




“Gravity and Cosmology 2024” 19th Feb. 2024

Circular polarization of gravitational waves passing through an axion domain wall

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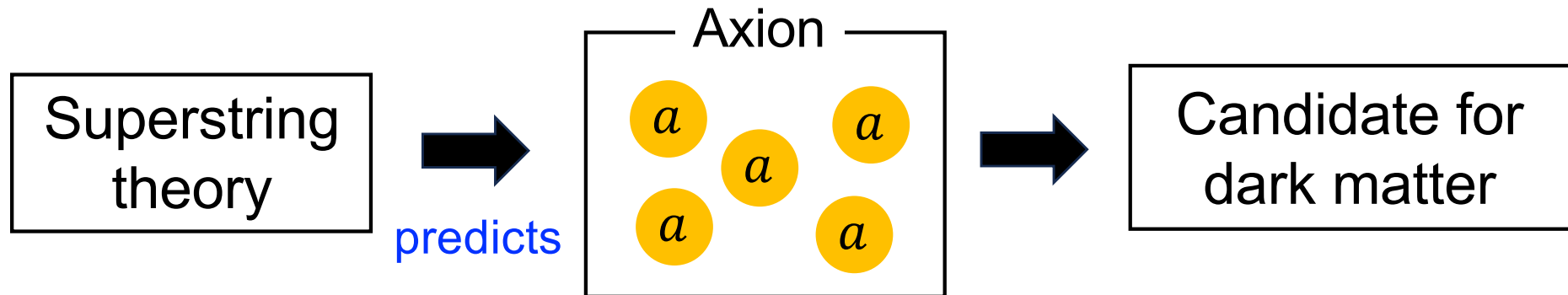
Based on PRD 108, 083525 (2023)



Axion and Chern-Simons gravity

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Why is axion important?



It leads to verifying superstring theory and dark matter.

Chern-Simons(CS) gravity coupling constant

$$S = \frac{M_{\text{p}}^2}{2} \int d^4x \sqrt{-g} R + \frac{M_{\text{p}} \ell^2}{8} \int d^4x \sqrt{-g} \phi R\tilde{R} \quad \left[R\tilde{R} = \frac{1}{2} \varepsilon^{\rho\sigma\alpha\beta} R^{\mu\nu}{}_{\alpha\beta} R_{\nu\mu\rho\sigma} \right]$$

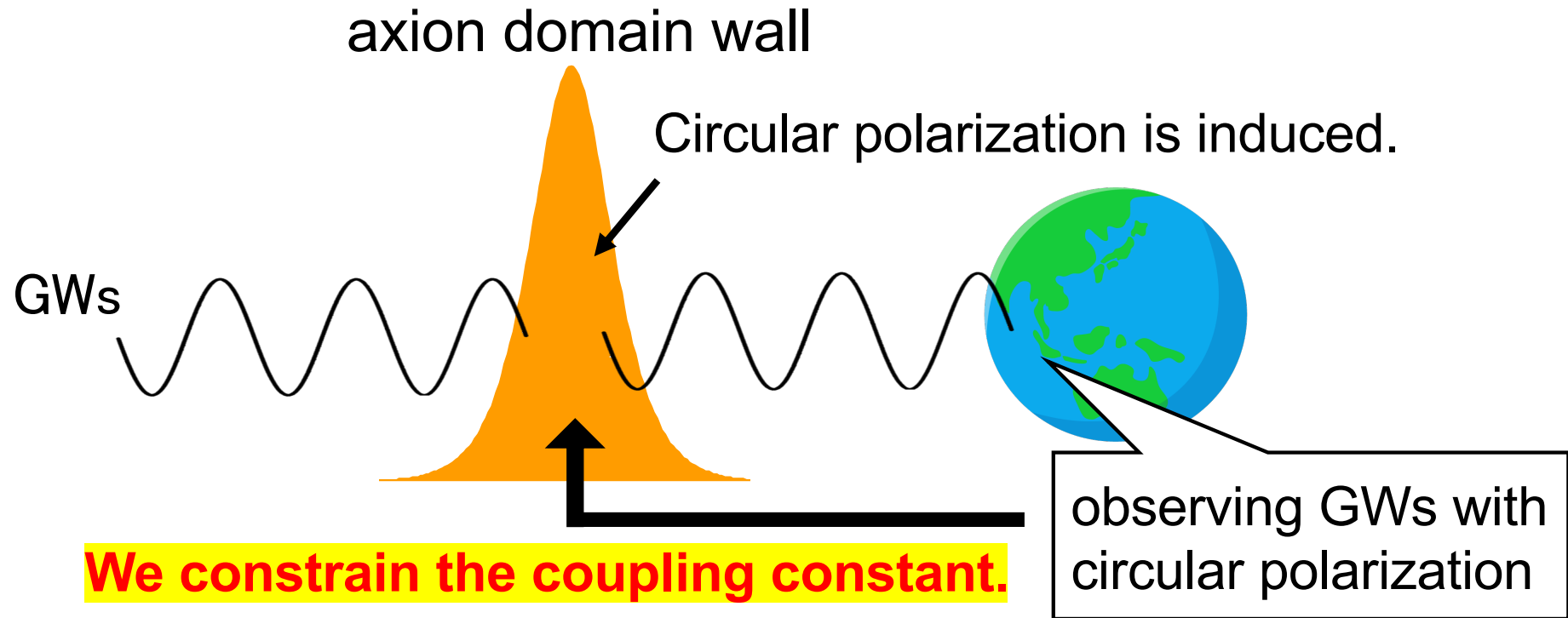
axion

CS term

CS term induces the circular polarization in GWs.

Purpose of this study

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What we investigated in this work

- Is the circular polarization large enough to be observable?
- How do we give the constraint on the coupling constant?

Axion domain wall

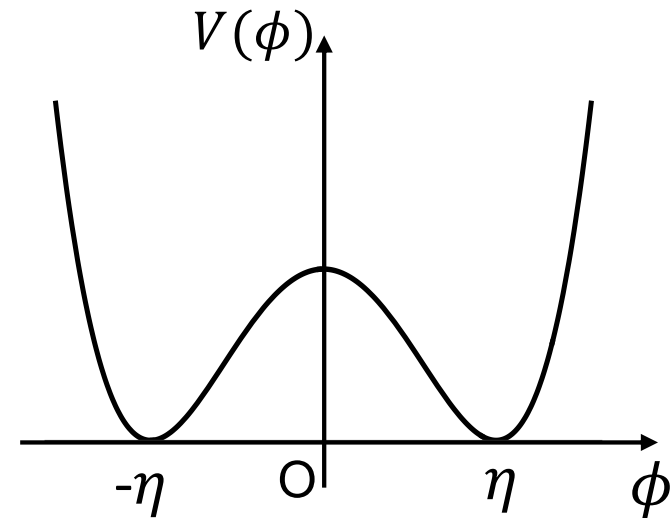
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action

$$S = - \int d^4x \sqrt{-g} \left[\frac{1}{2} \partial_\mu \phi \partial^\mu \phi + V(\phi) \right]$$

$$V(\phi) = \frac{\lambda}{4} (\phi^2 - \eta^2)^2$$

— self-coupling constant

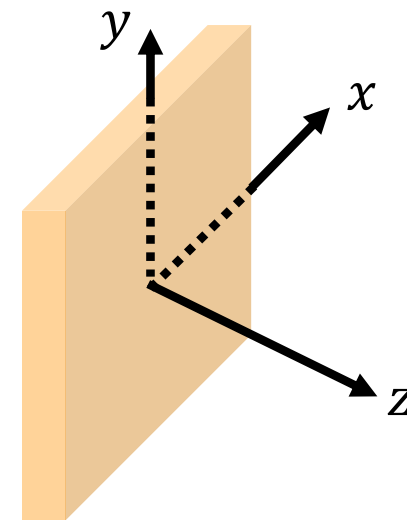


Assuming the domain wall is **static** and in **(x, y)-plane** gives

$$\frac{d^2 \phi(z)}{dz^2} = \lambda (\phi(z)^2 - \eta^2) \phi(z)$$

The boundary condition: $\phi(\pm\infty) = \pm\eta$

The solution: $\phi(z) = \eta \tanh \left(\sqrt{\frac{\lambda}{2}} \eta z \right)$



Axion domain wall

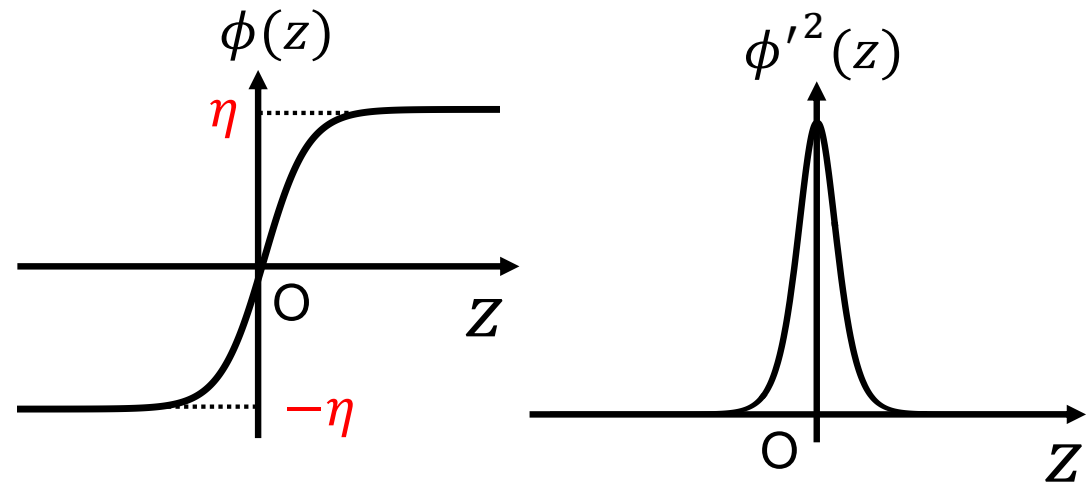
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The domain wall solution

$$\phi(z) = \eta \tanh \left(\sqrt{\frac{\lambda}{2}} \eta z \right)$$

$$\phi'(z) = \eta^2 \sqrt{\frac{\lambda}{2}} \operatorname{sech}^2 \left(\sqrt{\frac{\lambda}{2}} \eta z \right)$$

[prime: z - derivative]



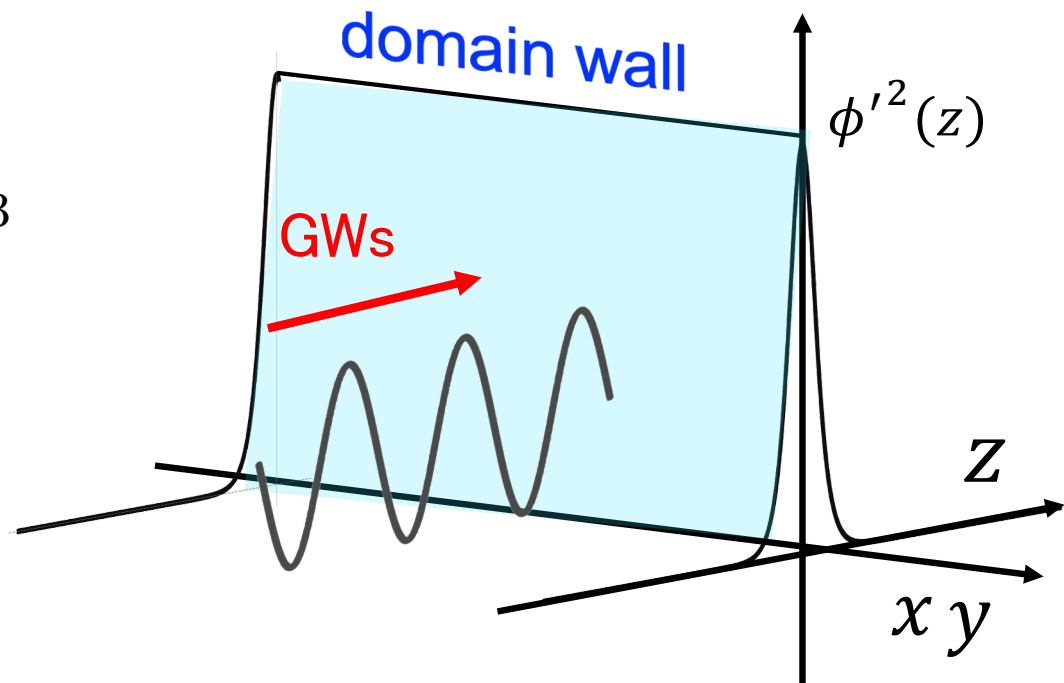
energy density

$$\sigma = \int dz \phi'^2(z) \propto \sqrt{\lambda} \eta^3$$

Constraint by CMB

$$\sigma < (0.93 \text{ MeV})^3$$

A. Lazanu (2015)



Equation of motion for CS-GWs

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metric

GWs

$$ds^2 = -dt^2 + (\delta_{ij} + h_{ij})dx^i dx^j$$

TT gauge $h^i_{j,i} = h^i_i = 0$

action

Coupling constant

$$S = \frac{M_p^2}{2} \int d^4x \sqrt{-g} R + \frac{M_p \ell^2}{8} \int d^4x \sqrt{-g} \phi R \tilde{R}$$

$$R \tilde{R} = \frac{1}{2} \epsilon^{\rho\sigma\alpha\beta} R^{\mu\nu}{}_{\alpha\beta} R_{\nu\mu\rho\sigma}$$

$$\epsilon^{0123} = \frac{1}{\sqrt{-g}}$$

CS term

$$= \frac{M_p^2}{8} \int d^4x \left[\dot{h}^{ij} \dot{h}_{ij} - h^{ij,k} h_{ij,k} \right.$$

$$\left. - \frac{\ell^2}{M_p} \epsilon^{izk} \partial_z \phi \left\{ \ddot{h}^m_i \dot{h}_{km} + \dot{h}_{im,\ell} (h^{\ell}_k{}^{,m} - h^m_k{}^{,\ell}) \right\} \right]$$


EOM

$$\square h^{ij} = \frac{\ell^2}{2M_p} \epsilon^{zik} [\phi' \square + \phi'' \partial_z] \dot{h}^j_k + \frac{\ell^2}{2M_p} \epsilon^{zjk} [\phi' \square + \phi'' \partial_z] \dot{h}^i_k$$

Equation of motion for CS-GWs

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$$\square h^{ij} = \frac{\ell^2}{2M_p} \varepsilon^{zik} [\phi' \square + \phi'' \partial_z] \dot{h}^j_k + \frac{\ell^2}{2M_p} \varepsilon^{zjk} [\phi' \square + \phi'' \partial_z] \dot{h}^i_k$$



$$h_{ij}(t, z) = h_R(t, z) e_{ij}^{(R)} + h_L(t, z) e_{ij}^{(L)}, \quad e_{ij}^{(R)} = \begin{pmatrix} 1 & i & 0 \\ i & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \quad e_{ij}^{(L)} = \begin{pmatrix} 1 & -i & 0 \\ -i & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

polarization tensor

$$\square h_{R/L}(t, z) = \pm i \frac{\ell^2}{M_p} [\phi' \square + \phi'' \partial_z] \dot{h}_{R/L}(t, z)$$



Using Fourier modes

$$h_{R/L}(t, z) = H_{R/L}(z) e^{i\omega t} \quad \omega : \text{frequency of GWs}$$

$$\left(1 \pm \frac{\omega \ell^2}{M_p} \phi'\right) H''_{R/L} \pm \frac{\omega \ell^2}{M_p} \phi'' H'_{R/L} + \omega^2 \left(1 \pm \frac{\omega \ell^2}{M_p} \phi'\right) H_{R/L} = 0$$

Equation of motion for CS-GWs

$$\square h^{ij} = \frac{\ell^2}{2M_p} \varepsilon^{zik} [\phi' \square + \phi'' \partial_z] \dot{h}^j_k + \frac{\ell^2}{2M_p} \varepsilon^{zjk} [\phi' \square + \phi'' \partial_z] \dot{h}^i_k$$

$$h_{ij}(t, z) = h_R(t, z) e_{ij}^{(R)} + h_L(t, z) e_{ij}^{(L)}, \quad e_{ij}^{(R)} = \begin{pmatrix} 1 & i & 0 \\ i & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \quad e_{ij}^{(L)} = \begin{pmatrix} 1 & -i & 0 \\ -i & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

Right-handed GWs H_R and left-handed GWs H_L follow different equations of motion!

→ Circular polarization

$$h_{R/L}(t, z) = \dots(z) e^{i\omega t} \quad \omega : \text{frequency of GWs}$$

$$\left(1 \pm \frac{\omega \ell^2}{M_p} \phi'\right) H''_{R/L} \pm \frac{\omega \ell^2}{M_p} \phi'' H'_{R/L} + \omega^2 \left(1 \pm \frac{\omega \ell^2}{M_p} \phi'\right) H_{R/L} = 0$$

How to solve the EOM

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EOM

$$\left(1 \pm \frac{\omega l^2}{M_p} \phi'\right) H''_{R/L} \pm \frac{\omega l^2}{M_p} \phi'' H'_{R/L} + \omega^2 \left(1 \pm \frac{\omega l^2}{M_p} \phi'\right) H_{R/L} = 0$$



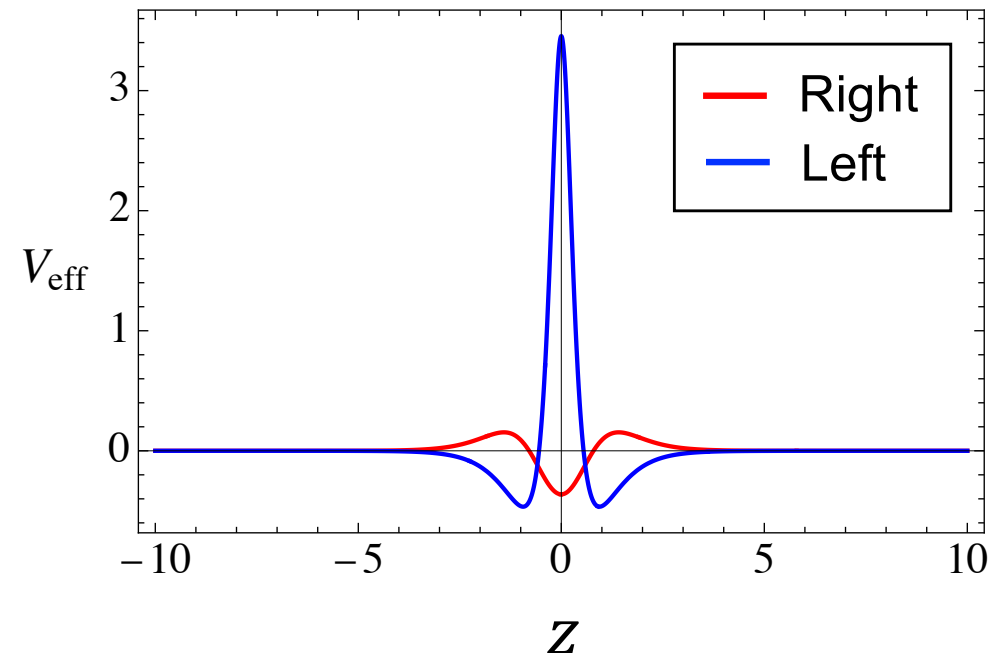
$$\tilde{H}_{R/L} = \sqrt{F(z)} H_{R/L} \quad F(z) = 1 \pm \frac{\omega l^2}{M_p} \phi'(z)$$

Effective potential

Schrödinger type Eq.:

$$\left[-\frac{d^2}{dz^2} + V_{\text{eff}}(z)\right] \tilde{H}_{R/L} = \omega^2 \tilde{H}_{R/L}$$

$$V_{\text{eff}} = -\frac{1}{4} \left(\frac{F'}{F}\right)^2 + \frac{1}{2} \frac{F''}{F}$$



How to solve the EOM

EOM

$$\left[-\frac{d^2}{dz^2} + V_{\text{eff}}(z) \right] \tilde{H}_{R/L} = \omega^2 \tilde{H}_{R/L}$$

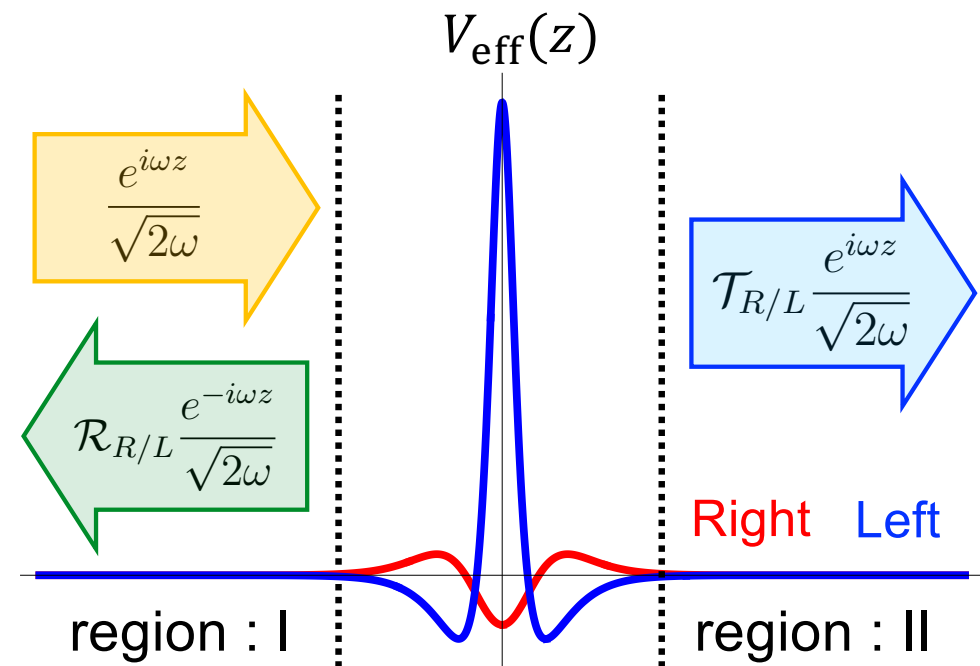
Scattering problem

region : I

$$\tilde{H}_{R/L}(z) = \frac{e^{i\omega z}}{\sqrt{2\omega}} + \mathcal{R}_{R/L} \frac{e^{-i\omega z}}{\sqrt{2\omega}}$$

region : II

$$\tilde{H}_{R/L}(z) = \mathcal{T}_{R/L} \frac{e^{i\omega z}}{\sqrt{2\omega}}$$



Degree of the circular polarization

$$\Pi = \frac{|\mathcal{T}_R|^2 - |\mathcal{T}_L|^2}{|\mathcal{T}_R|^2 + |\mathcal{T}_L|^2}$$

Just calculate $\mathcal{T}_{R/L}$ numerically as a scattering problem!

$$\left(1 \pm \frac{\omega l^2}{M_p} \phi'\right) H''_{R/L} \pm \frac{\omega l^2}{M_p} \phi'' H'_{R/L} + \omega^2 \left(1 \pm \frac{\omega l^2}{M_p} \phi'\right) H_{R/L} = 0$$

CS correction < 1

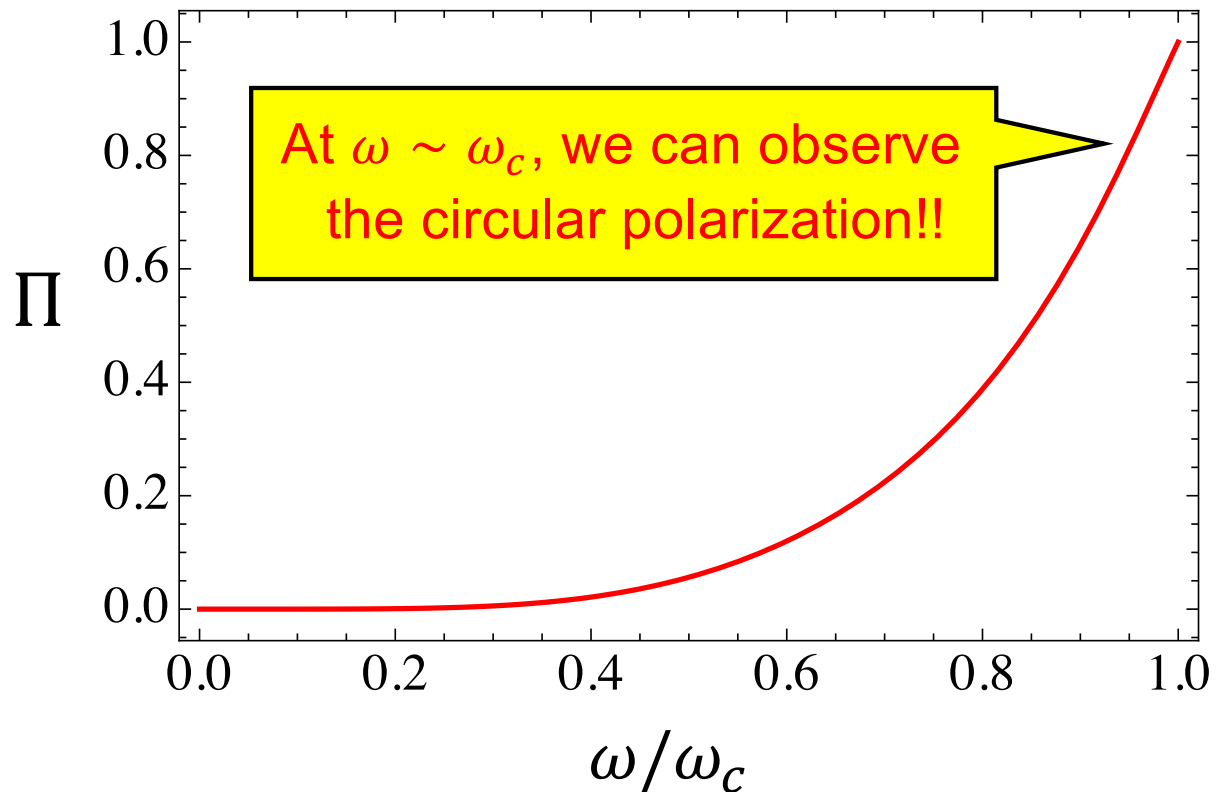
$$\frac{\omega l^2}{M_p} \phi' = \frac{\omega l^2}{M_p} \eta^2 \operatorname{sech}^2(\eta z)$$

$$\leq \frac{\omega l^2}{M_p} \eta^2 < 1$$

critical frequency

$$\omega_c \equiv \frac{M_p}{l^2 \eta^2}$$

The degree of circular polarization



Constraint on the coupling constant ¹⁰/₁₁

Constraint on ω and ℓ

the observation by Gravity Probe B

$$\frac{\omega \ell^2}{M_p} \eta^2 < 1 \quad \longrightarrow \quad \omega \ell^2 < 10^{15} \text{ (eV)}^{-1}$$

$$\ell < 10^8 \text{ km}$$

Y. Ali-Haimoud,
Y. Chen (2011)

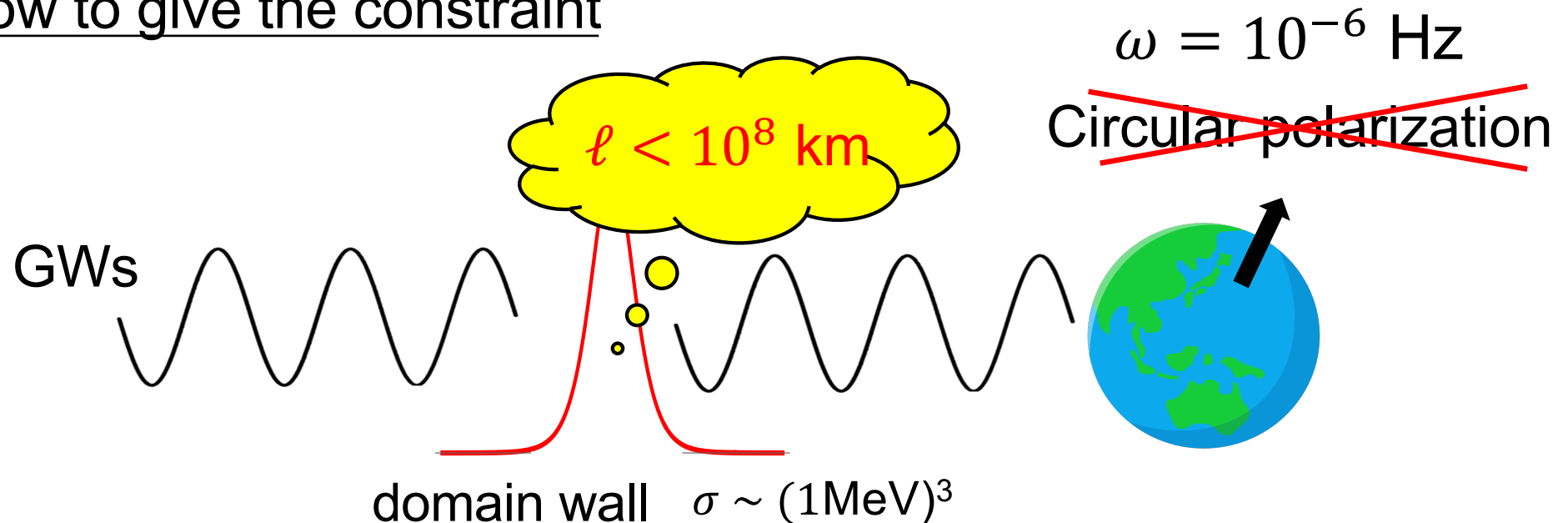
$$M_p = 10^{18} \text{ GeV}$$

$$\sigma \sim \eta^3 = (1 \text{ MeV})^3$$

$$\ell \sim 10^8 \text{ km} \quad \longrightarrow \quad \omega_c \sim 10^{-6} \text{ Hz}$$

CMB

How to give the constraint



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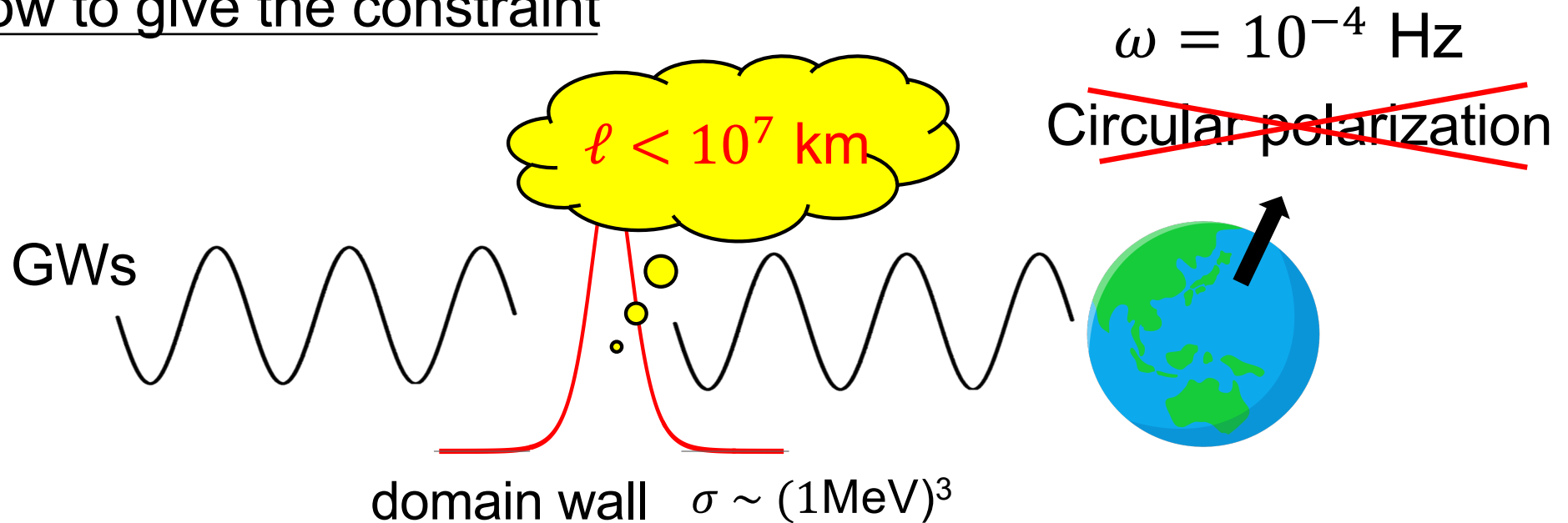
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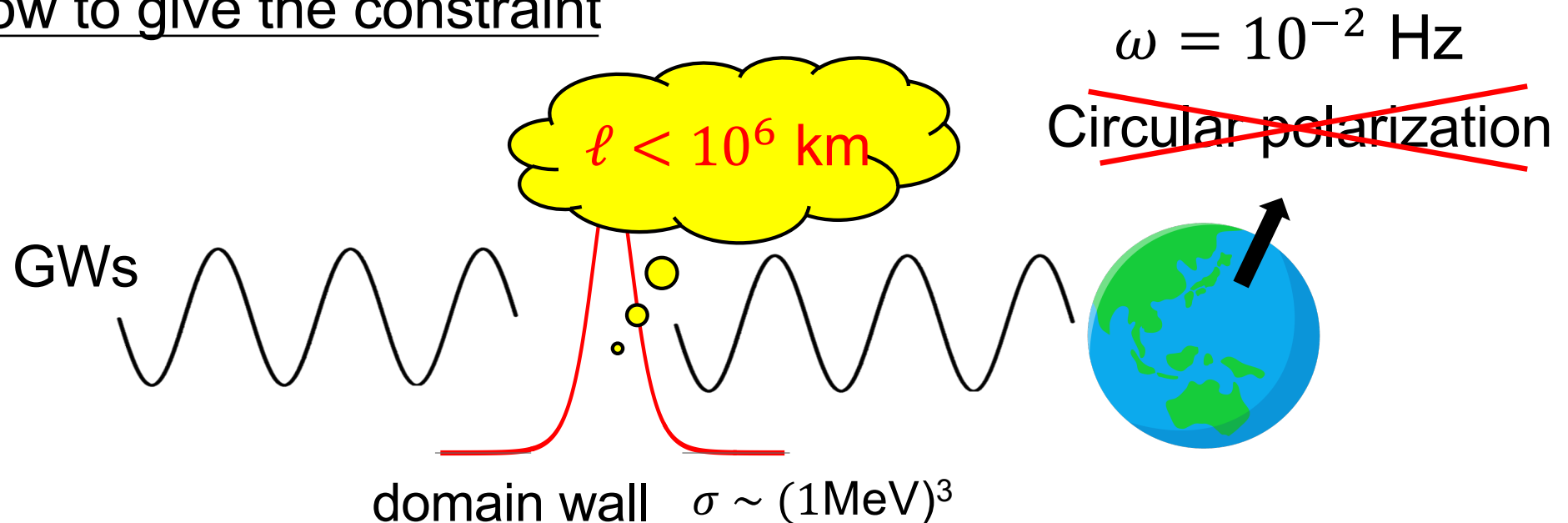
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CMB

$$\ell \sim 10^8 \text{ km} \quad \longrightarrow \quad \omega_c \sim 10^{-6} \text{ Hz}$$

How to give the constraint



- The degree of the circular polarization becomes large enough to be observed at near the critical frequency.

$$\omega_c \equiv \frac{M_p}{\ell^2 \eta^2}$$

M_p : Plank mass
 η : Energy of the domain wall
 ℓ : coupling constant between axion and gravity

- Observing the circular polarization with increasing frequency, we can obtain constraints on the coupling constant ℓ .

Future Work

Enabling the observation of GWs across a broad frequency!

(Especially high frequency)