

The origin of fast radio bursts implied from the cosmic stellar- mass density evolution



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YITP workshop Fast Radio Bursts and Cosmic Transients 7 June 2022

Outline

1. Introduction

2. FRB population analyses

TH+2022, MNRAS, 511, 1961

TH+2020c, MNRAS, 498, 3927

3. Testing general relativity using FRBs

TH+2021b, PhysRevD, 104, 124026

Sen, TH+2022, MNRAS, 509, 5636

4. A new telescope plan in Taiwan BURSTT

Lin et al including TH in prep.

Game changers!

FRBs per year

1000

>1000



600

CHIME

FAST

20

- Parkees
 - UTMOST
 - GBT
 - Arecibo
 - ASKAP
- 61 FRBs in total

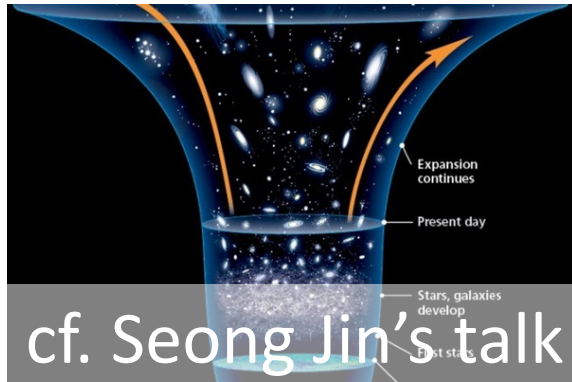
Keane 2018



2021
2022

Key sciences to be addressed by FRBs

Dark energy



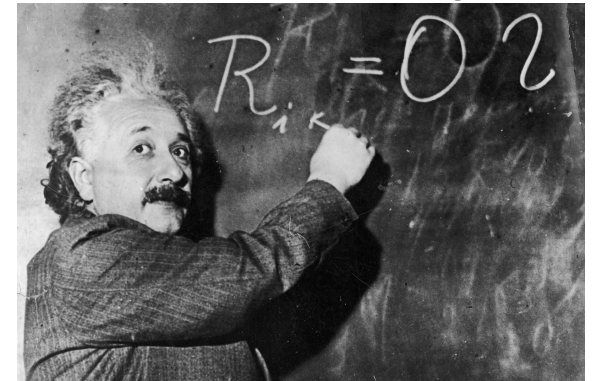
cf. Seong Jin's talk

Missing baryon problem



cf. Ilya/Ryuichi talks

General relativity

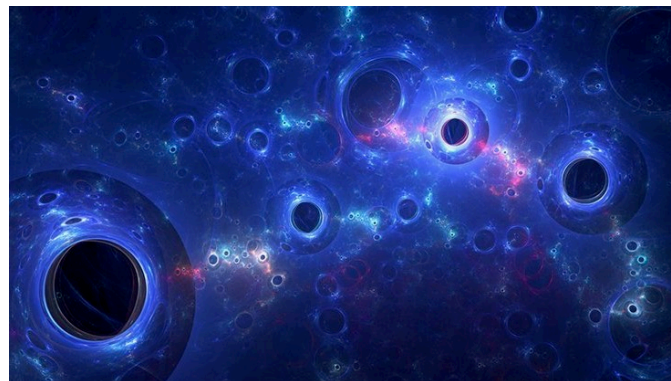


Cosmic reionization



cf. Prof. Kumar's talk

Dark matter

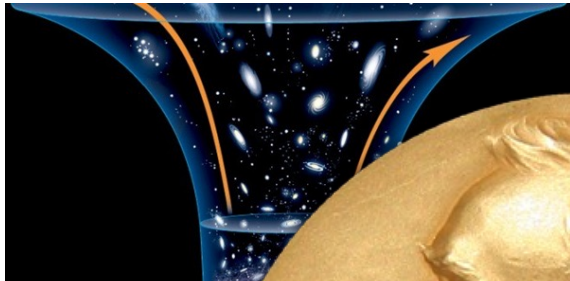


Key sciences to be addressed by FRBs

Dark energy

Missing baryon problem

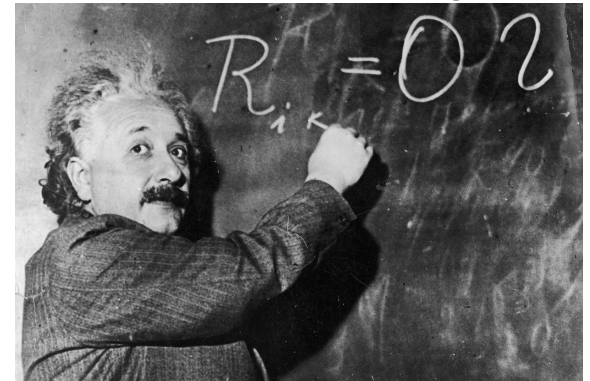
General relativity



cf. Sean



ichi talks

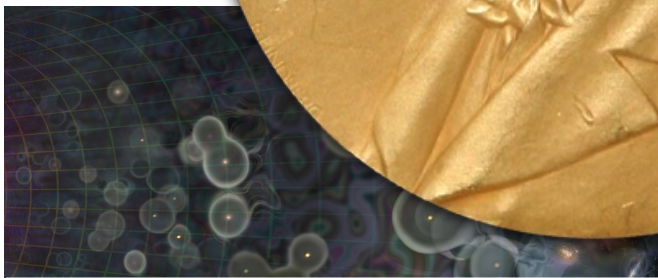


Co

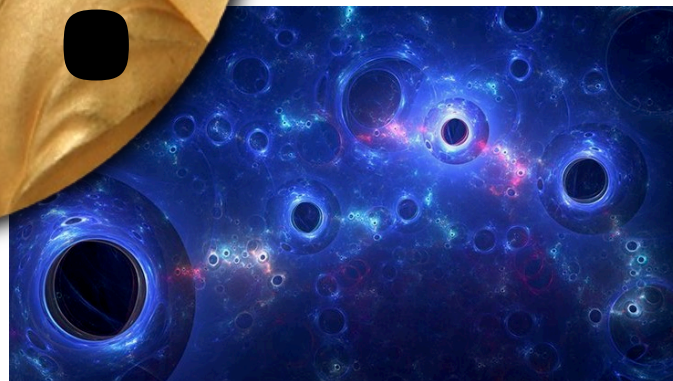


reion

dark matter

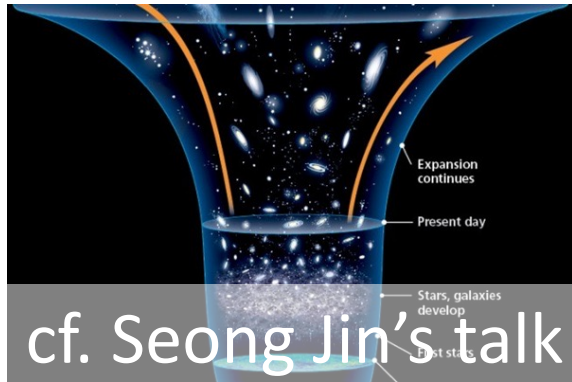


cf. Prof. Kumar's talk



Key sciences to be addressed by FRBs

Dark energy



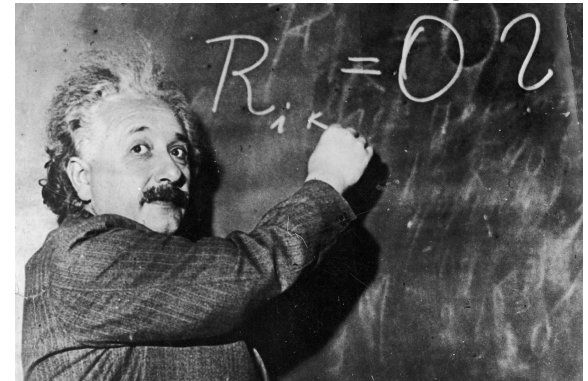
cf. Seong Jin's talk

Missing baryon problem

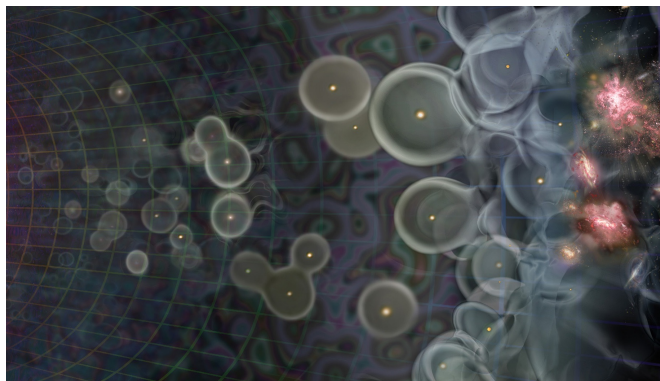


cf. Ilya/Ryuichi talks

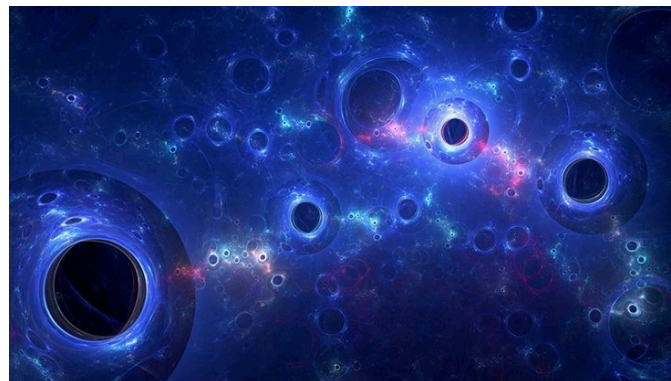
General relativity



Cosmic reionization



Dark matter



The origin of FRBs



Credit: Tetsuya Hashimoto/CICANTHU

Problem

Previous research: tried to 'localize' FRB positions in the sky to reveal their origins
→ didn't work well

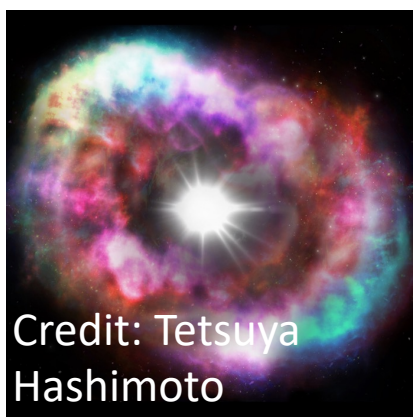
Solution (this work)

We changed the point of view
→ focus on the history of FRBs
(x10 more samples than before)



Introduction: possible FRB origins

White dwarf



Old neutron star



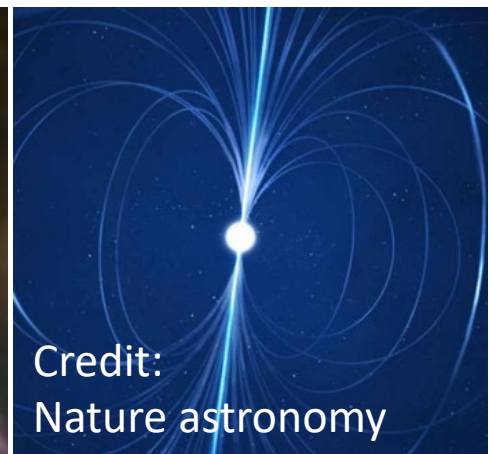
Old stellar-mass black hole (BH)



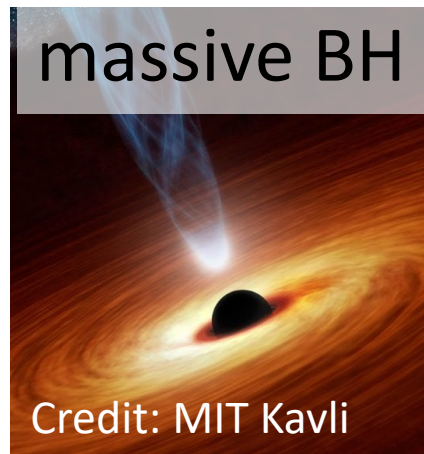
Magnetar



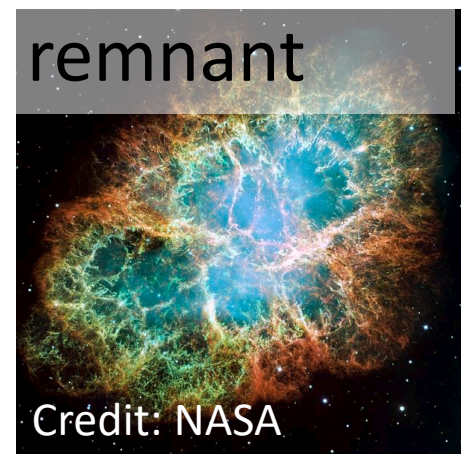
Young pulsar



Super massive BH



Supernova remnant



Introduction: possible FRB origins

White dwarf

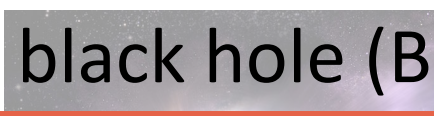


Old neutron



star

Old stellar-mass



black hole (BH)

Old objects \propto stellar mass

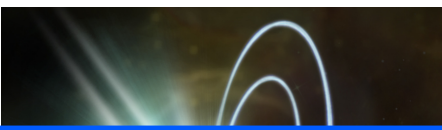
Credit: Tetsuya Hashimoto

Credit: Mark Garlick

Credit: B. Kiziltan/I. Karacan.



Magnetar



Young pulsar



Super



massive BH

Supernova



remnant

Young objects \propto star formation

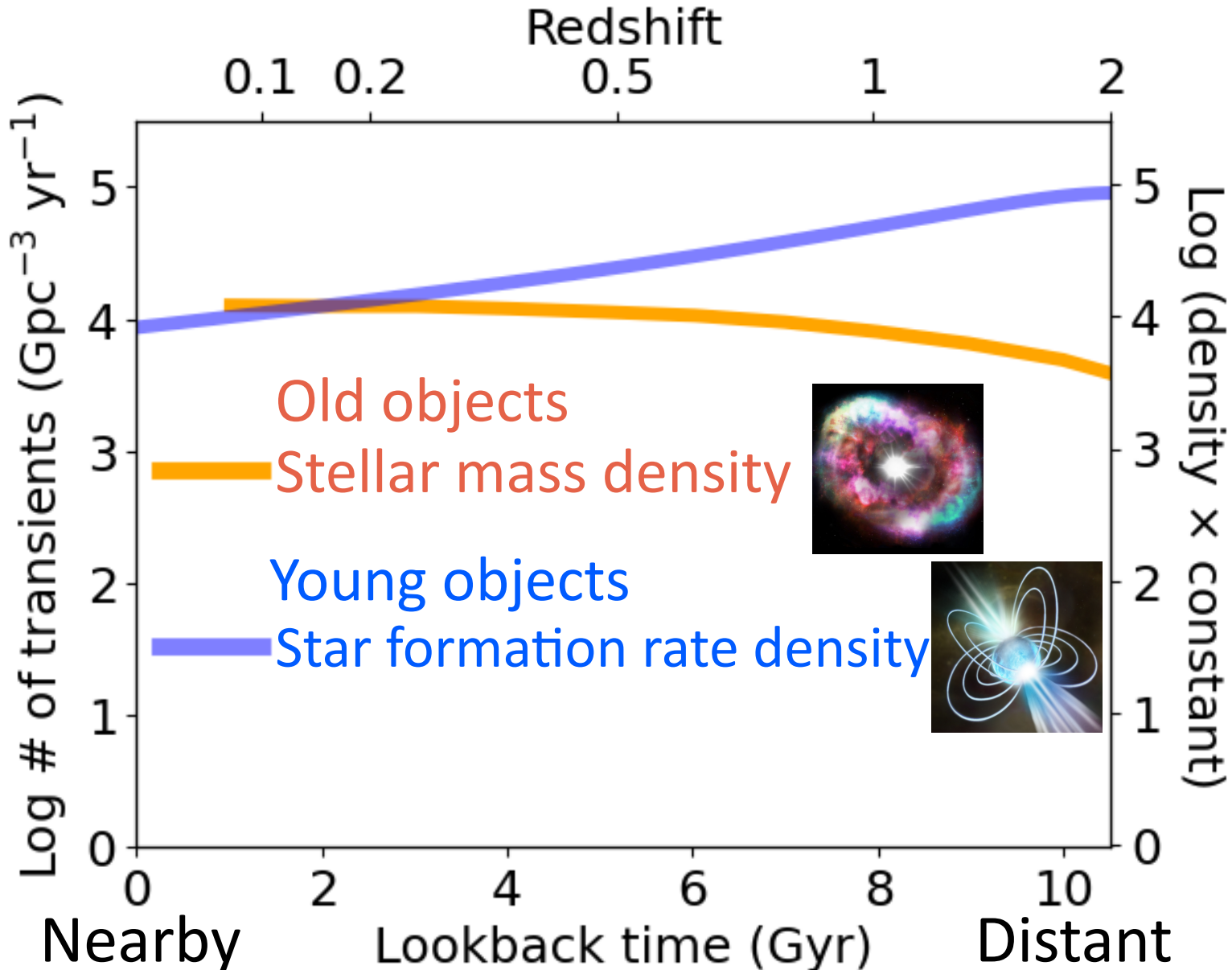
Credit: Tetsuya Hashimoto

Credit: Nature astronomy

Credit: MIT Kavli

Credit: NASA

Introduction: Old vs Young

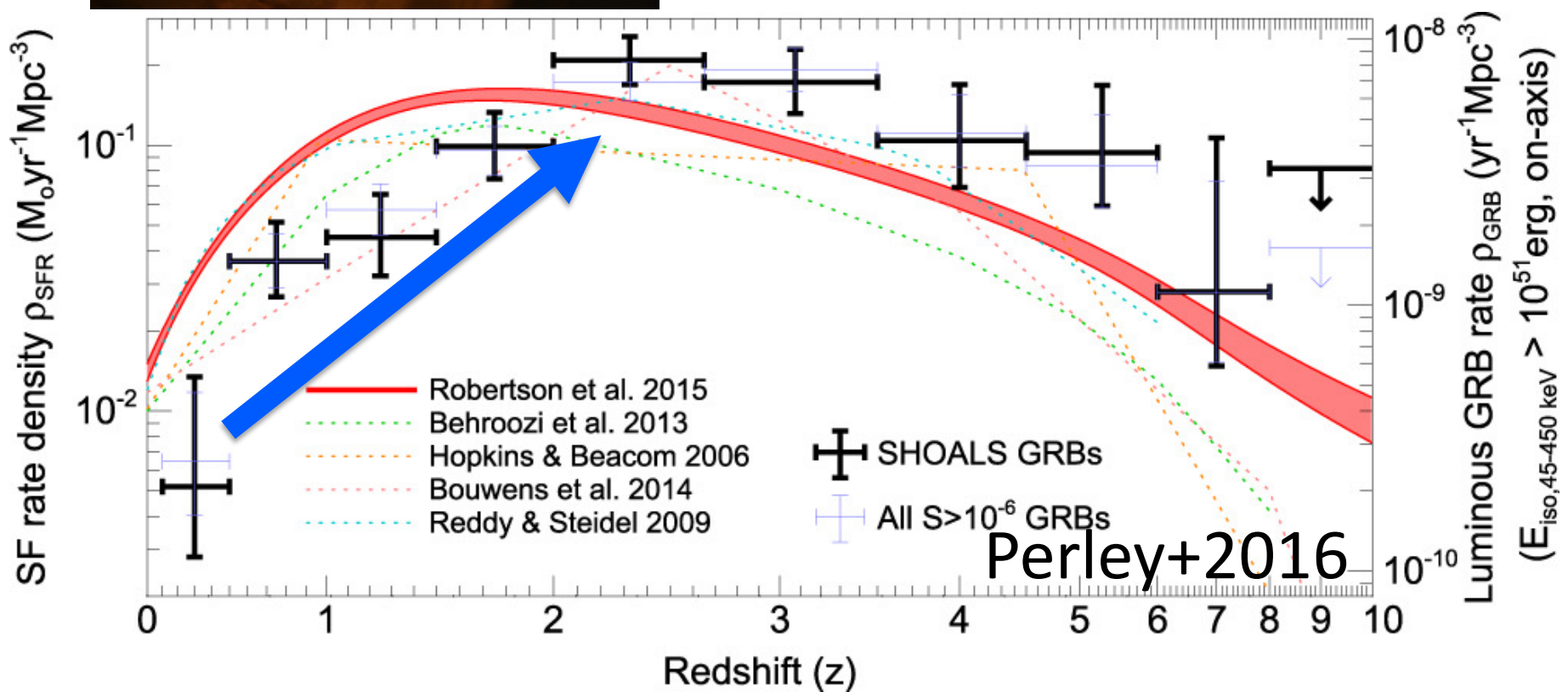


Introduction: Example (LGRBs)

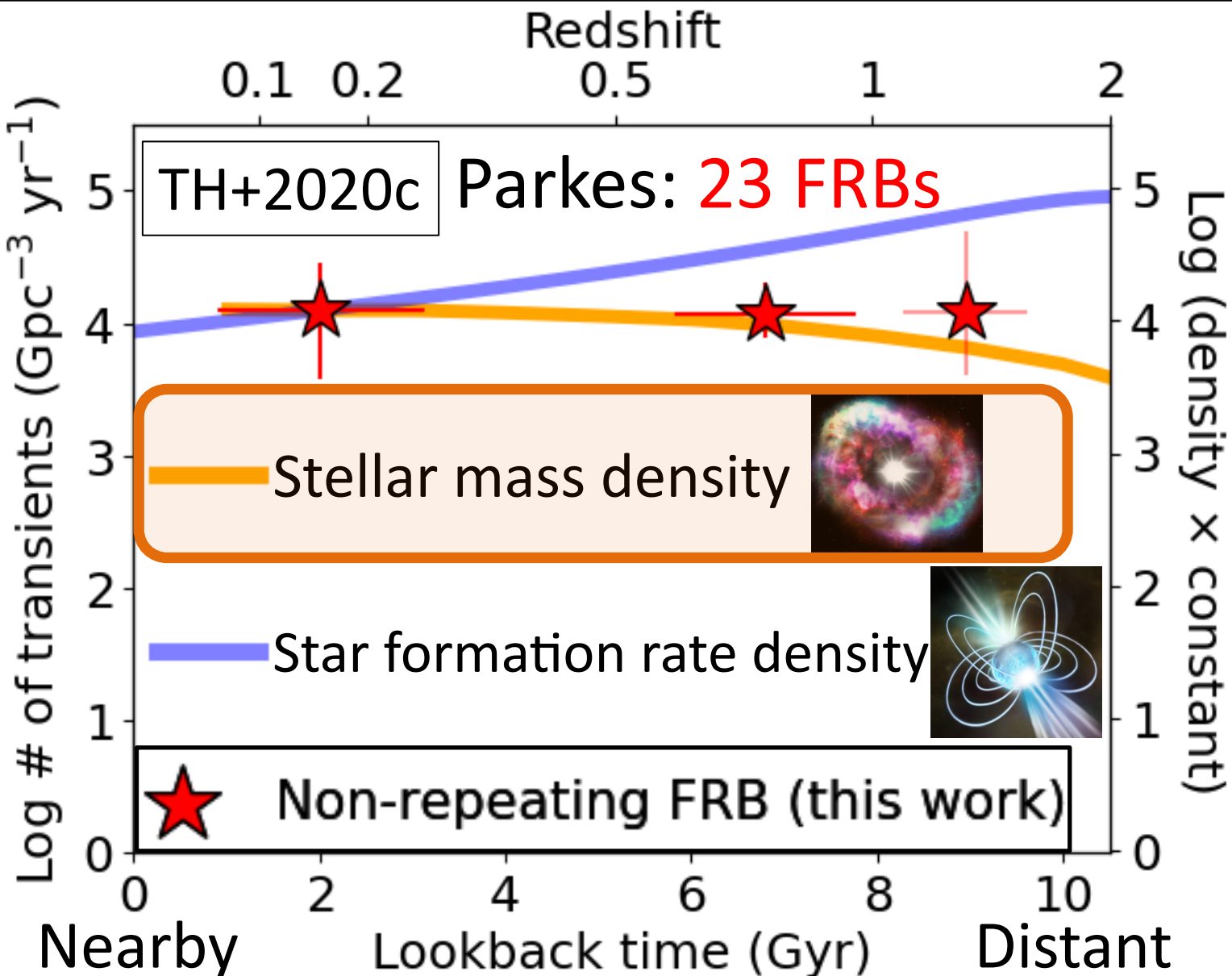


Credits: NASA, ESA and M. Kornmesser

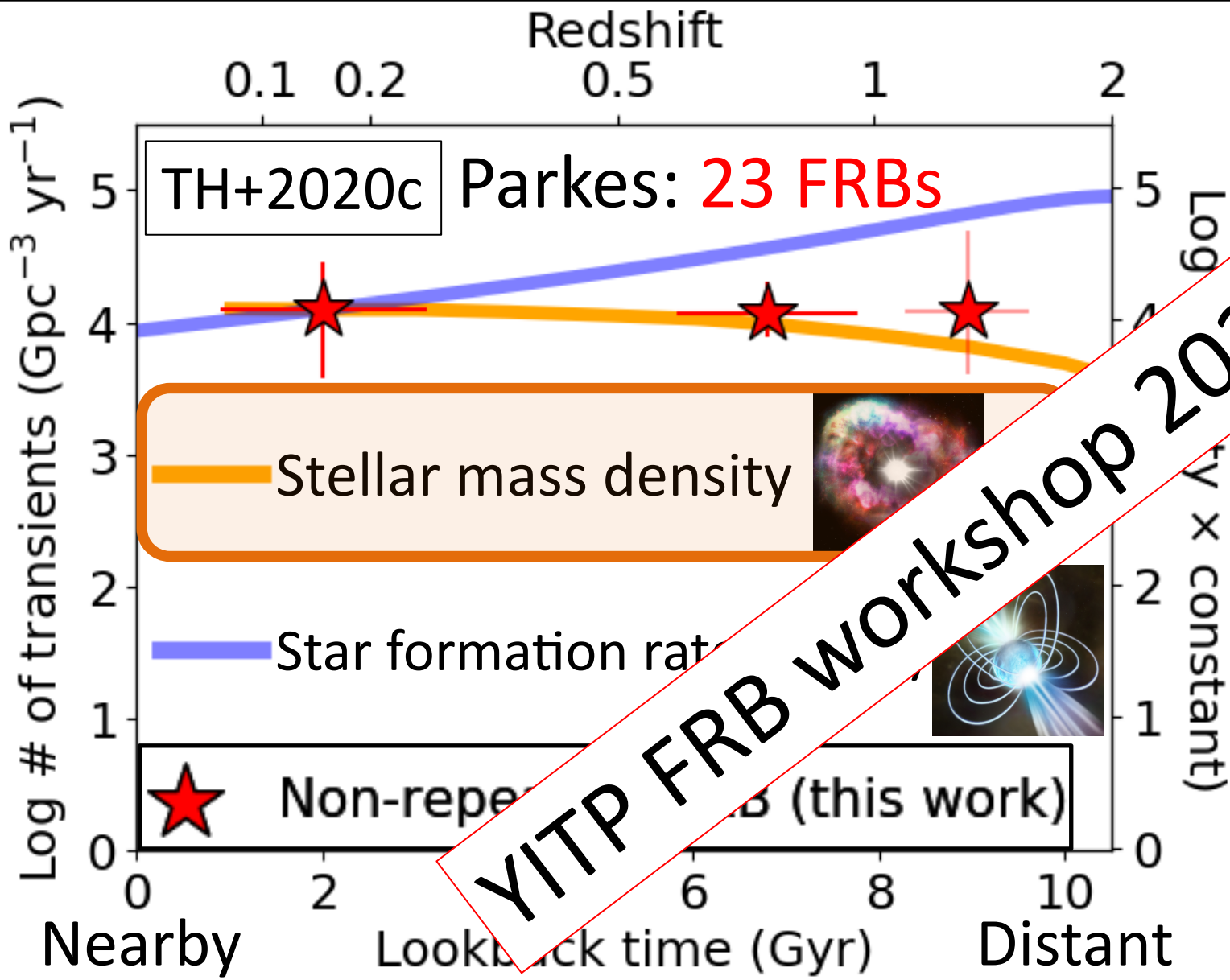
Long Gamma-ray bursts
 \propto star formation



Introduction: Non-repeating FRBs



Introduction: Non-repeating FRBs



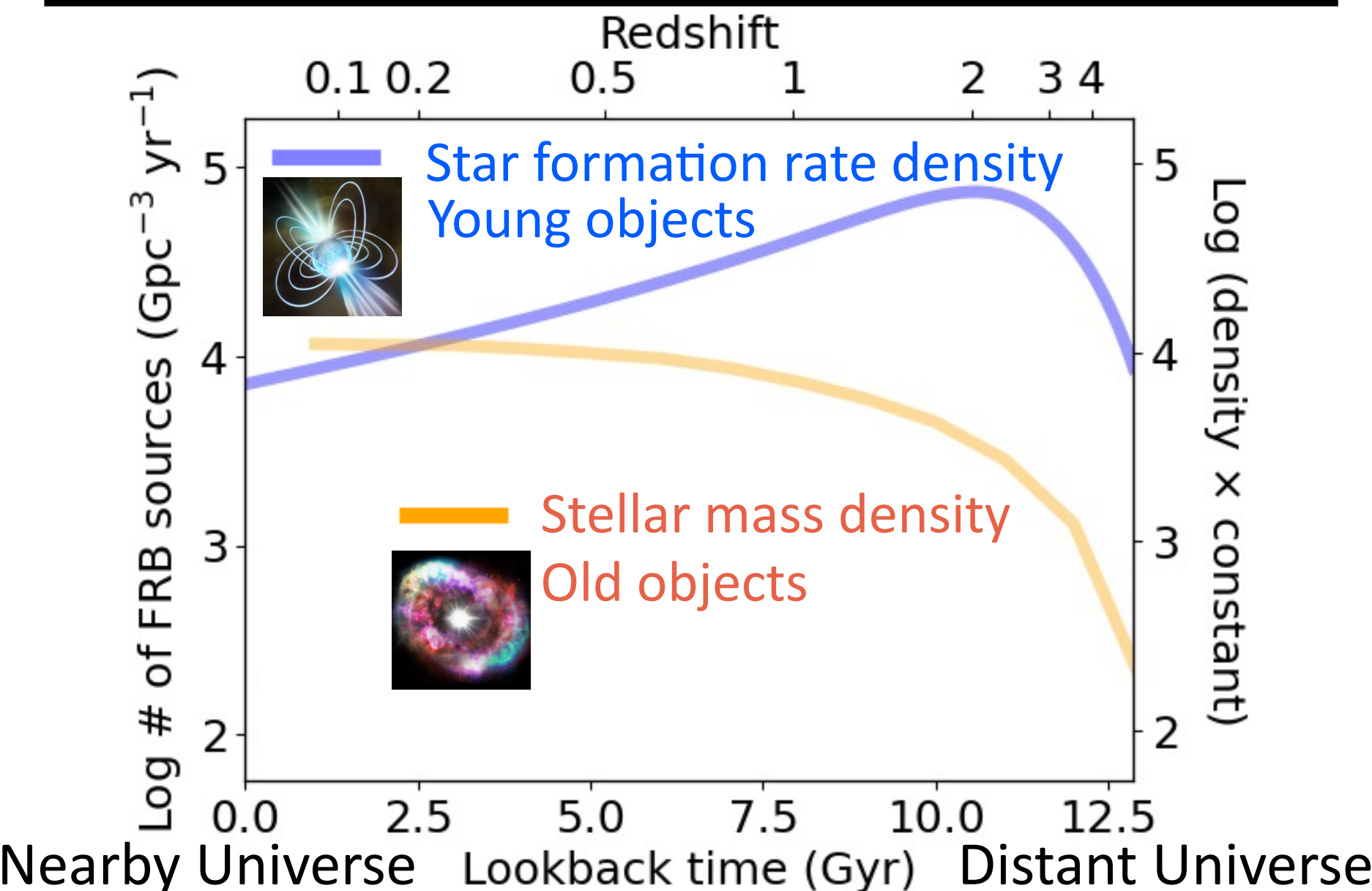
YITP FRB Workshop 2021

Let's see the answer!

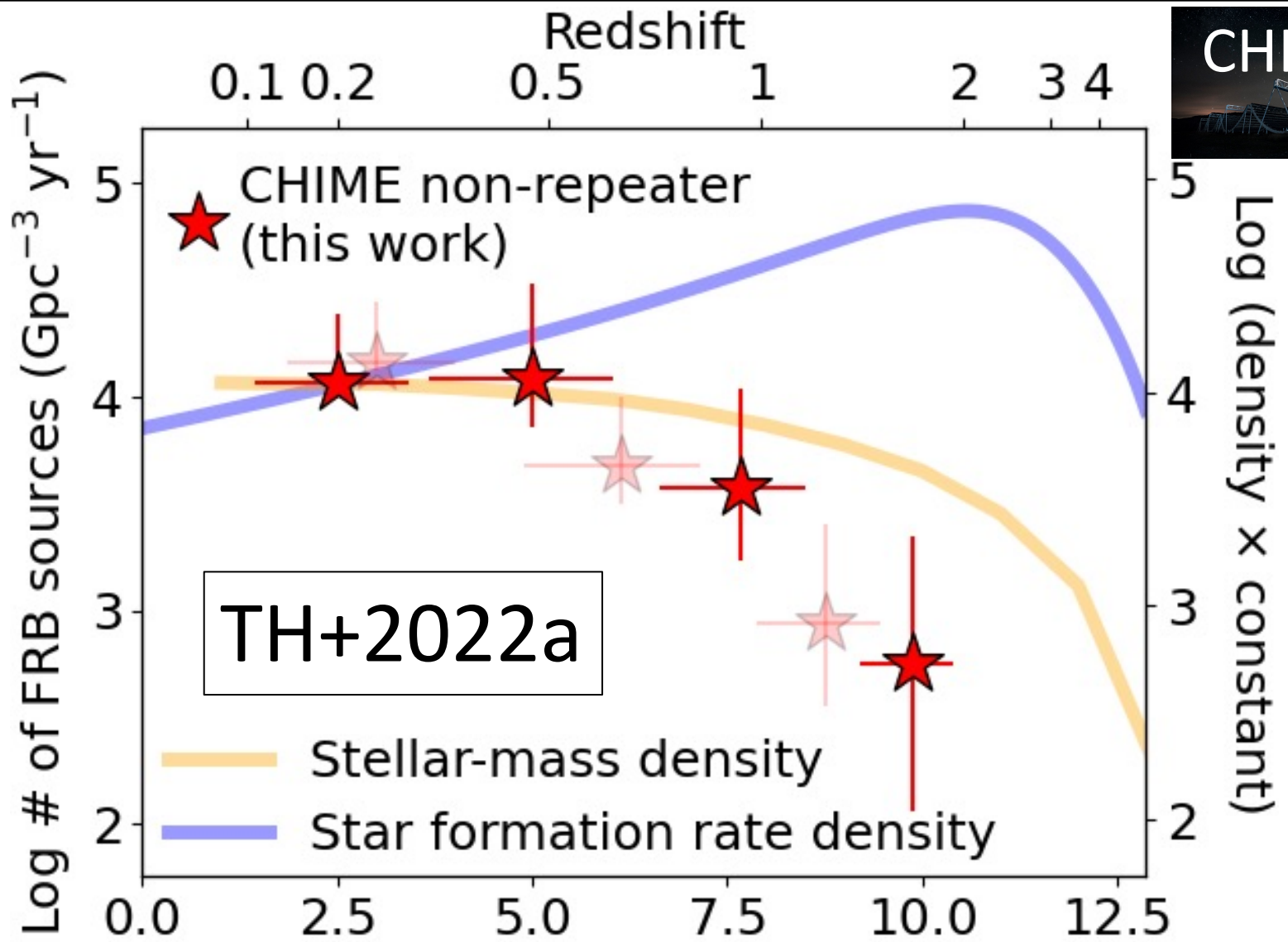


CHIME: ~500 non-repeating FRBs in 2021!

Introduction: Old vs Young



Our result: Non-repeater \rightarrow Old



Nearby Universe Lookback time (Gyr) Distant Universe

FRB population analyses

Author	free from the z-evolution assumption?	test an old population scenario?	sample	homogeneous sample?	conclusion
TH+2022a	Yes	Yes	CHIME	Yes	Old pop.
Zhang&Zhang 2022	No	Yes	CHIME	Yes	Old pop.
James+2022	No	No	ASKAP/ Parkes	No	Young pop.
Arcus+2021	No	Yes	ASKAP/ Parkes	No	Both young and no-evo pops.
Zhang+2021	No	Yes	ASKAP/ Parkes	No	Both pops.
TH+2020c	Yes	Yes	Parkes	Yes	Old. pop.



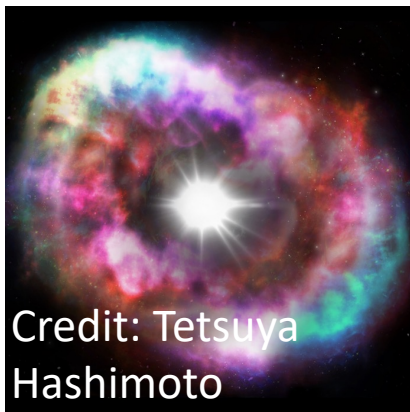
Conclusion

TH+2022a

x10 more samples than before

Non-repeater → Old objects

White dwarf



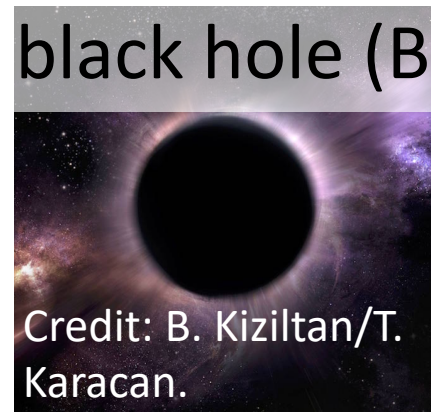
Credit: Tetsuya Hashimoto

Old neutron star



Credit: Mark Garlick

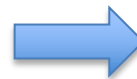
Old stellar-mass black hole (BH)



Credit: B. Kiziltan/T. Karacan.

Cautions

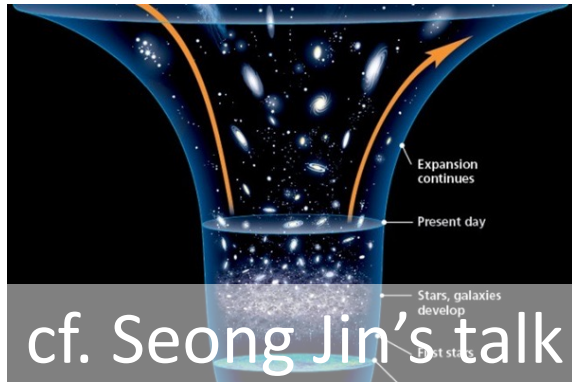
- high scattering
- faint bursts



missed by CHIME

Key sciences to be addressed by FRBs

Dark energy



Missing baryon problem



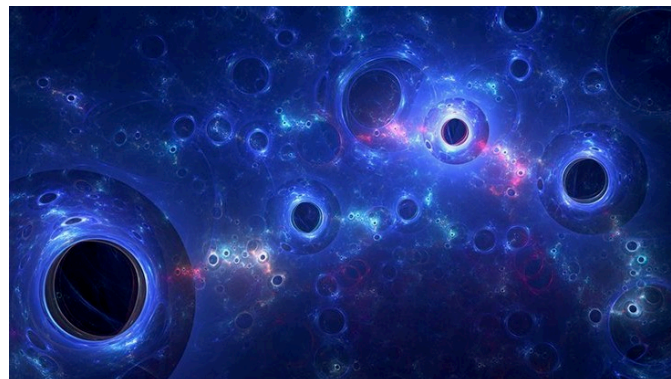
General relativity



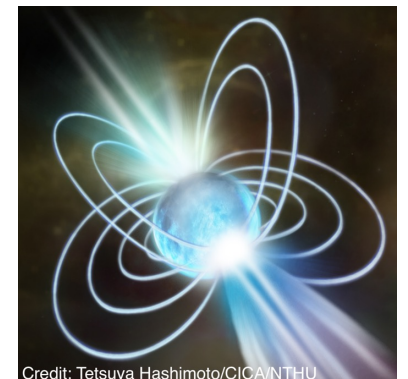
Cosmic reionization



Dark matter

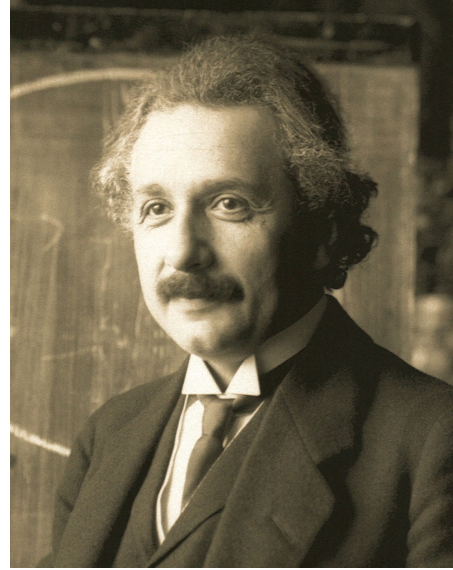


The origin of FRBs



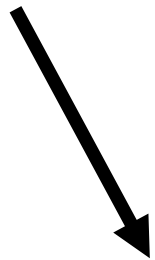
Newtonian mechanics

General relativity (GR)

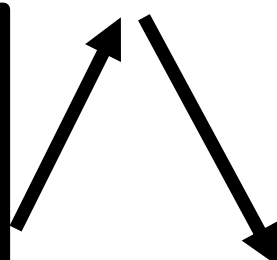


Isaac Newton

Albert Einstein



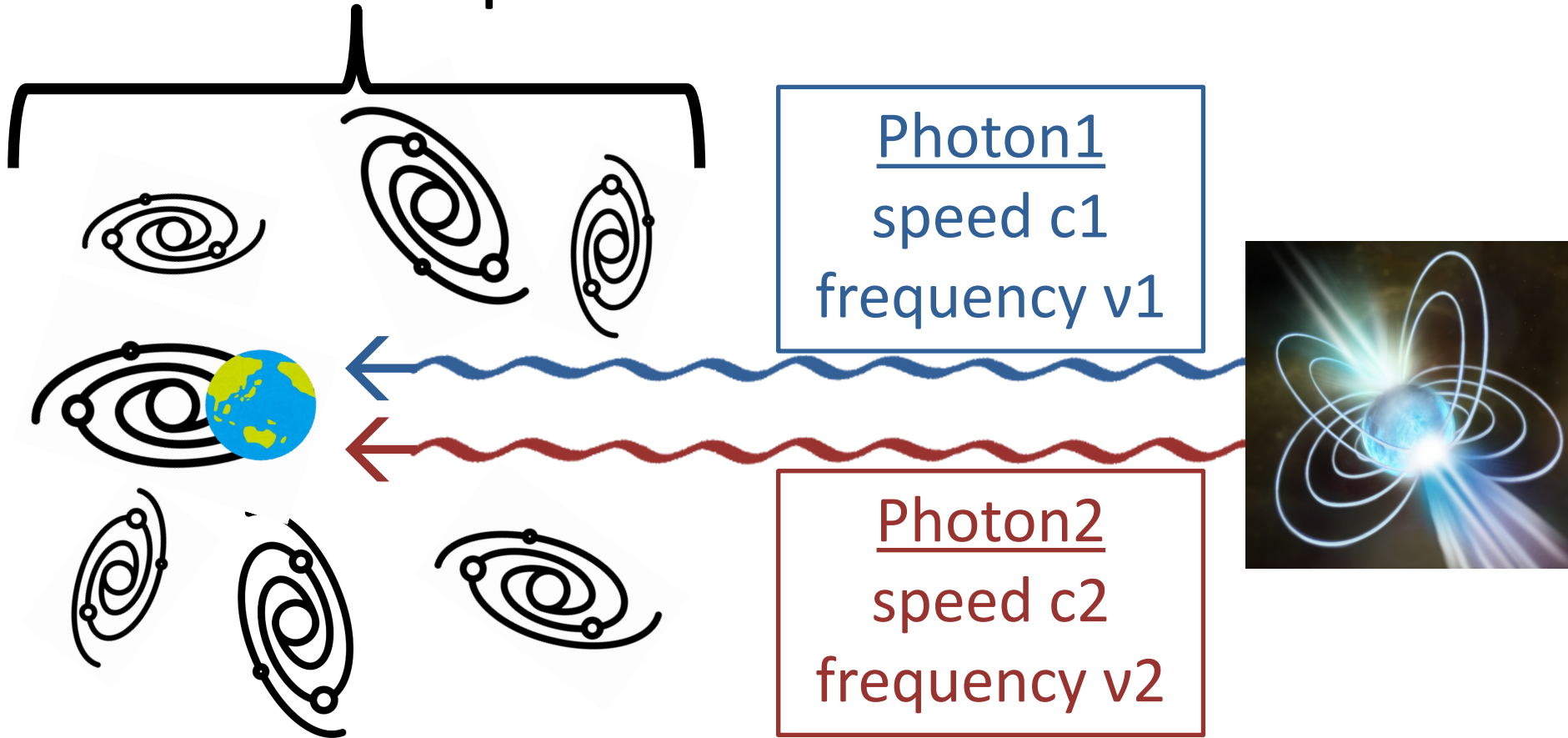
Michelson-Morley experiment
Speed of light = constant



Testing the speed of light with fast radio bursts (FRBs)

Einstein's weak equivalence principle (WEP)

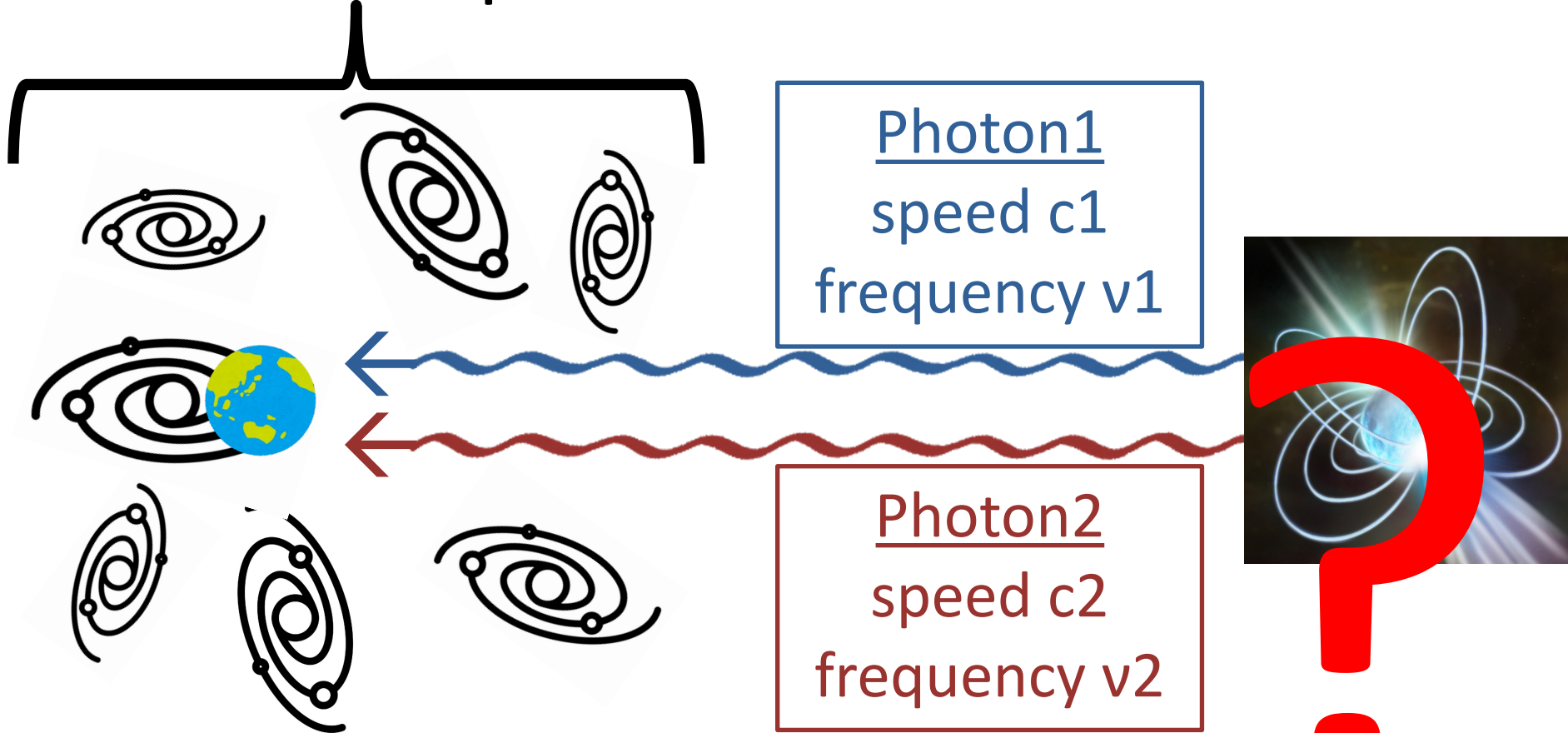
Gravitational potential



General relativity assumes $c_1 = c_2$

Einstein's weak equivalence principle (WEP)

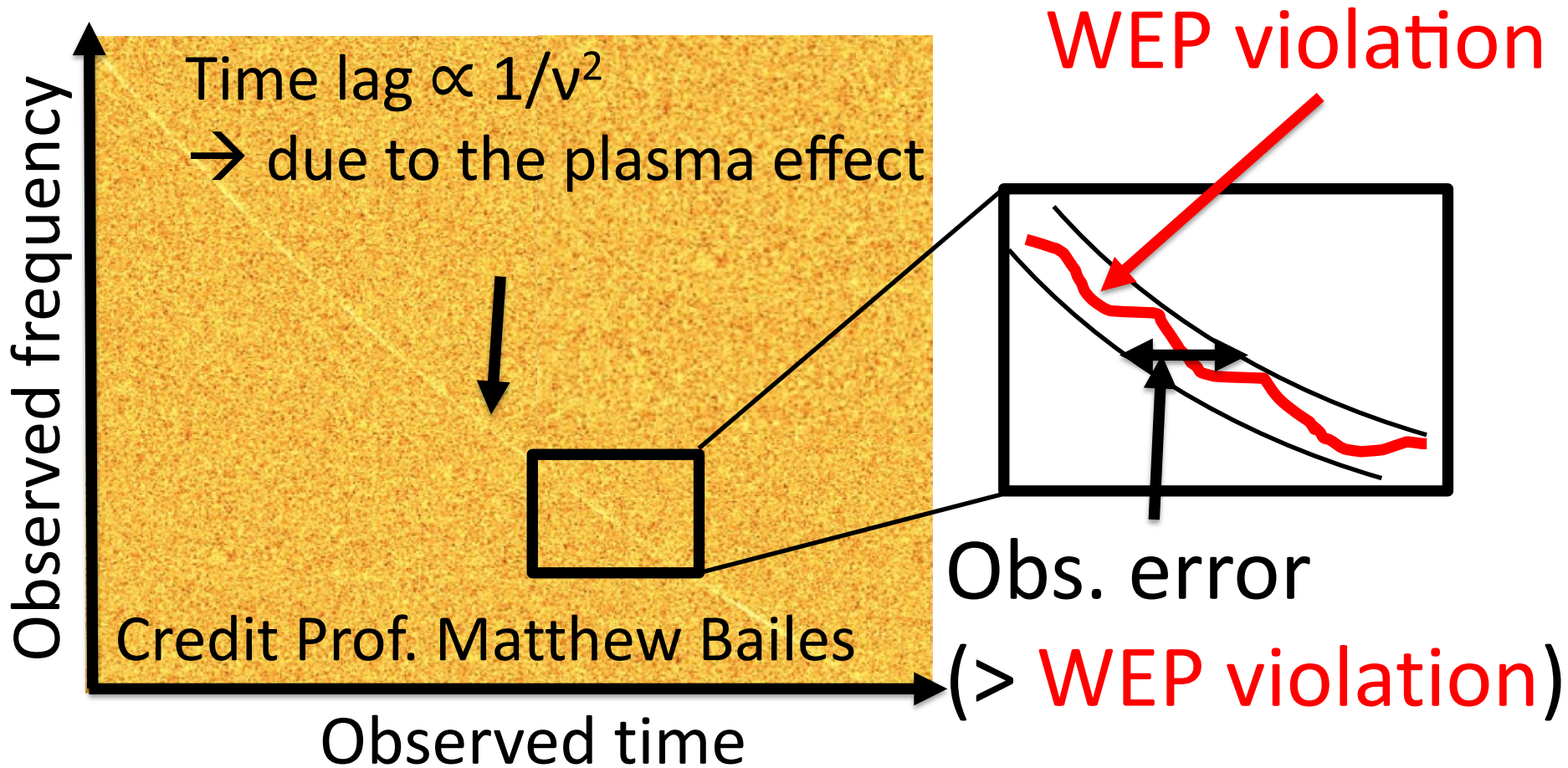
Gravitational potential



General relativity assumes $c1 = c2$

Introduction: FRB is the ideal laboratory

because they travel the cosmological distances, and we measure millisecond timescales eventually.



Introduction: FRB is the ideal laboratory

e.g. Minazzoli+2019

$$\Delta\gamma := \gamma_2 - \gamma_1 < \left(\underbrace{\Delta t_{\text{obs}}}_{\text{Observed time lag between } \nu_1 \text{ and } \nu_2} - \underbrace{\Delta t_{\text{DM}}}_{\text{Time lag due to the dispersion}} \right) \frac{6c^3}{\Omega_m H_0^2 \underbrace{d_S^3}_{\text{Distance to the source}}} \quad (7)$$

Observed time lag
between ν_1 and ν_2

Time lag due to the
dispersion

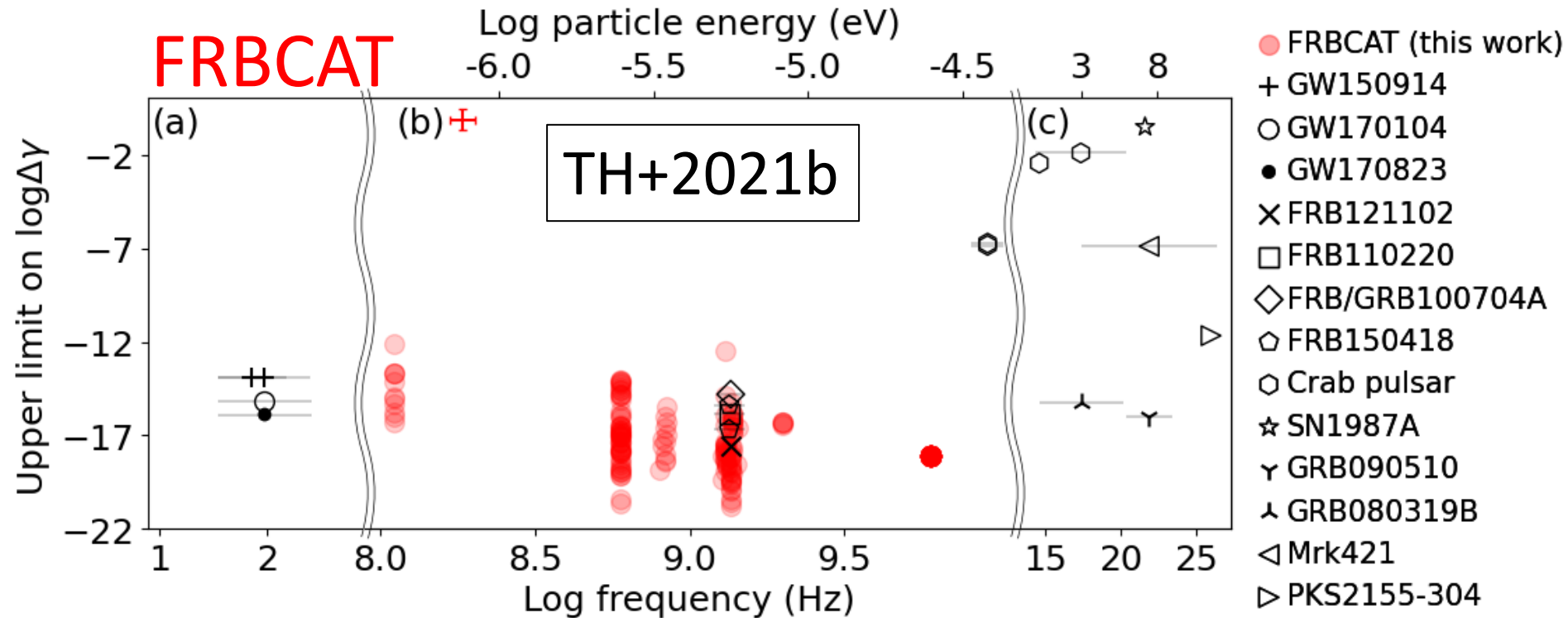
Distance to the
source

$$\Delta t_{\text{obs}} - \Delta t_{\text{DM}} < \underbrace{\delta \Delta t_{\text{DM}}}_{\text{Observational uncertainty of the arrival time due to the uncertainty of dispersion measure}} \quad (10) \quad \text{New}$$

Observational uncertainty of the arrival time
due to the uncertainty of dispersion measure

Results: the most accurate test than ever!

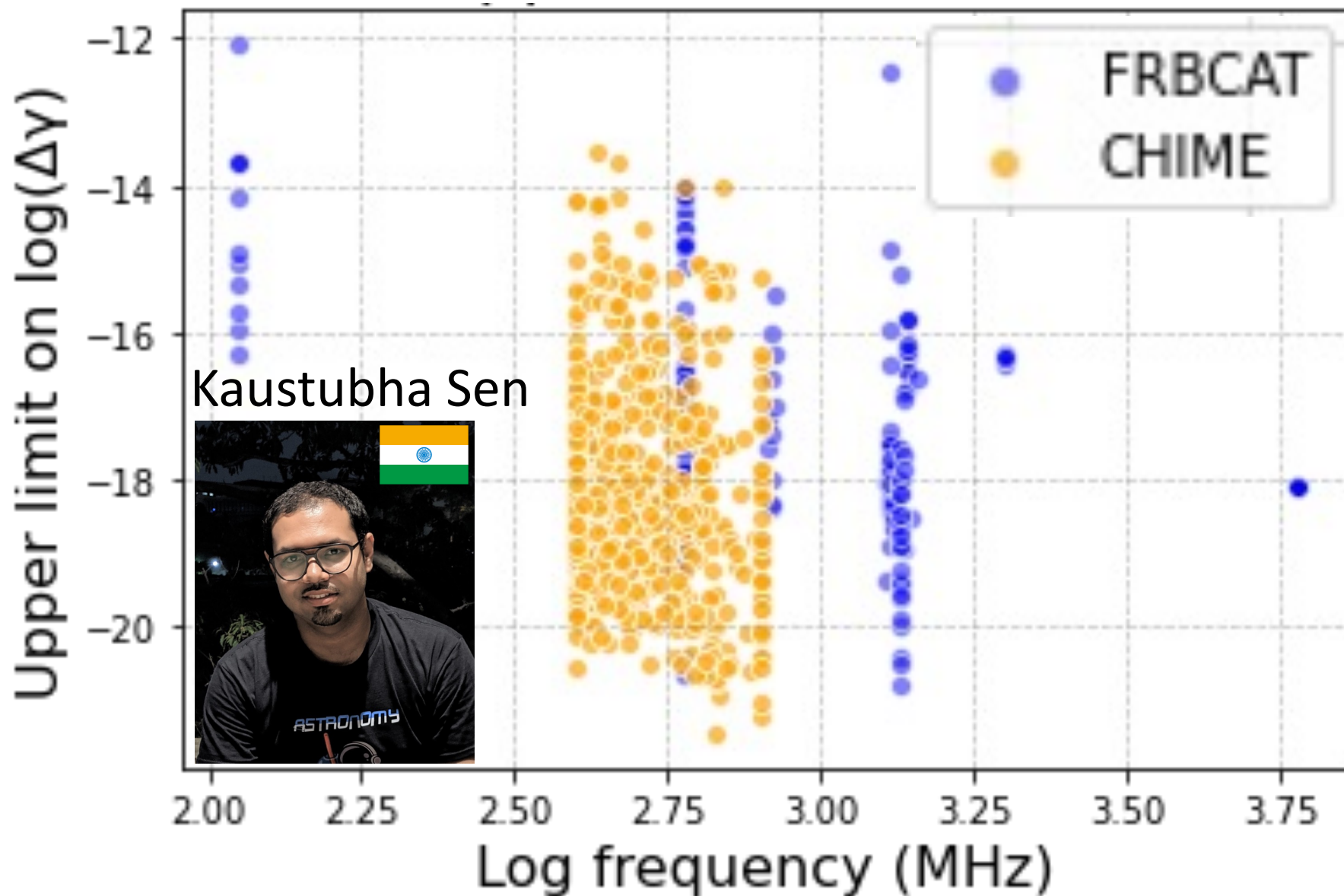
$\Delta\gamma = \gamma_1 - \gamma_2$ for photons with ν_1 and ν_2 , respectively
 $\Delta\gamma = \gamma_1 - \gamma_2 = 1 - 1 = 0$ if WEP is not violated



More than **three order of magnitude better**
than previous works

Sen, TH et al. 2022

WEP with CHIME FRBs



Sen et al

Testing Einstein's Work Using Observations of Fast Radio Bursts

Two recent papers use data from radio telescopes as a way of testing Einstein's Theory of General Relativity



Christopher Carroll Dec 5, 2021 · 5 min read ★



Bs



3.75

Upper limit on $\log(\Delta\gamma)$

-12

-14

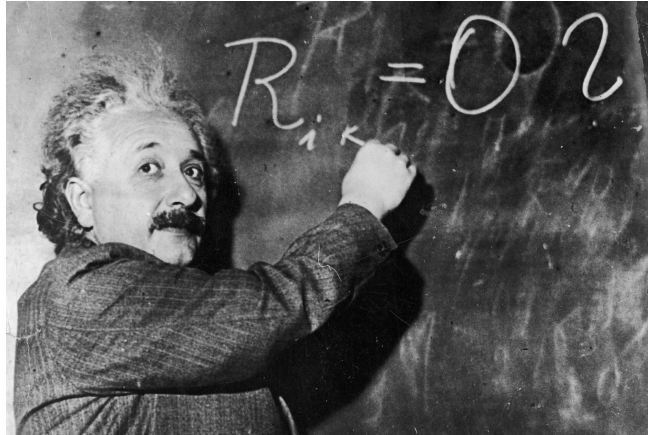
-16

-18

-20

2

Conclusion:



TH+2021b

Sen, TH et al. 2022

- **A new method** to constrain the WEP violation using **FRBs**
- **The most accurate test** than ever:
Old catalogue: three orders of magnitude better
New CHIME: even one order of magnitude better

A future FRB telescope in Taiwan

BURSTT

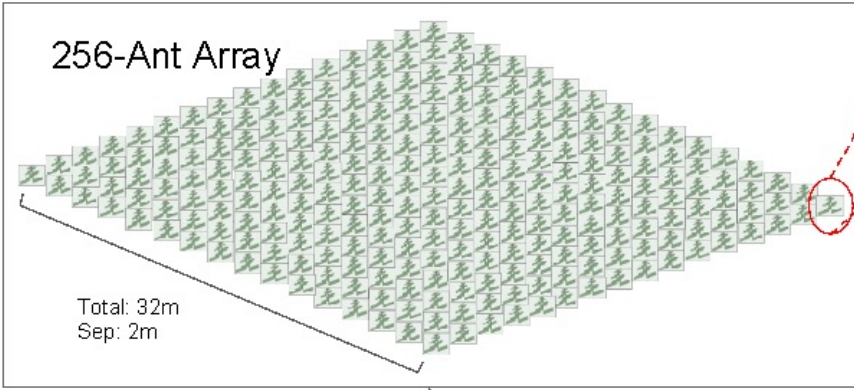
What are the bottlenecks?

- lack of localization capability
 - e.g., CHIME \sim arcmin
 - spec-z, host galaxy, progenitor etc.
- small FoV and low cadence
 - e.g., CHIME: 5-10 min per day (<1% of the day)
 - missing population of FRBs?
 - expensive for follow-up telescopes
- mismatched distance
 - GWs, neutrinos, high-time resolution follow-up
 - > nearby Universe

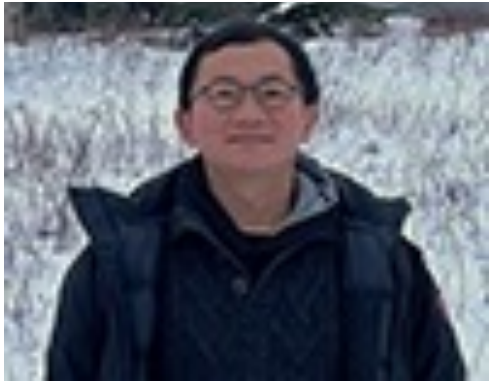
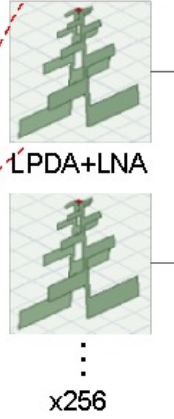
cf. Sota's and Yuu's talks

Future FRB telescope in Taiwan

Bustling Universe Radio Survey Telescope in Taiwan (BURSTT)



Signal chain

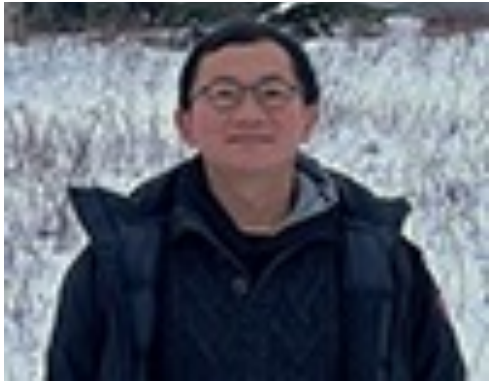
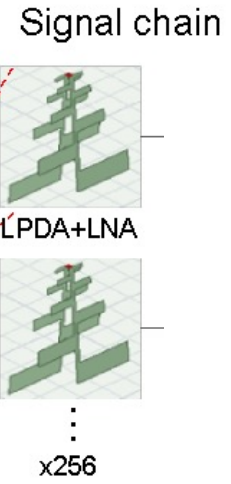
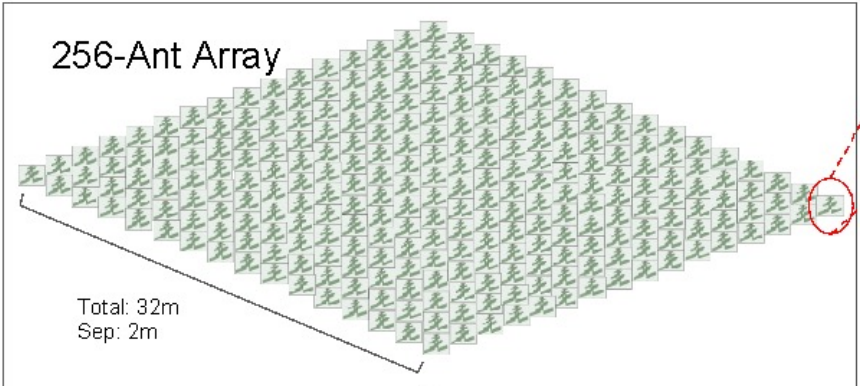


led by
Prof. Pen, Ue-Li



Future FRB telescope in Taiwan

Bustling Universe Radio Survey Telescope in Taiwan (BURSTT)



led by
Prof. Pen, Ue-Li



The world-best telescope to detect nearby FRBs

BURSTT will resolve the bottlenecks

- lack of localization capability
BURSTT: sub-arcsecond
e.g., CHIME \sim arcmin
- small FoV and low cadence
BURSTT: 25 times better than that of CHIME
e.g., CHIME: 5-10 min per day (<1% of the day)
- mismatch in distance
BURSTT: dedicated to the nearby Universe
GWs, neutrinos, high-time resolution follow-up
--> nearby Universe cf. Sota's and Yuu's talks

Predictions of BURSTT FRBs

Bustling Universe Radio Survey Telescope
in Taiwan (BURSTT)

Number prediction of
gravitationally lensed FRBs



Simon C.-C. Ho

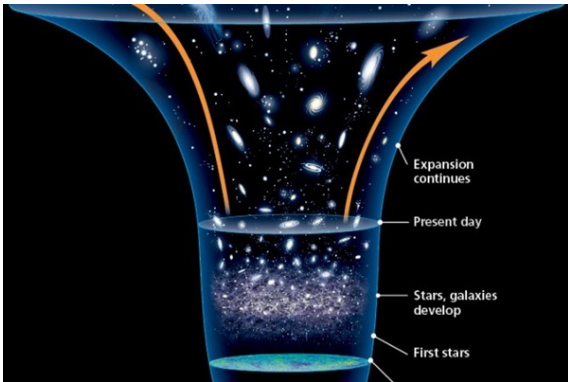
Number prediction of
Galactic FRB-like events



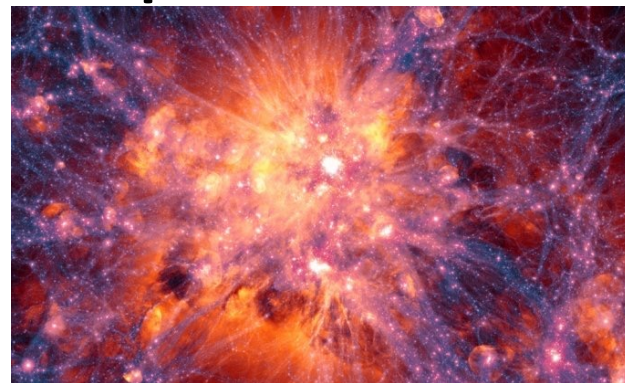
Decmend Lin

Conclusion: FRB science is exciting

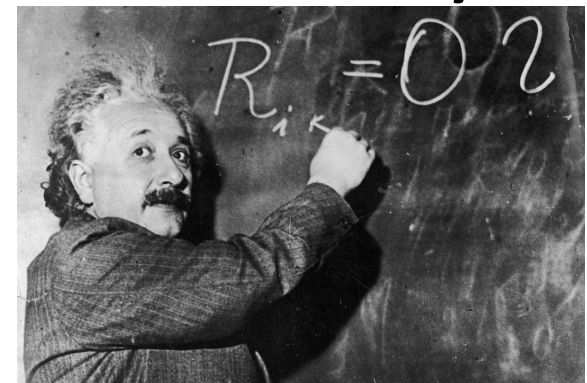
Dark energy



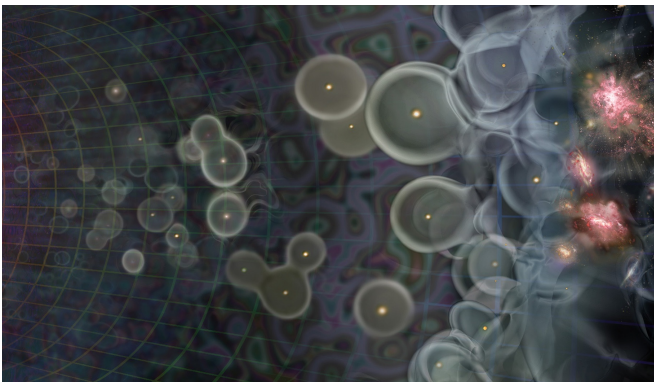
Missing baryon problem



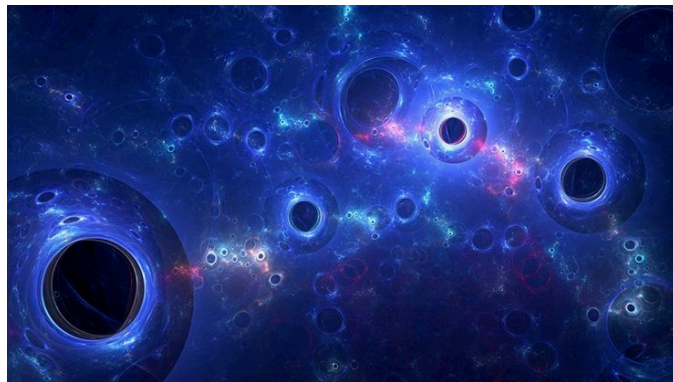
General relativity



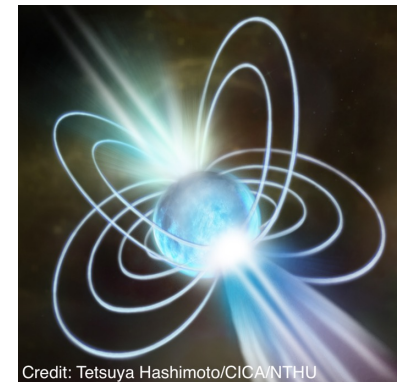
Cosmic reionization



Dark matter



The origin of FRBs



Credit: Tetsuya Hashimoto/CICANTHU

Backup slides

Redshifts derived from dispersion measures

