

Constraints on non-standard interaction in ν propagation from ν_{atm} & T2KK long-baseline experiment

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1. Introduction

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1. Introduction

Framework of 3 flavor ν oscillation

Mixing matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Functions of mixing angles θ_{12} , θ_{23} , θ_{13} , and CP phase δ

Information we have obtained so far:

ν_{solar} +KamLAND (reactor)

$$\theta_{12} \approx \frac{\pi}{6}, \Delta m_{21}^2 \approx 8 \times 10^{-5} \text{ eV}^2$$

ν_{atm} +K2K,MINOS(accelerators)

$$\theta_{23} \approx \frac{\pi}{4}, |\Delta m_{32}^2| \approx 2.5 \times 10^{-3} \text{ eV}^2$$

CHOOZ (reactor)

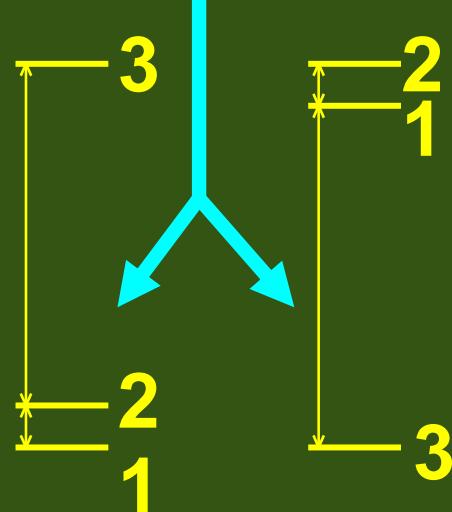
$$|\theta_{13}| \leq \sqrt{0.15}/2$$

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \simeq \begin{pmatrix} c_{12} & s_{12} & \epsilon \\ -s_{12}/\sqrt{2} & c_{12}/\sqrt{2} & 1/\sqrt{2} \\ s_{12}/\sqrt{2} & -c_{12}/\sqrt{2} & 1/\sqrt{2} \end{pmatrix}$$

• Both mass hierarchies are allowed

- θ_{13} :only upper bound is known
- δ :undetermined

Next task is to measure θ_{13} , sign(Δm^2_{31}) and δ .



normal hierarchy

$$\Delta m^2_{32} > 0$$

inverted hierarchy

$$\Delta m^2_{32} < 0$$

Future LBL exp. (on-going / under construction / proposed)

- (conventional) superbeam

T2K phase I (2009-, 0.75MW, E~1GeV, L=295km)

T2K phase II (4MW+HK, E~1GeV, L=295km)

T2KK (JAERI→HK&Korea, E~1GeV, L=295km&1000km)

NOvA (FNAL→ Ash River (MN), E~2GeV, L=810km)

LBNE (FNAL→Homestake(SD), E~3GeV, L=1300km)

SPL (CERN→Frejus, E~0.25GeV, L=130km)

- neutrino factory ($E_\nu \sim 25\text{GeV}$, L ~ 3000km+7500km)

- beta beam ($E_\nu = 0.5\text{-}1.5\text{GeV}$, L ~ 130km)

Future reactor experiments (E~4MeV, L~2km)

Double CHOOZ (France) , Daya Bay (China), Reno (Korea)

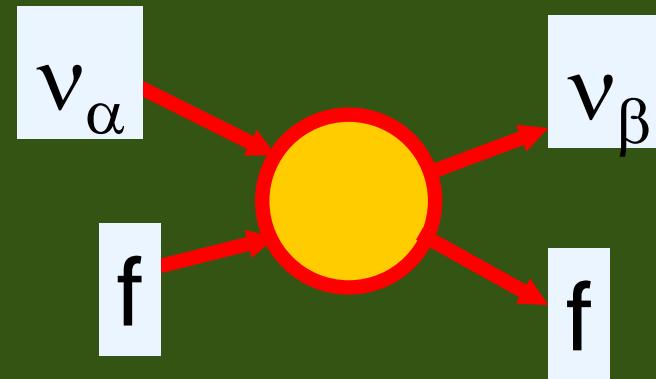
→ With these experiments, θ_{13} , sign(Δm^2_{31})
and δ are expected to be determined

Motivation for research on New Physics

High precision measurements of ν oscillation in future experiments can be used to probe **physics beyond SM** by looking at deviation from SM+ m_ν (like at B factories).
→ Research on **New Physics** is important.

Phenomenological New Physics
considered in this talk: 4-fermi
exotic interactions:

$$\mathcal{L}_{eff} = G_{NP}^{\alpha\beta} \bar{\nu}_\alpha \gamma^\mu \nu_\beta \bar{f} \gamma_\mu f'$$



neutral current
Non Standard
Interaction

● NSI in propagation (non-standard matter effect)

$$i \frac{d}{dt} \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \left[U \text{diag} \left(\frac{m_1^2}{2E}, \frac{m_2^2}{2E}, \frac{m_3^2}{2E} \right) U^{-1} + A \begin{pmatrix} 1 & \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{\mu e} & 1 & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{\tau e} & \epsilon_{\tau\mu} & 1 & \epsilon_{\tau\tau} \end{pmatrix} \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} \right]$$

$A = 2^{1/2} G_F N_e$

N_e = electron density

NSI

● Constraints on $\epsilon_{\alpha\beta}$

Davidson et al., JHEP 0303:011,2003; Berezhiani, Rossi, PLB535 ('02)
207; Barranco et al., PRD73 ('06) 113001; Barranco et al., arXiv:0711.0698

Biggio et al., JHEP 0908, 090
(2009) w/o 1-loop arguments

related to each
other by ν_{atm}

can be improved
by ν_{atm}

$$\left(\begin{array}{l} |\epsilon_{ee}| \lesssim 4 \times 10^0 \\ |\epsilon_{e\mu}| \lesssim 3 \times 10^{-1} \\ |\epsilon_{\mu\mu}| \lesssim 7 \times 10^{-2} \\ |\epsilon_{e\tau}| \lesssim 3 \times 10^0 \\ |\epsilon_{\mu\tau}| \lesssim 3 \times 10^{-1} \\ |\epsilon_{\tau\tau}| \lesssim 2 \times 10^1 \end{array} \right)$$

2. High energy behavior of ν_{atm} data & NSI

● Standard case with $N_\nu=2$

$$1 - P(\nu_\mu \rightarrow \nu_\mu) \sim \sin^2 2\theta_{\text{atm}} \sin^2 \left(\frac{\Delta m_{\text{atm}}^2 L}{4E} \right) \propto \frac{1}{E^2}$$

● Standard case with $N_\nu=3$

$$1 - P(\nu_\mu \rightarrow \nu_\mu) \simeq \left(\frac{\Delta m_{31}^2}{2AE} \right)^2 \left[\sin^2 2\theta_{23} \left(\frac{c_{13}^2 AL}{2} \right)^2 + s_{23}^2 \sin^2 2\theta_{13} \sin^2 \left(\frac{AL}{2} \right) \right] \propto \frac{1}{E^2}$$

● Deviation of $1 - P(\nu_\mu \rightarrow \nu_\mu)$ due to **NSI** contradicts with data

$$1 - P(\nu_\mu \rightarrow \nu_\mu) \simeq \mathbf{C_0} + \frac{\mathbf{C_1}}{E} + \frac{c_{20}L^2 + c_{21} \sin^2(c_{22}L)}{E^2}$$

→ High ν_{atm} data gives constraints on **NSI**:

$$|\mathbf{C_0}| \ll 1, |\mathbf{C_1}| \ll 1$$

$$1 - P(\nu_\mu \rightarrow \nu_\mu) \simeq \mathbf{C_0} + \frac{\mathbf{C_1}}{E} + \frac{c_{20}L^2 + c_{21}\sin^2(c_{22}L)}{E^2}$$

$|C_0| \ll 1 \rightarrow |\varepsilon_{e\mu}| \ll 1, |\varepsilon_{\mu\mu}| \ll 1, |\varepsilon_{\mu\tau}| \ll 1$

$|\varepsilon_{\mu\tau}| \ll 1$: Already shown by Fornengo et al. PRD65, 013010, '02;
Gonzalez-Garcia&Maltoni, PRD70, 033010, '04; Mitsuka@nufact08

$|\varepsilon_{\mu\mu}| \ll 1$: Already shown from other expts. by Davidson et al.
JHEP 0303:011, '03

$|\varepsilon_{e\mu}| \ll 1$: New observation (analytical consideration only)

$|C_1| \ll 1 \rightarrow |\varepsilon_{\tau\tau} - |\varepsilon_{e\tau}|^2 / (1 + \varepsilon_{ee})| \ll 1$

Already shown by
Friedland-Lunardini,
PRD72:053009, '05

● Summary of the constraints on $\varepsilon_{\alpha\beta}$

To a good approximation, we are left with 3 independent variables ε_{ee} , $|\varepsilon_{e\tau}|$, $\arg(\varepsilon_{e\tau})$:

$$A \begin{pmatrix} 1 + \varepsilon_{ee} & \varepsilon_{e\mu} & \varepsilon_{e\tau} \\ \varepsilon_{\mu e} & \varepsilon_{\mu\mu} & \varepsilon_{\mu\tau} \\ \varepsilon_{\tau e} & \varepsilon_{\tau\mu} & \varepsilon_{\tau\tau} \end{pmatrix} \quad \xrightarrow{\text{red arrow}} \quad A \begin{pmatrix} 1 + \varepsilon_{ee} & 0 & \varepsilon_{e\tau} \\ 0 & 0 & 0 \\ \varepsilon_{e\tau}^* & 0 & |\varepsilon_{e\tau}|^2 / (1 + \varepsilon_{ee}) \end{pmatrix}$$

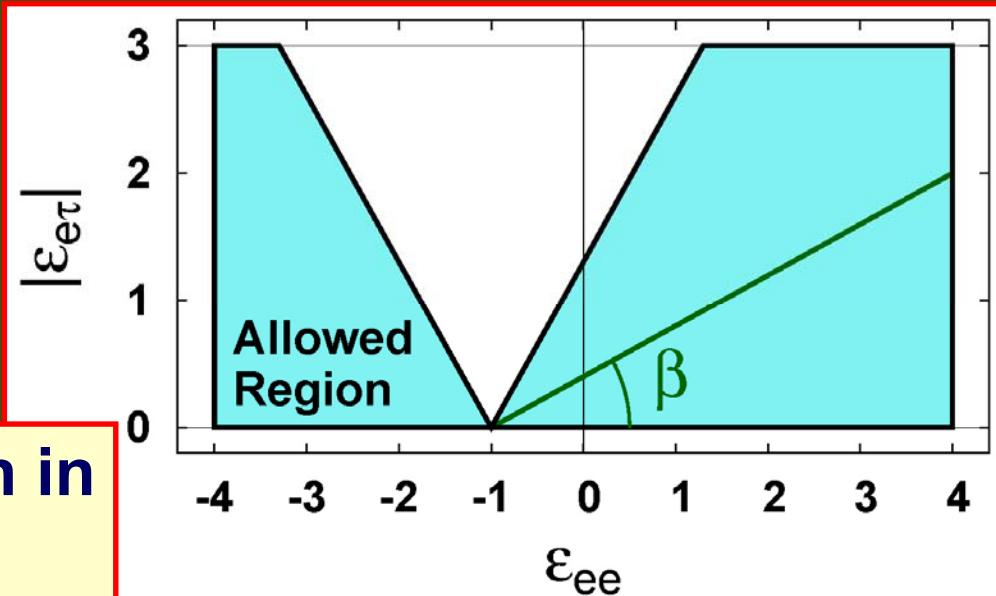
Furthermore, ν_{atm} data implies

$$\tan\beta = |\varepsilon_{e\tau}| / (1 + \varepsilon_{ee}) < 1.3$$

Friedland-Lunardini,
PRD72:053009, '05



**Allowed region in
(ε_{ee} , $|\varepsilon_{e\tau}|$)**

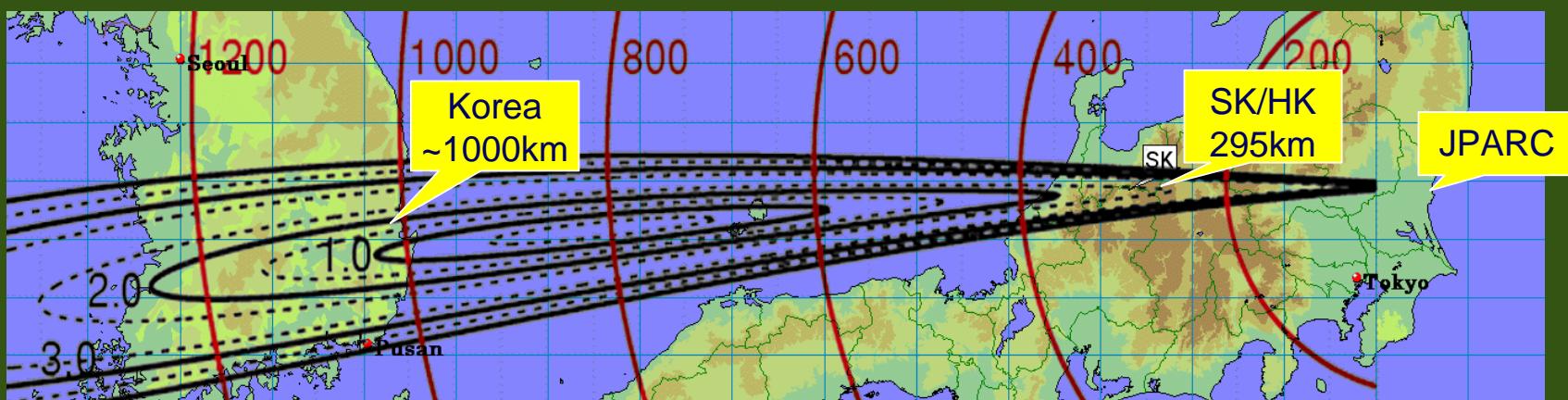


3. Sensitivity to NSI of propagation at T2KK

T2KK proposal with baselines L=295km, 1050km
→ L=1050km is sensitive to the matter effect

	$ \Delta m^2_{31} L/4E$	$ \Delta m^2_{21} L/4E$	$AL/2$
L=295km	~1	~0.04	~0.06
L=1050km	~5	~0.1	~0.3

dependence on A & Δm^2_{21} at L=1050km is non-negligible



Outline of our Analysis

$$A \equiv \sqrt{2}G_F n_e$$

Our ansatz

$$i \frac{d}{dt} \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \left\{ U^{-1} \text{diag} \left(\frac{m_1^2}{2E}, \frac{m_2^2}{2E}, \frac{m_3^2}{2E} \right) U + A \begin{pmatrix} 1 + \mathcal{E}_{ee} & 0 & \mathcal{E}_{e\tau} \\ 0 & 0 & 0 \\ \mathcal{E}_{e\tau}^* & 0 & \frac{|\mathcal{E}_{e\tau}|^2}{1 + \mathcal{E}_{ee}} \end{pmatrix} \right\} \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}$$

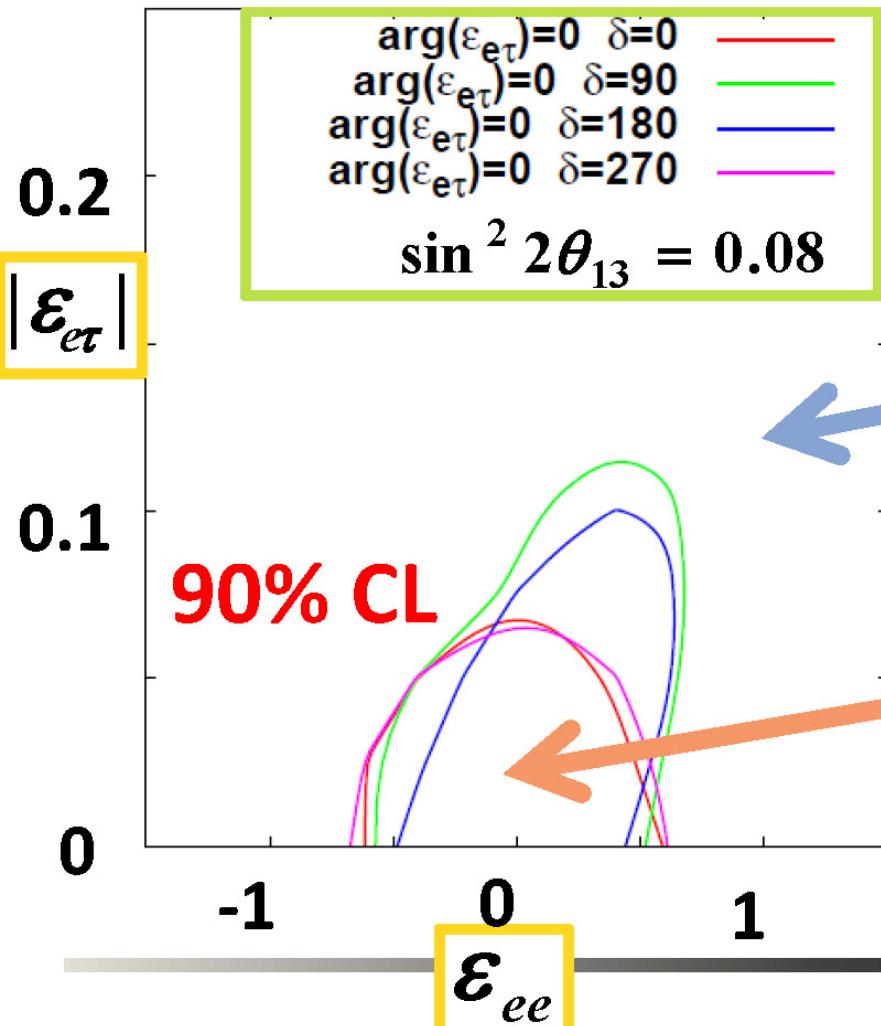
Black : standard Red : non-standard

$$\Delta\chi^2(\text{NSI}) = \min_{\substack{\text{std} \\ \text{parameters}}} \sum_i \frac{(N_i^0(\text{NSI}) - N_i(\text{std}))^2}{\sigma_i^2} + \Delta\chi^2_{\text{prior}}$$

$$\Delta\chi^2_{\text{prior}} = \frac{(\sin^2 2\theta_{23} - \sin^2 2\theta_{23}^{\text{best}})^2}{(\delta \sin^2 2\theta_{23})^2} + \frac{(\Delta m_{31}^2 - \Delta m_{31}^{2\text{best}})^2}{(\delta \Delta m_{31}^2)^2} + \frac{(\sin^2 2\theta_{13} - \sin^2 2\theta_{13}^{\text{best}})^2}{(\delta \sin^2 2\theta_{13})^2}$$

$\Delta\chi^2 > 4.6$: Deviation of NSI from vSM is significant compared with errors (at 90% CL of 2 degrees of freedom \mathcal{E}_{ee} , $|\mathcal{E}_{e\tau}|$)

Sensitivity to ε_{ee} , $|\varepsilon_{e\tau}|$



Marginalized over θ_{13} ,
 θ_{23} , $|\Delta m_{31}^2|$, $\text{sign}(\Delta m_{31}^2)$

- Outside of the curves : Effects of NSI can be distinguished from the standard case.
- Inside of the curves : Effects of NSI are not significant.

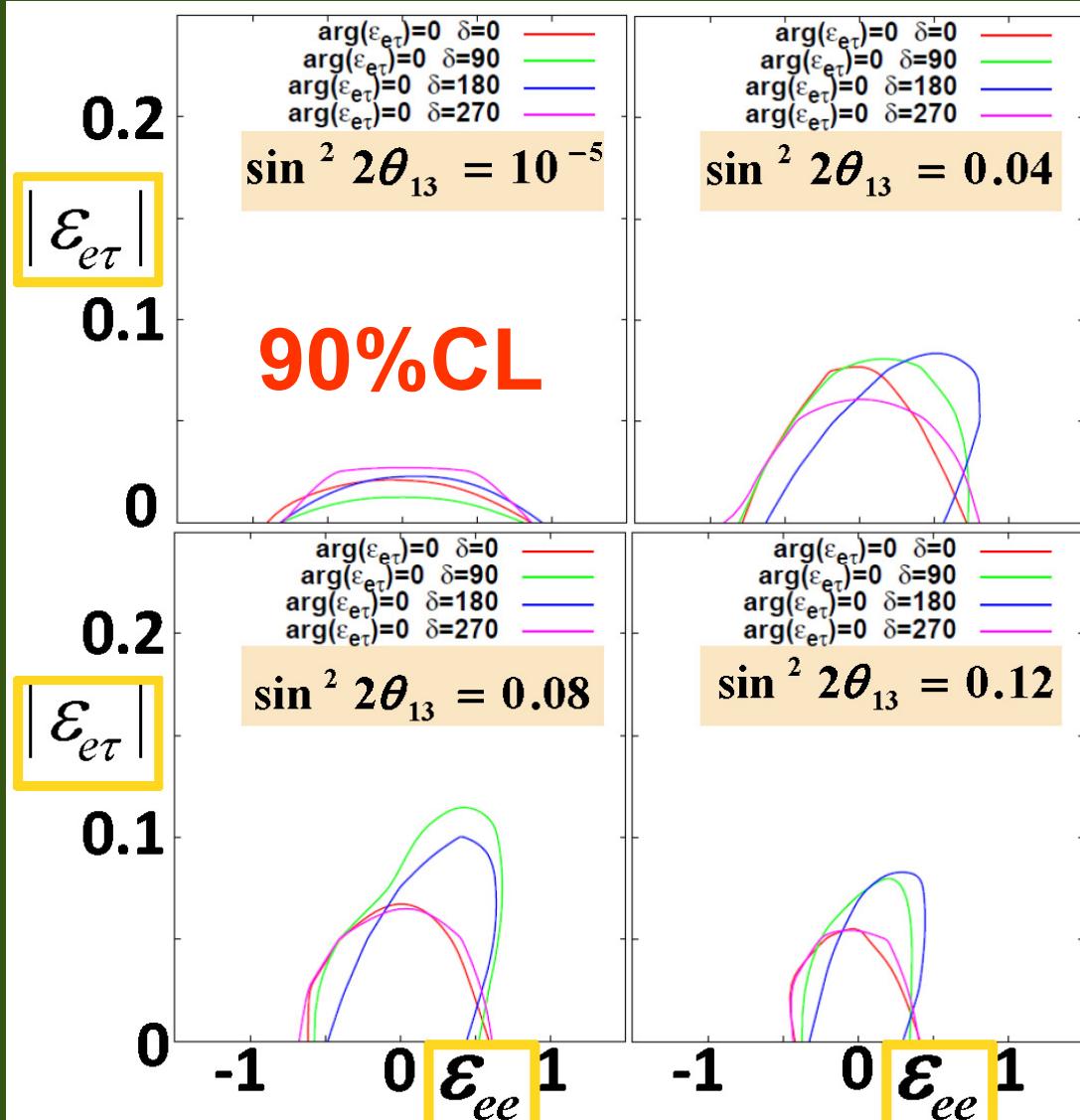
● Dependence on phases δ and $\arg(\varepsilon_{e\tau})$

$$P(\nu_\mu \rightarrow \nu_e) \cong P_0(\nu_\mu \rightarrow \nu_e) \Big|_{\Delta m_{21}^2 = 0} + \Delta m_{21}^2 P_1(\nu_\mu \rightarrow \nu_e)$$

Function of
 $\delta + \arg(\varepsilon_{e\tau})$

Approximately
function of
 $\arg(\varepsilon_{e\tau})$ only

Sensitivity to ϵ_{ee} , $|\epsilon_{e\tau}|$ for various θ_{13}



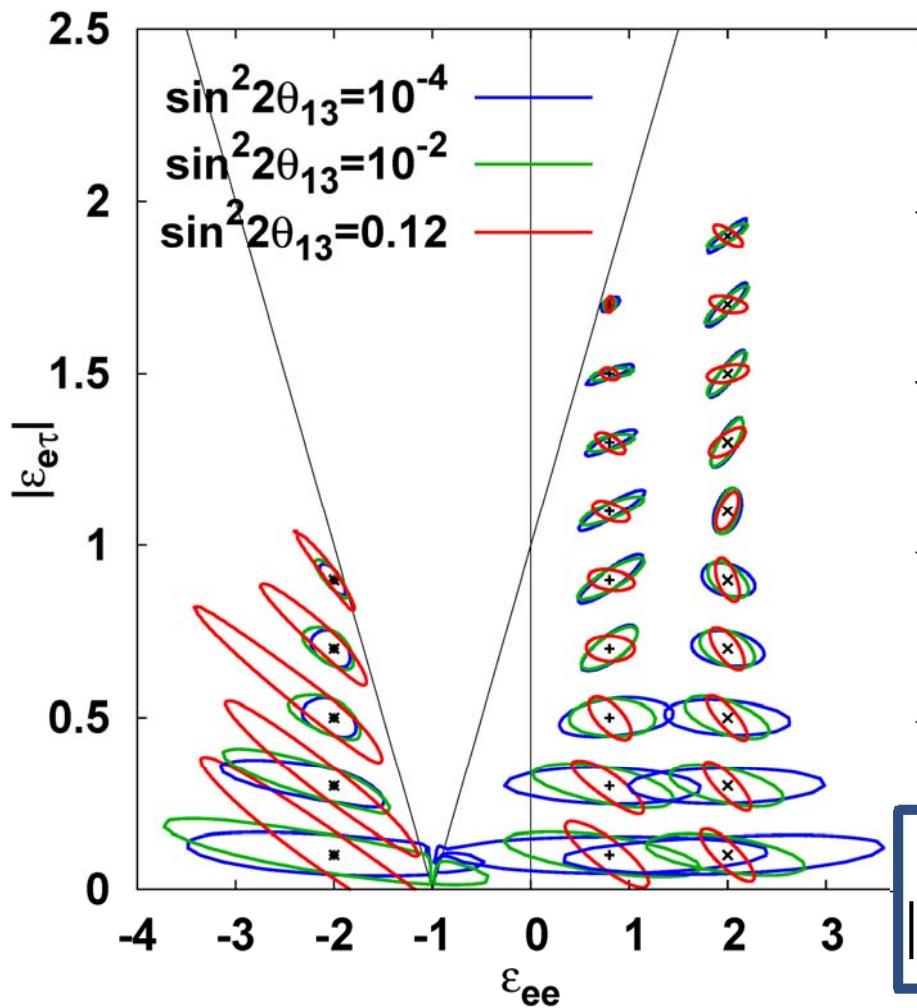
regions depend on
 $\theta_{13}, \delta, \arg(\epsilon_{e\tau})$

$|\epsilon_{ee}| < 4$
 $|\epsilon_{e\tau}| < 3$

↓

$|\epsilon_{ee}| \lesssim 1$
 $|\epsilon_{e\tau}| \lesssim 0.2$

Precision of ε_{ee} , $|\varepsilon_{e\tau}|$



$|\varepsilon_{e\tau}| \geq 0.5$
 $\varepsilon_{ee}, |\varepsilon_{e\tau}|$
determined
separately



ε_{ee} not determined
 $|\varepsilon_{e\tau}|$ determined (if > 0.1)

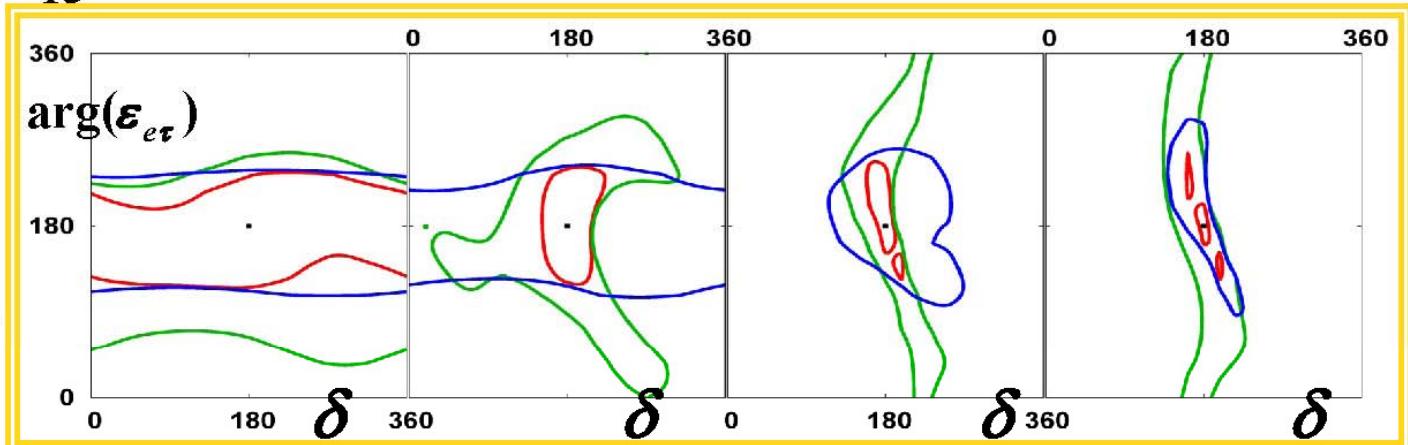
Sensitivity to δ and $\arg(\epsilon_{e\tau})$



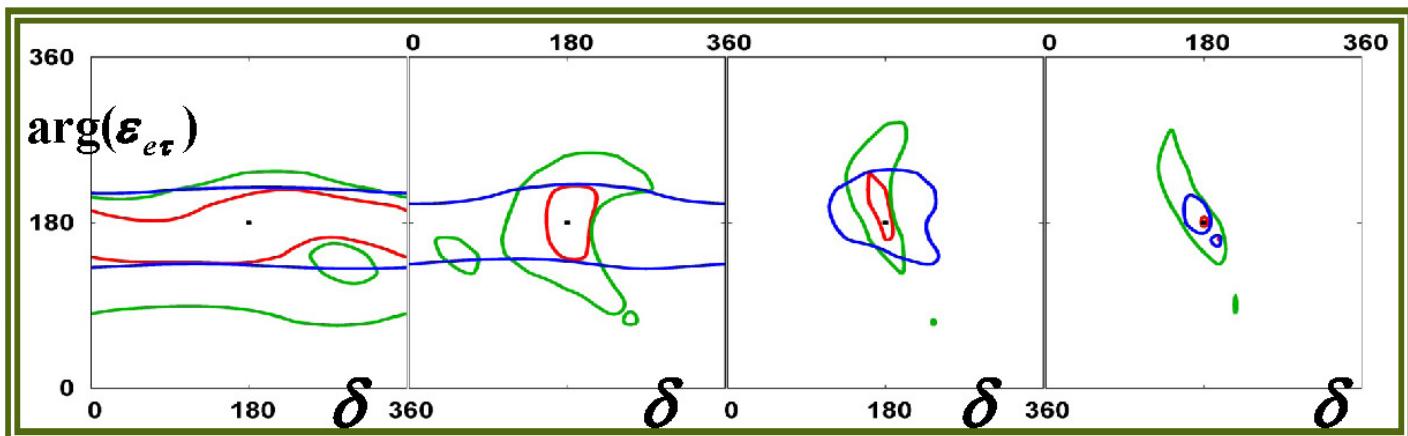
- Correlation of measured $\arg(\epsilon_{e\tau})$ and δ

$$\sin^2 2\theta_{13} = 10^{-4} \quad 10^{-3} \quad 10^{-2} \quad 0.12$$

$$\epsilon_{ee} = 0.8 \\ |\epsilon_{e\tau}| = 0.2$$



$$\epsilon_{ee} = 0.8 \\ |\epsilon_{e\tau}| = 0.4$$



- If $(\epsilon_{ee}, |\epsilon_{e\tau}|, s_{13})$ are large, we can determine $\arg(\epsilon_{e\tau}), \delta$

4. Summary (1)

- We provided an analytical argument on the oscillation probability for high energy atmospheric neutrinos that $|\epsilon_{\mu\alpha}| \ll 1$ & $|\epsilon_{\tau\tau} - |\epsilon_{e\tau}|^2/(1+\epsilon_{ee})| \ll 1$ must hold.

$$A \begin{pmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{\mu e} & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{\tau e} & \epsilon_{\tau\mu} & \epsilon_{\tau\tau} \end{pmatrix} \rightarrow A \begin{pmatrix} 1 + \epsilon_{ee} & 0 & \epsilon_{e\tau} \\ 0 & 0 & 0 \\ \epsilon_{e\tau}^* & 0 & |\epsilon_{e\tau}|^2/(1 + \epsilon_{ee}) \end{pmatrix}$$

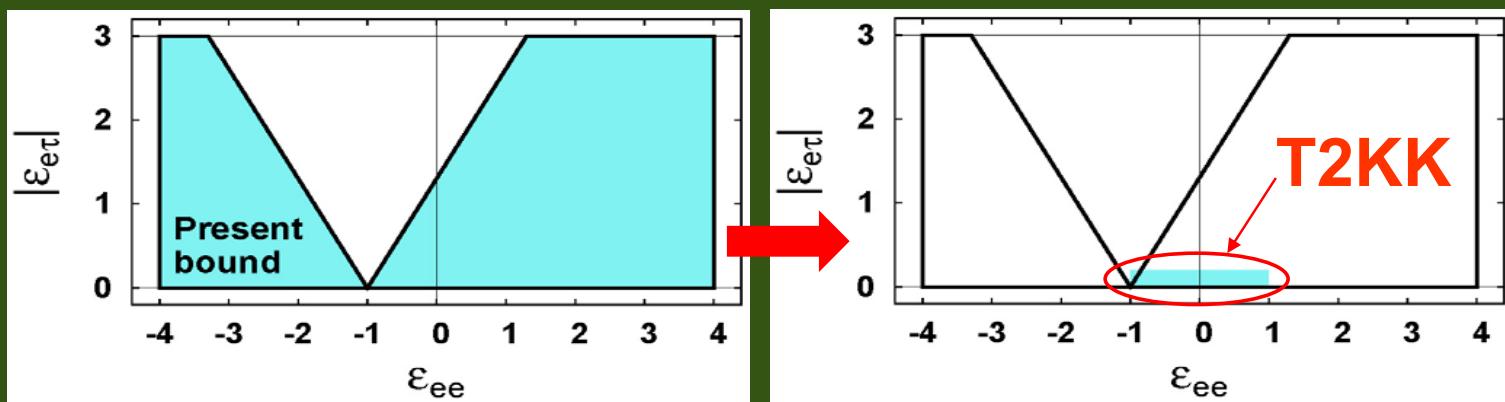
- Deviation of $1 - P(\nu_\mu \rightarrow \nu_\mu)$ from the standard case in high energy ν_{atm} data gives strong constraints on Non Standard Interaction.
→ It would be great if we can constrain/determine c_0, c_1, c_{2j} ($j=0,1,2$) in high energy ν_{atm} data:

$$1 - P(\nu_\mu \rightarrow \nu_\mu) \simeq \mathbf{C_0} + \frac{\mathbf{C_1}}{E} + \frac{c_{20}L^2 + c_{21}\sin^2(c_{22}L)}{E^2}$$

Is it possible at SK, IceCube, HK?

4. Summary (2)

- We studied phenomenologically sensitivity to NSI in propagation of the T2KK proposal.
- Under the assumptions $\varepsilon_{e\mu} = \varepsilon_{\mu\mu} = \varepsilon_{\mu\tau} = 0$ & $\varepsilon_{\tau\tau} = |\varepsilon_{e\tau}|^2 / (1 + \varepsilon_{ee})$, we found that T2KK can restrict the NSI parameters: $|\varepsilon_{ee}| \lesssim 1$, $|\varepsilon_{e\tau}| \lesssim 0.2$



- If ε_{ee} , $|\varepsilon_{e\tau}|$ and s_{13} are large, then T2KK can determine ε_{ee} , $|\varepsilon_{e\tau}|$, δ and $\arg(\varepsilon_{e\tau})$ separately.