

磁化された電子陽電子プラズマ中 における誘導散乱のプラズマ粒子 シミュレーション

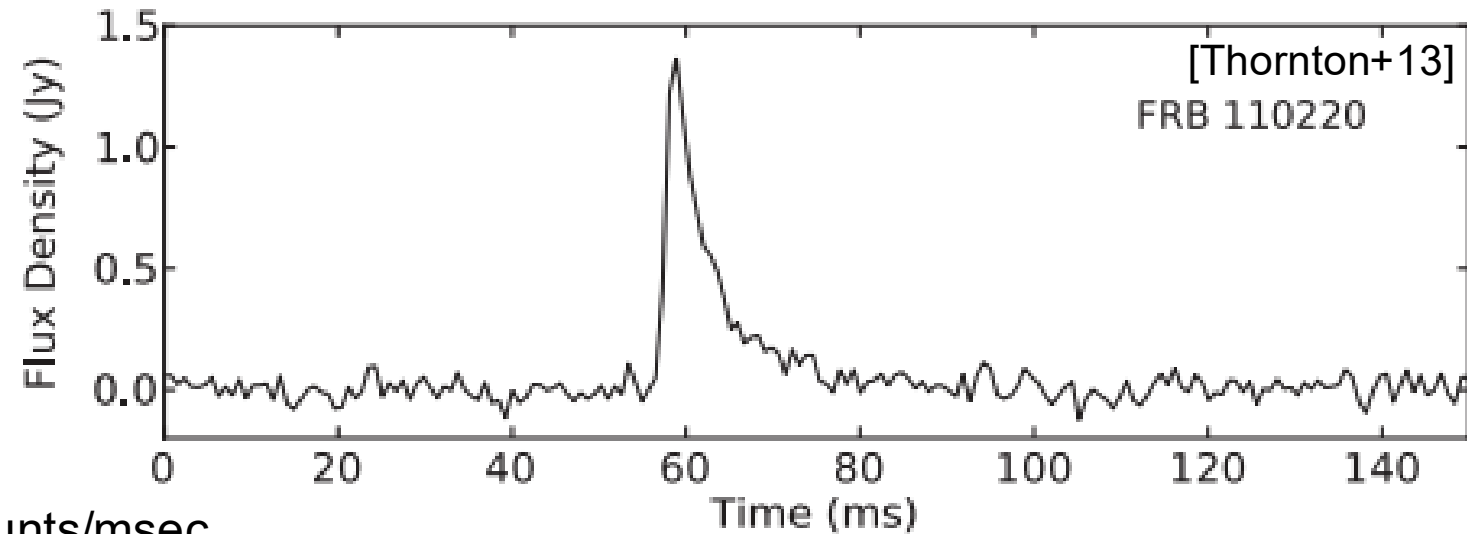
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Fast Radio Burst (FRB)

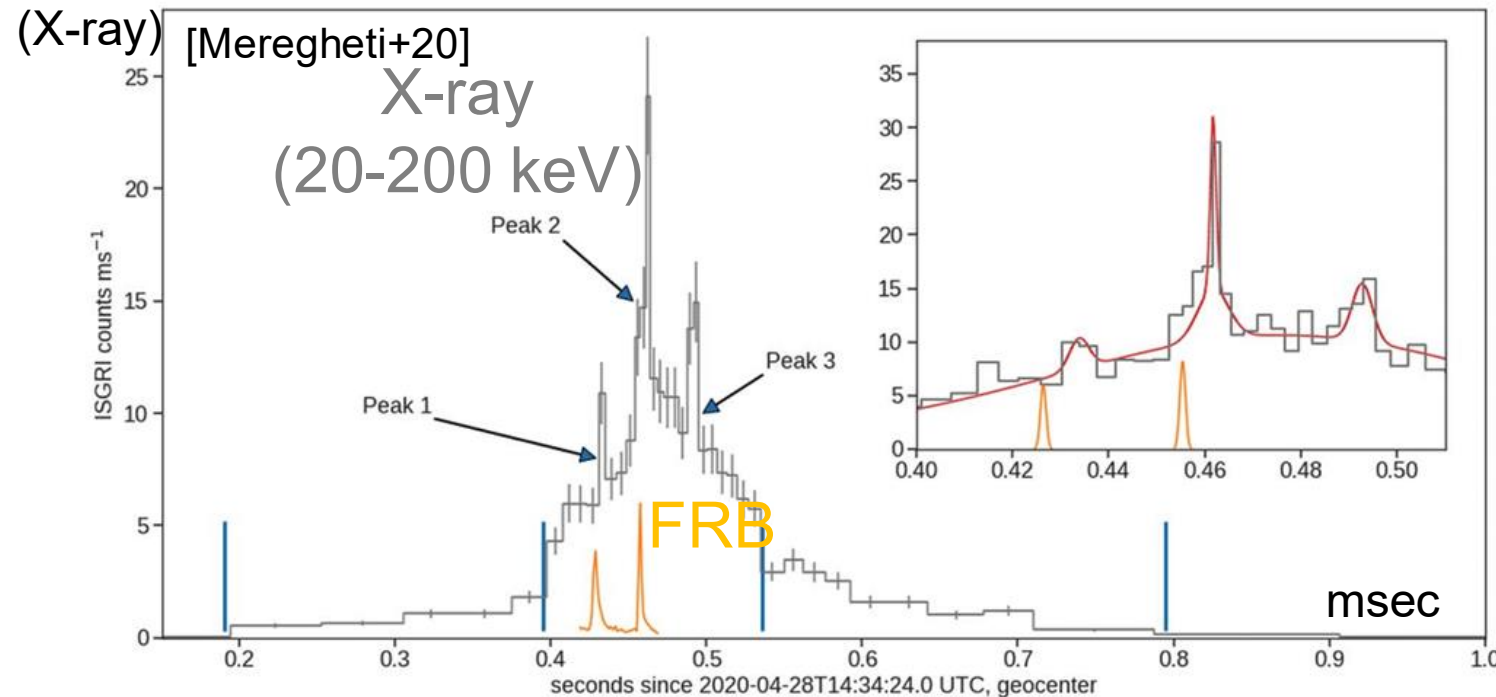
- ❑ Most luminous radio transient
- ❑ Large dispersion measure
→ Cosmological
- ❑ Frequency: $\nu \sim \mathcal{O}(\text{GHz})$
- ❑ Duration: $\Delta t \sim \mathcal{O}(\text{msec})$
- ❑ Flux density: $S_\nu \sim \mathcal{O}(\text{Jy})$ @GHz
- ❑ High Brightness temperature:
→ Coherent emission
- ❑ FRB from Galactic magnetar is observed in 2020.

One of the origins of FRBs is a magnetar.



counts/msec

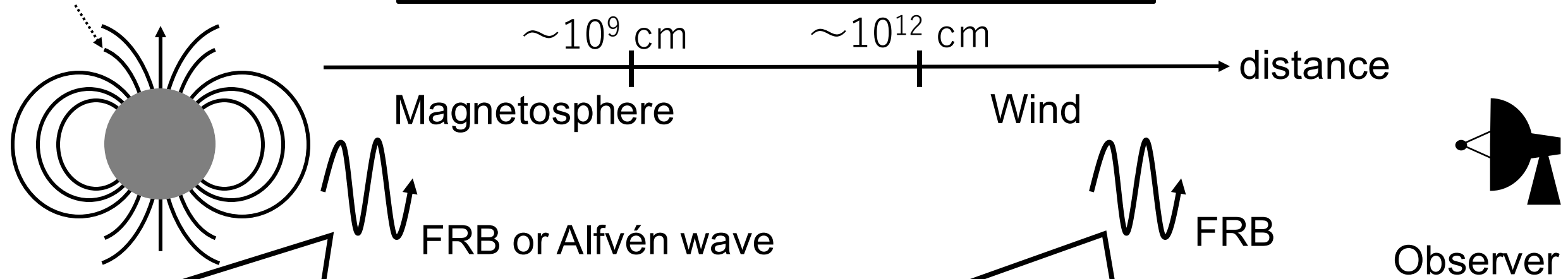
(X-ray)



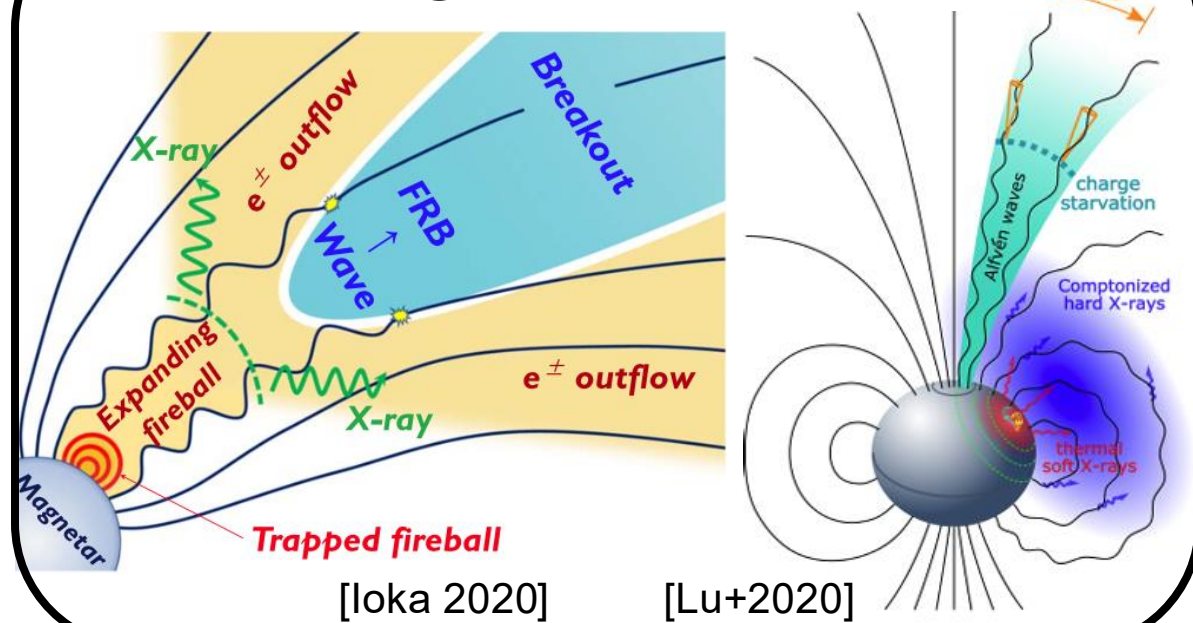
Magnetar Model

Magnetized electron-positron plasma

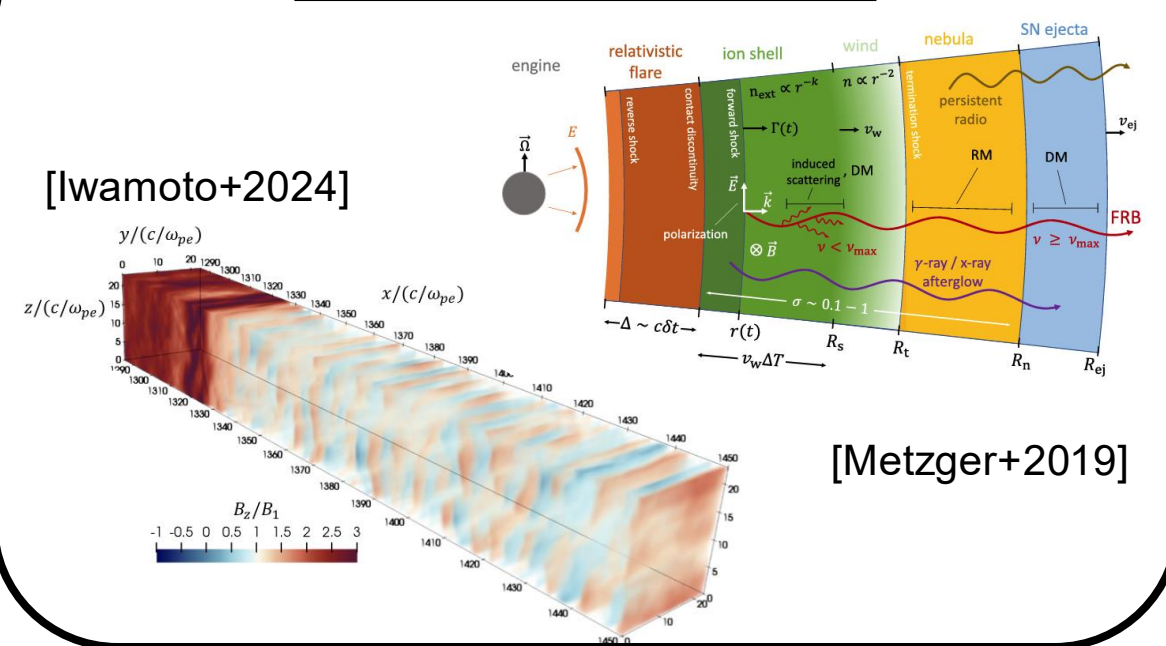
Magnetar



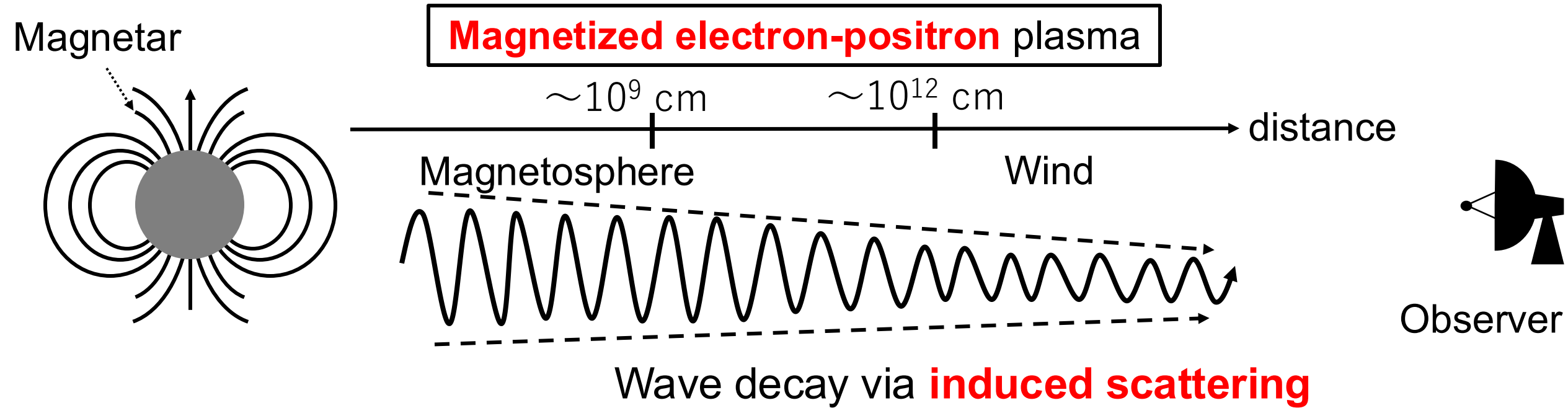
Inner Magnetosphere Model



Outer Shock Model



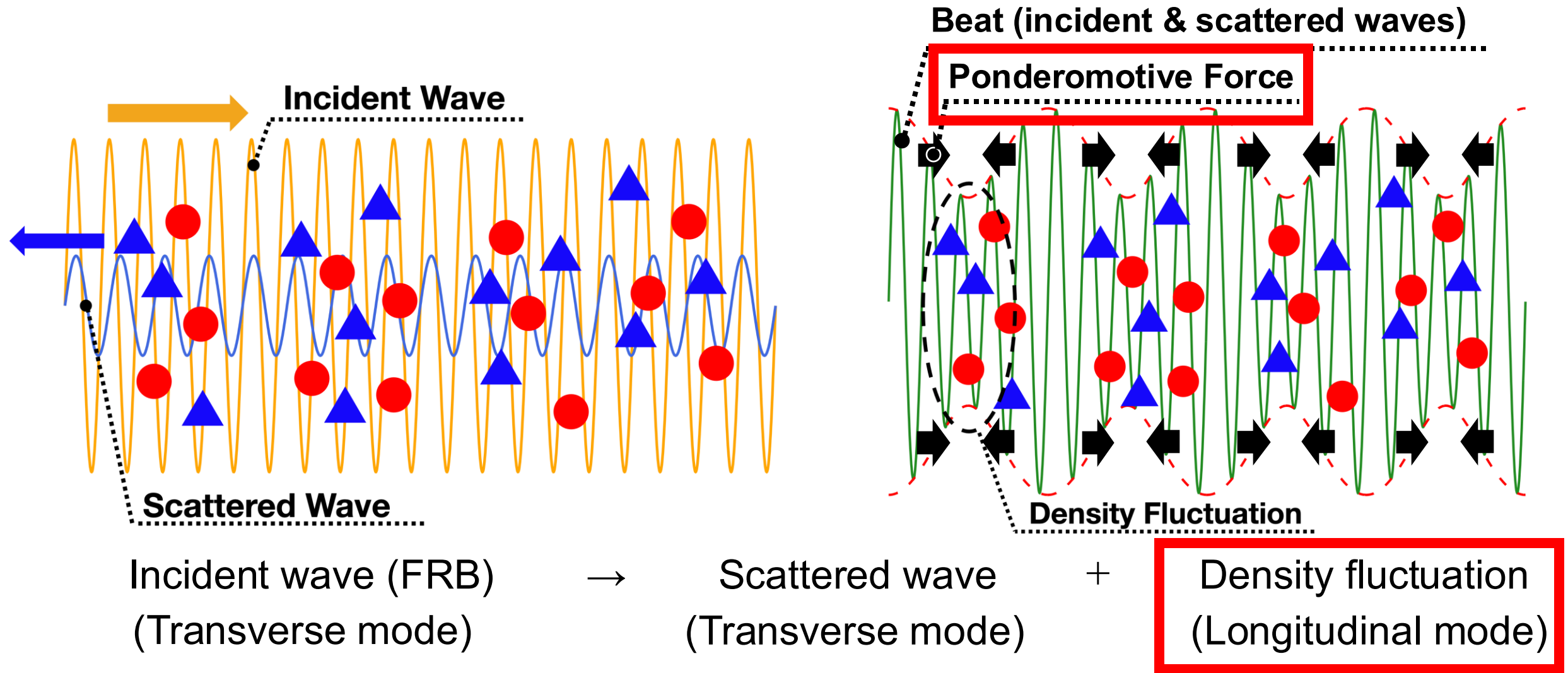
Magnetar Model



The wave propagation in magnetized plasma is common for both models.
Induced scattering (parametric instability) is important in plasma.
→ This could limit the emission region and mechanisms.

**We investigate induced scattering (parametric instability)
in magnetized electron-positron pair plasma.**

Induced Scattering (Parametric Instability)



Incident waves (FRB) lose their energy while propagating in plasma.

Ponderomotive Force w/ Background B-field

[e.g. Klima 1966, 1968, Lee & Parks 1983, 1996]

$$\vec{F}_{\text{pond}} = -\frac{e^2}{4m} \nabla \left[\underbrace{-\frac{|E_x|^2 + |E_z|^2}{\omega_c^2 - \omega_0^2}}_{\text{blue underline}} \underbrace{\pm i \frac{\omega_c(E_z^* E_x - E_x^* E_z)}{\omega_0(\omega_c^2 - \omega_0^2)}}_{\text{red underline}} \right]$$

+ : Positron

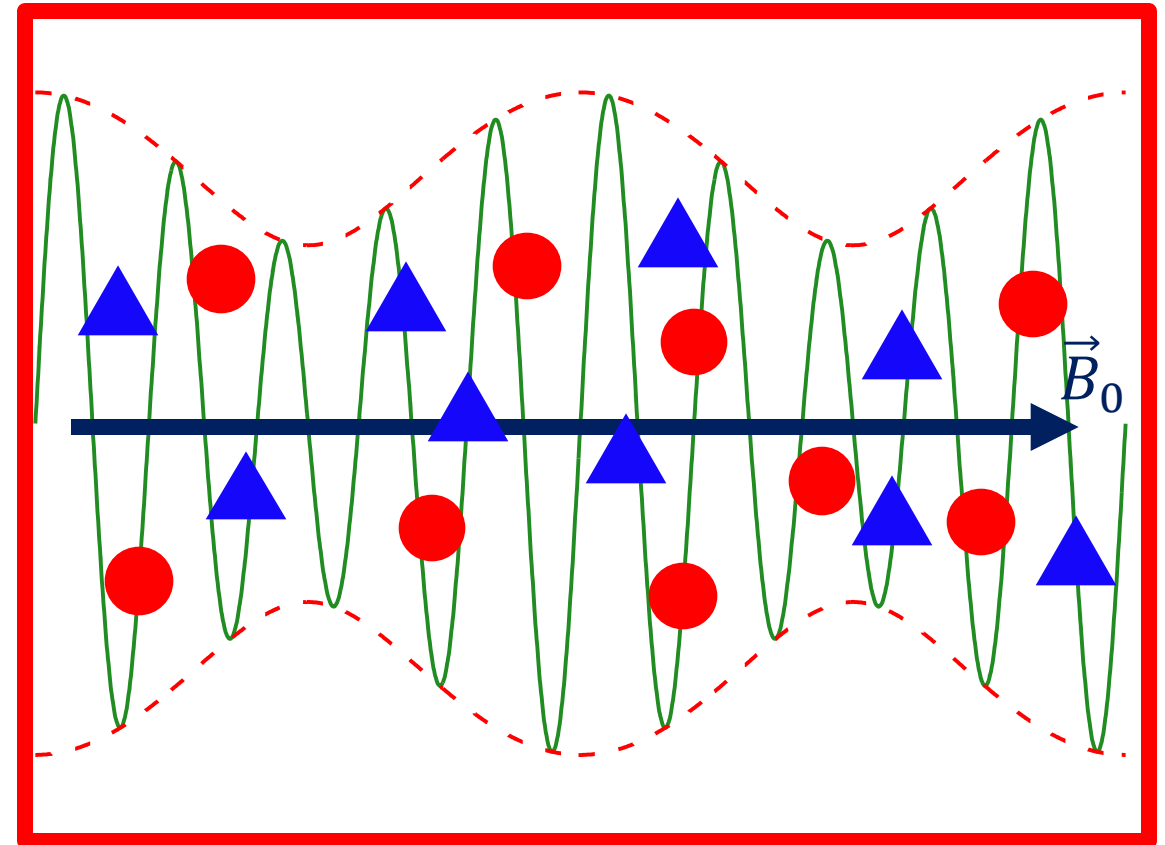
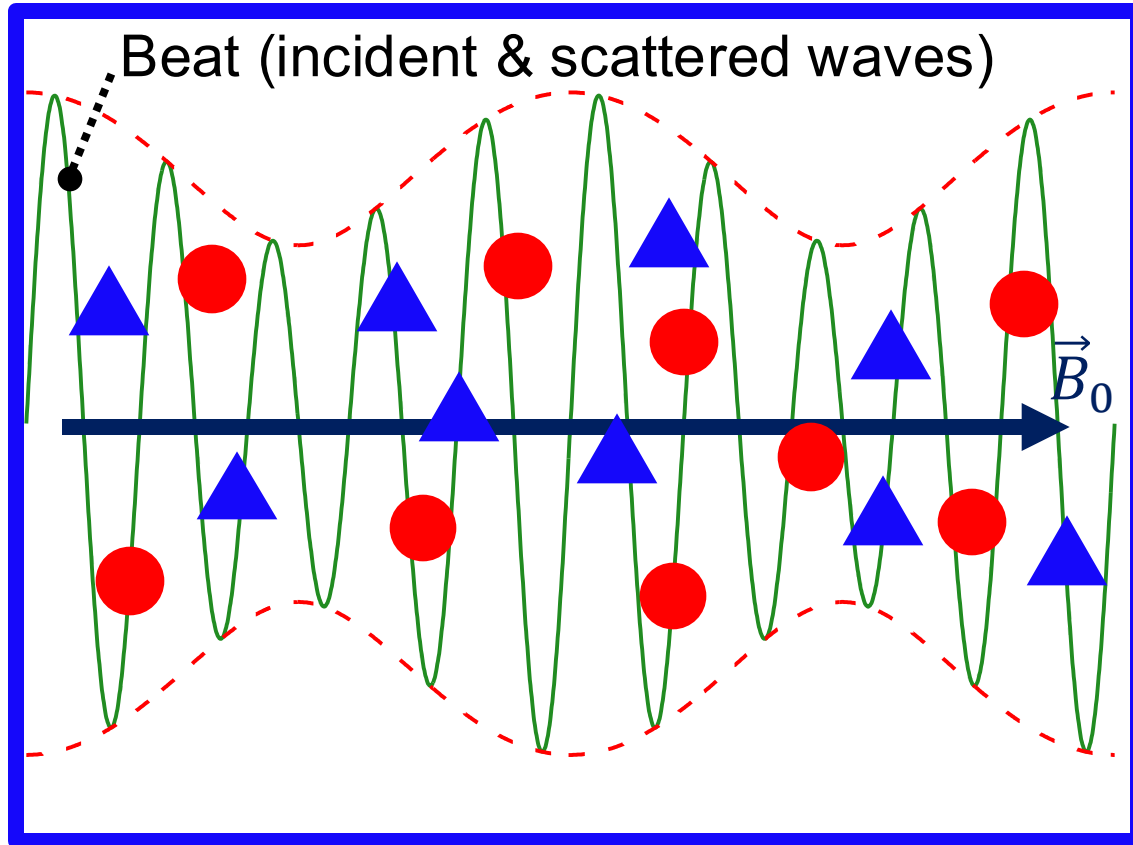
− : Electron

$\vec{B}_0, \vec{k}_0 \parallel y$

● : Positron

▲ : Electron

$\omega_c = eB_0/(m_e c)$: cyclotron freq. $> \omega_0$: incident wave freq.



Induced Compton Scattering Rate w/ Background B-field

[Nishiura & Ioka 2024, Nishiura, **SFK**+2025]

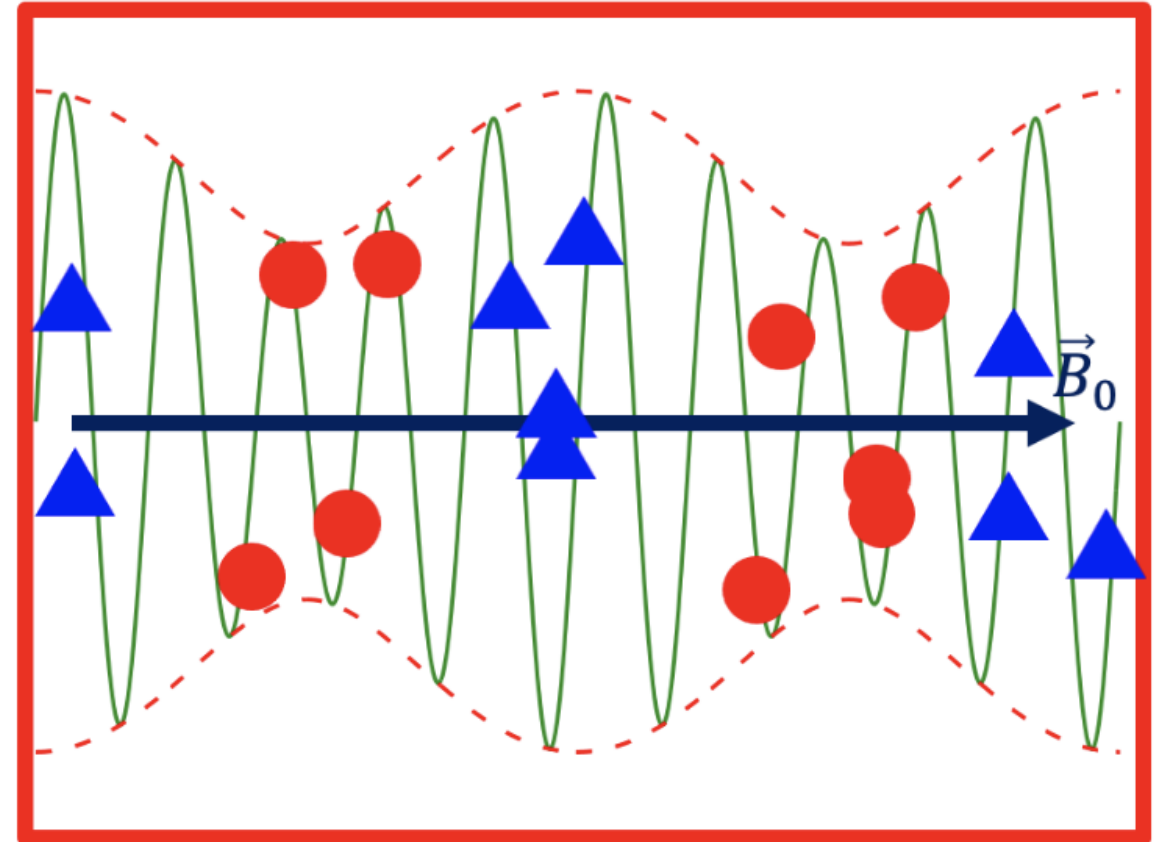
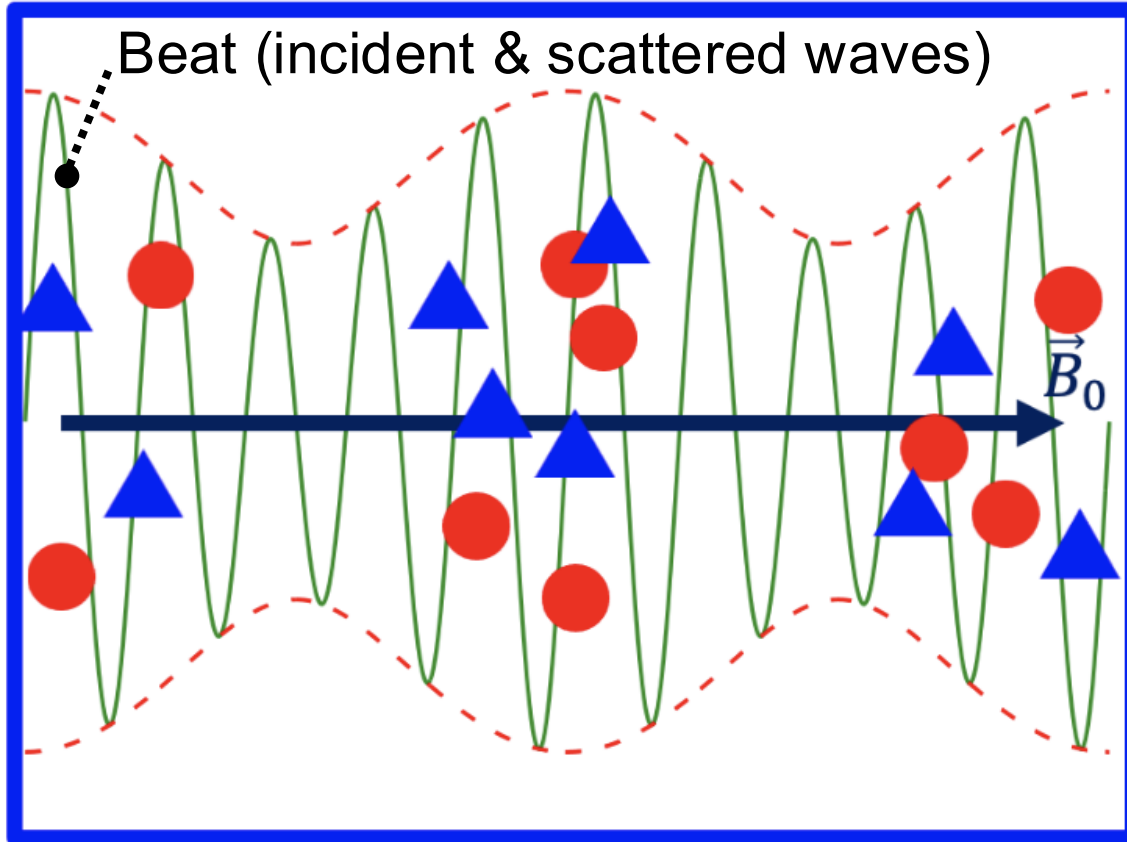
$$\Gamma_{\max}^{\text{w/ } B_0} = \Gamma_{\max}^{\text{w/o } B_0} \max \left[\underbrace{\left(\frac{\omega_0}{\omega_c} \right)^4}_{\text{Neutral Mode}}, \underbrace{\frac{32e}{\pi} \left(\frac{\omega_0}{\omega_c} \right)^2 \left(\frac{\omega_0}{\omega_p} \right)^4 \left(\frac{k_B T_e}{m_e c^2} \right)^2}_{\text{Charged Mode}} \right]$$

● : Positron

Neutral Mode

Charged Mode

▲ : Electron $\omega_c = eB_0/(m_e c)$:cyclotron freq. $> \omega_p$: plasma freq. $> \omega_0$:incident wave freq.



Parametric Instability for FRB

□ Growth rate of ICS w/ background B-field

$$\Delta t^{-1} = 10^3 \text{ s}^{-1}$$

(Inverse of the burst duration)

$$\begin{aligned} (t_{\text{charged}}^{\text{broad}})^{-1} &= \frac{\pi}{4} \left(\frac{\omega_0}{\omega_c} \right)^2 \frac{\omega_p^2 a_e^2}{\omega_0} \left(\frac{8k_B T_e}{m_e c^2} \right)^2 \left(\frac{\omega_0}{\omega_p} \right)^4 \left(\frac{\omega_0}{\Delta\omega} \right)^2 \\ &= 4.6 \times 10^2 \text{ s}^{-1} \frac{P_{\text{sec}} r_8^7 B_{p,14} L_{38} T_{80\text{keV}}^2}{\mathcal{M}_6 \nu_9^2 B_{p,14}^3 R_{\text{NS},6}^9} \left(\frac{\Delta\nu/\nu_0}{1} \right)^{-2} \sim \Delta t^{-1} \end{aligned}$$

$$\begin{aligned} (t_{\text{neutral}}^{\text{broad}})^{-1} &= \pi \frac{\omega_p^2 a_e^2}{\omega_0} \left(\frac{\omega_0}{\omega_c} \right)^4 \left(\frac{\omega_0}{\Delta\omega} \right)^2 \\ &= 1.8 \times 10^{-2} \text{ s}^{-1} \frac{\mathcal{M}_6 L_{38} \nu_9^2 r_8^7 T_{80\text{keV}}^2}{P_{\text{sec}} B_{p,14}^3 R_{\text{NS},6}^9} \left(\frac{\Delta\nu/\nu_0}{1} \right)^{-2} \ll \Delta t^{-1} \end{aligned}$$

Waves can escape.

□ Growth rate of ICS w/o background B-field

[Drake+1974, Ghosh+2022]

$$\left(t_{\text{c,max}}^{\text{w/o } B_0} \right)^{-1} = \sqrt{\frac{\pi}{32e}} \frac{\omega_p^2 a_e^2}{\omega_0} \frac{m_e c^2}{k_B T_e} \sim 10^{20} \text{ s}^{-1} \frac{B_{p,14} \mathcal{M}_6 L_{38} R_{\text{NS},6}}{P_{\text{sec}} r_8^5 \nu_9^3 T_{80\text{keV}}} \gg \Delta t^{-1}$$

Growth Rate in Magnetized Plasma

[Nishiura & Ioka 2024, Nishiura+ in prep.]

Linear growth rate of induced Compton scattering w/ background B-field

$$\Gamma_{\max}^{\text{w/ } B} = \Gamma_{\max}^{\text{w/o } B} \max \left[\underbrace{\left(\frac{\omega_0}{\omega_c} \right)^4}_{\text{Neutral Mode}}, \underbrace{\frac{32e}{\pi} \left(\frac{\omega_0}{\omega_c} \right)^2 \left(\frac{\omega_0}{\omega_p} \right)^4 \left(\frac{k_B T_e}{m_e c^2} \right)^2}_{\text{Charged Mode}} \right]$$

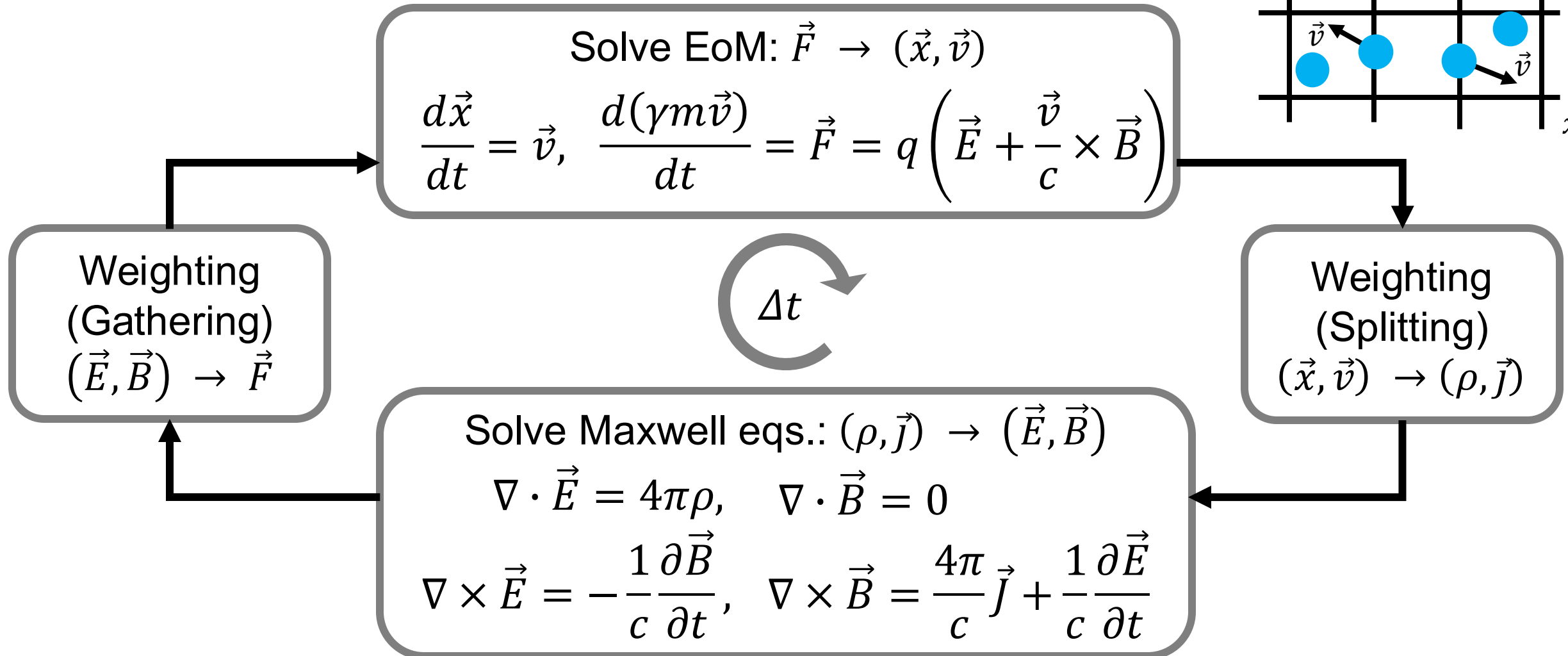
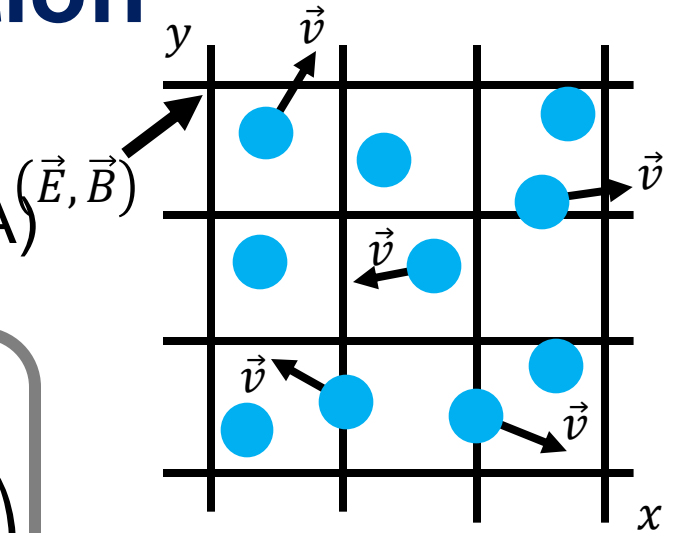
Neutral Mode **Charged Mode**

We investigate the linear and the nonlinear stages of neutral and charged modes of induced Compton scattering by using Particle-in-Cell (PIC) simulations.

$$\omega_0 < \omega_p < \omega_c$$

Particle-in-Cell (PIC) Simulation

- ❑ WumingPIC2D (public code) [Matsumoto+2024]
- ❑ Flow (Nagoya U.), Yukawa-21 (YITP), XC50, XD2000 (CfCA)



Setup

□ Right-handed circular pol. Alfvén wave (incident wave) $\Delta x = \Delta y = \Delta t = 1, m_e = 1, c = 1$

$$\vec{B}_p = B_p [-\sin \phi_0 \hat{x} + \cos \phi_0 \hat{z}], \phi_0 = k_0 y - \omega_0 t$$

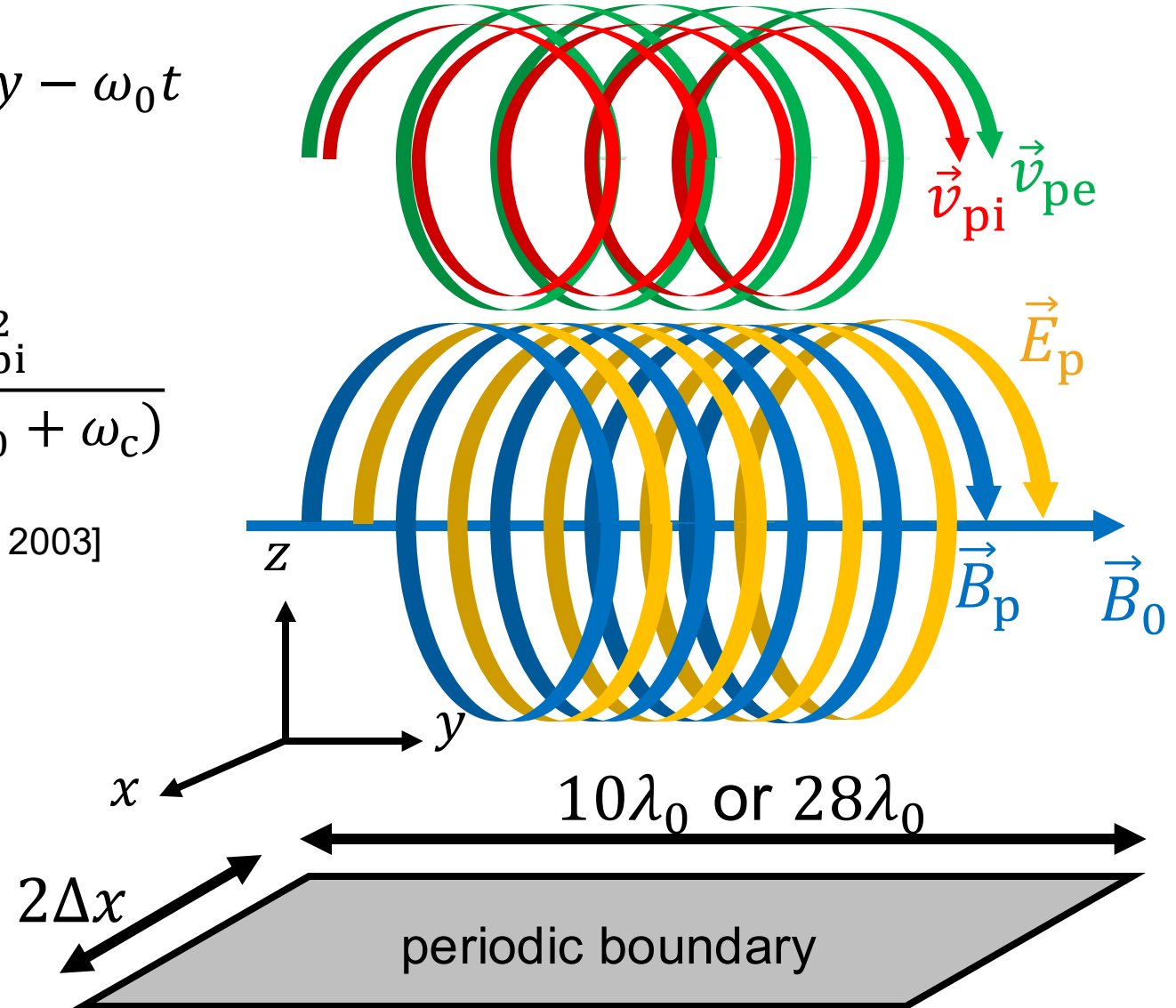
$$\vec{E}_p = -\frac{\omega_0}{ck_0} B_p [\cos \phi_0 \hat{x} + \sin \phi_0 \hat{z}]$$

$$\left(\frac{ck_0}{\omega_0}\right)^2 = 1 - \frac{\omega_{pe}^2}{\omega_0(\gamma_e \omega_0 - \omega_c)} - \frac{\omega_{pi}^2}{\omega_0(\gamma_i \omega_0 + \omega_c)}$$

□ initial $e^\pm$ plasma velocity [Matsukiyo & Hada 2003]

$$\frac{\vec{v}_{e(i)}}{c} = -\frac{\omega_0}{ck_0} \frac{\eta \omega_c}{\omega_c \pm \gamma_{e(i)} \omega_0} \frac{\vec{B}_p}{B_p}$$

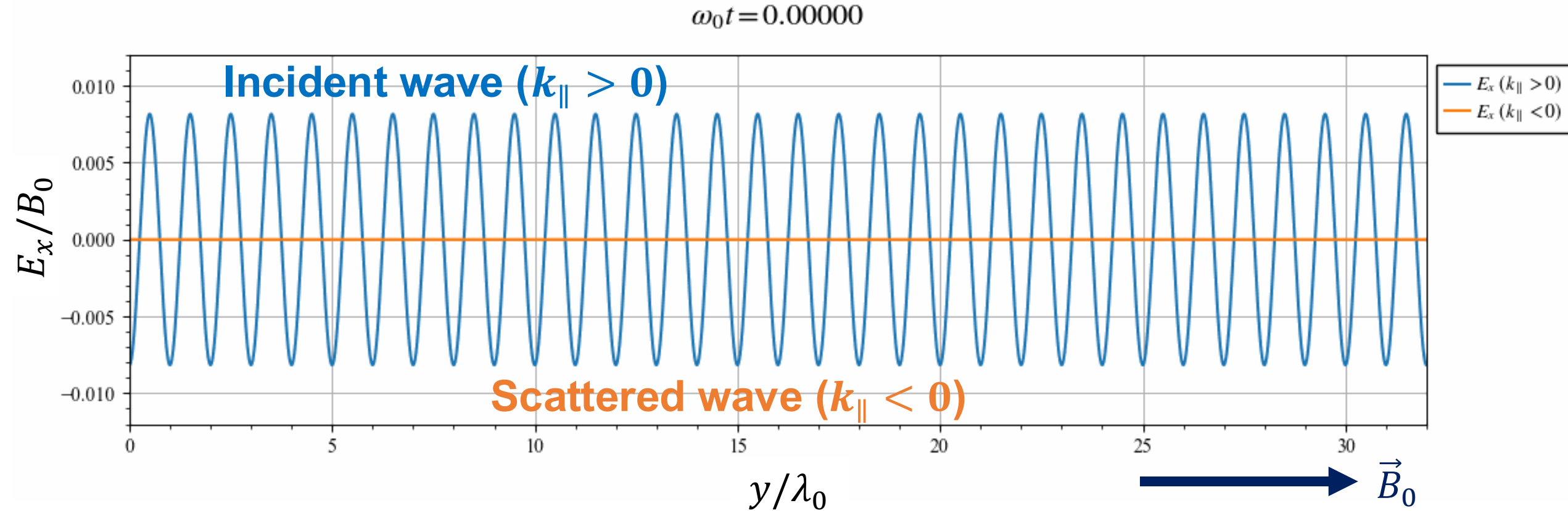
$$\gamma_{e(i)} = \frac{1}{\sqrt{1 - \left(\frac{v_{e(i)}}{c}\right)^2}} \quad \eta = \frac{B_p}{B_0} = a_0 \frac{\omega_0}{\omega_c}$$



Simulation Setup

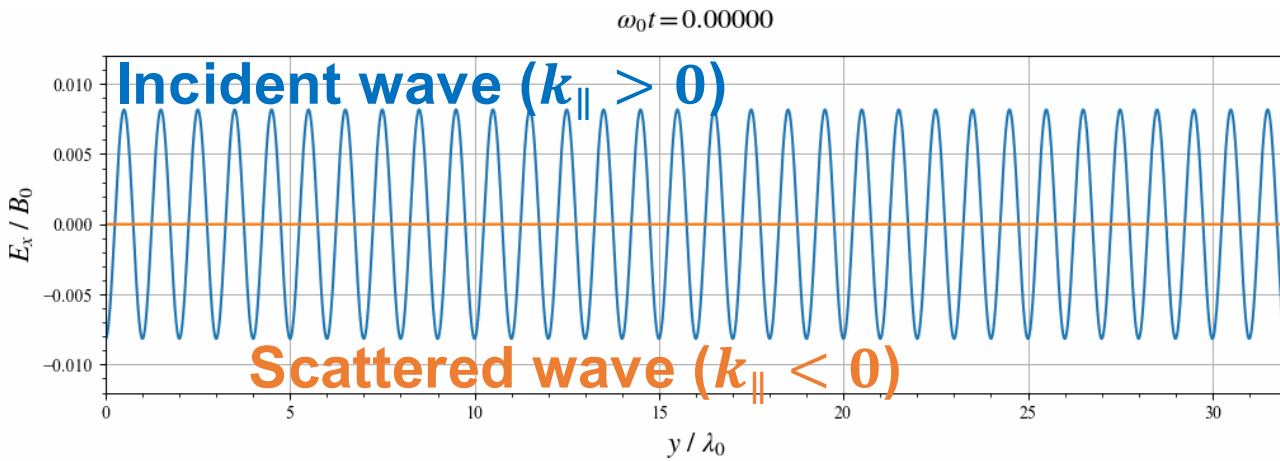
	Charged	Neutral	Neutral
mass ratio: m_r	1		
frequency ratio: $\frac{\omega_0}{\omega_{pe}}$ $(\omega_0 \Delta t < \omega_{pe} \Delta t < \omega_c \Delta t = \sqrt{\sigma_e} \omega_{pe} \Delta t < 0.1)$	0.9		0.1
electron sigma: $\sigma_e = \frac{B_0^2}{4\pi n m_e c^2} = \left(\frac{\omega_c}{\omega_{pe}}\right)^2$	2500		4
thermal velocity: $\frac{v_{th,e}}{c} = \sqrt{\frac{k_B T_e}{m_e c^2}} = \frac{v_{th,i}}{c}$	0.21	0.04	0.04
amp. of incident wave: $\frac{B_p}{B_0}$	0.1, 0.141, 0.316, 0.562		0.01, 0.0141, 0.0316
wavelength of incident wave: $\lambda_0 / \Delta x$	3200		1200
The number of particles per cell: n [/cell]	100		
simulation box: $L_x \times L_y$	$2\Delta x \times 10\lambda_0$	$2\Delta x$	$2\Delta x \times 32\lambda_0$

Time evolution of waves



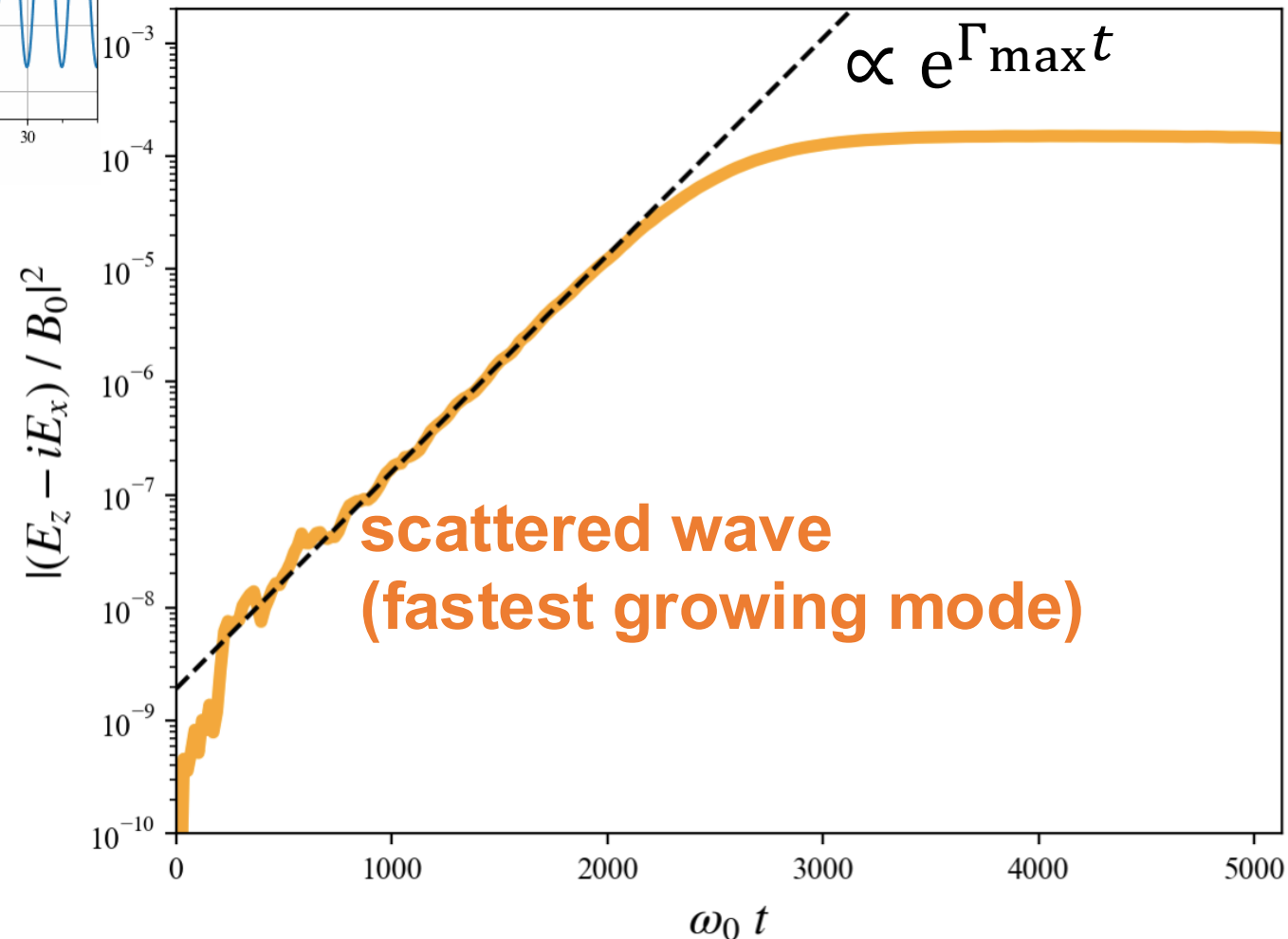
The growth rate is estimated from the time evolution of the power of the scattered wave.

growth rate from simulations

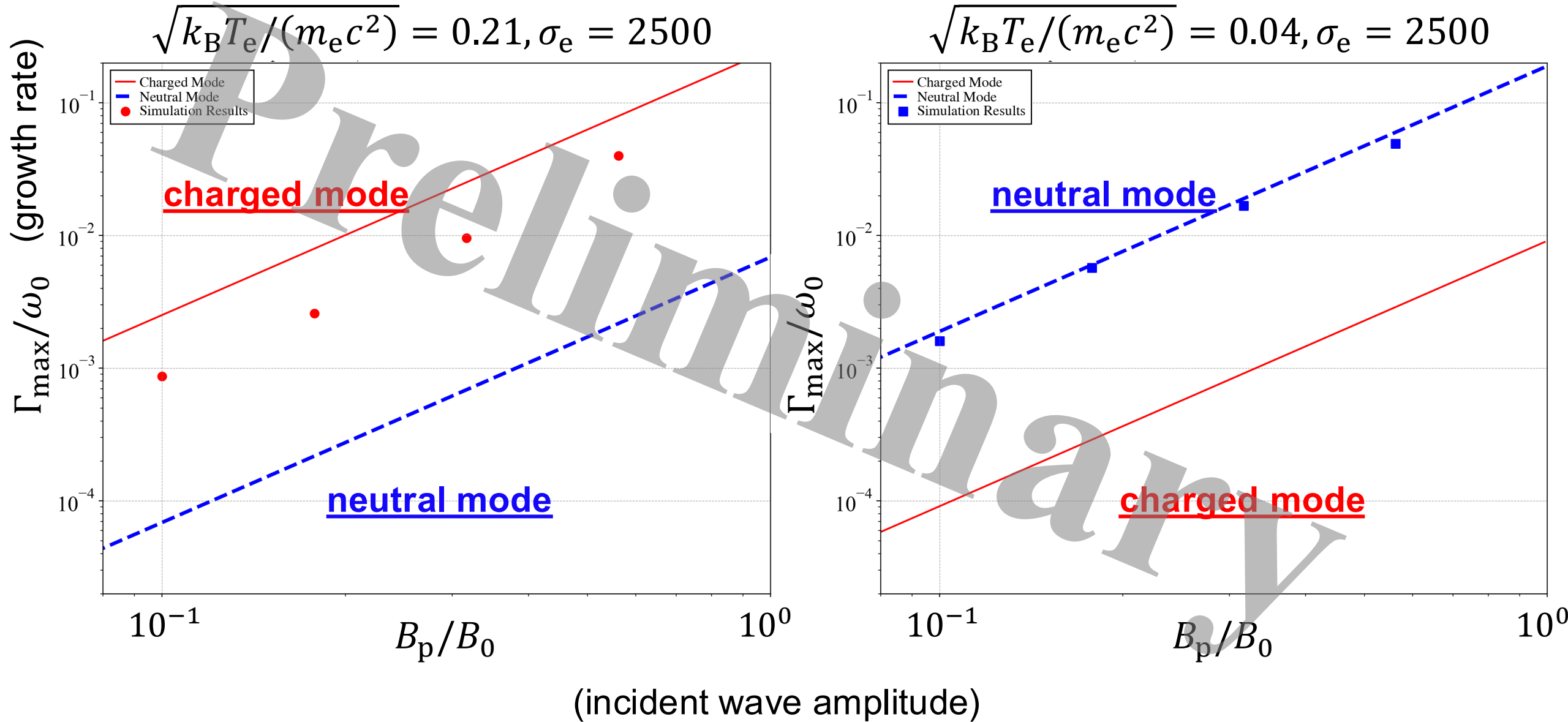


We extract the fastest growing mode of the scattered waves.

$\Gamma_{\max}$ is given from simulations.

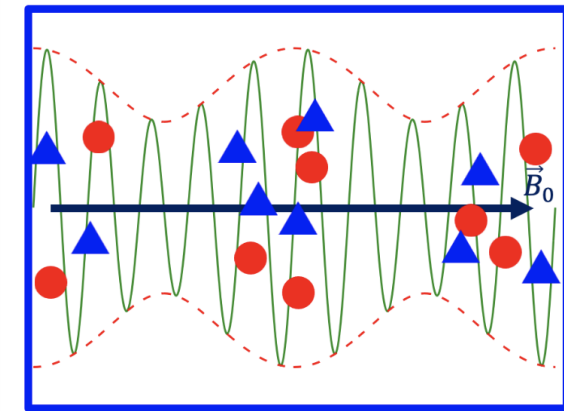
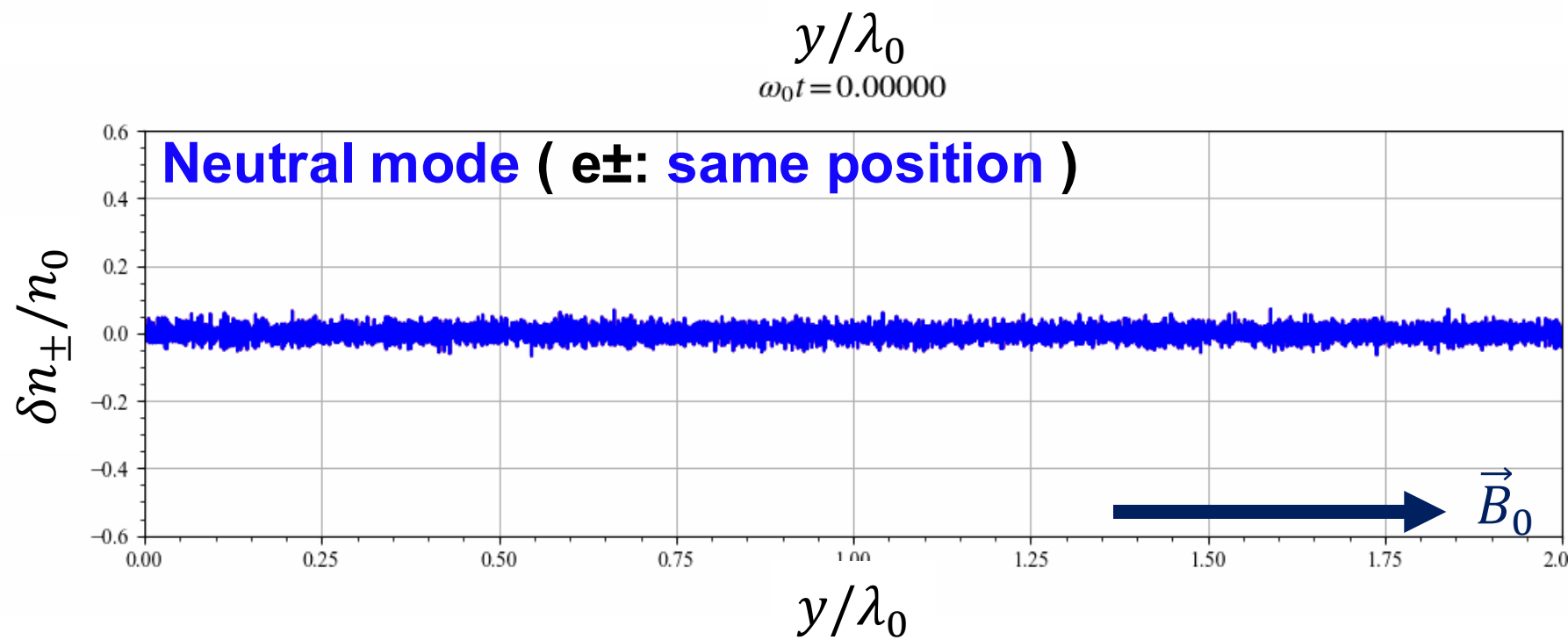
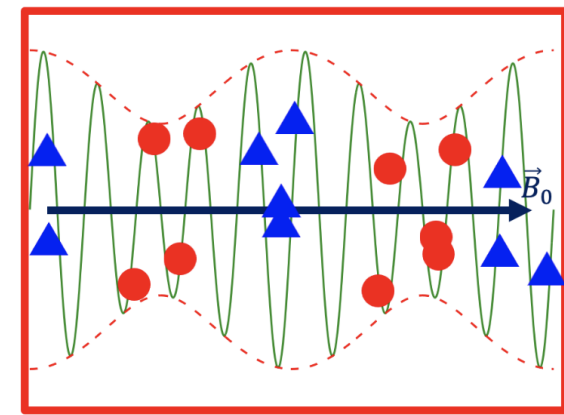
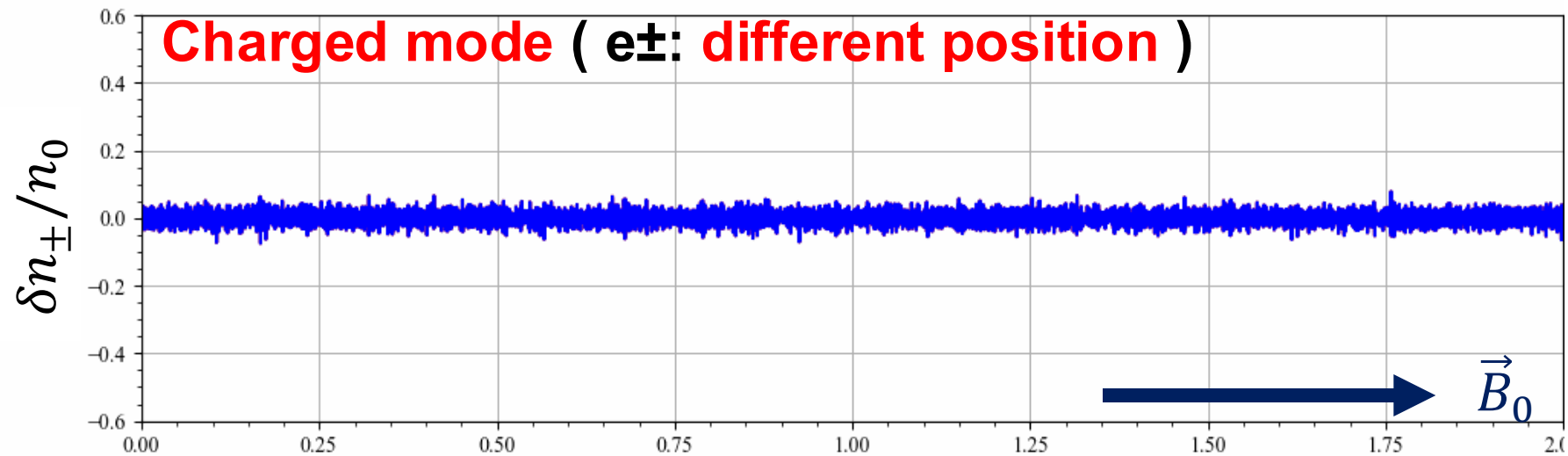


Growth rate of induced Compton scattering



Density fluctuation

$$\omega_0 t = 0.00000$$



density fluctuations of
electrons and positrons

Summary & Future Work

- ❑ We investigate propagation of Alfvén waves in magnetized pair plasma by using Particle-in-Cell simulations.
- ❑ Simulation results are almost in good agreement with the theoretical growth rate of induced Compton scatterings (charged & neutral modes)
- ❑ Incident wave: plane wave $\rightarrow$ pulse, circular pol. $\rightarrow$ linear pol.
- ❑ Dependency of other parameters