, $\delta g_{13}=(1-z^2)h_0\partial_z Y_{l0}$

• Introduce Lagr. multipl. $Q = \dot{h}_1 - h_0' + 2h_0/r$

Result

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Result

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- lacksquare No ghost for $AF-2BA'\xi'>0$
- Radial speed

$$c_{\text{odd}}^2 = \frac{1}{A} \frac{AF - 2BA'\xi'}{F - 2B'\xi' - 4B\xi''}$$

 $\delta g_{00}=-AH_0Y_{l0}$, $\delta g_{01}=H_1Y_{l0}$, $\delta g_{rr}=H_2Y_{l0}/B$, $\delta g_{rz}=lpha\partial_z Y_{l0}$, $F=F(r)+\delta F\,Y_{l0}$, $\xi=\xi(r)+\delta \xi\,Y_{l0}$

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- Define $\delta F=v_1-4(1-B)\delta\xi/r^2$, $\alpha=\alpha(v_0,H_2,v_1')$
- ullet Integrate out H_2 and H_0

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• $\det K_{ij} \to 0$, as $B \to 1 + Cr^2$: (a)dS.

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- ullet Require that, with $ar{G}\equiv 16r_s^2/r^6$

$$-1 < \tanh 2b = -\frac{2\sqrt{\bar{G}}U_{,F\xi}}{U_{,\xi\xi} + \bar{G}U_{,FF}} < 1$$

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- If $M_{
 m cutoff} \sim L_{
 m exp}^{-1}$ viable if

$$|9\bar{G}\det(f_{,ij})| = |m_{+}^{-2}m_{-}^{-2}| \ll L_{\exp}^{4}$$
$$|3(f_{,RR} - \bar{G}f_{,GG})| = |m_{+}^{-2} + m_{-}^{-2}| \ll L_{\exp}^{2}$$

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- $m_+^2 = \mathcal{O}(10^{22.7-2.7p} \mathrm{cm})^{-2} \gg (1 \mathrm{AU})^{-2}$, if $p \ge 4$

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- If $|f_{,RR}|<|\bar{G}f_{,GG}|$, problem: $m_+\leftrightarrow m_-$, $m_-\ll m_+$

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