### Space-time random tensor networks and holographic duality

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### Outline

- Background: tensor networks and holographic duality
- Definition of space-time random tensor networks (RTN)
- Entanglement properties of space-time RTN --HRT formula with quantum corrections
  - --Quantum error correction properties of the holographic mapping
- Gauge fixing and finite D correction Reference



- Zhao Yang, XLQ, in preparation
- Closely related previous work on spatial RTN Patrick Hayden, Sepehr Nezami, XLQ, Nathaniel Thomas, Michael Walter, Zhao Yang, arxiv: 1601.01694

#### Tensor networks

Spatial tensor networks: many-body entangled states made from few-qubit states. (PEPS, F. Verstraete, J.I. Cirac, 04')





Space-time tensor network: a discretization of path integral.



#### Tensor networks and holographic duality



Ryu&Takayanagi '06, Swingle '09



- The tensor network proposal: geometry emerges from the entanglement structure of quantum states
- Various tensor network proposals (Nozaki et al '12, XLQ '13, Pastawski et al '15, Yang et al '15, Hayden et al '16)
- Entanglement properties can be studied for tensors with good properties (Pastawski et al '15, Yang et al '15)

Random tensor networks Hayden et al arxiv: 1601.01694

• Tensor networks with a random tensor at each vertex



- In D → ∞ limit, many interesting properties emerge ---RT formula and its quantum correction ---Error correction properties of bulk-boundary operator correspondence ---Scaling dimension gap
- Problem: Need to pick a time slice. Hard to talk about dynamics. Solution: going to space-time networks.

#### Space-time tensor networks

discretize

 Space-time tensor networks are discretized version of path integral.

• Or 
$$Z = \int D\phi e^{-S(\phi)}$$



$$Z = \int D\phi e^{iS(\phi)}$$

$$\langle TO_1O_2O_3 \dots O_n \rangle = \frac{1}{Z} \times$$

#### Space-time tensor networks

• Example: Ising model  $S = -J \sum_{\langle xy \rangle} s_x s_y$ 





Space-time tensor network for holography

#### • <u>Wanted:</u>

---a tensor network representation of the (d + 1)-dimensional bulk geometry

---The tensor network defines  $Z_{bulk} = Z_{boundary}$ , a d-dimensional QFT.



#### Space-time tensor network for holography

- Our proposal:
- 1. Introduce a space-time tenor network in the bulk
- This defines a bulk QFT, which we called the "parent theory". (Can be defined for either Euclidean or Lorentzian time)
- 2. Introduce a random projection at each bulk link
- Key result: n-th Renyi entropy $\Rightarrow$  $S^n$  discrete gauge theory
- Subtlety: gauge fixing (will discuss later)



### Holographic duality from space-time tensor networks *i*

- Random projection at each link
- Correlation function  $\langle TO_1O_2O_3 \dots O_n \rangle =$





### Second Renyi entropy of random tensor networks

• Second Renyi entropy definition.  $tr(\rho_A^2) = tr(\rho \otimes \rho \cdot$ 





#### Second Renyi entropy



• The random average leads to the partition function of a  $Z_2$  gauge theory.

• 
$$\overline{tr(\rho_A^2)} = \sum_{g_{xy}=I \text{ or } X} e^{-\mathcal{A}[g_{xy}]}$$

- The choice of boundary region A determines the boundary condition of gauge field  $g_{xy}$ .
- Gauge invariance = permutation symmetry between two replica

#### Second Renyi entropy

• The  $Z_2$  gauge field is minimally coupled to two copies of the parent theory



#### Second Renyi entropy

• For example, if the parent theory is an Ising model

• 
$$S_P = -\sum_{\langle xy \rangle} s_x s_y$$

• The random average gives ( $\alpha = 1,2$ )

• 
$$S_{bulk}[s_x^{\alpha}, g_{xy}] = -\sum_{\langle xy \rangle} s_x^{\alpha} g_{xy}^{\alpha\beta} s_y^{\beta}$$

- $[g_{xy}^{\alpha\beta}] = I$  or  $\sigma_x$  is the  $Z_2$  connection
- Schematically, if we could take continuum limit,
- $e^{-\mathcal{A}_{eff}[g]} \equiv \int D\phi e^{-\mathcal{A}_{P}[\phi^{1},\phi^{2},g]}$ (or corresponding form in real time)

### HRT formula

- The dynamics of the gauge field is induced by integrating out the bulk "parent theory".
- Consider a particular example: the valence bond solid (VBS) state



- Gauge field action:  $S = -\frac{1}{2} \log D \sum_{p \in \Box} F_p$
- $F_p = \prod_{\langle xy \rangle \in p} g_{xy}$  is the flux in plaquette p.



#### HRT formula

- For the VBS state,  $Z = \sum_{\gamma = flux \ config.} e^{-\log D|\gamma|}$
- In large D limit,  $Z \simeq e^{-\log D|\gamma_A|}$ , and  $S_2 \simeq \log D |\gamma_A|$
- $|\gamma_A| = \min |\gamma|$  is the area of minimal co-dimension-2 surface bounding A.
- In general, when the bulk parent theory is a massive theory, the effective action of gauge theory has a similar Maxwell form, leading to HRT formula (Hubeny Rangamani Takayanagi '07).



#### HRT formula for Lorentzian time

- The bulk parent theory can be defined with either Euclidean or Lorentzian signature.
- For Lorentzian bulk theory, the effective action of gauge field is also Lorentzian.
- Lorentzian Maxwell action of a  $\mathbb{Z}_2$  gauge theory (continuum limit)

• 
$$S = -\frac{1}{4g^2}F_{\mu\nu}F^{\mu\nu} + \rho_s(\partial_\mu\theta - 2A_\mu)(\partial^\mu\theta - 2A^\mu)$$

- In the classical limit  $(\frac{1}{g^2}, \rho_s \to \infty), Z \simeq e^{i\alpha |\gamma_A|}$
- The area of saddle point surface  $|\gamma_A|$  is imaginary (real) for surface with Euclidean (Lorentzian) signature.  $\alpha \propto \log D$

#### HRT formula for Lorentzian time

- In our approach, HRT surface can be defined even for time-like regions.
- Renyi entropies are generalized to multipoint functions of twist operators.
- $e^{-S_n(n-1)} = \langle TX_n(x_2, t_2)X_n(x_1, t_1) \rangle = \sum_{\gamma} e^{i\mathcal{A}(\gamma)}$
- HRT formula can be generalized if  $\mathcal{A}(\gamma) \propto (n-1)|\gamma|$  and if the saddle point approximation applies.



#### Quantum corrections to HRT

 Consider a bulk parent theory with VBS⊗low energy field theory





- Int. out VBS leads to  $S = -\frac{1}{2}\log D\sum_{p}F_{p}$  for gauge field.
- Gauge field is coupled to the low energy theory with dimension  $D_b \ll D$
- $S_2(A) \simeq \mathcal{A}_{cl} = \log D |\gamma_A| + S_{2bulk}(E_A)$
- Both terms are entanglement entropy in the bulk



Consistent with Faulkner et al '13

## Generalization to higher Renyi entropies

- The random average can be generalized to higher Renyi entropies  $tr(\rho_A^n) = tr(\rho^{\otimes n}X_{An})$
- $\overline{\phi_{i_1}^* \phi_{j_1} \phi_{i_2}^* \phi_{j_2} \dots \phi_{i_n}^* \phi_{j_n}} = \frac{1}{C_n} \sum_{g \in S^n} g_{i_1 i_2 \dots i_n}^{j_1 j_2 \dots j_n}$
- Summation over all permutation group elements



### Generalization to higher Renyi entropies

• Random average of n copies of the network  $\Rightarrow S^n$ gauge theory minimal coupled to n replica of the parent theory

• 
$$e^{-\mathcal{A}[g^{(n)}]} \equiv \int D\phi^{(i)} e^{-\mathcal{A}_P[\phi^{(1)},\phi^{(2)},...,\phi^{(n)},g^{(n)}]}$$

• For short-range entangled states such as VBS, in large *D* limit one obtains

 $S_n \simeq \frac{1}{n-1} \mathcal{A}_{cl} = \log D |\gamma_A| + S_{n \ bulk}(E_A)$ with  $S_{n \ bulk}(E_A)$  the bulk low energy field contribution

• Leading order Renyi entropies are *n* independent, different from CFT results. Reason: absence of back reaction. (c.f. Xi Dong 1601.06788)

#### Operator correspondence

- Definition of Bulk local operators:
- Bulk operators  $\phi_x$  acting on low energy subspace have nontrivial effect to the boundary after random projection.





• Does each bulk *low energy* operator correspond to a boundary operator? If yes, can the bulk operator be represented in a region of the boundary?

#### Operator correspondence



- Different operators create different states in the parent theory  $|\psi_{1P}\rangle = \prod \phi_i(x_i, t_i) |G\rangle$
- Correspondingly there are different boundary states after the random projection.
- Generically each state gives different gauge field action.

#### Error correction properties

- A "code subspace" can be defined as the subspace of bulk states with the same gauge field action (i.e. quantum corrections to HRT can be ignored)
- For such states in  $D \to \infty$ ,  $tr_P(\rho_{1P}(E_A)\rho_{2P}(E_A)^n) = tr(\rho_1(A)\rho_2(A)^n) \forall n. \Rightarrow S(\rho_{1P}(E_A)|\rho_{2P}(E_A)) = S(\rho_1(E_A)|\rho_2(E_A))$

Random

projection

 $\Rightarrow$  (Jaffer et al '15, Harlow et al '16)

• For each bulk operator  $\phi$  acting in  $E_A$ ,  $\exists$  boundary operator  $O_A$ , s.t.  $\langle \psi_1 | O_A | \psi_2 \rangle = \langle \psi_{1bulk} | \phi | \psi_{2bulk} \rangle$  for any two states in the code subspace.

#### Gauge redundancy and gauge fixing

- The random average leads to a <u>gauge field</u> partition function with gauge redundancy  $Z_n(A) =$  $\sum_{\{g_{xy} \in S^n\}} e^{-\mathcal{A}[g_{xy}]} = Z_{S^n}(A) \cdot \Omega$
- $\Omega$  is the volume of gauge orbit.
- Not a problem if we calculate ratios such as  $\frac{Z_n(A)}{Z_n(\emptyset)}$
- However, problem happens when we consider fluctuations:

• 
$$\overline{\delta Z_2^2} = \overline{tr(\rho^2)^2} - \overline{tr(\rho^2)}^2 = Z_4(\emptyset) - Z_2(\emptyset)^2 = \Omega_4 Z_{S^4}(\emptyset) - \Omega_2^2 Z_{S^2}(\emptyset)^2$$

• The fluctuation is large because  $\Omega_4 \gg \Omega_2^2$ . We need to remove the gauge redundancy.

#### Gauge redundancy and gauge fixing

- For discrete gauge theories, gauge fixing can be done by directly fixing  $g_{xy}$  on some links, without constraining any gauge flux.
- This corresponds to picking a spanning tree in the bulk, and only impose random projection on links off the tree.
- Fluctuations are suppressed in large *D* limit after gauge fixing. (Similar to spatial RTN Hayden et al '16)



#### Summary



• "Low energy states" actually means states that share a common entanglement structure in large D limit.

#### Open questions

- How to start from the boundary and construct the bulk theory?
- How to take into account of the back reaction and describe correct entanglement properties of CFTs?
- How to obtain the bulk geometry equation (Einstein equation)?
- A formalism in the continuum limit?
- Does the space-time RTN helps us to understand of the Black hole information paradox?
- Does this approach allow us to define holography in flat and positively curved space?





# More details about finite D fluctuations

• If we find an upper bound for

$$\frac{\overline{Z_{nA}^2}}{e^{-\mathcal{A}_A^{(2n)}}} - 1 \leq f(D)$$

and  $f(D) \ll 1$  in the large D limit, then in the limit  $f(D) \ll 1$ ,

$$\operatorname{Prob}\left(\left|S_n(A) - S_n^{RT}(A)\right| \le \delta\right) \ge 1 - \frac{32}{\delta^2} f(D)$$

•  $f(D) = \frac{e^{\lambda \Omega}}{D}$  gives a very loose bound that can be prooved.

- $f(D) = \frac{\Omega}{D^{\frac{\Delta_{2n}}{2}}}$  gives a better bound that's less rigorous.
- $\Omega$  is spacetime volume.