

BOREXINO: THE DETECTOR
AND
THE PHYSICS PROGRAM

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TOKYO, JUNE 12, 1998

NEW ERA IN NEUTRINO PHYSICS

- DESCRIPTION OF THE DETECTOR
- PHYSICS PROGRAM
- CTF RESULTS

BOREXINO COLLABORATION

GERMANY:

MAX-PLANCK-INSTITUT FÜR KERNPHYSIK HEIDELBERG
TECHNISCHE UNIVERSITÄT MÜNCHEN

HUNGARY:

KFKI-RMKI BUDAPEST

ITALY:

UNIVERSITA' DI GENOVA
LAB. NAZ. DEL GRAN SASSO
UNIVERSITA' DI MILANO
UNIVERSITA' DI PAVIA
UNIVERSITA' DI PERUGIA

POLAND:

JAGELLONIAN UNIVERSITY KRAKOW

RUSSIA:

JINR DUBNA

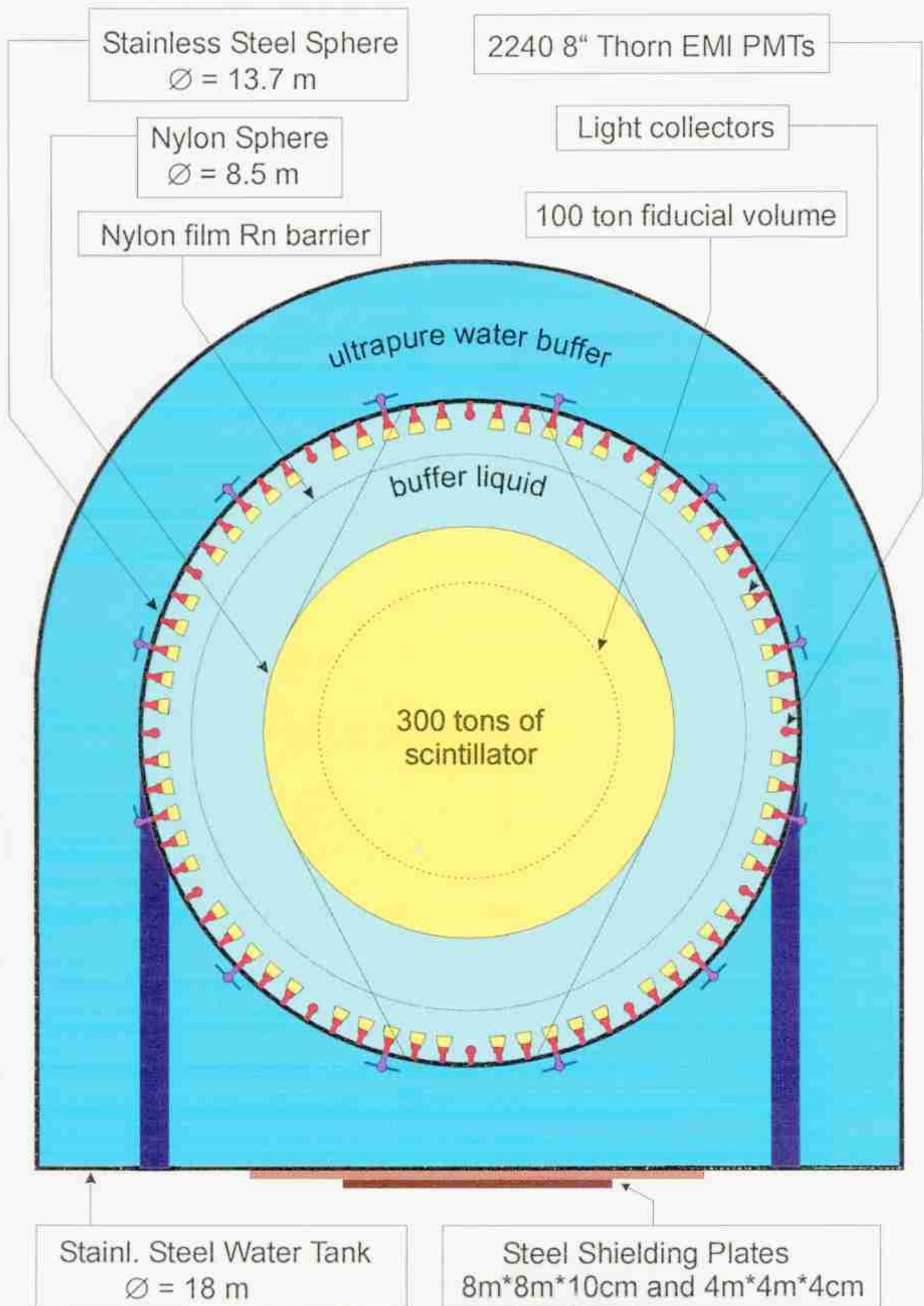
USA:

AT&T BELL LABORATORIES
MASSACHUSETTS INSTITUTE OF TECHNOLOGY
PRINCETON UNIVERSITY

BOREXINO

A REAL TIME, CALORIMETRIC,
SCINTILLATION DETECTOR FOR
LOW ENERGY SOLAR NEUTRINOS,
TO BE INSTALLED AT THE GRAN SASSO
UNDERGROUND LABORATORY, AIMED
AT THE DETECTION OF THE ${}^7\text{Be}$
NEUTRINO LINE.

Borexino Detector Design



SCINTILLATOR

SOLVENT : PSEUDOCUMENE

SOLUTE : PP0 (1.5 g/l)

LIGHT YIELD : $\sim 11,000$ PH / MeV

ATT. LENGTH (@ 420 nm) : ~~30~~ m

SCATT. LENGTH (@ 420 nm) : ~~7~~ m

DECAY TIME (FAST COMPONENT) : 3.5 ms

GOOD α/β DISCRIMINATION PROPERTIES

PHOTOMULTIPLIERS

8" THORN EMI 9351

P/V = 2.5

TRANSIT TIME SPREAD 1 MS

DARK COUNT RATE 1 KHZ

AFTERPULSING < 3%

LOW RADIOACTIVITY GLASS AND INTERNAL PARTS

LIGHT CONCENTRATORS

TRUNCATED STRING CONE DESIGN

OPTIMIZED TO COLLECT THE LIGHT FROM THE INNER VESSEL AND 20 CM BEYOND IT

DETECTOR PERFORMANCES

COVERAGE: 34 %

PHOTOELECTRON YIELD (~ 450 PE/MEV)

ENERGY RESOLUTION (@ 1 MeV) : $\sim 8\%$

SPATIAL RESOLUTION (@ 1 MeV) : 12 cm

CALIBRATIONS

A VARIETY OF CALIBRATION AND MONITORING SYSTEMS ARE ENVISAGED:

- LASER FEEDING FIBERS CONNECTED TO EACH TUBE
 - TIMING CALIBRATION
 - GAIN ADJUSTEMENT VIA DETECTION OF THE SINGLE PHOTOELECTRON PEAR
- EXTERNAL SOURCES (TH) LOCATED NEAR TO THE S.S.S.
 - CHECK OF THE STABILITY IN TIME OF THE OVERALL DETECTOR RESPONSE
-
- INTERNAL SOURCES (RN) INSIDE THE SCINTILLATOR
 - TIME (AND SPACE) CALIBRATION

- INTERNAL SOURCES (RN) IN THE BUFFER

- STUDY OF SCINTILLATION IN THE BUFFER

- LASER BEAM THROUGH THE BUFFER AND SCINTILLATOR

- "ON SITE" VERIFICATION OF OPTICAL PROPERTIES OF BUFFER AND SCINTILLATOR VIA PHOTOEXCITATION

- DIFFUSING BALL IN THE CENTER OF THE DETECTOR

- ALTERNATIVE MEAN OF PMT + ELECTRONIC TIMING CALIBRATION

- TAGGING OF INTERNAL BACKGROUNDS VIA CORRELATED EVENTS

- ABSOLUTE ENERGY SCALE DETERMINATION

PLANTS

- PURIFICATION SYSTEM (SCINT.)

- WATER EXTRACTION

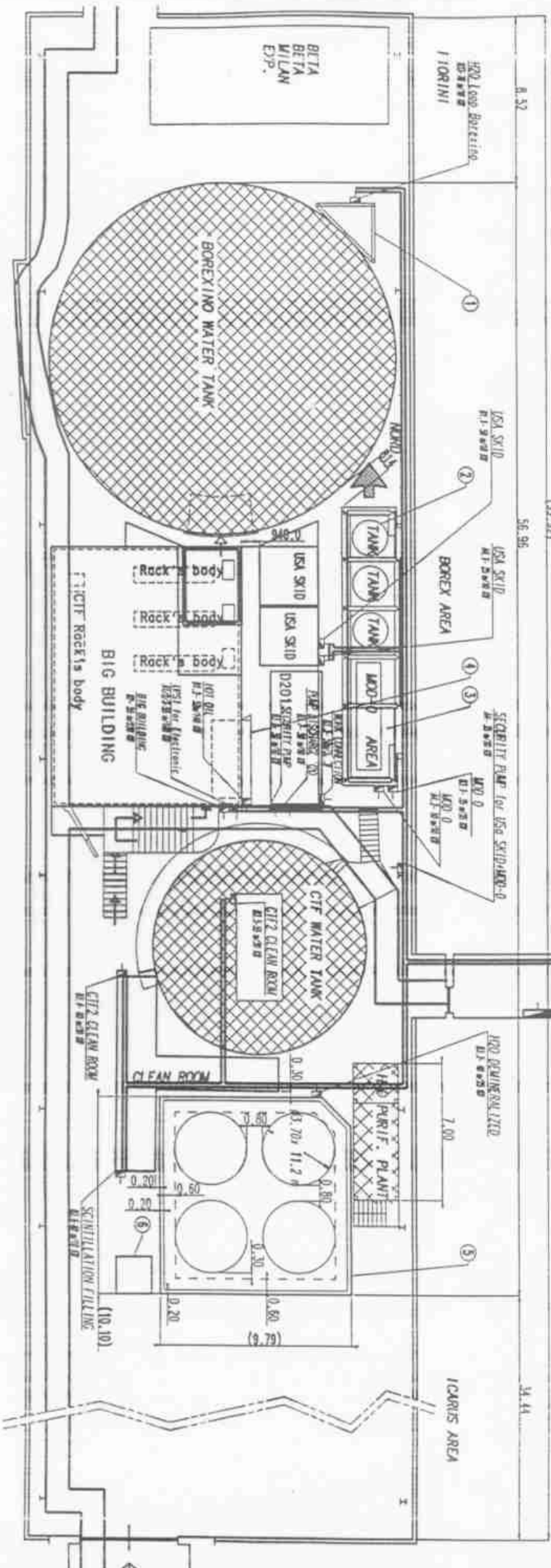
- DISTILLATION

- NITROGEN FLUSHING

- WATER PURIFICATION SYSTEM

- LIQUID HANDLING

- STORAGE VESSELS



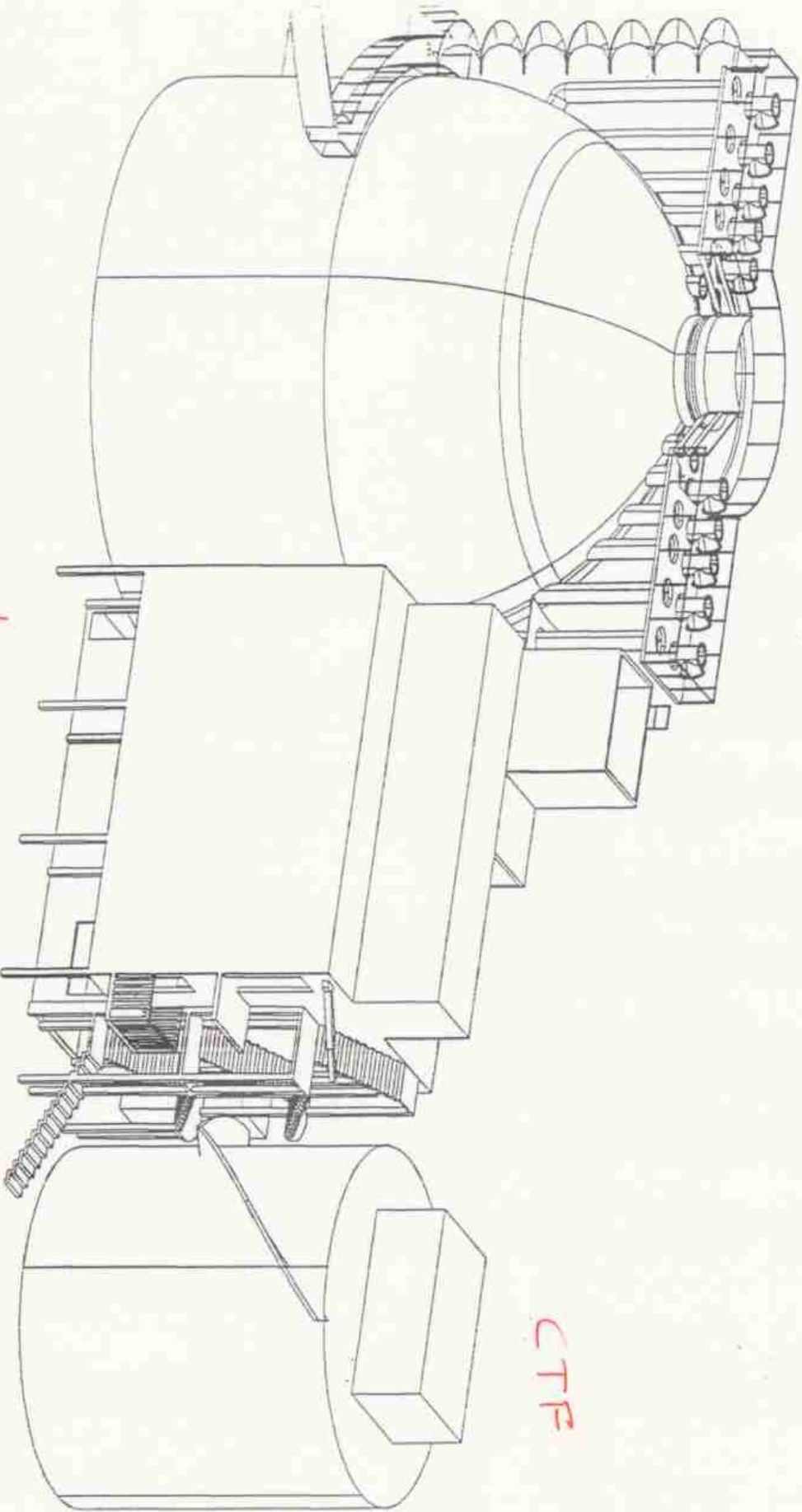
General Switchboard

IPS1

IPS2

Hall-C-Layout-14-05-98.dwg

BOREXINO



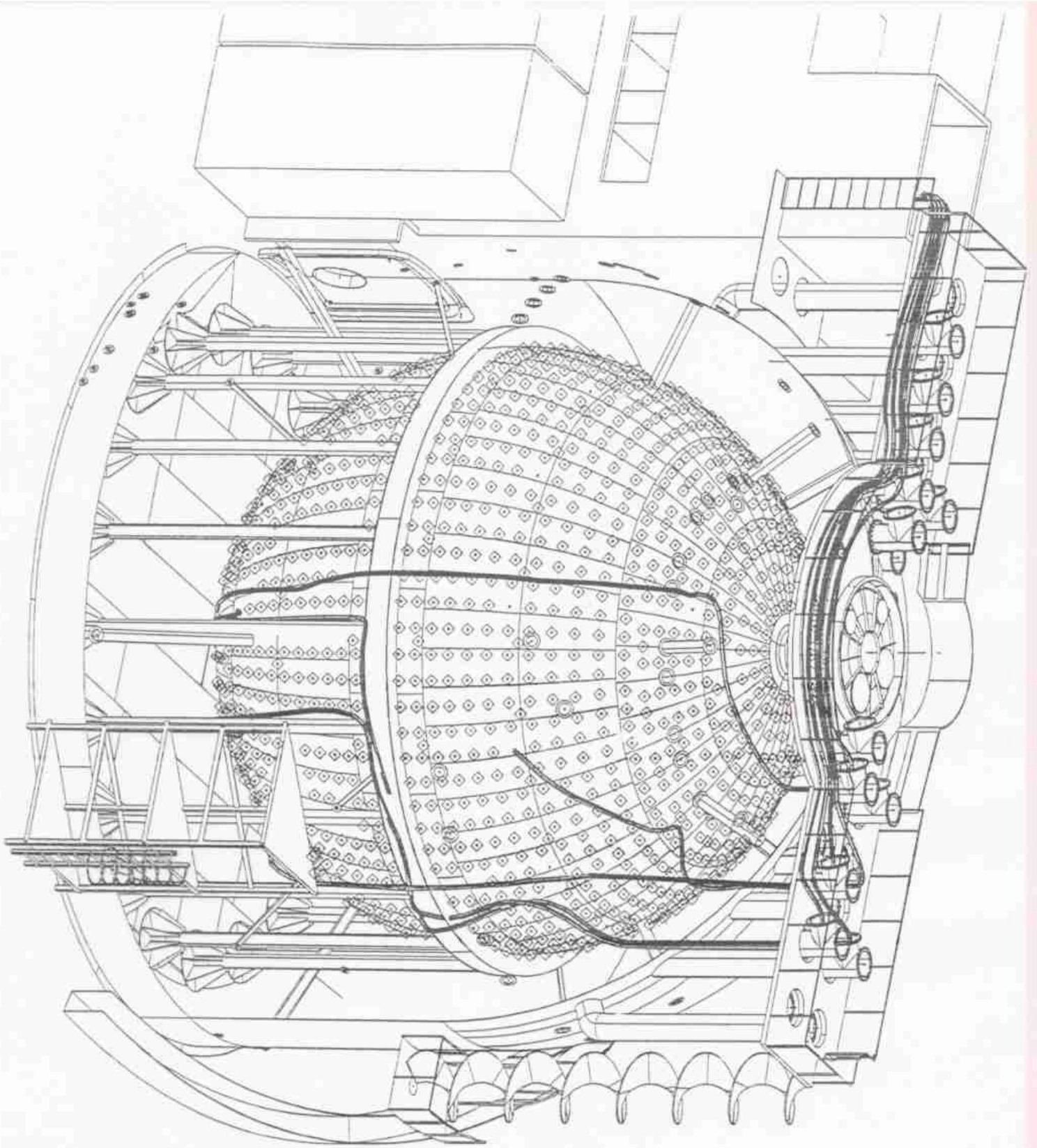
CTF

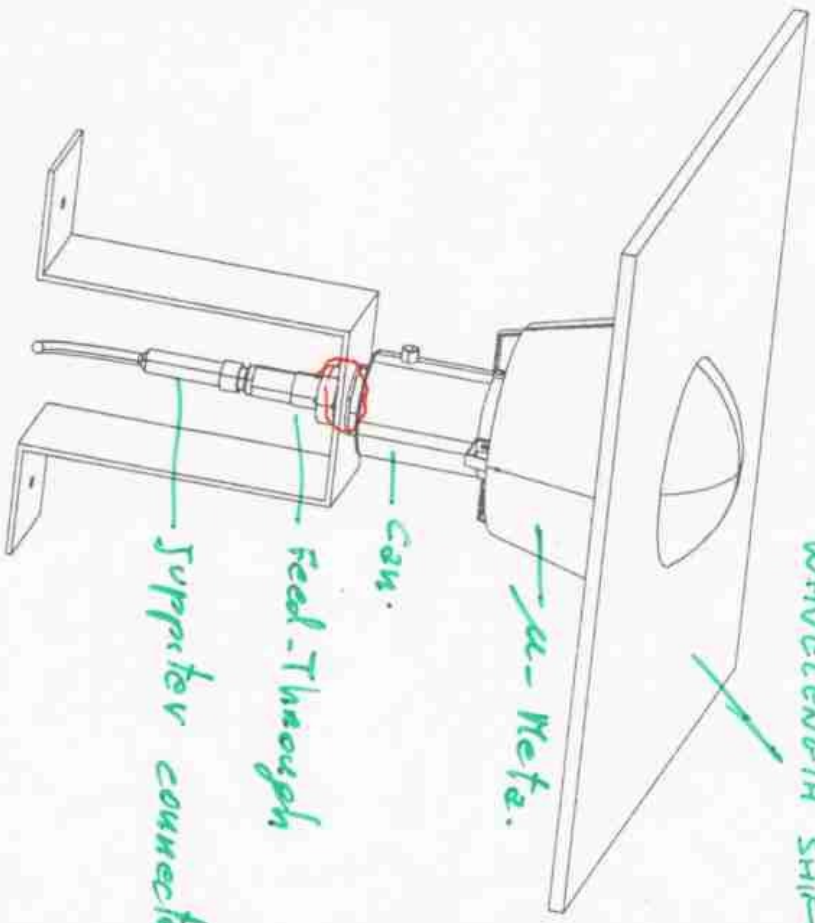
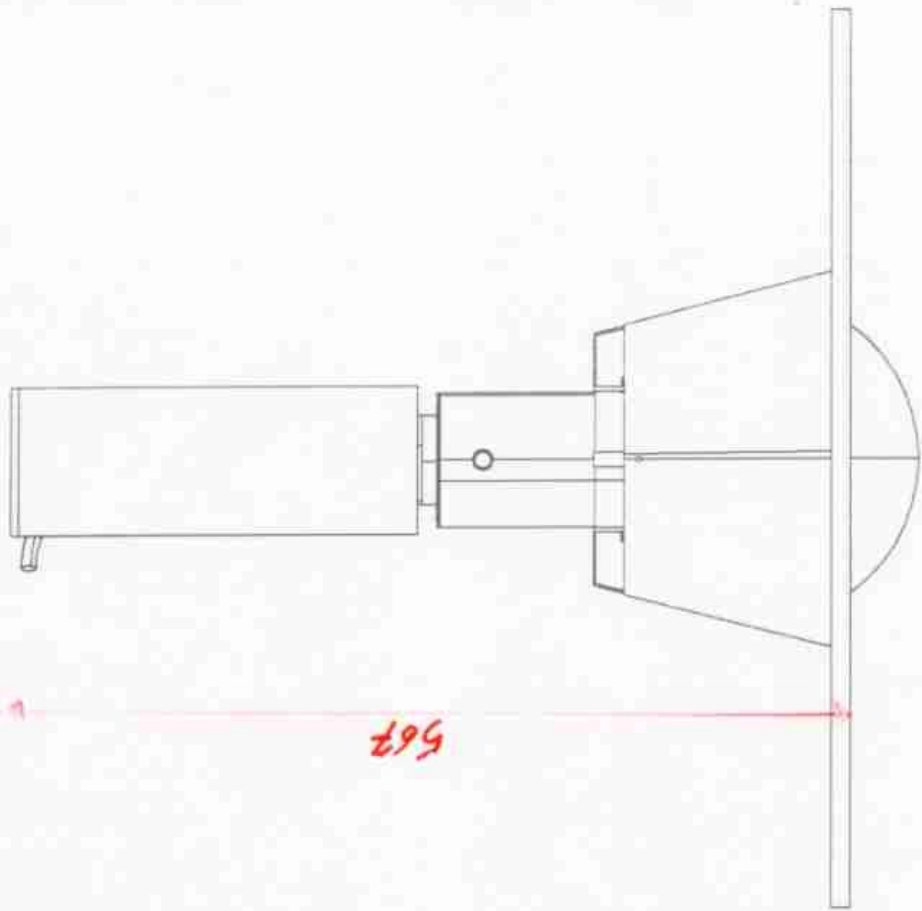
1st FLOOR : CLEAN ROOM

2nd FLOOR : CHEMISTRY AND R_N MEAS. LABS

3rd FLOOR : COUNTING ROOM







μ-Meta-PMT

ACRYLIC PLATE WITH WAVELENGTH SHIFTER

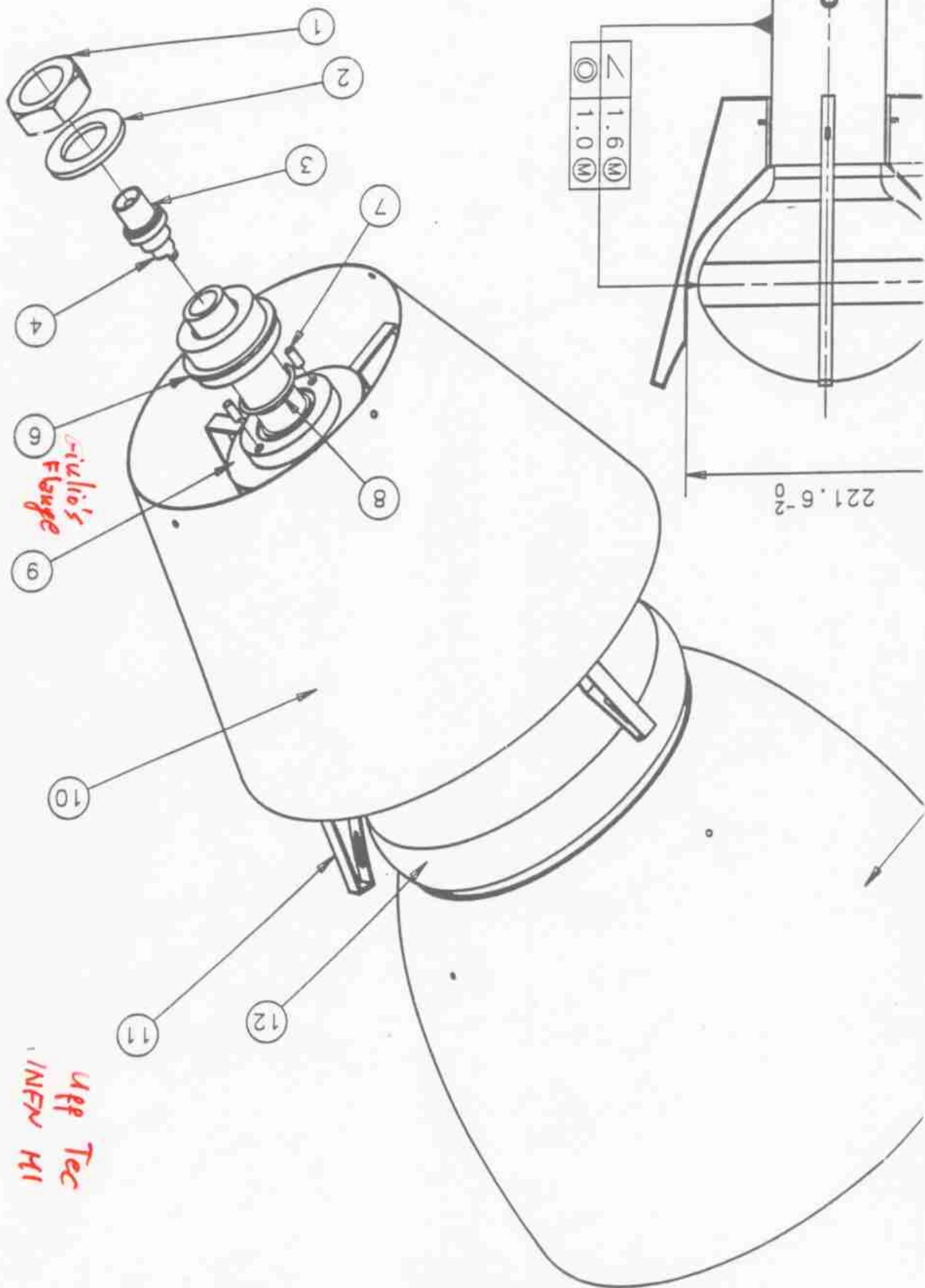
μ-Meta.

Can.

Feed-Through

Supporter connector

Utt Tec
INFN-MI

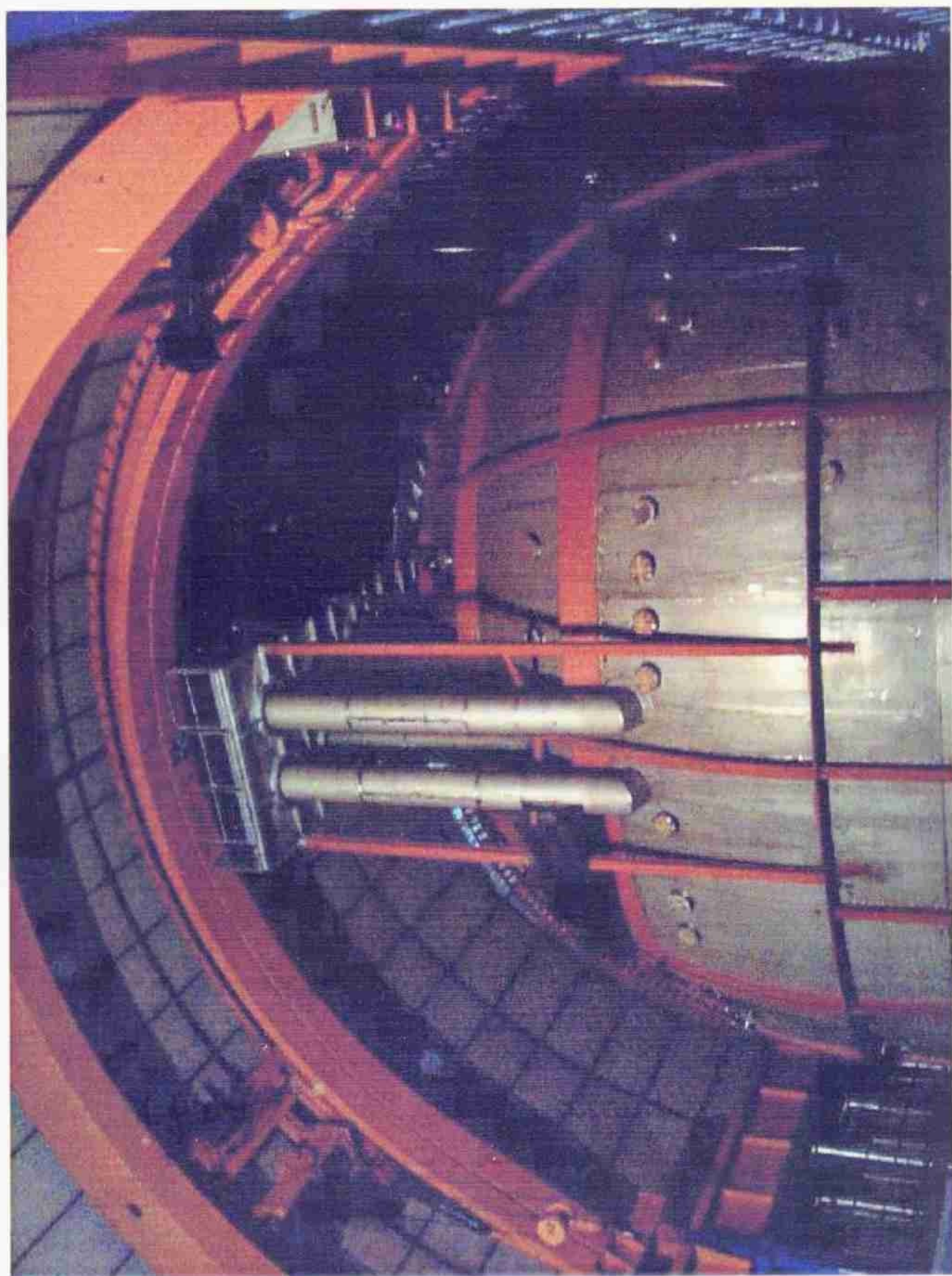


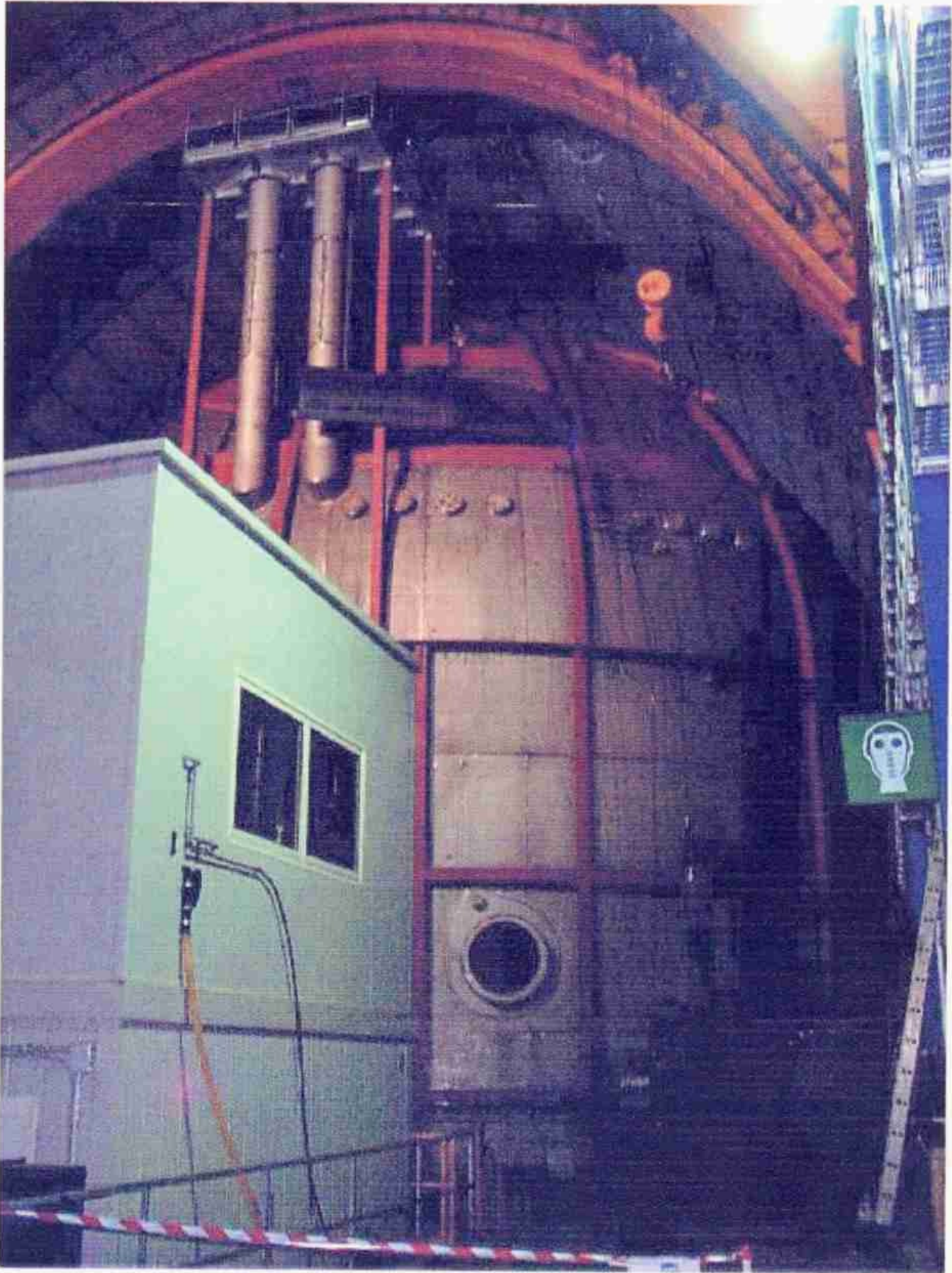
∠	1.6M
⊙	1.0M

221.6-2/0

UPP Tec
INFV M1







NEUTRINO DETECTION IN BOREXINO

DETECTION THROUGH THE
SCATTERING REACTION



OFF THE ELECTRONS OF THE SCINTILLATOR

- HIGH LUMINOSITY

AND

- HIGH PURITY

OF THE SCINTILLATOR LEAD TO A

LOW DETECTION THRESHOLD

250 KeV

⇒ IT IS POSSIBLE TO DETECT THE RECOIL

ELECTRONS PRODUCED BY THE

MONOENERGETIC (0.861 MeV) ${}^7\text{Be}$ ν .

MAXIMUM RECOIL ENERGY

0.66 MeV

S.S.M. PREDICTION ~ 55 ev/d FOR 100 T F.M.

Standard Solar Model Neutrino Fluxes and Energy Acceptances of the Experiments

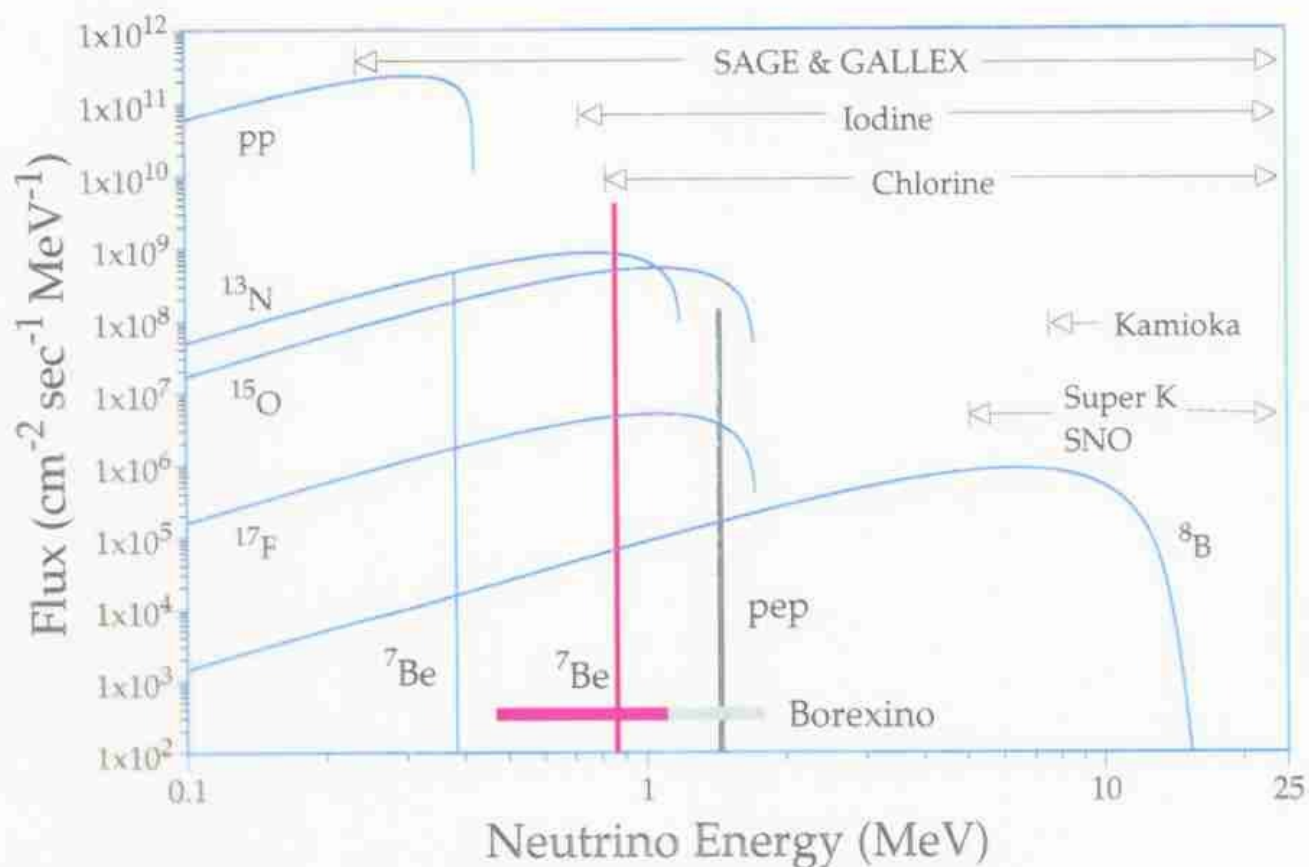
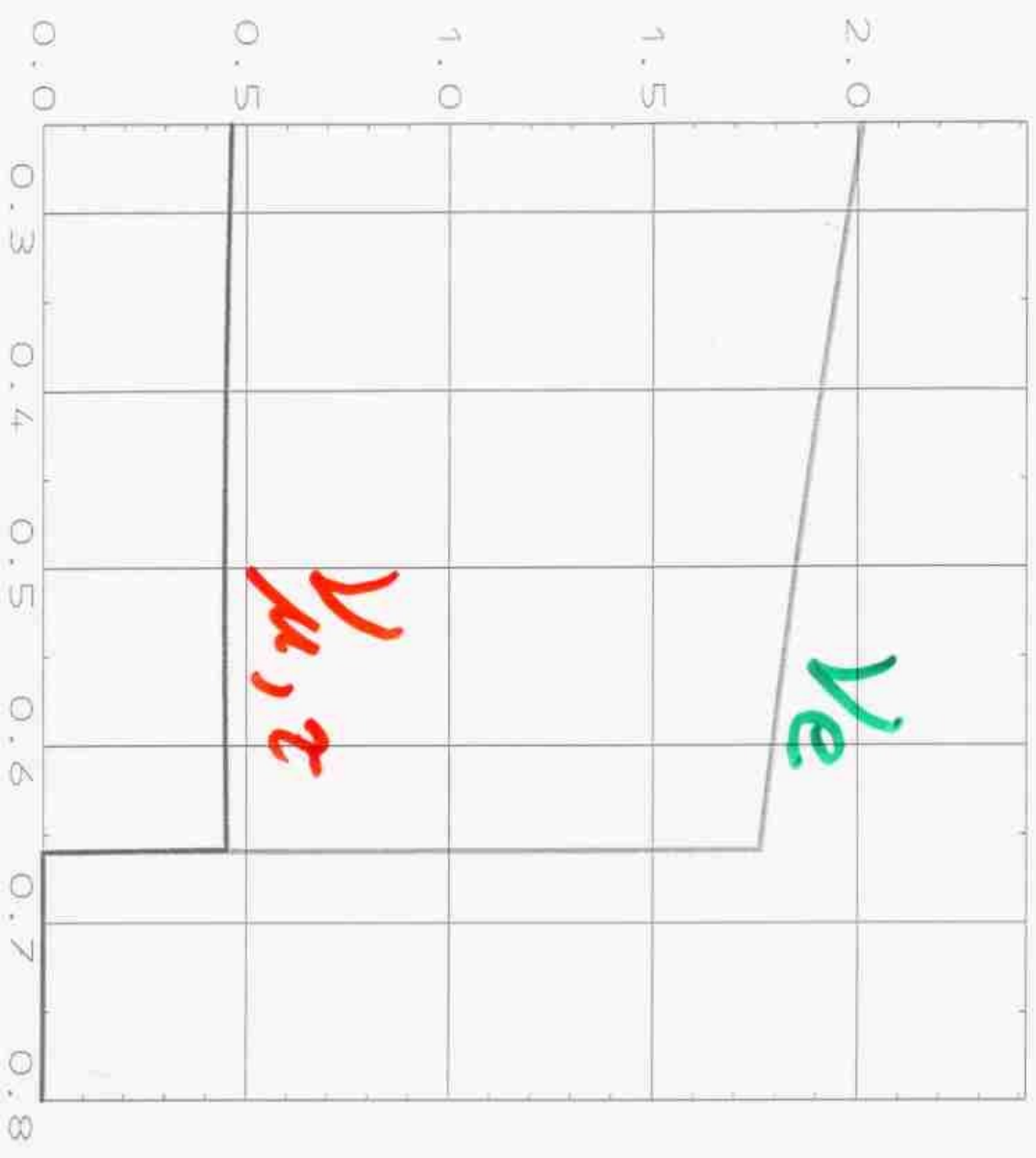


Figure 2: The energy spectrum of solar neutrinos produced by the *pp* and *CNO* cycles in the Standard Solar Model [9]. The monoenergetic line fluxes have units $\text{cm}^{-2}\text{s}^{-1}$.

Recoil energy spectrum ${}^7\text{Be} - \nu_0$



recoil energy / MeV

$$\nu_0 = \nu_e \quad (55M)$$

$$\sim \underline{\underline{55/day}}$$

$$\nu_0 = \nu_{\mu, \tau} \quad (MSW)$$

$$\sim \underline{\underline{15/day}}$$

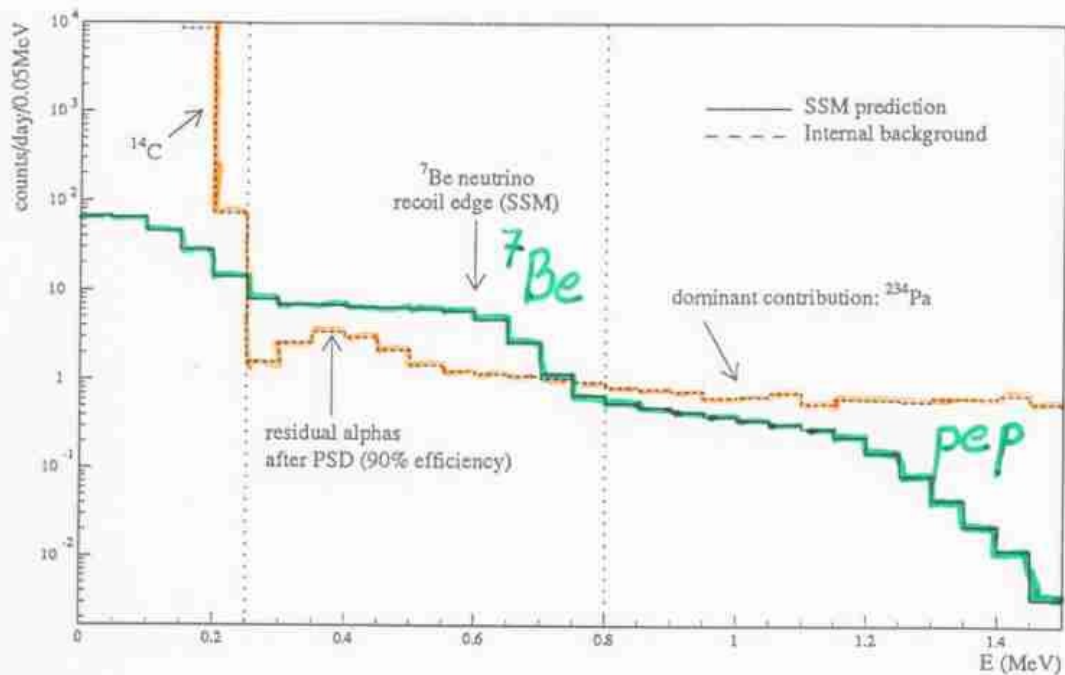


Figure 15: Estimate of internal background spectrum and SSM neutrino signal in a 100-ton fiducial volume, assuming 300 photoelectrons/MeV and an energy resolution given by: $\Delta E/E = 1/\sqrt{N_{pe}}$.

Energy spectrum ν -events
and internal bg.

Requirements:

$$^{238}\text{U} : 10^{-16} \text{ g/g}$$

$$^{232}\text{Th} : 10^{-16} \text{ g/g}$$

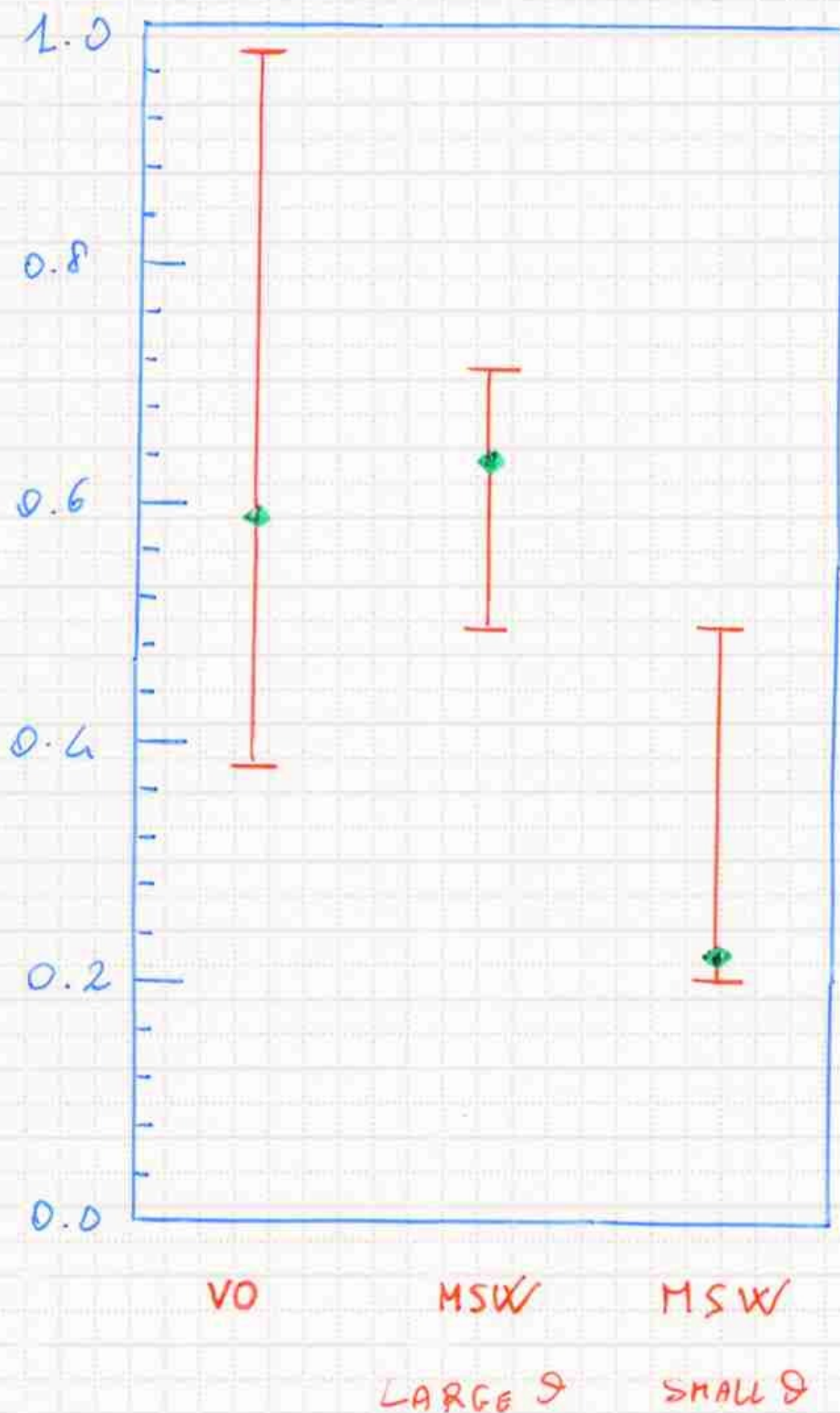
$$^{14}\text{C}/^{12}\text{C} : 10^{-18}$$

$$K_{\text{nat}} : 10^{-14} \text{ g/g}$$

BOREXINO CAPABILITIES

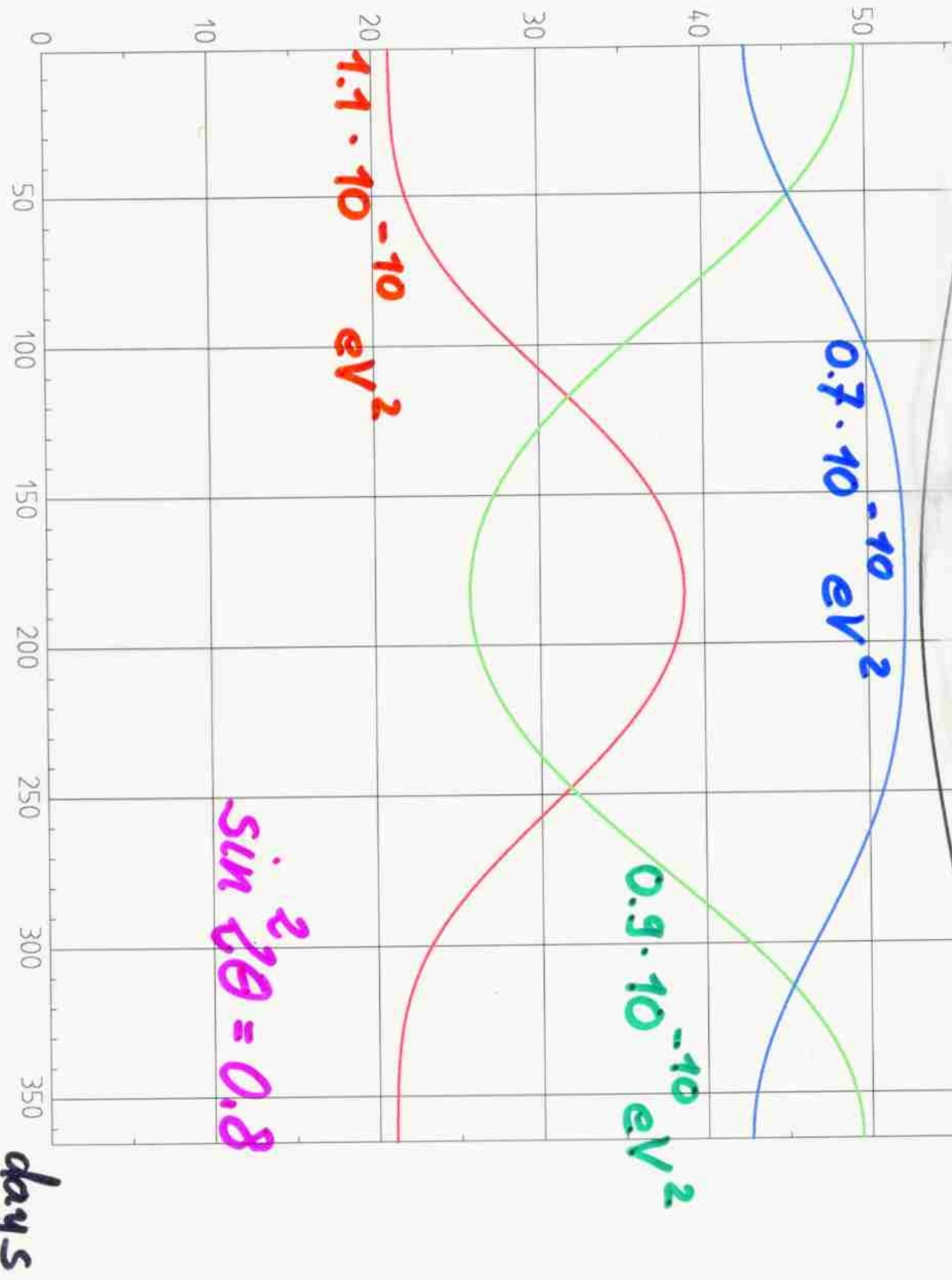
- FIRST DIRECT MEASUREMENT OF γ_{B_e} FLUX
- PROBING VACUUM OSCILLATIONS VIA SEASONAL VARIATION OF THE FLUX
- IN ABSENCE OF OTHER TIME VARIATIONS DEMONSTRATION OF THE SOLAR ORIGIN OF THE SIGNAL THROUGH THE $\gamma\%$ VARIATION DUE TO THE EARTH-SUN DISTANCE VARIATION DURING THE YEAR

PREDICTED FLUXES AT 90% CL IN THE THREE ALLOWED REGIONS



Vacuum Oscillations ${}^7\text{Be} - \nu_0$

counts
day



days

RADIOPURITY OF THE SCINTILLATOR

- MAIN ISSUE FOR THE FEASIBILITY OF THE EXPERIMENT
- PURITY REQUIREMENT FOR ^{238}U AND ^{232}Th $\sim 10^{-16}$ g/g
- LABORATORY MEASUREMENTS ON 1L SAMPLES $\sim 2 \div 3 \times 10^{-15}$ g/g

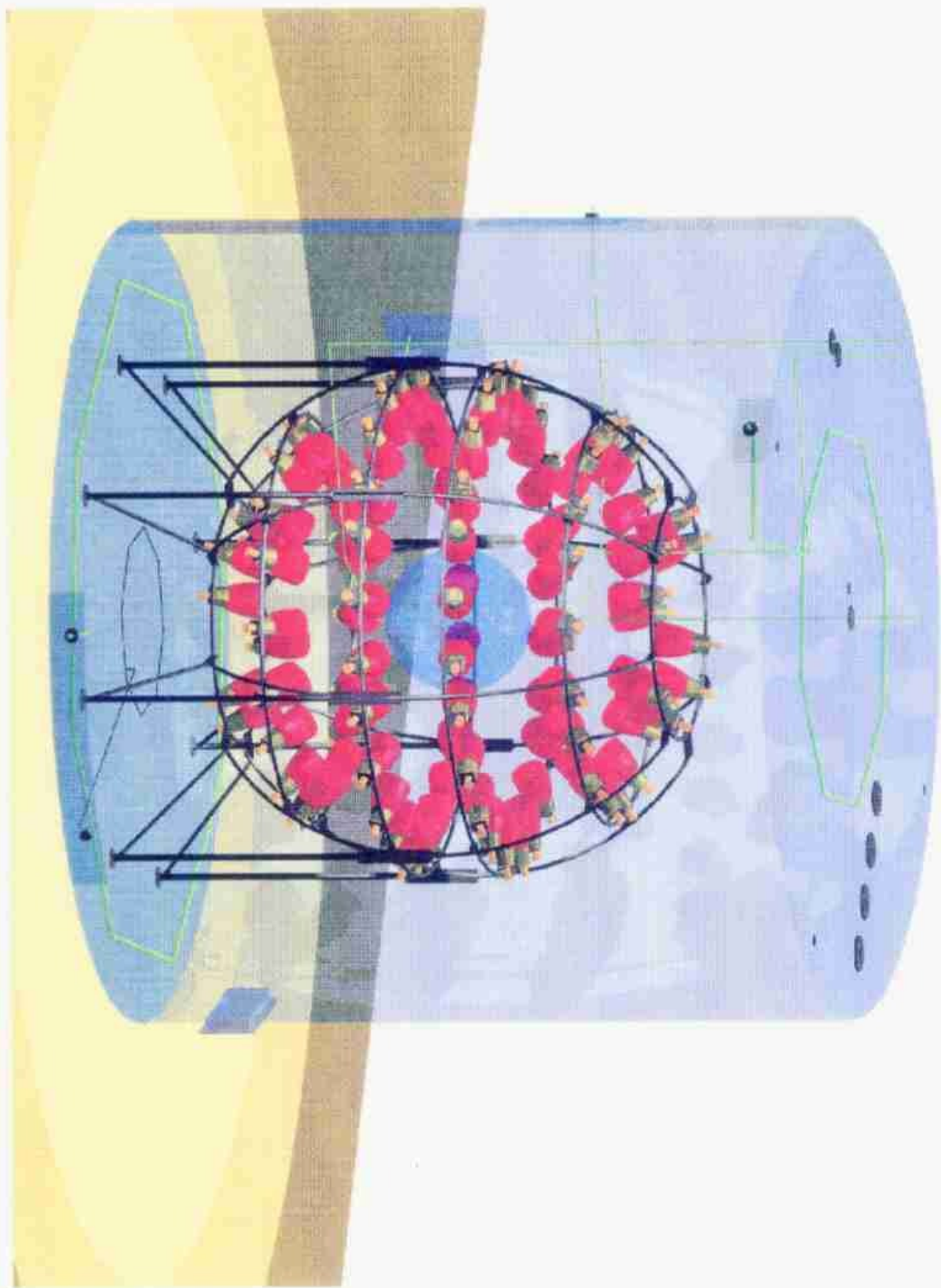
MAINLY LIMITED BY IMPURITIES LEACHED IN FROM THE WALL OF THE VESSELS

- NEEDED A DIRECT MEASUREMENT ON SOME TONS OF SCINTILLATOR WITH A SENSITIVITY LEVEL OF AT LEAST 5×10^{-16} g/g



CTF

ООО "Техстек" ИНН 77-07-0018277

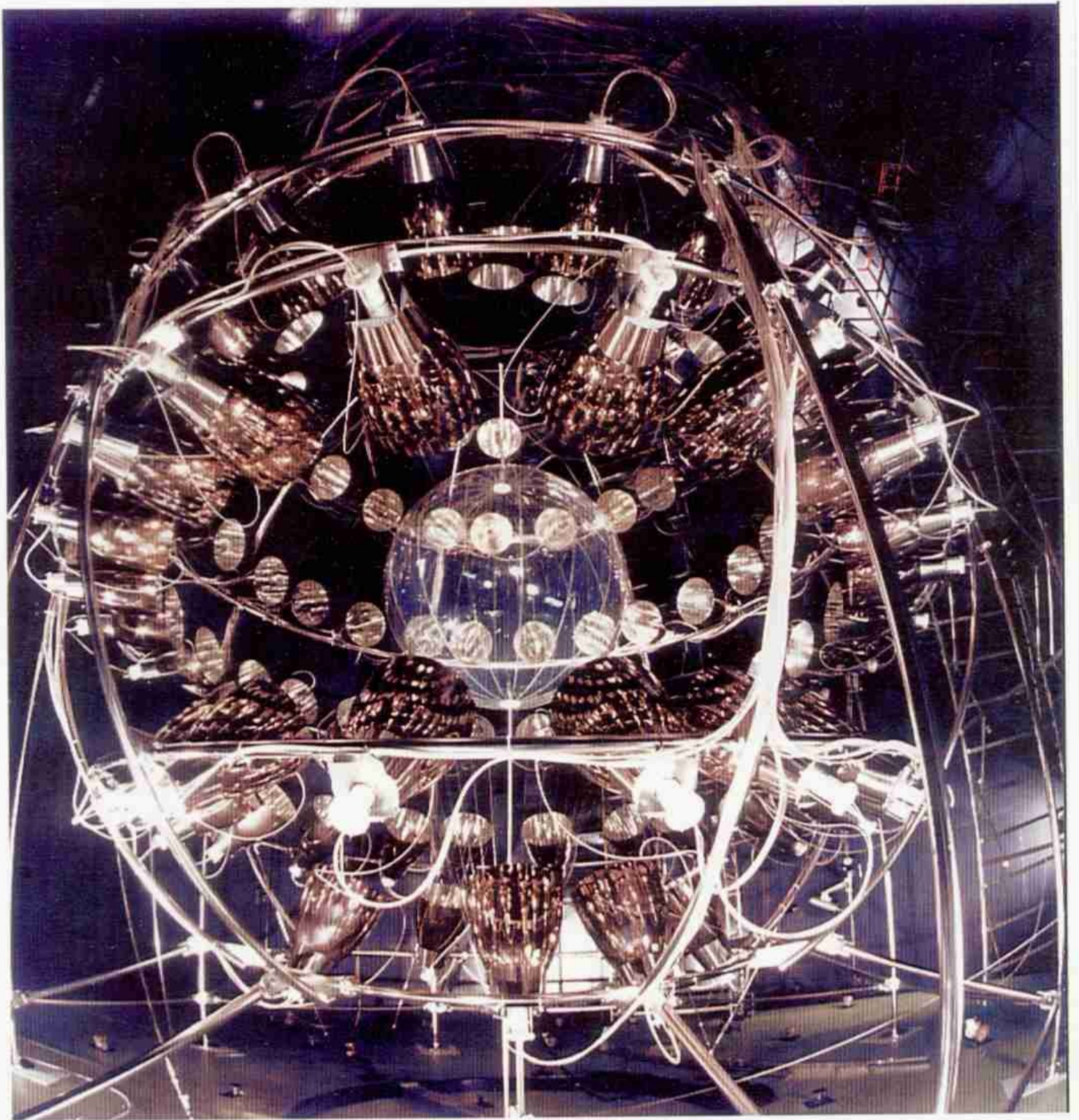


Feed This Side Up

Feed This Side Up

3M

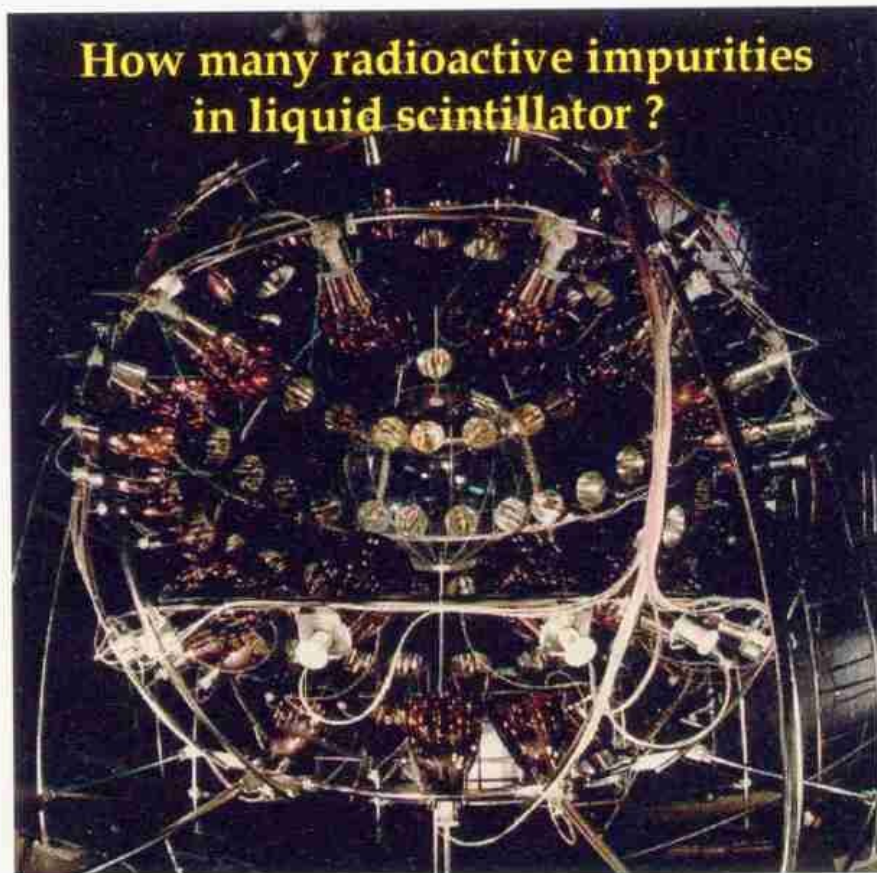
Counting Test Facility



of **Borexino** at Gran Sasso

Results of the CTF

(Counting Test Facility for Borexino)



$$^{14}\text{C}/^{12}\text{C} = (1.94 \pm 0.09) * 10^{-18}$$

$$^{232}\text{Th} < (4.4 \pm 1.5) * 10^{-16}$$

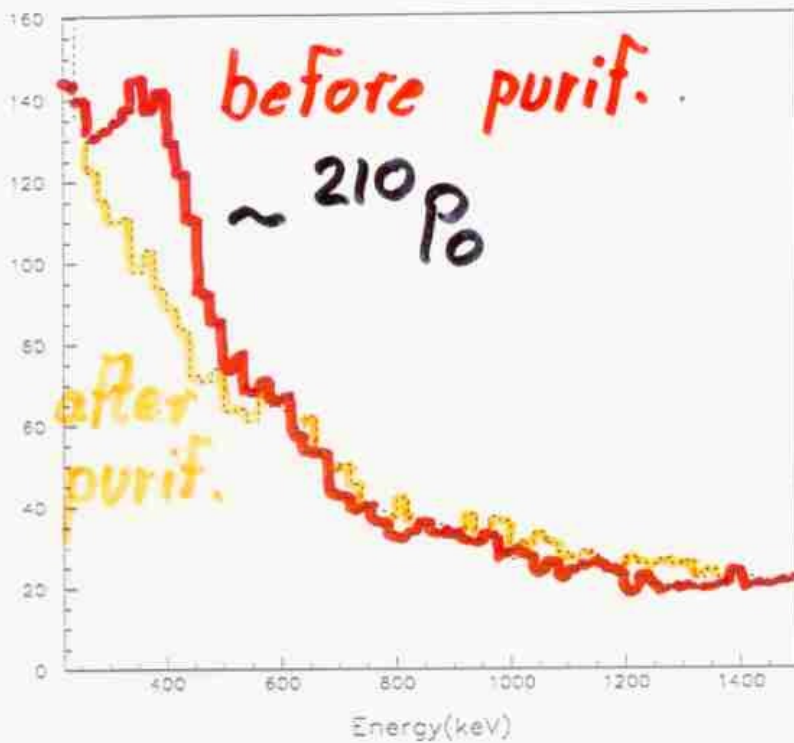
$$^{238}\text{U} < (3.5 \pm 1.3) * 10^{-16}$$

BOREXINO Collab., *Astroparticle Physics* 8 (1998) 141
BOREXINO Collab., *Phys. Lett. B* 422/1-4 (1998) 349

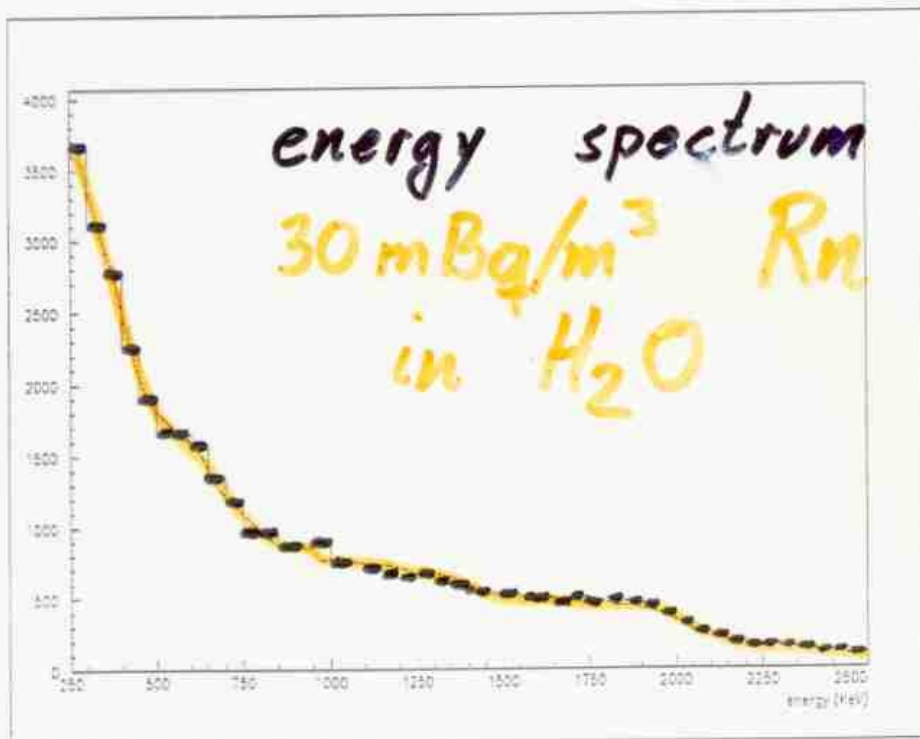
Neutron Activation Analysis:

$$^{238}\text{U} < 2 * 10^{-16}$$

Scintillator purification



- N_2 bubbling
- H_2O extraction
- distillation



single counting rate dominated by external background (Rn in H₂O)

CONCLUSIONS

- UNPRECEDENTED PURITY LEVELS IN SCINTILLATOR HAVE BEEN DEMONSTRATED IN CTF
- ON THIS BASIS BOREXINO HAS BEEN APPROVED AND THE CONSTRUCTION ALREADY STARTED
- THE DETECTOR ASSEMBLY IS EXPECTED TO BE COMPLETED IN THE YEAR 2000