

The Smarr Way of Doing Things

Bert Vercocke

Universiteit van Amsterdam

Based on work with:

1507.01022 Diego Cohen-Maldonado, Juan Diaz and Thomas Van Riet
1511.07453

previous work Souvik Banerjee, Iosif Bena, Borun Chowdhury, [Paul de Lange](#),
Daniel Mayerson, [Andrea Puhm](#), Amitabh Virmani

YITP Kyoto University, Nov 25 2015

Upshot

Non-extremal black holes

- ▶ Microstates – can we leap ahead?
 - Isolated stationary examples

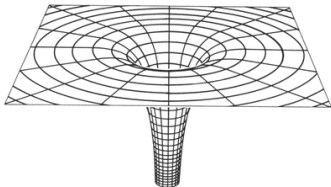
- ▶ Use a powerful analogy in string theory
 - Black holes \leftrightarrow Flux vacua

Upshot

- ▶ Two systems in string theory:

Vacua: $e^{2A(y)} ds_4^2 + \underbrace{g_{mn}(y) dy^m dy^n}_{6d}$

Black holes: $-e^{2U(x)} dt^2 + \underbrace{g_{mn}(x) dx^m dx^n}_{3d + 6d}$



- ▶ Break SUSY? Anti-branes in warped flux backgrounds!

Outline

1. Introduction: SUSY vs. non-SUSY
2. Probe branes in flux backgrounds
3. Smarr relations and polarization
4. Looking ahead

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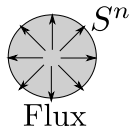
Chern-Simons couplings in String Theory

- Action: topological terms $\int (\star F \wedge F + A \wedge F \wedge F)$

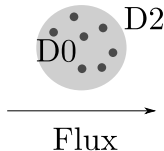
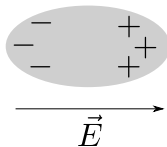
$$\text{IIB:} \quad d \star F_5 = H_3 \wedge F_2$$

$$\text{5d SUGRA:} \quad d \star F_2 = F_2 \wedge F_2$$

1. Charge dissolved in flux

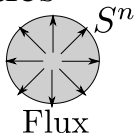


2. Dipole effect



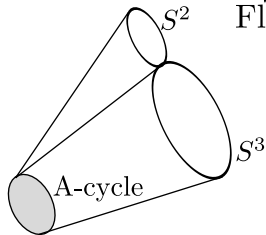
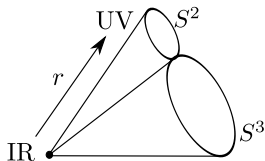
Supersymmetry = bye bye singularities

Charge dissolved in flux: **Topology and new IR phases**



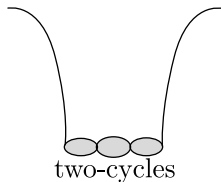
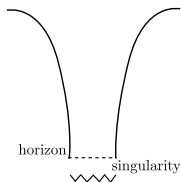
- ▶ RG flow: deformed conifold

[Klebanov-Strassler '00]



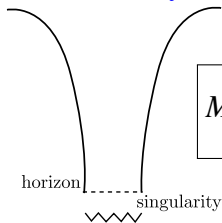
- ▶ Black holes & microstate geometries

[Denef, de Boer, Lunin-Mathur, Bena, most of the audience...]

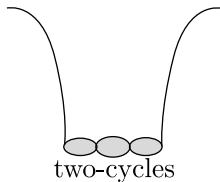


No Supersymmetry = ?

No stationary solitons without horizons or **topology** [Gibbons-Warner 13]



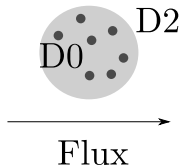
$$M_{ADM} = \int_{horizon} (\dots) + \int_{space} H_2 \wedge F_2$$



- ▶ 5d/6d/11d [Gibbons-Warner 13; Kunduri-Lucietti 13] [de Lange, Mayerson, BV 15]
[Haas 14]

Examples

- ▶ Very few supergravity solutions [JMART + extensions: Bena, Bossard, Katmadas, Turton '15; Running-Bolt Bena, Giusto, Ruef, Warner '09]
- ▶ Large classes at probe level [Bena, Puhm, BV 11-12]
→ dipole effect



Outlook for talk (and research)

| Black holes | | Flux Vacua |
|--------------------|---|-----------------------|
| in progress | ← | back-reaction: debate |
| ↕ | | ↕ |
| Smarr relation | → | this talk |

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Probe branes in flux backgrounds

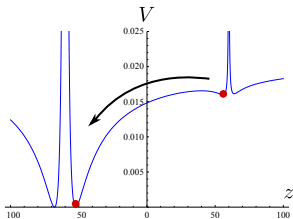
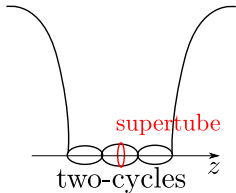
1. Compare probe constructions
2. Review anti-brane back-reaction debate

Compare probe constructions

Microstate geometries

M-theory $M_5 \times T^6$

polarization: M2-M2 \rightarrow M5

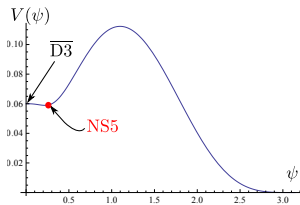
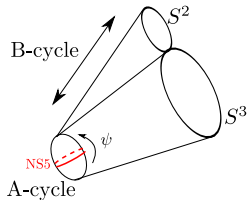


[Bena, Puhm, BV 11-12]

Klebanov Strassler

IIB $M_4 \times X_6$

polarization: $\overline{D3} \rightarrow$ NS5



[Kachru, Pearson, Verlinde '01]

Importance

Used for:

- ▶ Klebanov-Strassler and friends:
 - Metastable SUSY breaking
 - de Sitter vacua (KKLT) [Kachru, Kallosh, Linde, Trivedi '03]
- ▶ Black hole microstates
 - Alternative to collapse

Backreaction for p anti-branes unknown!

Review of backreaction debate

Short statement for p anti-branes:

- ▶ Does pathology-free supergravity solution exist ($g_s p \gg 1$)?
Many no-go results [Saclay,Uppsala,Leuven,Hannover '09-'15...]
- ▶ Does a string theory solution exist ($g_s p$ small)?
“Yes-go” result from EFT [Polchinski and friends '14-'15]

Note: probe limit is $p/M \ll 1$ ($M = \int_A F_3$ background flux)

Black hole microstate geometries?

Supergravity: many no-go results

$g_{sp} \gg 1$ (supergravity) [Saclay,Uppsala,Leuven,Hannover...]

H_3 -density has an unphysical singularity at brane

$$|H_3|^2 \rightarrow \infty$$

- ▶ Tachyonic instabilities
- ▶ Time-dependence?
- ▶ Singularity not cloaked by a horizon

Caveats:

- ▶ Mostly for smeared solutions
- ▶ NS5 polarization for $p/M \ll 1$? (the true metastable state!)

String theory: positive arguments

$g_{sp} \ll 1$ (single anti-brane)

- ▶ EFT: match string perturbation theory

[Michel,Mintun,Polchinski,Puhm,Saad]

- ▶ Anti-D3 on orientifold: nilpotent superfields

[Kallosh+ Quevedo, Uranga; Bergshoef,Freedman, Van Proeyen, Wrase]

Caveats:

- ▶ EFT for NS5 polarization?
- ▶ Validity of 4d supergravity model with nilpotent superfield?

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Our work: Match near and far solutions

IR/UV matching:

► No-go results:

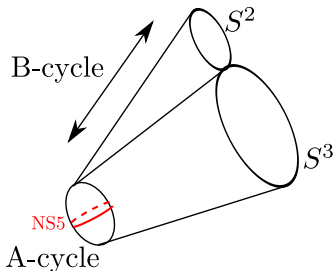
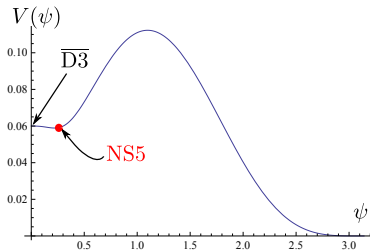
- IR/UV matching: singular anti-D3 brane, even at finite T

[Gautason, Junghans, Zagermann '13] [Blaback, Danielsson, Van Riet, Vargas '14]

► Our result:

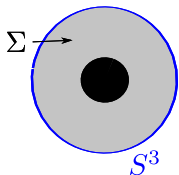
- $|H_3|^2$ singularity could become self-energy of NS5!

[Cohen-Maldonada, Diaz, Van Riet, BV '15]



Smarr relation in five dimensions

- ▶ K timelike Killing vector \rightarrow ADM mass in flat space:



$$\int_{\infty} \star dK = \int_H \star dK + \int_{\Sigma} d \star dK.$$

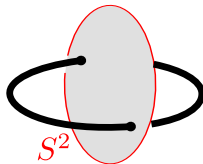
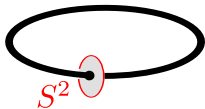
$$M_{ADM} = TS + \Phi Q + \phi q + \int_{\Sigma} H_2 \wedge F_2$$

[Empanan '04, Copsey-Horowitz '05, Gibbons-Warner '13, Kunduri-Lucietti '13]

- ▶ Electric monopole charge Q , dipole charge q :

$$Q = \int_{S^3} \star F$$

$$q = \int_{S^2} F$$



- ▶ Dipole charge can support mass: NS5 polarization?

Smarr relation for anti-D3 in KS

- ▶ Ansatz $ds_{10}^2 = e^{2A} ds_4^2 + ds_6^2$
 $C_4 = \alpha \text{vol}_4$
 $e^{-\phi} \star_{10} H_3 = C_4 \wedge F_3 + dB_6.$

- ▶ Boundary term from EOM [Gautason, Junghans, Zagermann '13]

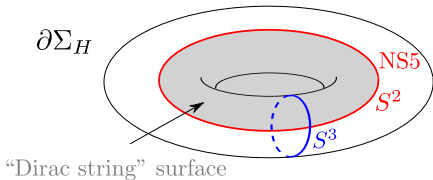
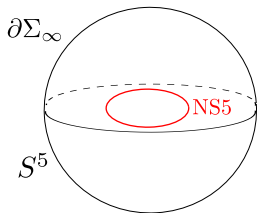
$$\int_{\infty} \mathcal{B} = \int_H \mathcal{B}, \quad \mathcal{B} = C_4 \wedge F_5 - B_2 \wedge dB_6$$

- ▶ Asymptotically: mass! [Blaback, Danielsson, Van Riet, Vargas '14]

$$M = \int_{\infty} \mathcal{B} \quad \text{with } M \propto p$$

- ▶ **NO-GO:** If anti-D3, $e^{-\phi} |H_3|^2 \rightarrow \infty$

Smarr relation for NS5 polarization



$$\int_{\infty} \mathcal{B} = \int_H \mathcal{B},$$

$$\mathcal{B} = C_4 \wedge F_5 - B_2 \wedge dB_6$$

- ▶ Just like in flat space:

$$\boxed{M = \Phi Q_{D3} + \phi q_{NS5} V_{S^2}}, \quad Q_{D3} = \int_{S^5} F_5, \quad q_{NS5} = \int_{S^3} H_3,$$

- ▶ Near-horizon

$$C_4 = \Phi \text{vol}_4, \quad B_6 = \phi \text{vol}_4 \wedge \text{vol}_{S^2}$$

Interpret singularity as NS5 self-energy?

- ▶ Near solution for NS5: specific divergence

$$e^{-\phi}|H_3|^2 \sim e^{-2A} \quad \Rightarrow \quad e^{-\phi} \star_{10} H_3 = dB_6 + C_4 \wedge F_3 \sim e^{4A}$$

- ▶ With $C_4 = \Phi \text{vol}_4$, $B_6 = \phi \text{vol}_4 \wedge \text{vol}_{S^2}$:

$$d\phi + \Phi \rightarrow 0$$

- ▶ Mass is number of anti-branes:

$$M = \Phi Q_{D3} + \phi q_{NS5} V_{S^2} > 0$$

Bye-bye no-go theorem

- ▶ If anti-D3: $\phi = \text{const.}$, $Q_{NS5} = 0 \rightarrow$ inconsistency
- ▶ If NS5: no problem to avoid singularity!

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Summary

Vacua with hierarchies and black holes very similar

- ▶ Warped throats
- ▶ Topology and flux resolves singularities
- ▶ Polarized branes \leftrightarrow metastable solutions

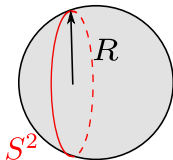
Smarr relation: fate of polarized state!

Outlook for vacua

For vacua:

- ▶ Confront NS5 result with probe: discrepancy!

$$R \sim \sqrt{p/M} \quad \text{vs} \quad R \sim p/M$$



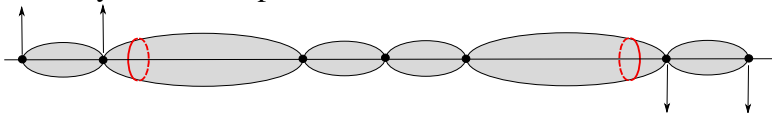
- ▶ Redo EFT for NS5? [Michel-Mintun-Puhm-Polchinski-Saad '14]
- ▶ Cosmological fuzzballs?

Outlook for non-extremal microstates

- ▶ Back-reaction of probe constructions

$$M_{ADM} = \Phi Q + \phi q + \int_{\Sigma} H_2 \wedge F_2$$

- ▶ Stability \leftrightarrow time-dependence?



[Bena, Pasini '15]

- ▶ Are microstate geometries property of AdS_3/CFT_2 ?