

THE SNO COLLABORATION

E.P. Bonvin, M.G. Boulay, E.D. Earle, L. Erhardt, H.C. Evans, G.T. Ewan, R.J. Ford,
A.L. Hallin, A. Hamer, J.D. Hepburn, C.J. Jillings, H.W. Lee, J.R. Leslie, H-B. Mak,
A.B. McDonald, W. McLatchie, B.A. Moffat, T.J. Radcliffe, B.C. Robertson, P. Skensved,
R.L. Stevenson, B. Sur

Queen's University, Kingston, Ontario, Canada

W.F. Davidson, F. Dalnoki-Veress, J. Farine, C.K. Hargrove, I. Levine, K. McFarlane,
T. Noble, M. O'Neill, M. Shatkay, D. Sinclair,
Centre for Research in Particle Physics, Ottawa, Ontario, Canada

M.C. Chon, P. Jagam, J. Law, R. Ollerhead, J.J. Simpson, N. Tagg, J.-X. Wang
University of Guelph, Guelph, Ontario, Canada

J. Bigu, E.D. Hallman, R.U. Haq, G. Jonkmans, A. Roberge, E. Saettler, R. Tafirout,
C.T. Virtue
Laurentian University, Sudbury, Ontario, Canada

J. Heise, R. Helmer, R. Komar, C. Nally, A. Poon, C. Waltham
University of British Columbia, Vancouver, BC, Canada

E.W. Beier, D. Cowen, W. Frati, J.R. Klein, P. Keener, D. McDonald, M. Neubauer,
M. Newcomer, R. Van Berg, R. Van de Water, P. Wittich
University of Pennsylvania, Philadelphia, PA, USA

J. Boger, R.L. Hahn, J.K. Rowley
Brookhaven National Laboratory, Upton, LI, NY, USA

T.J. Bowles, M.M. Fowler, A. Goldschmidt, A. Hime, J.B. Wilhelm, J.M. Wouters,
Los Alamos National Laboratory, Los Alamos, NM, USA

Y.D. Chan, M.P. Isaac, K.T. Lesko, E. Norman, C. Okada, A. Schuelke, A.R. Smith,
R. Stokstad
Lawrence Berkeley Laboratory, Berkeley, CA, USA

J.C. Barton, S. Biller, J.R.N. Cameron, X. Chen, G. Doucas, H.D. Heron, N.A. Jelley,
A.B. Knox, M.D. Lay, W.J. Locke, J.J. Lyon, M.E. Moorhead, N.W. Tanner, R.K. Taplin,
M. Thorman, P.T. Trent, D.L. Wark, N. West
University of Oxford, Oxford, United Kingdom

Q.R. Ahmad, M.C. Browne, P.J. Doe, C.A. Duba, S.R. Elliott, J.V. Germani, K.M. Heeger,
R. Meijer-Drees, R.G.H. Robertson, T.D. Steiger, P. Thornewell, J.F. Wilkerson,
University of Washington, Seattle, WA, USA

H.H. Chen,
University of California at Irvine, Irvine, CA, USA

Sudbury Neutrino Observatory Anticipated Rates

Experiment	Detection Rates (not SSM)
Chlorine	~25/year
Kamiokande	~45/year
SAGE/GALLEX	~50/year each
SuperK	~5000/year (>6.5 MeV)

SNO 1000 tons of D₂O



(Bahcall 95)

BP98

CC: 15000 events/year SSM 11000

NC: 6000 events/year SSM 4700

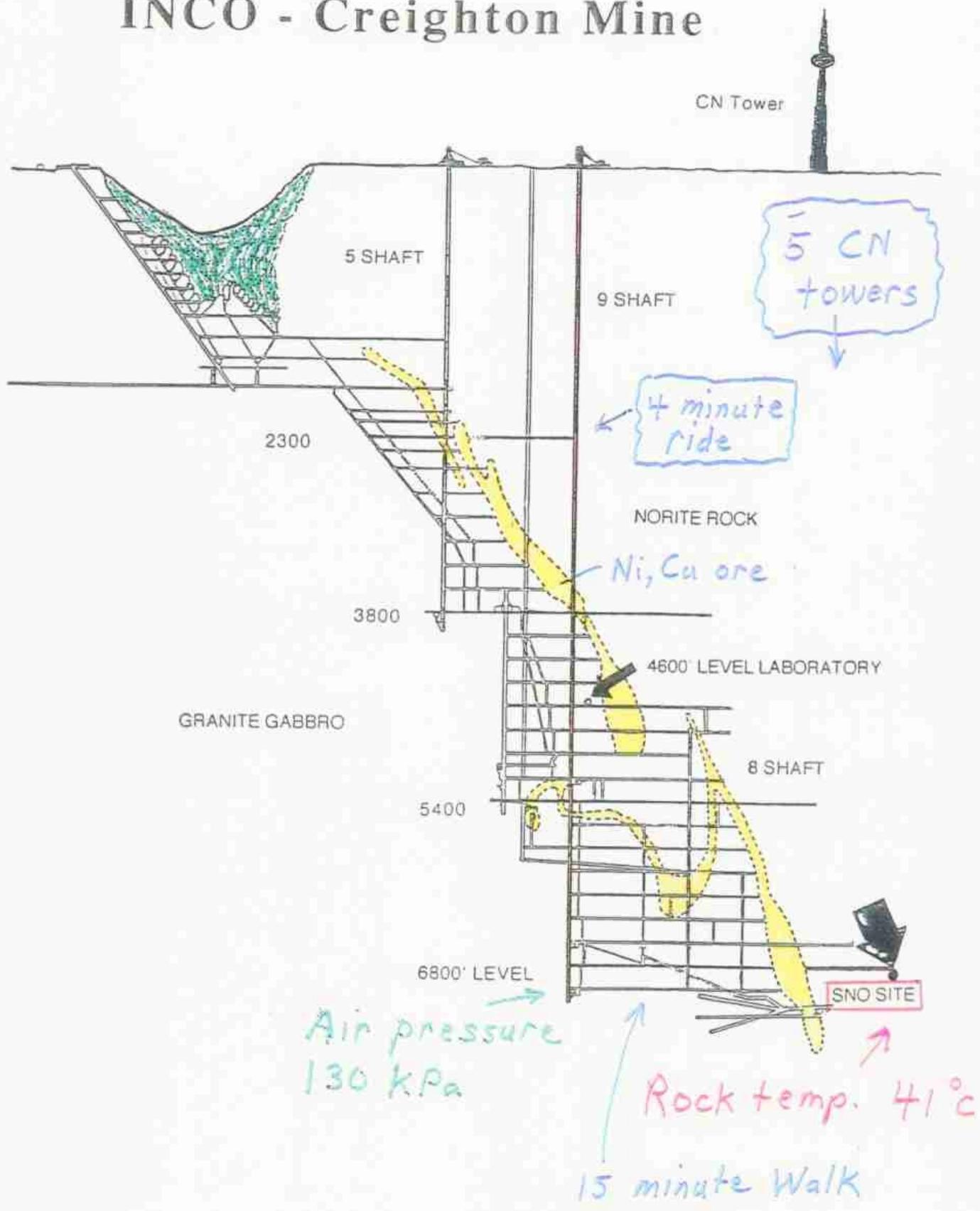
ES: 1000 events/year SSM

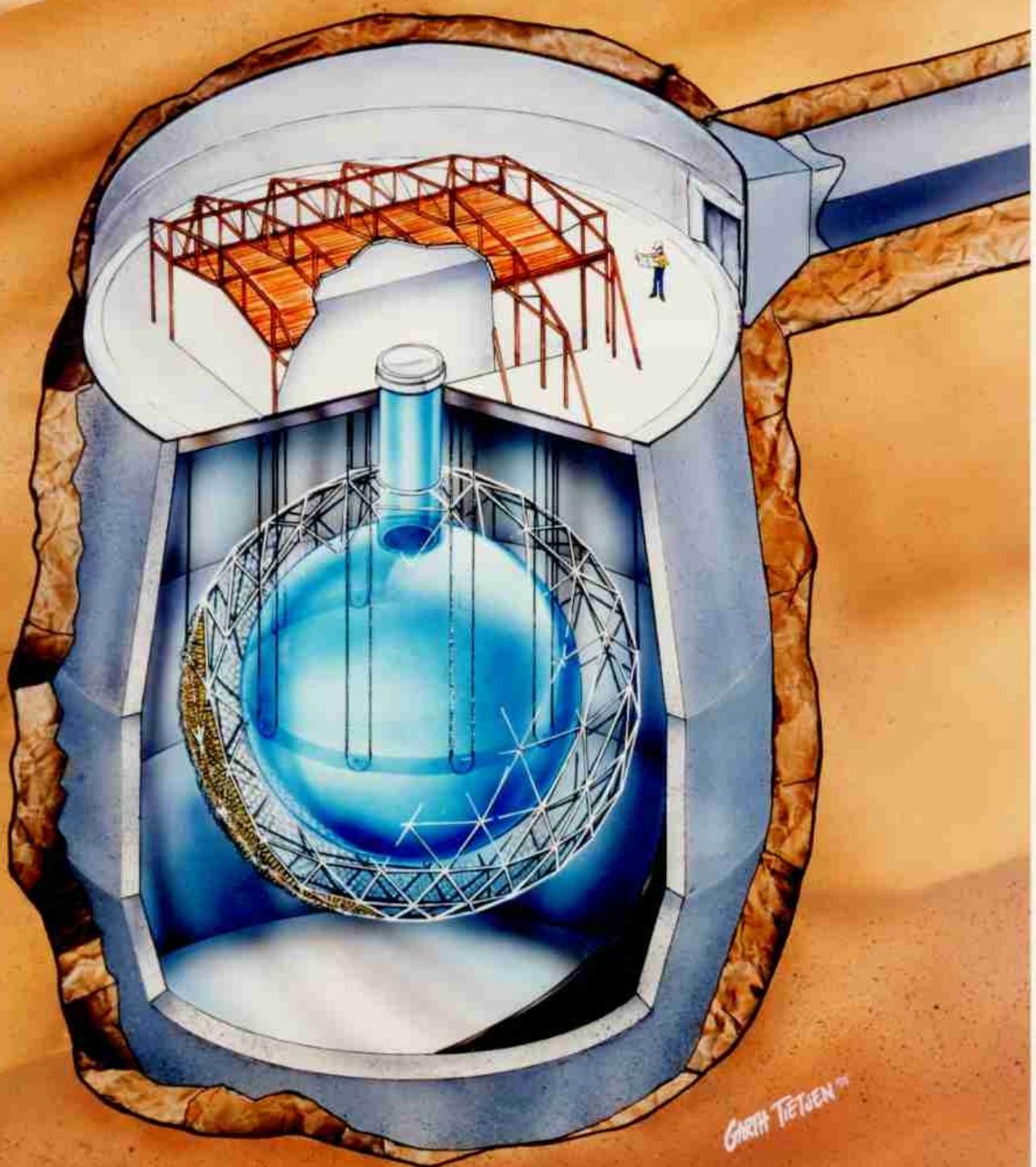
(With Threshold)

CC above 6 MeV and at SK meas.
Rate ~3300 events/year

Figure 2

INCO - Creighton Mine

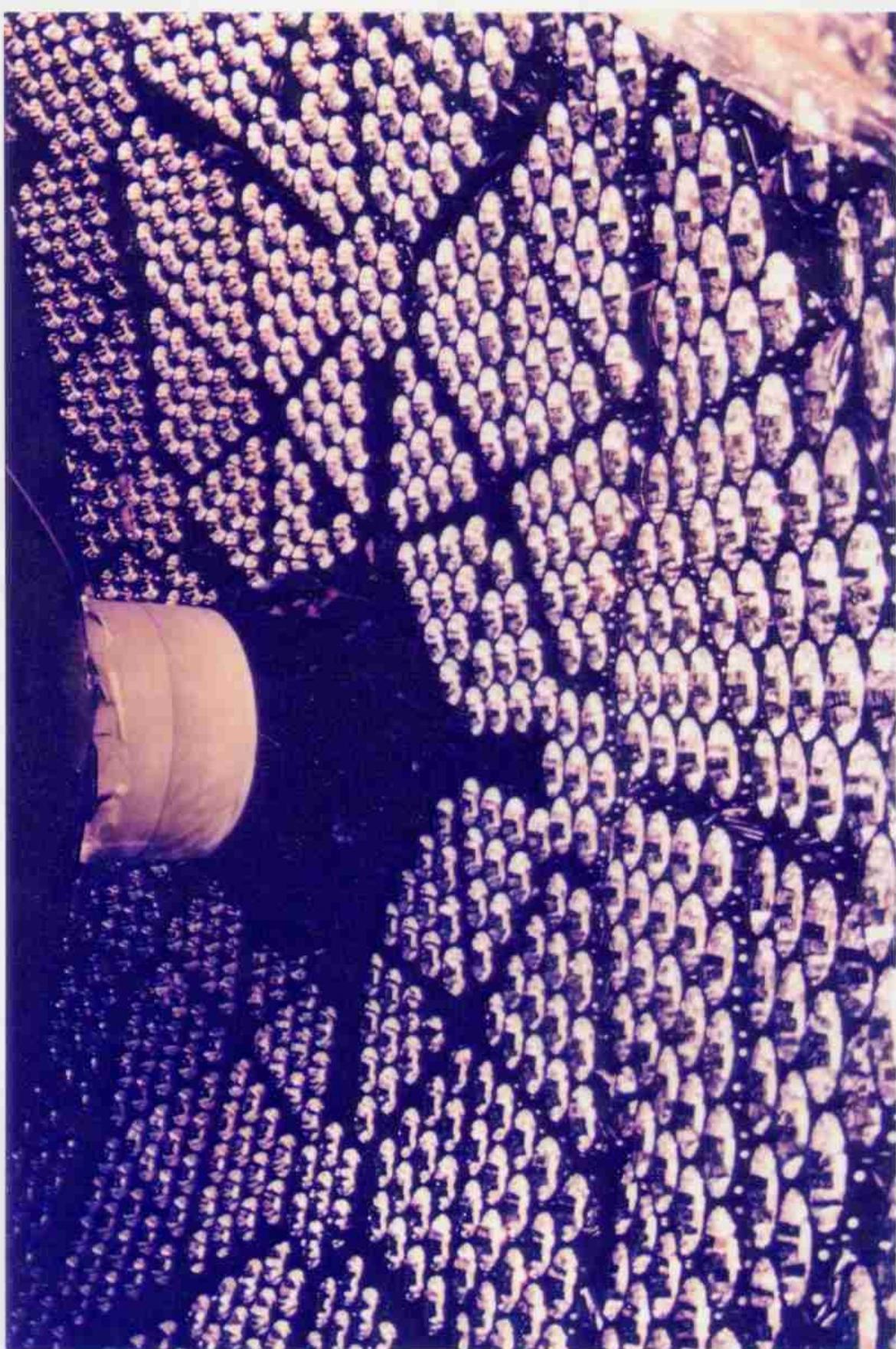


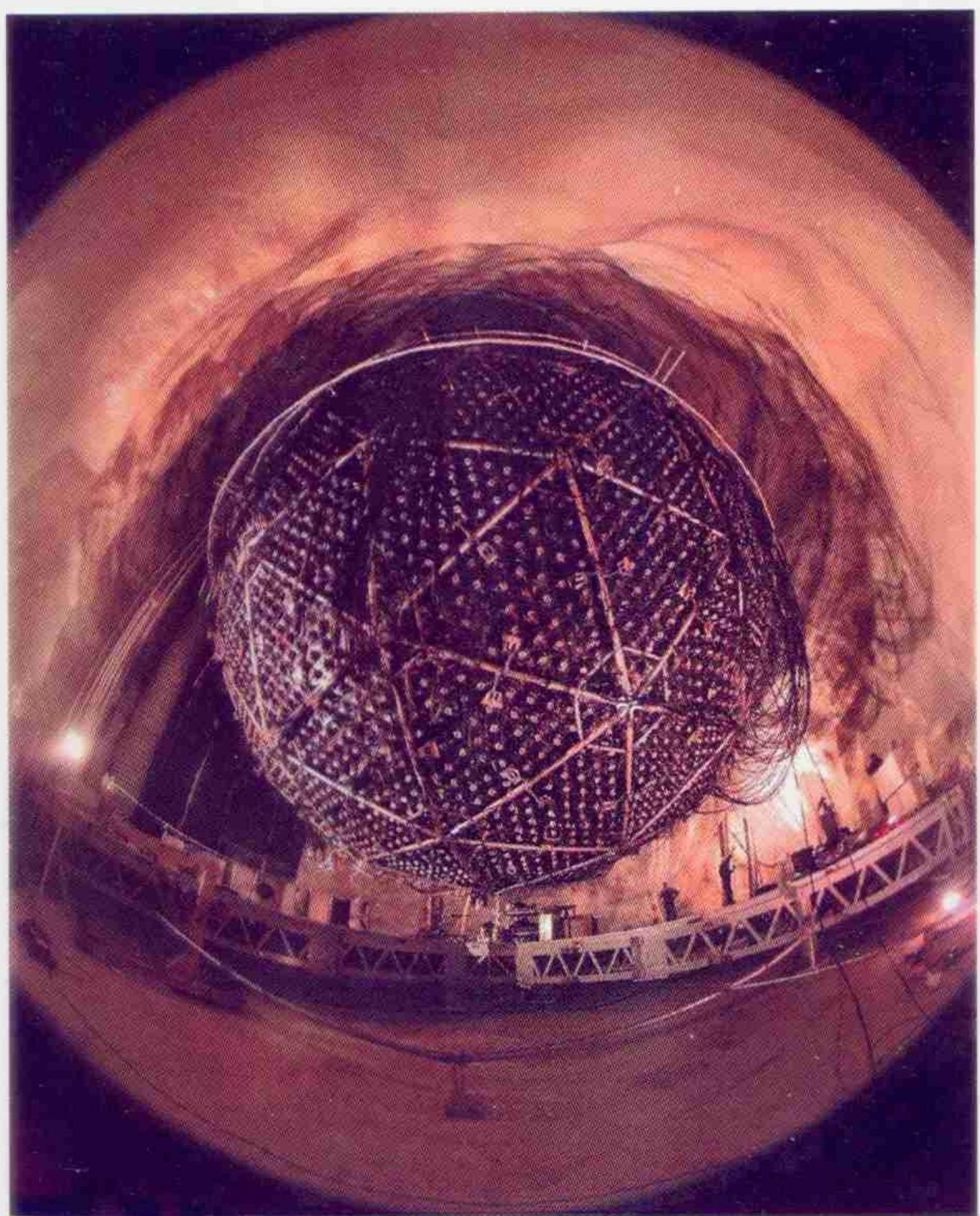


GERTH TIELEMANS

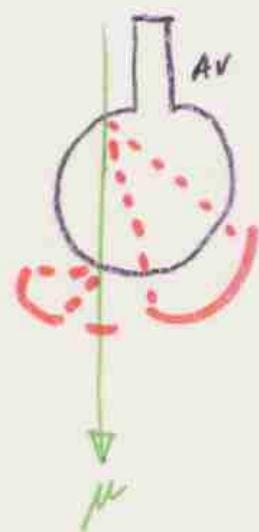
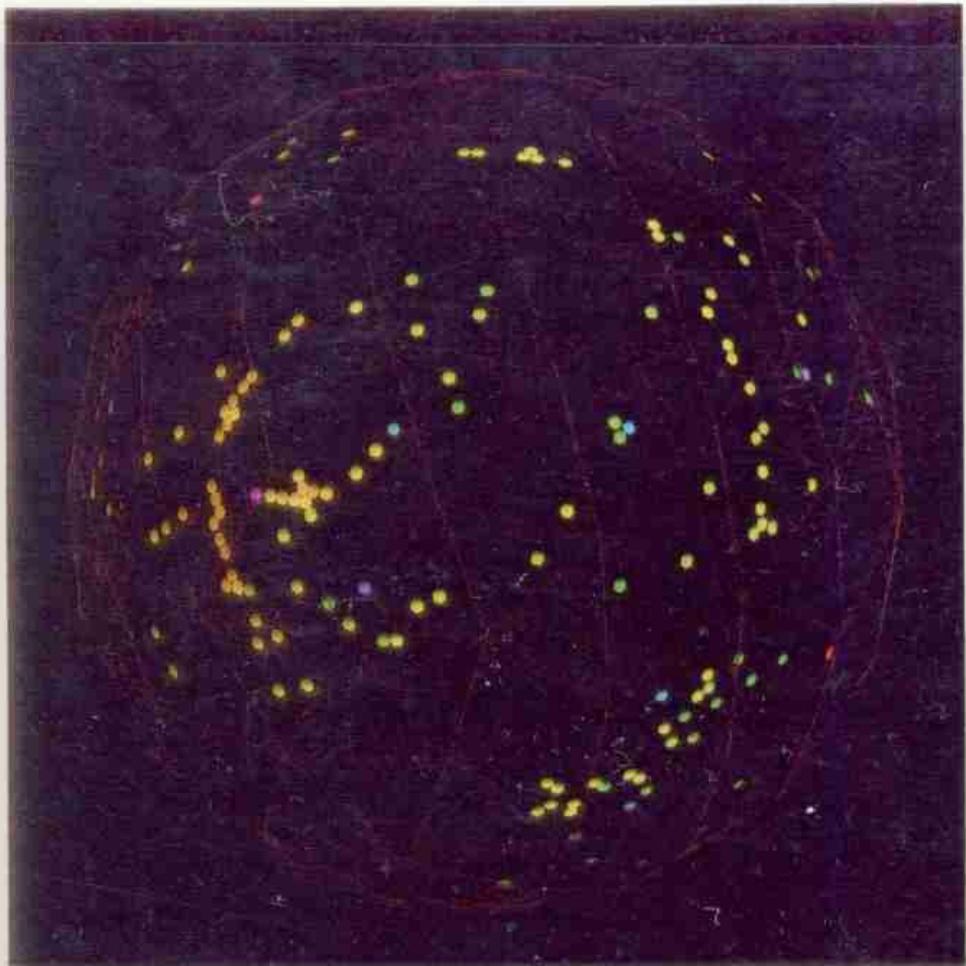




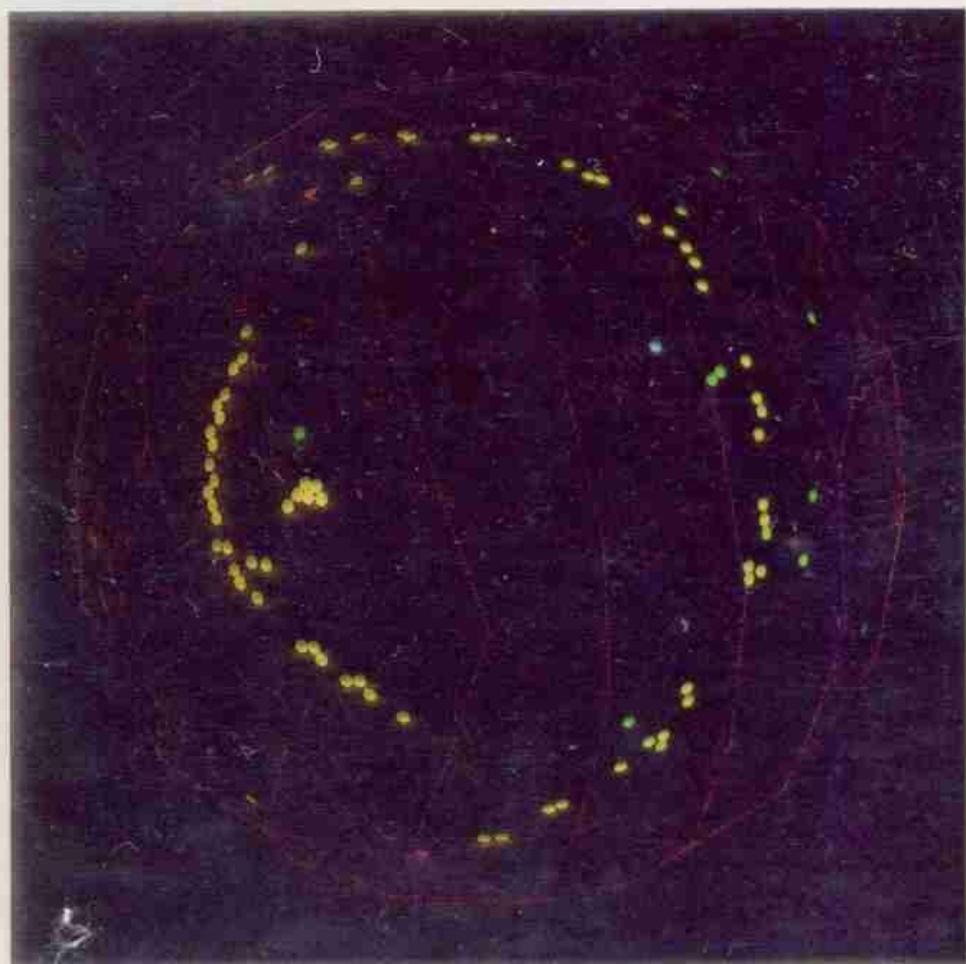








REAL



SIMULATED

PARTIAL
CERENKOV
CONES:
CUT OFF
BY TOTAL
INTERNAL
REFLECTION
IN ACRYLIC

Neutral Current Detection at SNO

Use two techniques with different systematic uncertainties

* Mg Cl₂

advantage:

just add material to D₂O

disadvantage:

CC and NC signals are intermingled within PMT response

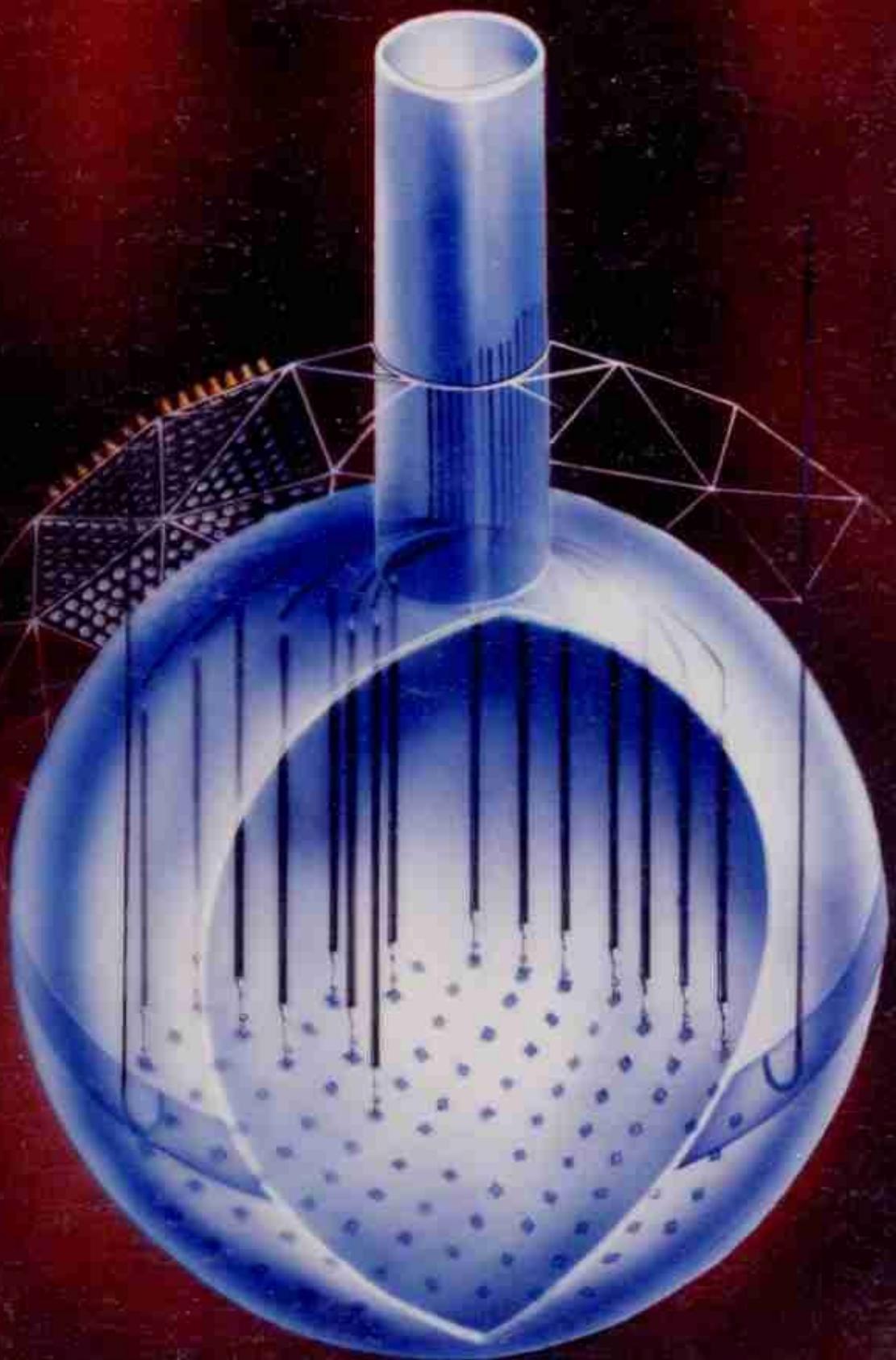
* ³He-filled proportional counters:

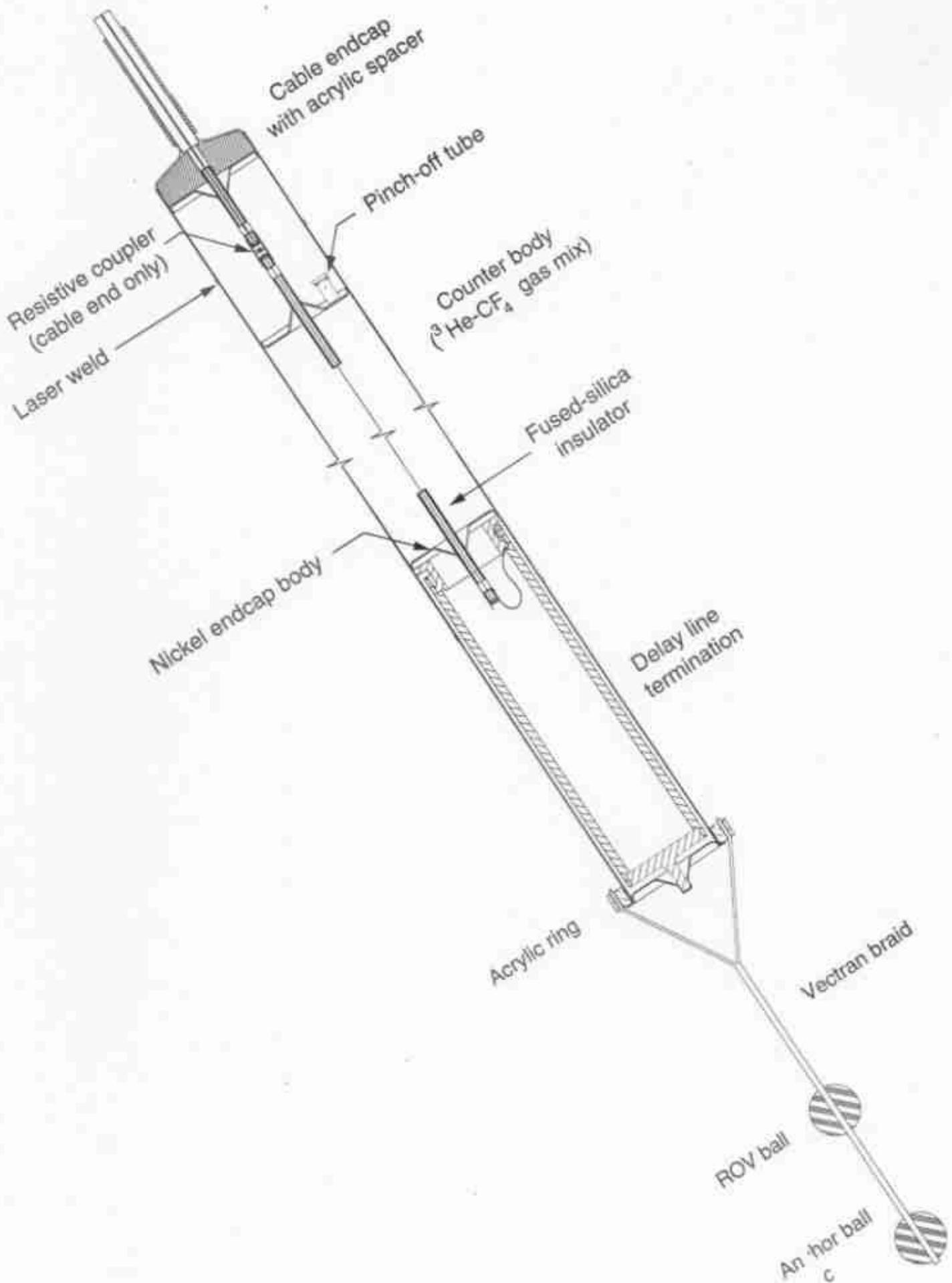
advantage:

CC and NC signals detected with separate systems. Hence they are distinguished on an event-by-event basis.

disadvantage:

complicated hardware





Background Issues at SNO

Any photon greater than 2.2 MeV can breakup the deuteron. This is a background for the NC detection.

For Example: 0.53 micrograms of Th distributed evenly through the heavy water would produce 1% SSM equivalent in neutrons: ~50/year

Th is usually present in most materials at some level:

Material	Mass for 1%	
CVD Ni	450 kg	Very Clean
Nylon	600 g	Rather Dirty
Polishing Compound	~10 Sand Grains	Filthy

Background Issues at SNO

Any photon greater than 2.2 MeV can breakup the deuteron. This is a background for the NC detection.

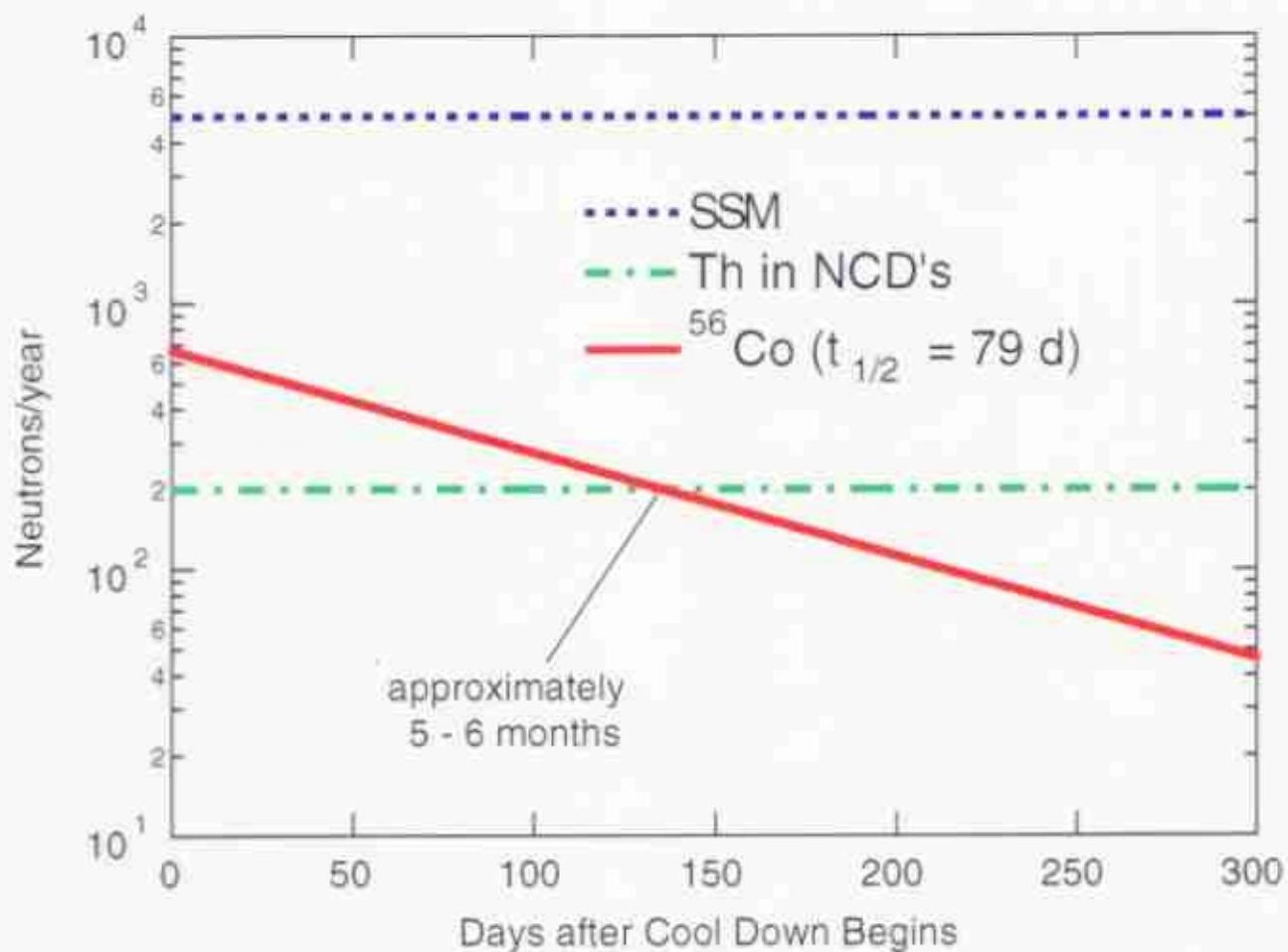
For Example: 1.3 micrograms of Th distributed evenly through the heavy water would produce 1% SSM equivalent in neutrons: ~50/year.

Th is usually present in most materials at some level. Some examples:

<u>Material</u>	<u>mass for 1%</u>	
CVD Ni	450 kg	it is clean
Nylon	600 g	not so clean
Polishing compound	~10 sand size grains	Filthy

- double purification of CVD Ni.

Photodisintegration Neutrons from Cosmogenic ^{56}Co



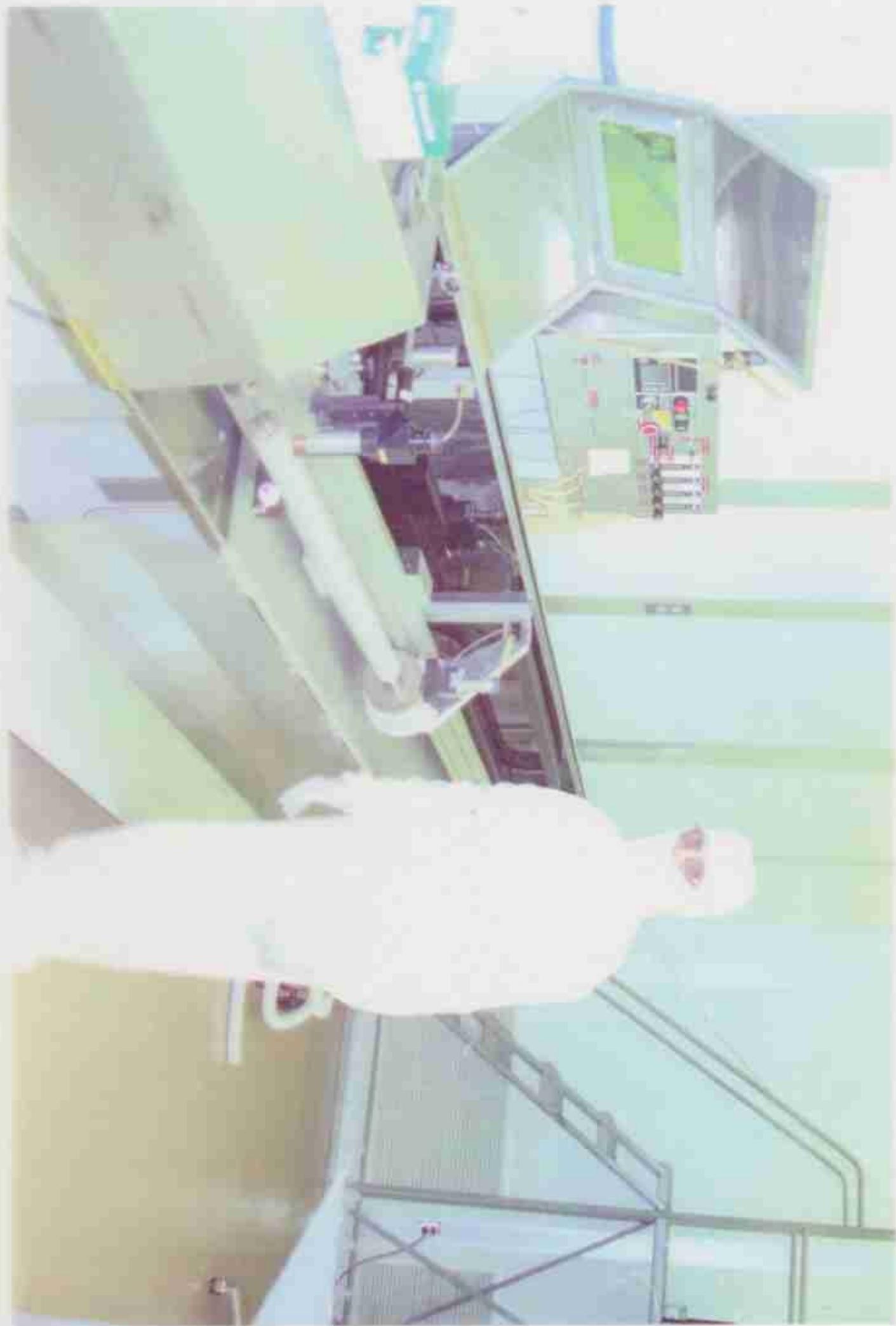
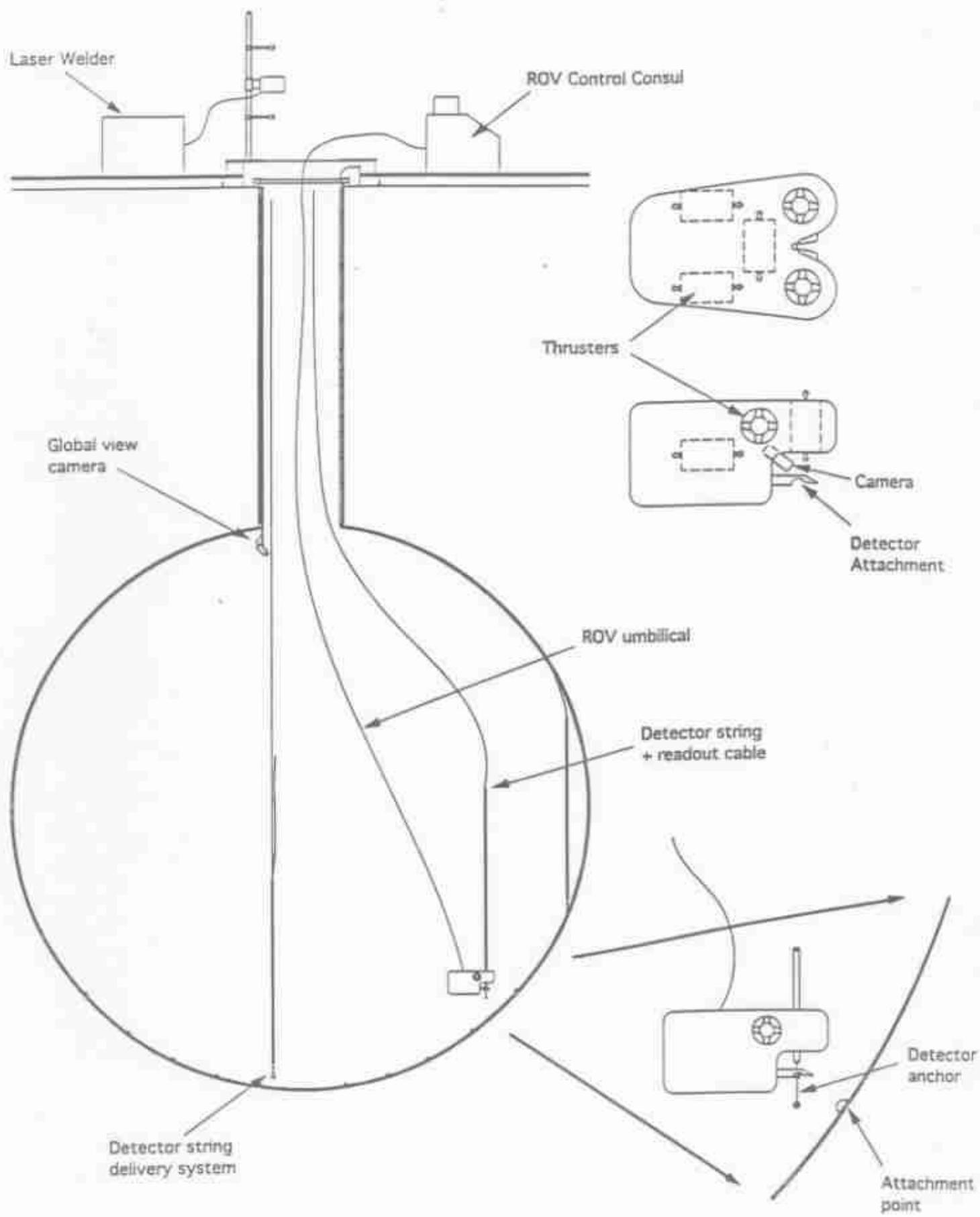
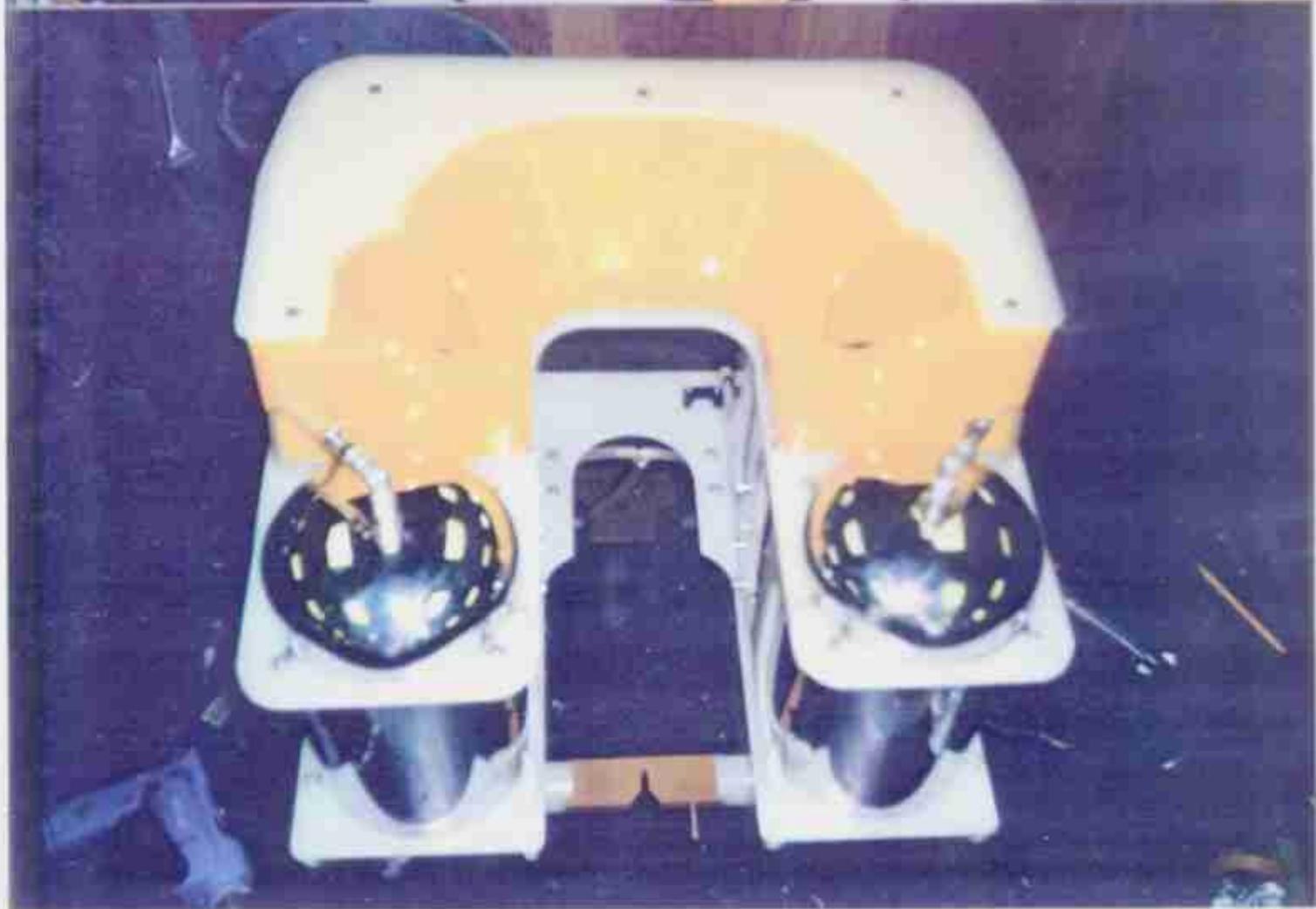




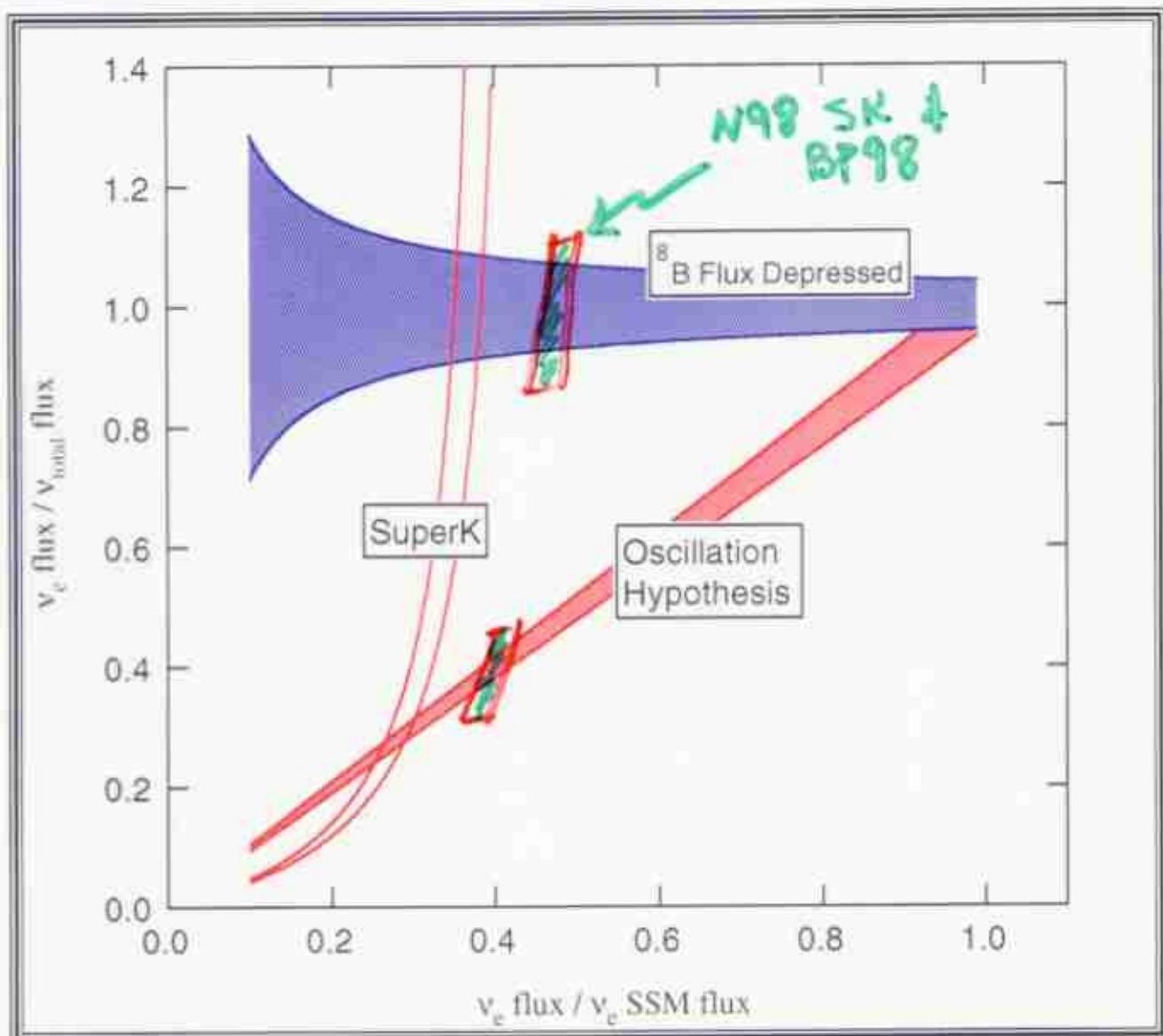
FIGURE 5. SCHEMATIC OF ROV DEPLOYMENT HARDWARE







CC/NC Ratio 1 Year



photodisintegration neutrons: 900/year

CC efficiency: 0.61

NC efficiency: 0.22

SuperKamiokande result from
Preprint Dated April 14, 1998