

Forecasting $z \sim 2$ direct galaxy-cosmic web alignment for Subaru-PFS

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Overview

- Review recent results by Codis et al. 2018 ([arXiv:1809.06212v1](https://arxiv.org/abs/1809.06212v1)) on direct galaxy shape/spin-cosmic web alignment in the Horizon-AGN hydrosim
- Corresponding study of direct alignments in IllustrisTNG hydrosim @ $z=2$
- Forecast of expected spin, shape alignment signal detected by upcoming Subaru Prime Focus Spectrograph (PFS) survey

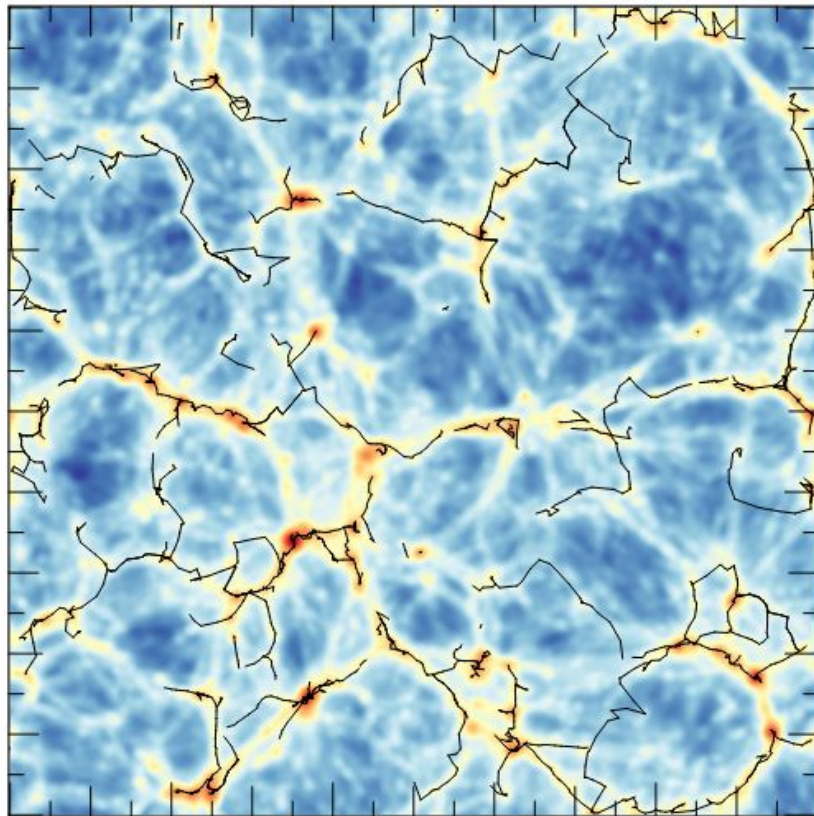
Codis et al. 2018: “Galaxy orientation with the cosmic web across cosmic time”

Data

- Horizon-AGN cosmological hydrodynamic simulation
 - Box length: 100 Mpc/h
 - Baryons, stars, stellar feedback
 - Cosmological redshifts: $z = 2, 1.5, 1.0, 0.5, 0$
- Galaxy sample
 - >50 star particles
 - Angular momentum vector (**spin**), minor axis of inertia tensor ellipsoid (**minor-axis shape**) calculated from star particles only
 - Stellar mass range: $\sim 10^8 - 10^{11-12} M_{\odot}$

Cosmic web characterization

- DisPerSE density ridge-extractor code ([arXiv:1009.4015v1](https://arxiv.org/abs/1009.4015v1))
 - “Skeleton” line segments along ridges in density, but using galaxy distribution directly instead of continuous density fields



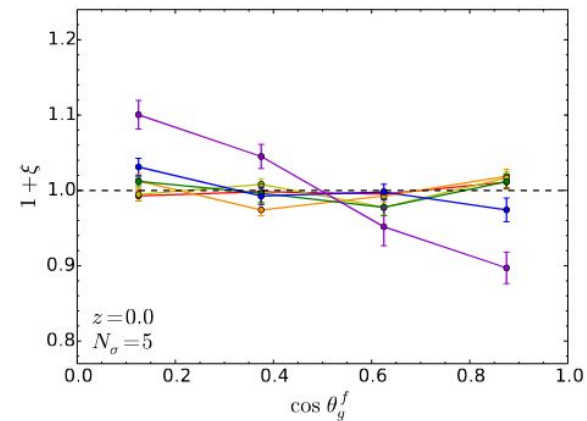
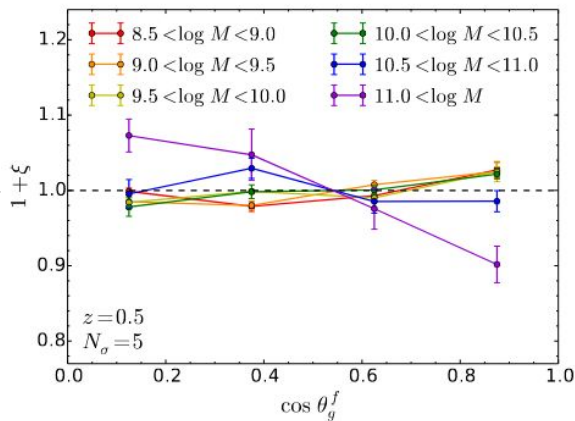
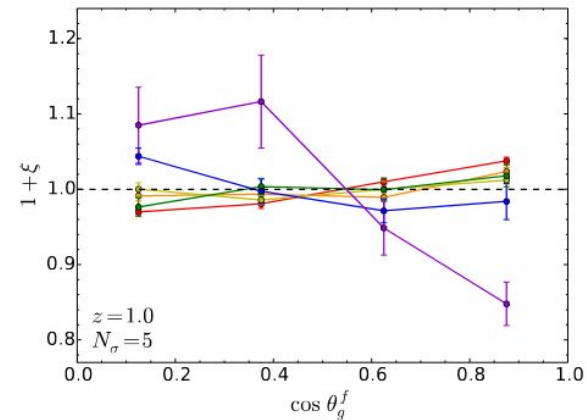
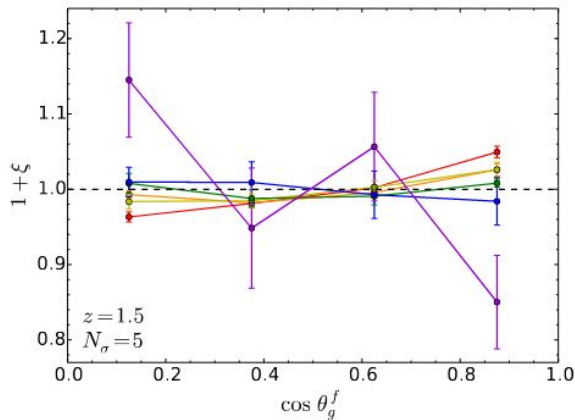
Codis et al. 2018, Figure 2

Methods

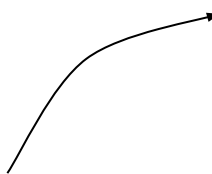
- Alignment metric for each galaxy
 - Absolute-valued dot product of galaxy's spin/minor-axis shape vector and average direction of 2 closest DisPerSE line segments
 - $|\cos \theta|$ (**alignment**)
 - $[0, 1]$ with 0 = perpendicular, 1 = parallel
- **Mean alignment** of collection of galaxies: $\langle |\cos \theta| \rangle$
 - Null case: $\langle |\cos \theta| \rangle = 0.5$ (2 vectors with independent uniformly random direction)
- Alignment error estimation
 - Error on mean alignment computed separately on 8 sub-cubes of simulation

Spin-Filament Results

- Clear spin-flip at transition mass
- Alignment strength decreases slightly with cosmic time

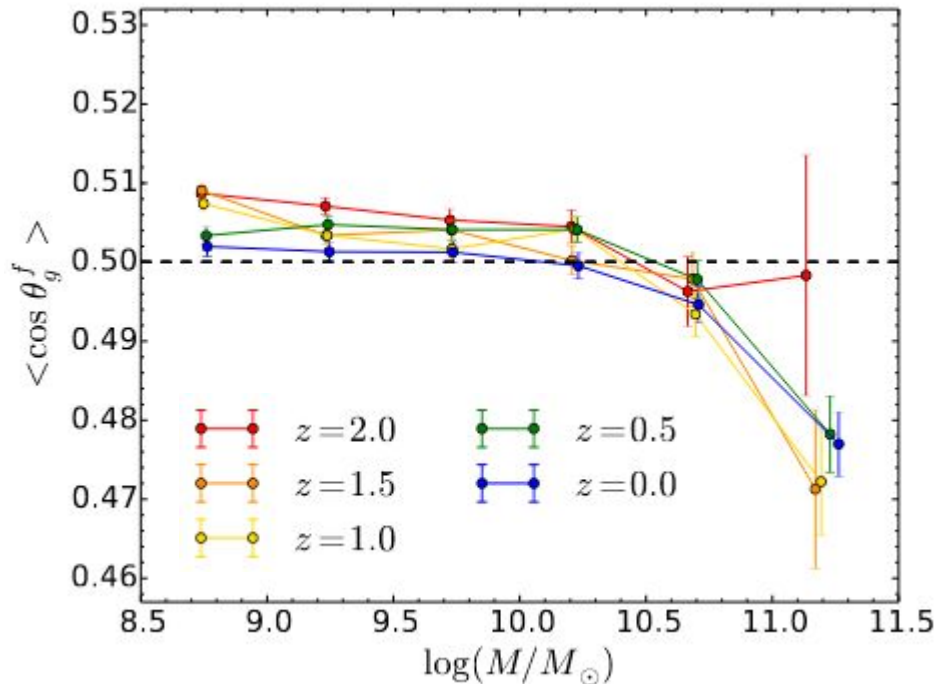


PDF of $|\cos \theta|$



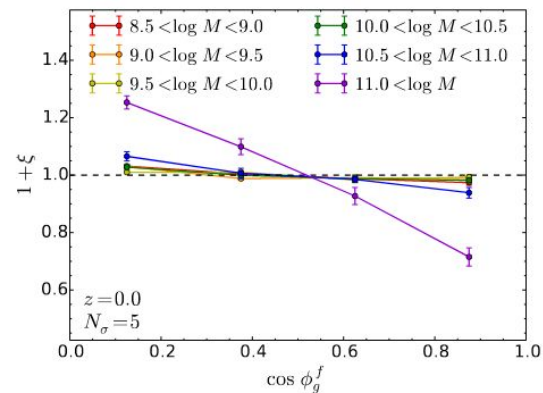
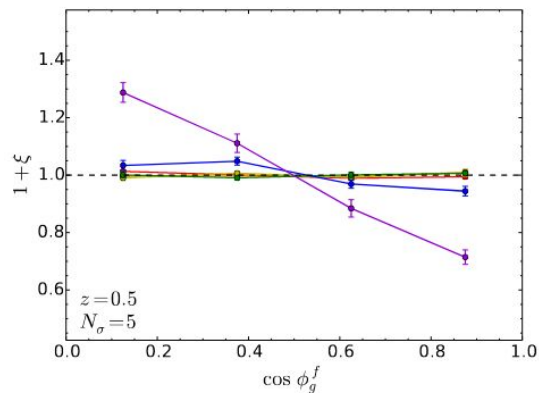
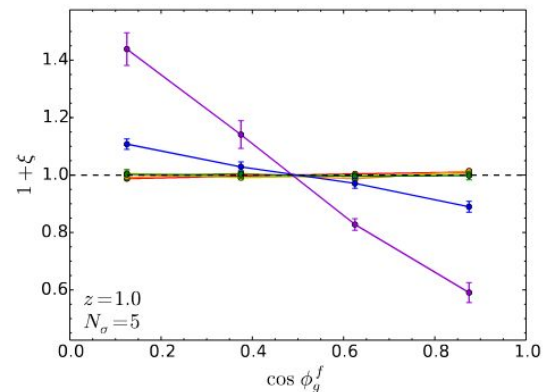
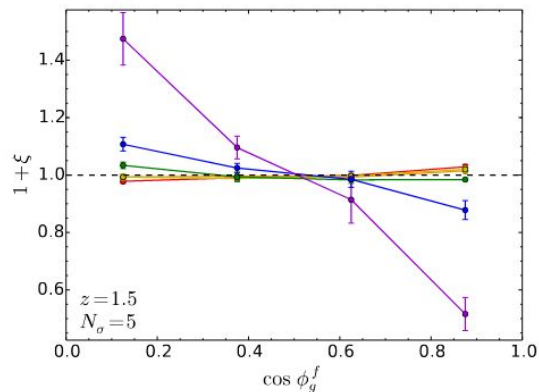
Spin-Filament Results contd.

- Transition mass at $\log_{10}(M_*/M_\odot) = 10.1 \pm 0.3$
- No clear redshift dependence for transition mass (likely due to galaxy sample size)



Minor-axis Shape-Filament Results

- Monotonic mass dependence
- Minor positive alignment for low-mass bins @ $z=1.5$
 - But unclear results for lower redshifts
- Suggestive of positive major axis-filament alignment trends



IllustrisTNG shape/spin alignments

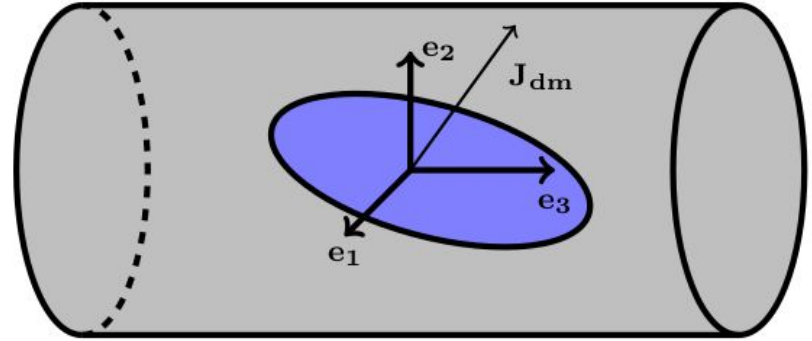
Data

- IllustrisTNG 300-1 cosmological hydrodynamical simulation
 - Box length: 205 Mpc/h
 - “Full-physics”: DM, baryons, stars, supermassive black holes
 - $z=2$ simulation snapshot
- Galaxy major-axis shape sample: Shi et al. 2021
 - Contrast with Codis et al.: reduced inertia-tensor’s **major** axis
 - Stellar mass $\geq 10^9 M_{\odot}$
- Galaxy spin sample
 - ≥ 50 *total* particles
- Simulation resolution comparison: TNG100-1
 - 75 Mpc/h: greater mass/physics resolution
 - Galaxy spins only; same cuts as 300-1

Cosmic Web Characterization

- **Deformation tensor**

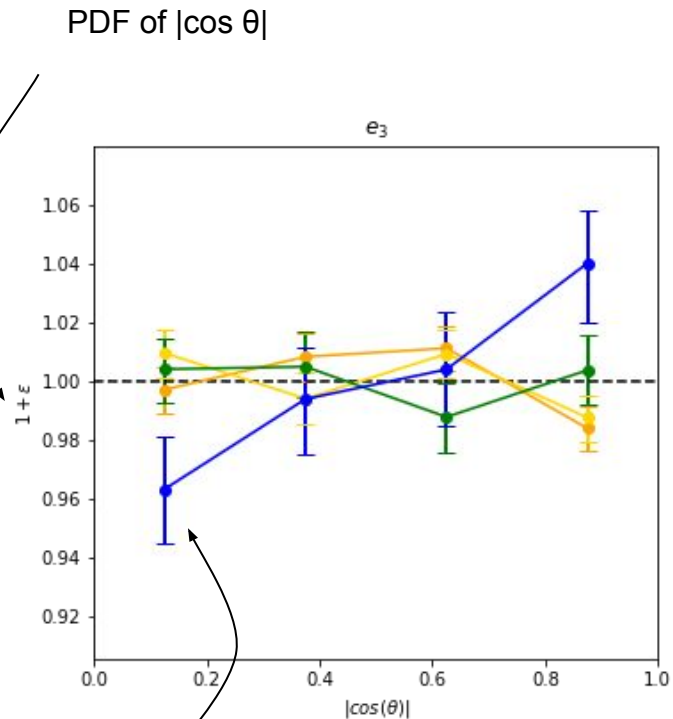
- Calculated from binned DM density field: Hessian of gravitational potential at each point
- Ordered eigenvalues of tensor $e_1 \geq e_2 \geq e_3$: matter collapse order in Zel'dovich approximation, along corresponding eigenvectors \mathbf{e}_1 , \mathbf{e}_2 , \mathbf{e}_3
- \mathbf{e}_3 filament direction + wall plane-parallel: generalized “cosmic web direction” for every point
- \mathbf{e}_1 also encodes useful information; \mathbf{e}_2 less so



Veena et al. 2020, Figure 4

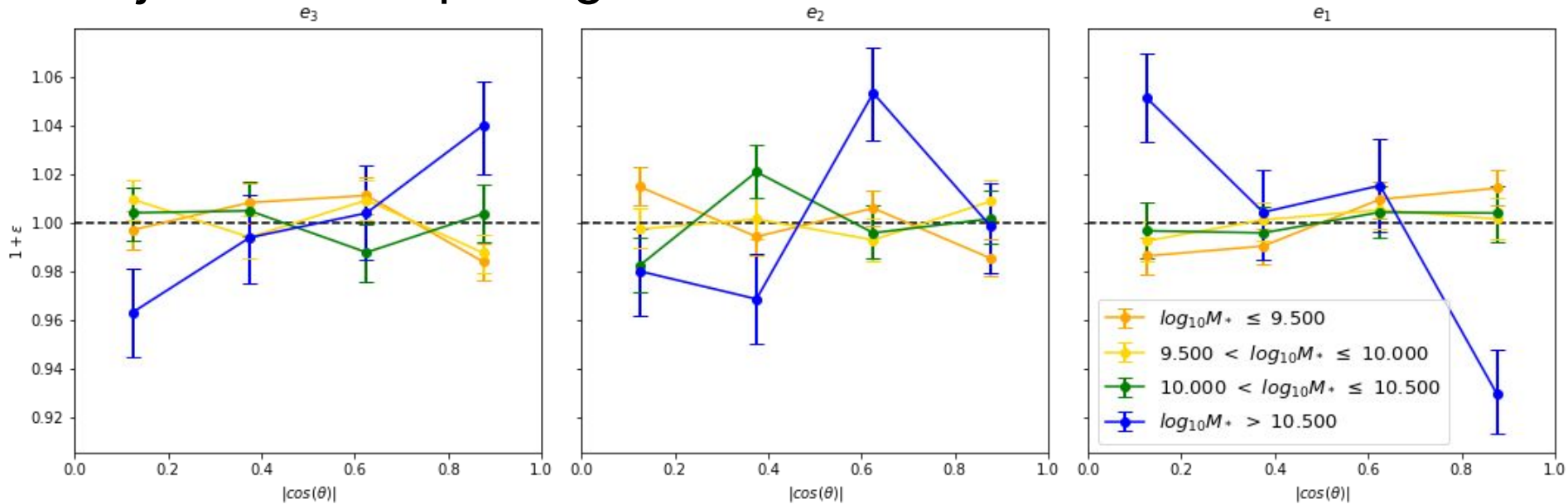
Methods

- Dark matter density binned to 512^3 grid, smoothed to 2 Mpc/h, deformation tensor calculated at each point
- Closest tensor to each galaxy found, $|\cos \theta|$ calculated between each eigenvector and major-axis shape/spin



1σ errors bootstrapped over (sub)sample

Major-Axis Shape Alignment Results

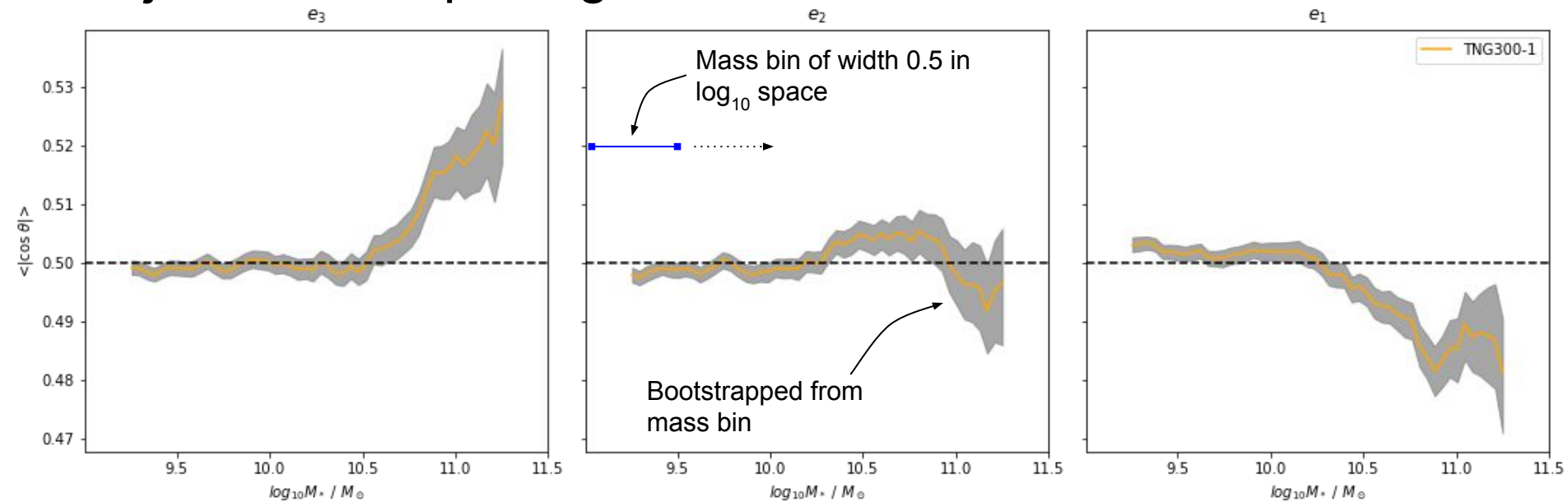


- Slightly negative trend until highest stellar mass bin
- Qualitatively agrees with Codis et al. minor-axis vs. filament

- No clear trends; less information about cosmic web encoded

- Monotonically more negative trend with mass

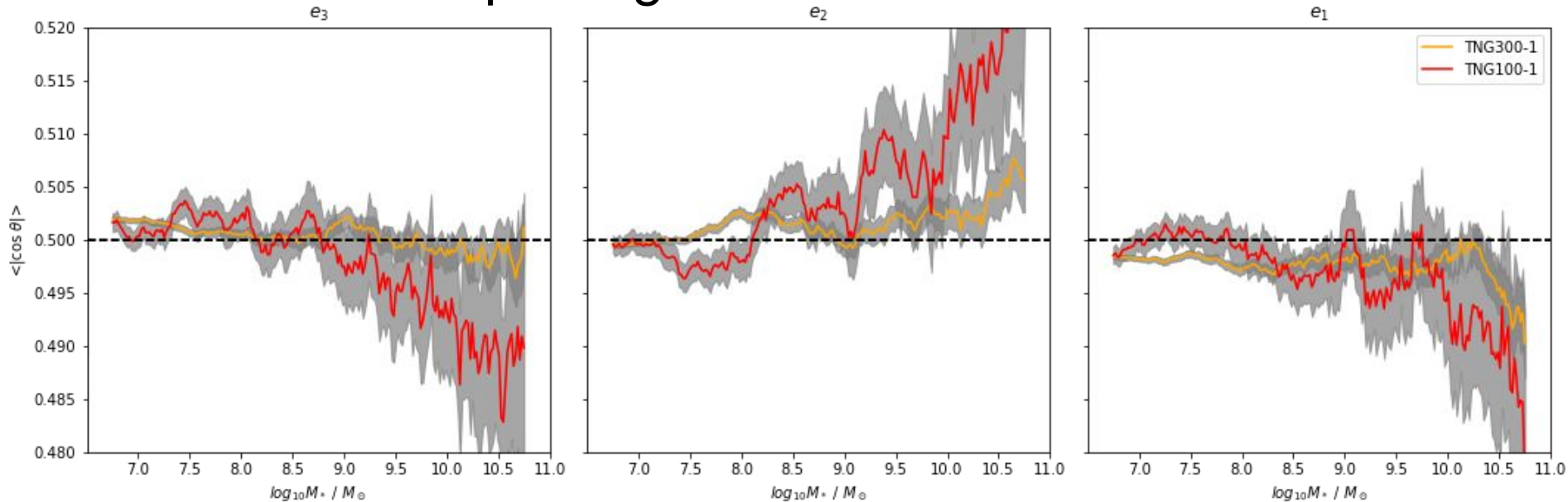
Major-Axis Shape Alignment Results contd.



- Clear transition mass at $M_* \sim 10^{10.5} M_\odot$!
- After transition mass, major-axis shape alignment strength monotonically increases

- Transition mass at $M_* \sim 10^{10.25} M_\odot$
- Negative signal after transition; similar amplitude to e_3

TNG300 & 100 Spin Alignment Results



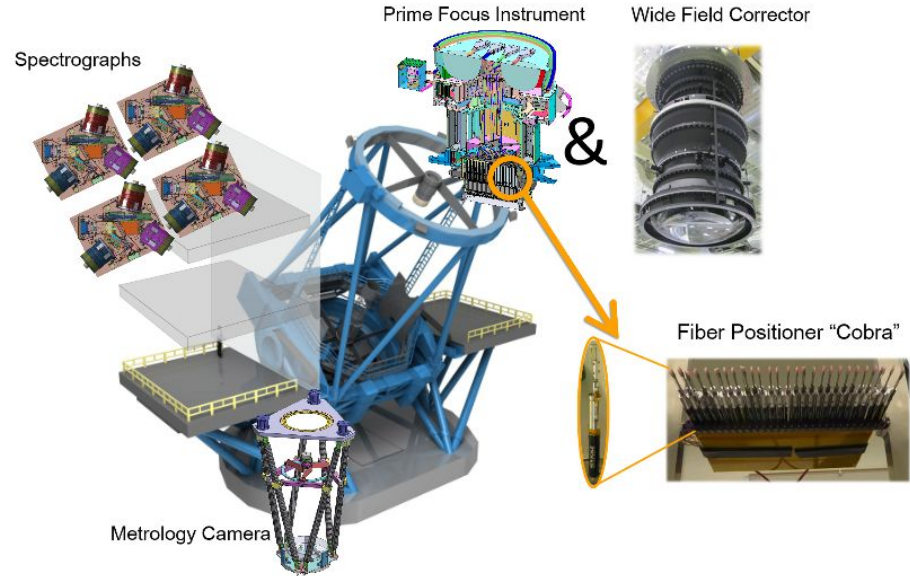
- Relatively small galaxy sample for TNG100 from $\sim 10^{8.5} M_\odot$ on
- For e_1 & e_3 , 1σ bootstrap errors of TNG100 agree with TNG300 above $\sim 10^8 M_\odot$
- e_3 spin-flip transition mass $\sim 10^{8.75}$ for TNG100, $\sim 10^{9.4}$ for TNG300
 - Disagrees with Codis et al. ($\sim 10^{10.4}$). Different cosmic web formalism / sim physics
 - Wang et al. 2018 ([arXiv:1810.04581v1](https://arxiv.org/abs/1810.04581v1)) reports $M_{\text{tr}} \sim 10^8$ for same cosmic web formalism on Illustris (75 Mpc/h)

Subaru-PFS Alignment Signal Forecast

How well can we measure this signal in PFS?

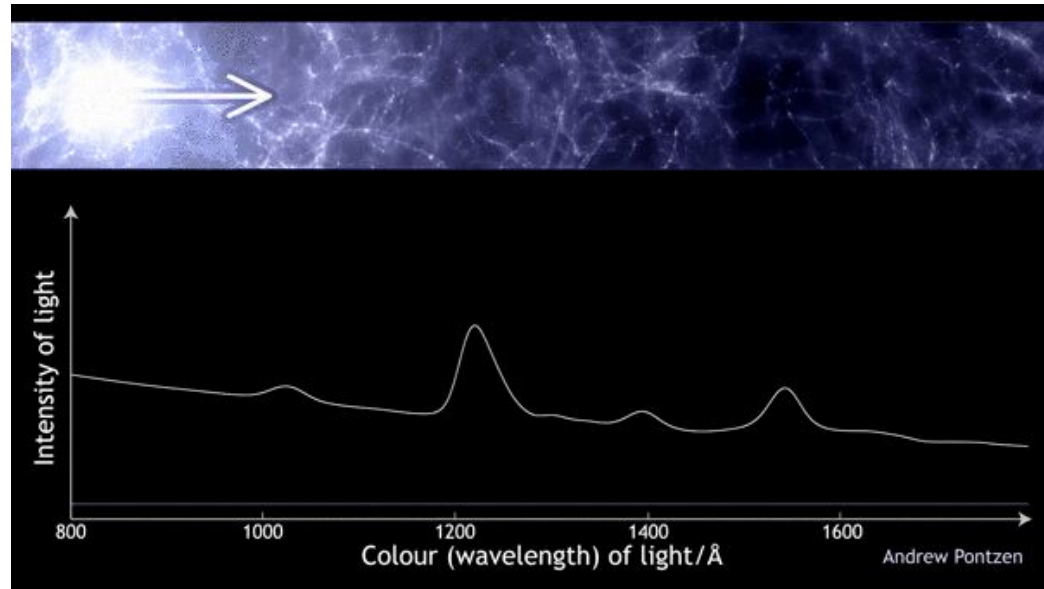
Subaru Prime Focus Spectrograph Survey

- Starting mid-2023
- Spectrographic redshifts / spectra from $0.6 < z < 7$
- Our focus is on “cosmic noon” subprogram @ $z \sim 2.3$
 - Spectrographic redshifts for galaxies within volume $2.7 * 10^7 h^{-3} \text{ Mpc}^3$
 - 15,000/30,800 galaxies for program, depending on exact distribution of observing time



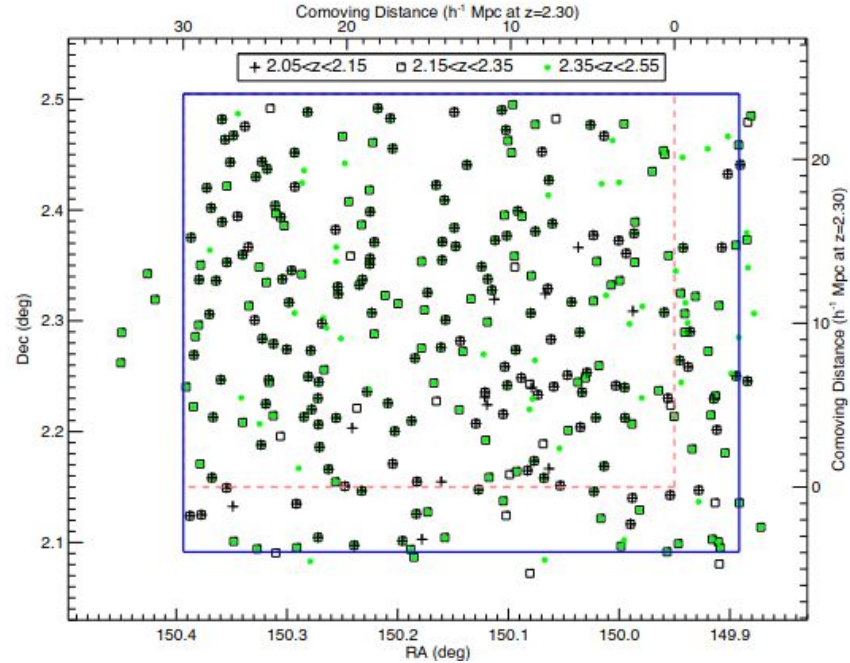
IGM Tomography

- Cosmic web reconstruction at high-z hard: few galaxies!
 - COSMOS-level of study needed to attempt: see Ata et al. 2020 ([arXiv:2004.11027v2](https://arxiv.org/abs/2004.11027v2))
- IGM tomography offers direct probe of cosmic web
 - Neutral H produces redshifted absorption lines (Lyman-alpha forest) in spectrum of background objects
- CLAMATO survey: $4.1 * 10^5 h^{-3} \text{ Mpc}^3$
 - PFS to probe 2 orders of magnitude higher volume!



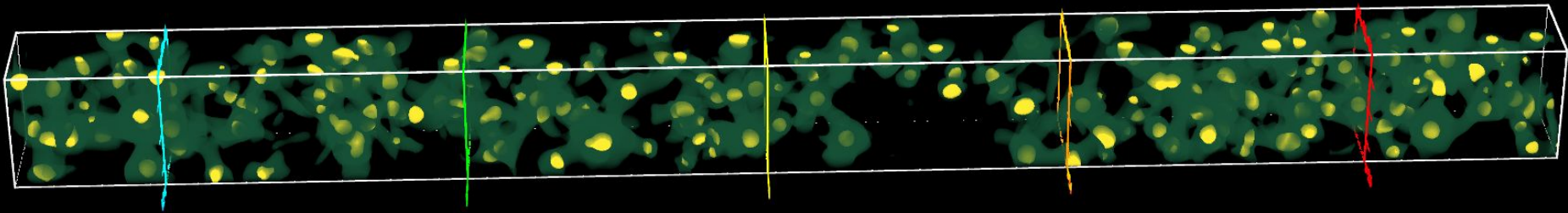
Source: [UCL Mathematical & Physical Sciences](https://www.ucl.ac.uk/mathematical-physics)

IGM Tomography contd.

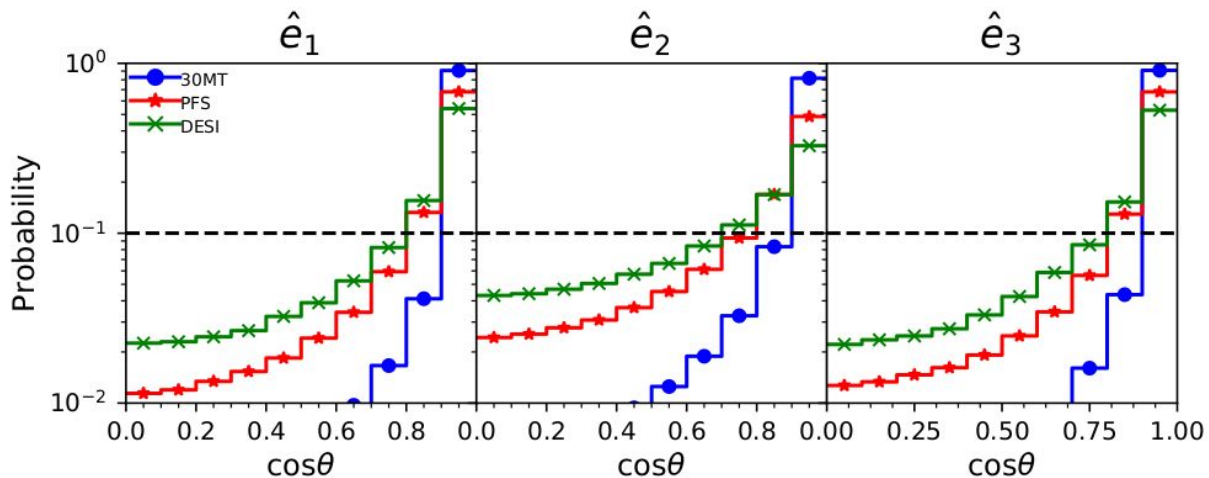


↑ Source: CLAMATO DR2 (Horowitz et al. 2021, [arXiv:2109.09660v1](https://arxiv.org/abs/2109.09660v1))

IGM Tomography contd.



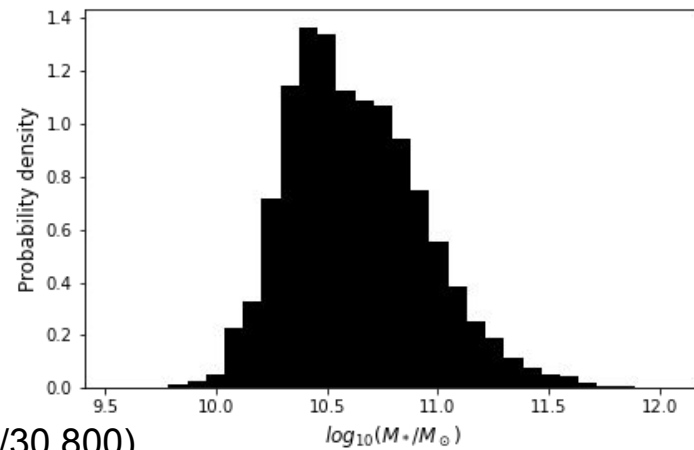
↑ Source: CLAMATO DR2 (Horowitz et al. 2021, [arXiv:2109.09660v1](https://arxiv.org/abs/2109.09660v1))



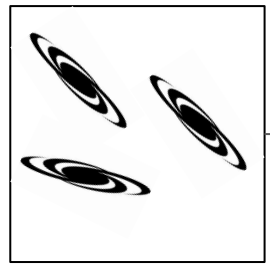
↑ Alignment between reconstructed eigenvec and true eigenvec. Source: Horowitz et al. 2019 ([arXiv:1903.09049v3](https://arxiv.org/abs/1903.09049v3))

Observational Data

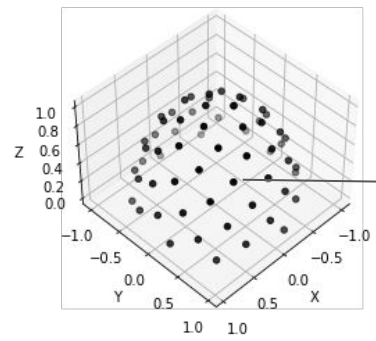
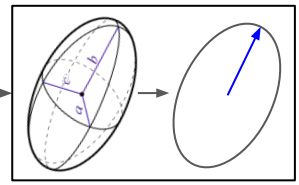
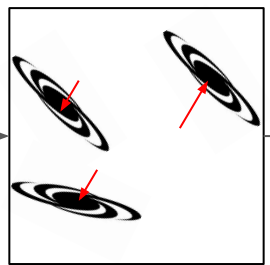
- Illustris TNG300 simulation @ $z=2$
- Galaxy sample magnitude-limited
 - Observed J-band magnitude (no dust)
 - Match expected number of PFS galaxy spec-z's (15,000/30,800)
 - Select dispersion-dominated (spheroid) galaxies via $\kappa_{\text{rot}} < 0.5$ ($\kappa_{\text{rot}} = K_{\text{rot}} / K$)
 - $M_* = [10^{10}, 10^{11.5}] M_{\odot}$
- For 64 viewing angles on half-sphere:
 - Displace galaxies by peculiar velocities along viewing angle LoS: redshift-space distortion (RSD)
 - Construct mock PFS-like IGM tomography survey along viewing angle LoS
 - Reconstruct cosmic web deformation tensor using TARDIS-II code (Horowitz et al. 2021, [arXiv:2007.15994v1](https://arxiv.org/abs/2007.15994v1))
 - Project shape ellipsoid onto viewing angle plane, take longest axis of ellipse as major-axis shape
 - Project deformation tensor eigenvectors onto viewing angle plane



Observational galaxy sample



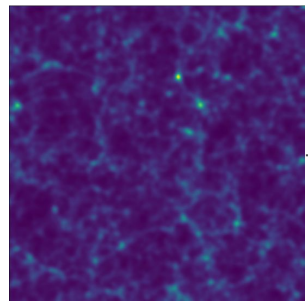
RSD



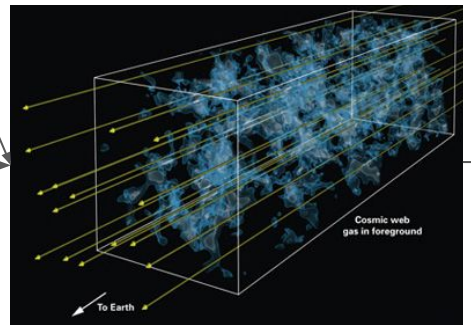
Viewing angle ℓ

Eigenvector projection onto viewing angle

True DM density

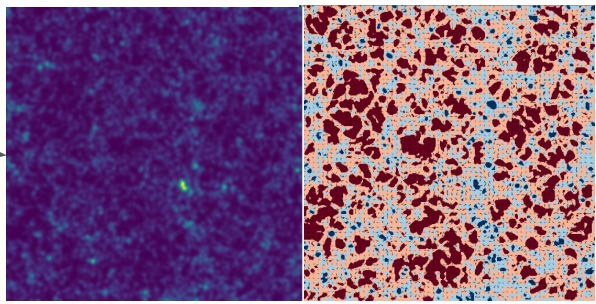


Mock IGM tomo. survey



TARDIS-II

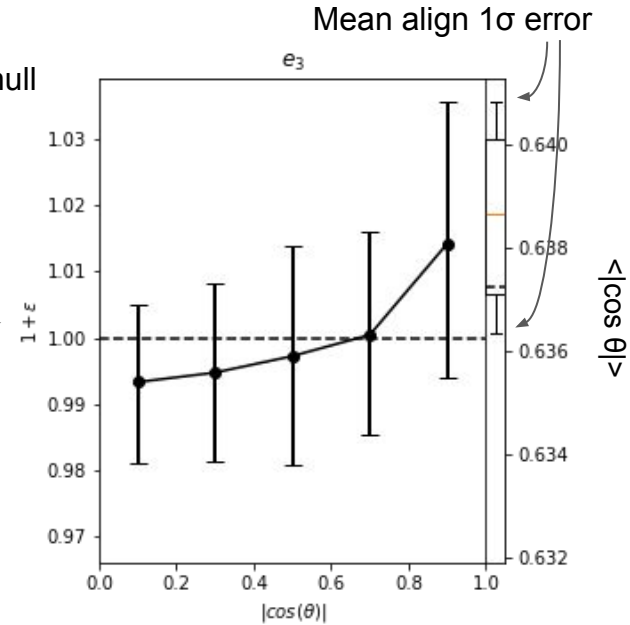
Reconstructed DM density + deformation tensor



2D Alignment

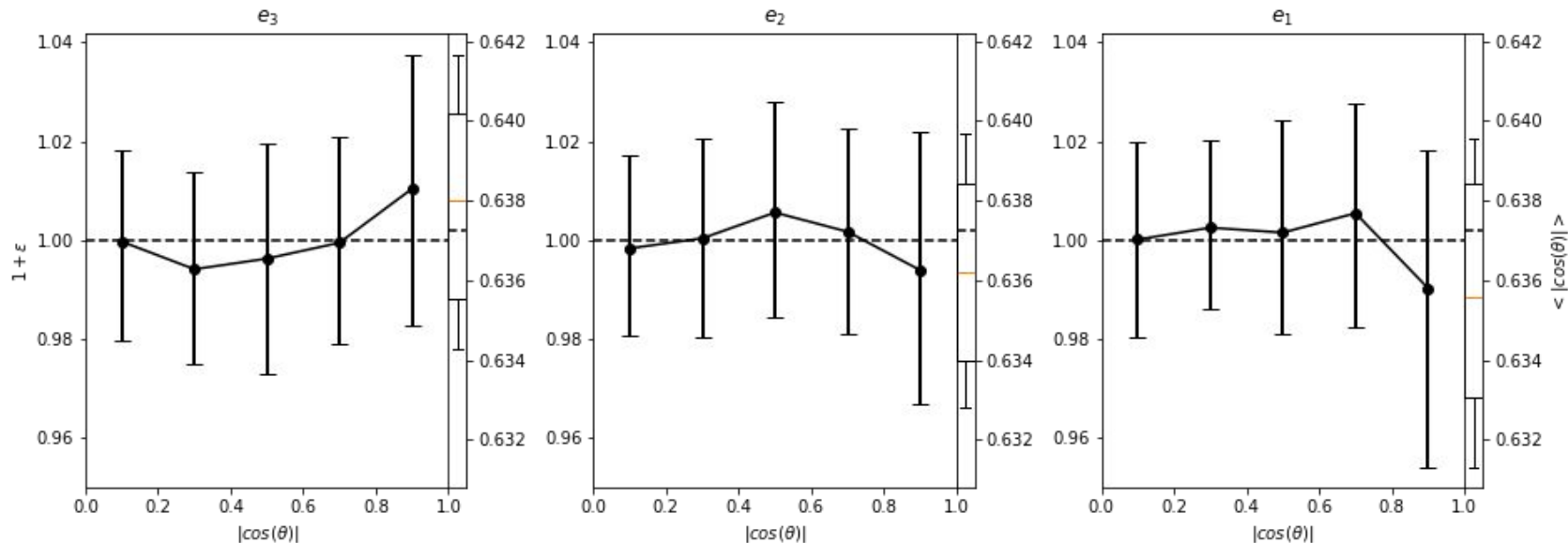
- Alignment null distribution after projection (2D): upper half of Beta[$\alpha = \beta = 0.5$]
 - Random 3D vectors after projection \rightarrow random 2D vectors
- Histogram bins calculated over alignments from all viewing angles
- For each bootstrap iteration, select alignments from random viewing angle first

Excess prob.
density over null
dist.



Fiducial Observation Time (N = 15,000)

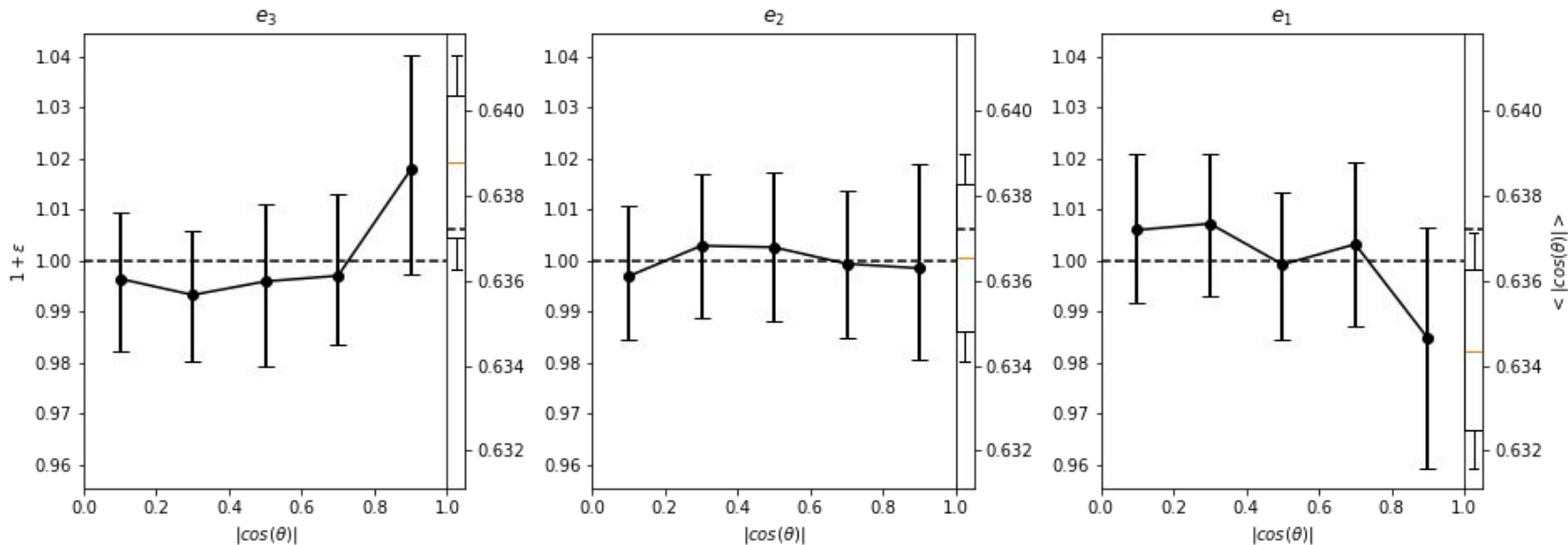
Cosmic web reconstructed with IGM tomography
Major-axis shape projected on 64 viewing angles



e_3	e_2	e_1
+0.21 σ	-0.30 σ	-0.41 σ

Extra Observations (N = 30,800)

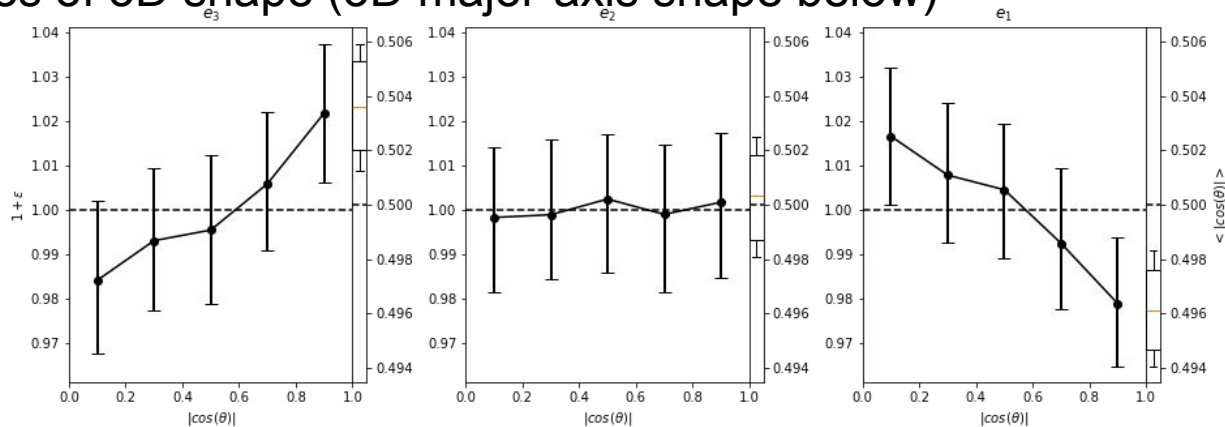
Cosmic web reconstructed with IGM tomography
Major-axis shape projected on 64 viewing angles



e_3	e_2	e_1
+0.58 σ	-0.27 σ	-1.02 σ

Avenues for Improvement

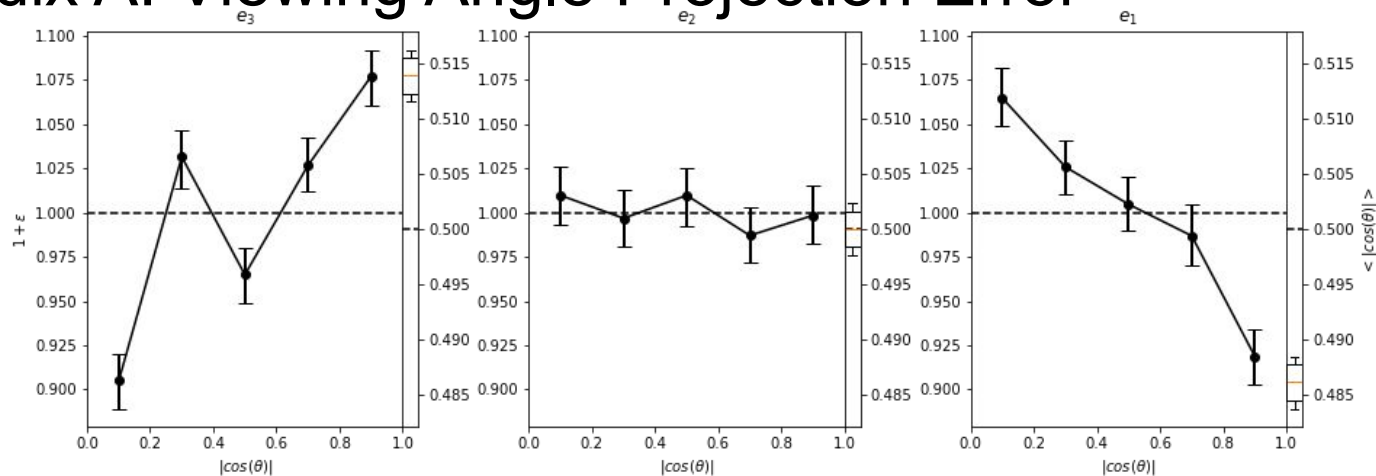
- Increased volume for next-gen survey
 - Approximate overall significance σ_T as $\sqrt{(\sigma_3^2 + \sigma_1^2)}$.
 - $\sigma_T = 0.46 \mid 1.17$, for $N=15,000 \mid N=30,800$
 - Assuming same magnitude distribution and errors $\propto N^{0.5}$, need 3x volume ($N=90,000$) for $\sigma_T=2$
- Improved IGM tomography reconstruction
- Estimates of 3D shape (3D major-axis shape below)



Appendices

Appendix A: Viewing Angle Projection Error

3D



2D

