

# Phenomenology of neutrino oscillation

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Elementary particles, Interaction of elementary particles

## Part II. Physics beyond the Standard Model

$\nu$  oscillation, Atmospheric  $\nu$  + Accelerator  $\nu$ , Solar  $\nu$  + Long baseline reactor  $\nu$ , Short baseline reactor  $\nu$ , 3 flavor  $\nu$  oscillation, Future plans, Beyond the standard scenario

## Summary

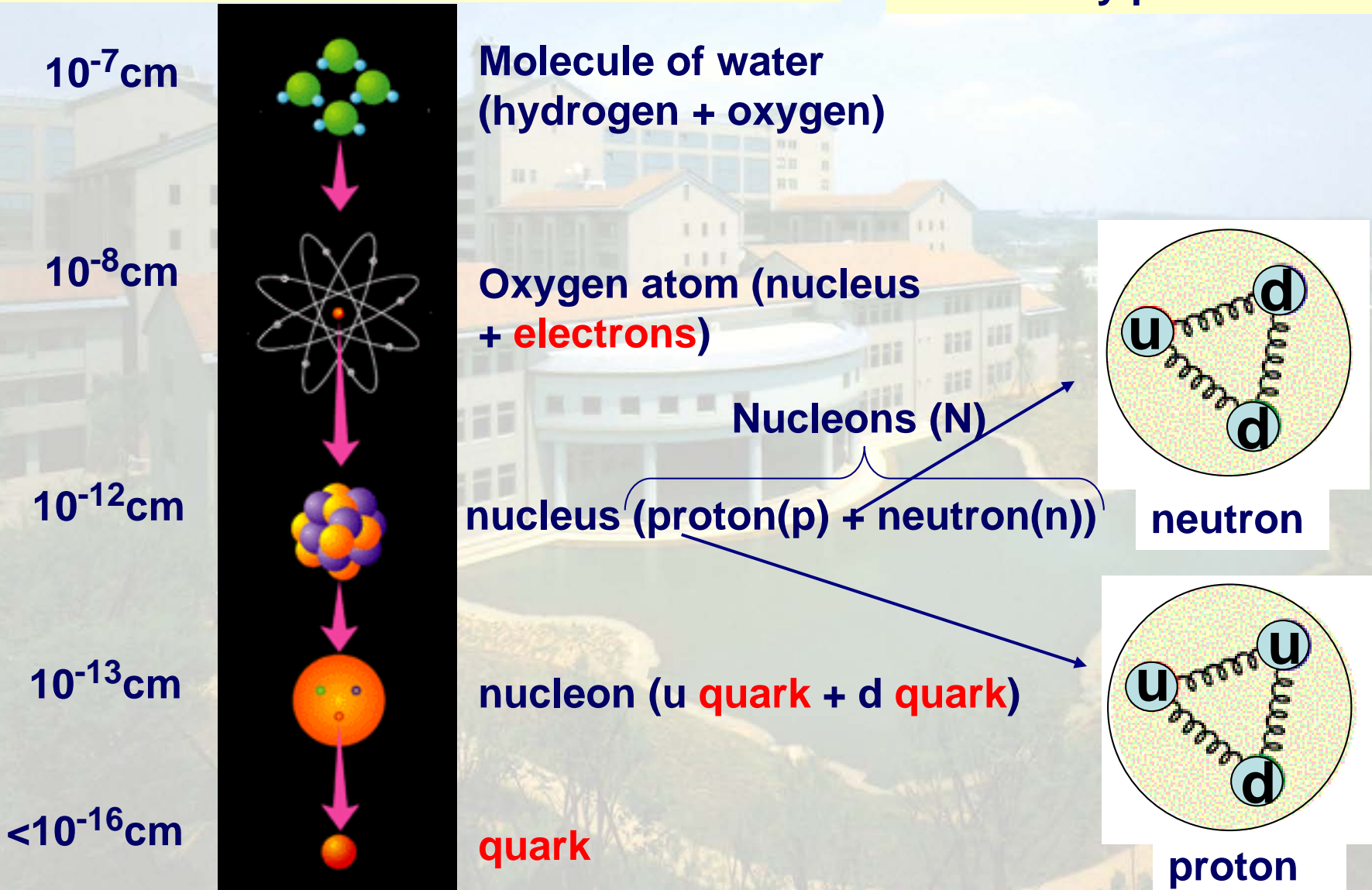
The background of the slide is a faded, high-angle photograph of a university campus. It shows several multi-story buildings with light-colored facades and red-tiled roofs. In the foreground, there is a large, irregularly shaped lake or pond surrounded by greenery and a winding path. The sky is blue with some light clouds.

# **Part I. The Standard Model**

## **(1) Elementary particles (the Standard Model)**

# Elementary particles: what cannot be divided further

At present, **electrons** and **quarks** are regarded as elementary particles.




# Neutrino ( $\nu$ )

Elementary particles predicted in 1933

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neutron  $\rightarrow$  proton + electron

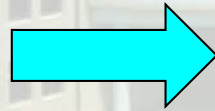
This process does not satisfy energy + momentum conservation

 Neutral particle called **neutrino** was introduced:  
neutron  $\rightarrow$  proton + electron + (anti-)neutrino

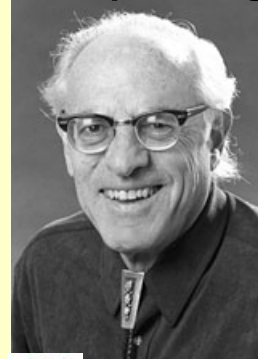


 Pauli

$\nu$  was first discovered in 1955



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 Reines

www.aps.org



Cowan

Until 1998 neutrino was regarded as **massless**.

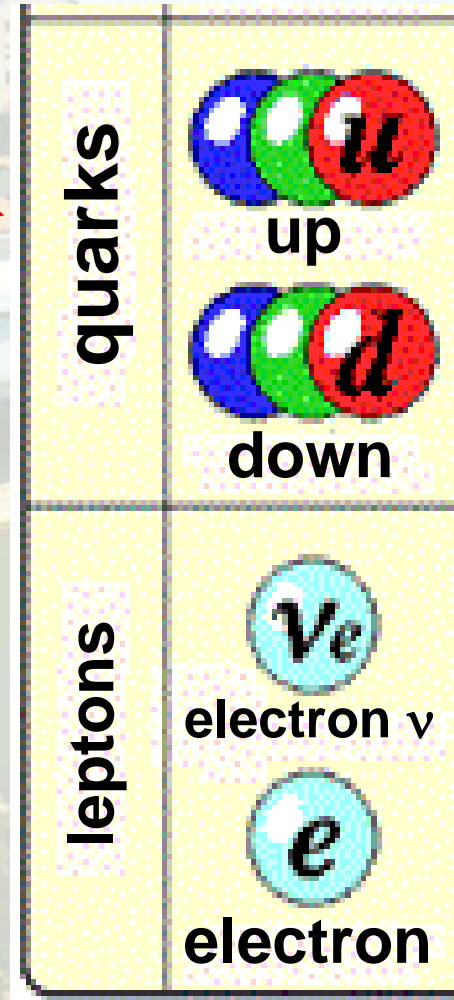
Discovery of neutrinos in 1955  
(neutrinos from a reactor)

# Summary (1): elementary particles

## Matter consists of quarks & leptons

**Quarks** constitute composite particles (e.g., protons, neutrons) by attractive force between quarks

**Leptons** have properties different from quarks, and do not constitute composite particles



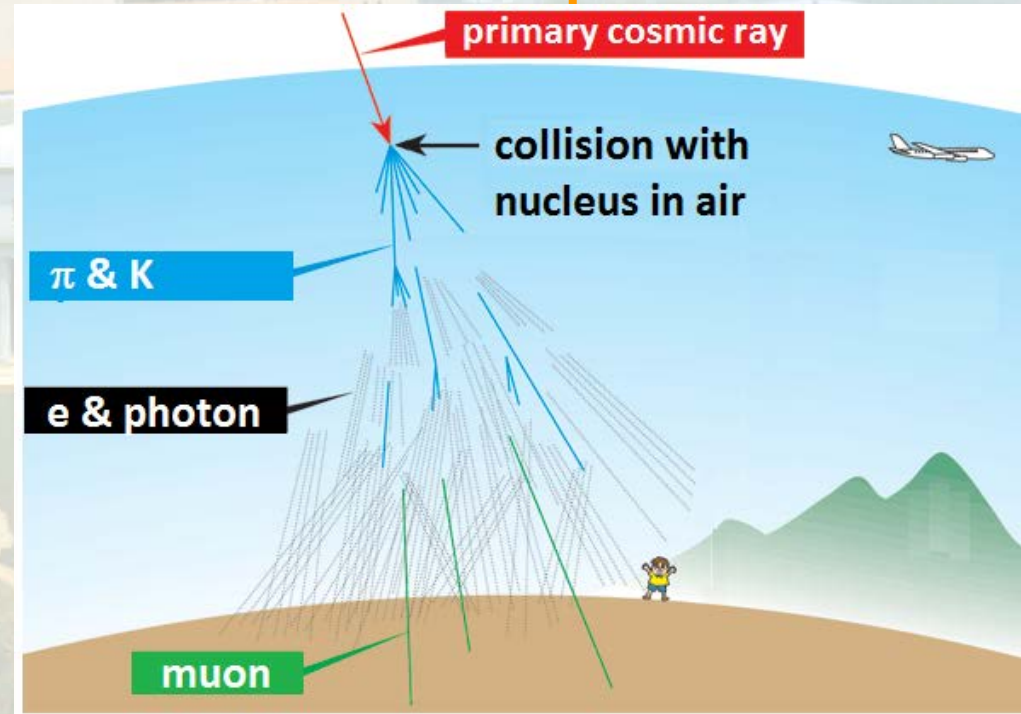
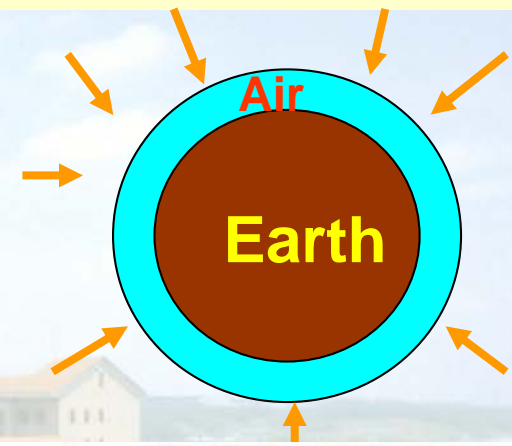
# Cosmic rays

It is known that so-called cosmic rays are falling down on Earth.

- Primary cosmic rays collide with nuclei in the air to produce particles which are called secondary cosmic rays.

- The major components of secondary cosmic rays are **muons** which have almost the same properties as electrons except their mass ( $m_{\mu} = 200m_e$ )













Primary cosmic rays (p, He)



**muons** : elementary particle of 2<sup>nd</sup> generation

# Summary (2): elementary particles

- There are 3 generations of elementary particles.
- Neutrinos are **massless** in the Standard Model of Elementary particles.

	1st	2nd	3rd
quarks	 up	 charm	 top
	 down	 strange	 bottom
leptons	 electron $\nu$	 mu $\nu$	 tau $\nu$
	 electron	 muon	 tauon

	1st	mass[MeV]
quarks	u	3
	d	6
leptons	e	0.5
	$\nu_e$	0

	2nd	mass[MeV]
quarks	c	1,200
	s	120
leptons	$\mu$	106
	$\nu_\mu$	0

	3rd	mass[MeV]
quarks	t	174,300
	b	4,000
leptons	$\tau$	1777
	$\nu_\tau$	0

Higher mass for higher generation

$E=mc^2$  tells us we need much energy to produce heavy particles



We need special device to produce particle of 2<sup>nd</sup> or 3<sup>rd</sup> generation



# Anti-particle: Particle with the same mass and opposite electric charge

1930: Dirac equation (Relativity+Quantum mechanics)

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 Dirac

↓  
**Positron(=anti-particle of electron) was theoretically predicted.**  
↓

	Mass	Electric charge
electron	0.5MeV	-e
positron	0.5MeV	+e

1932: Discovery of positron

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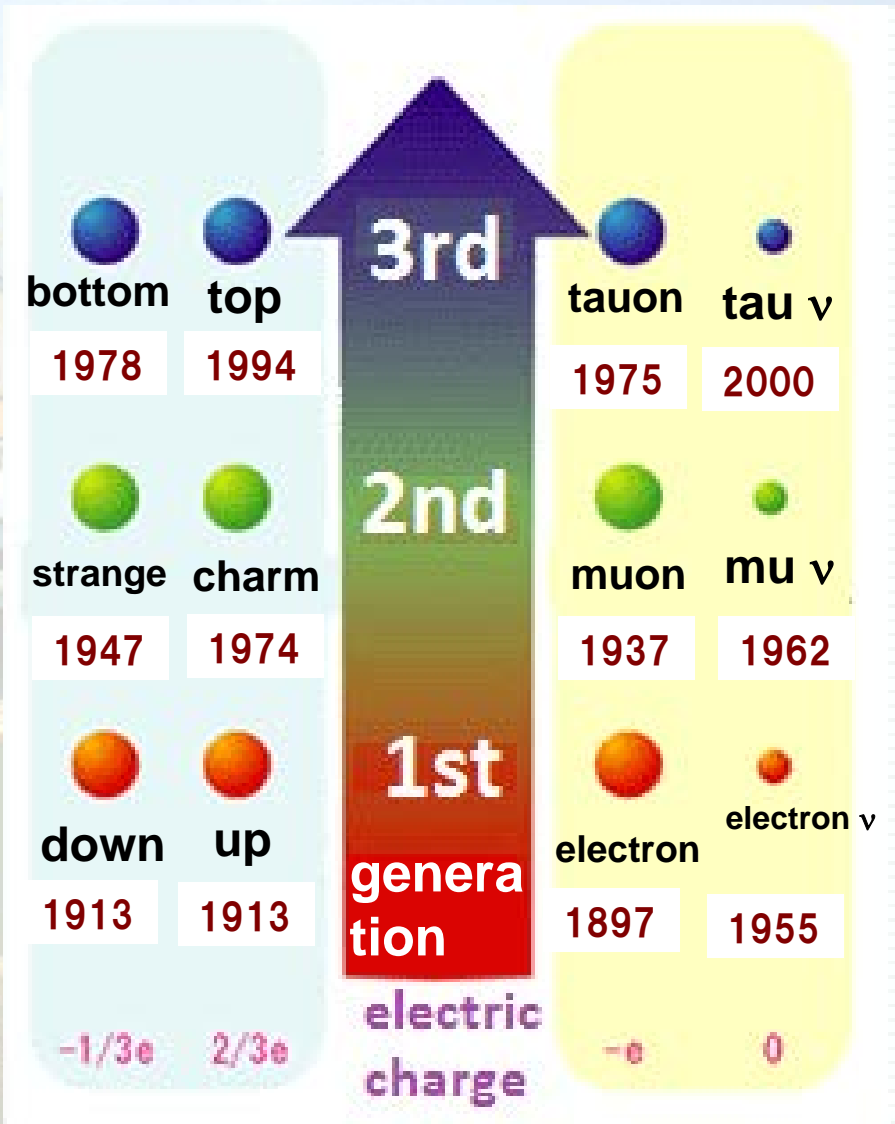
 Anderson

**In general, particles (3 generation of quarks & leptons) have their own anti-particles.**

# Actually 3 generation of quarks were theoretically predicted!


## 1972 Kobayashi-Maskawa

From motivation for so-called **CP violation**, 3 generation of quarks were theoretically predicted.



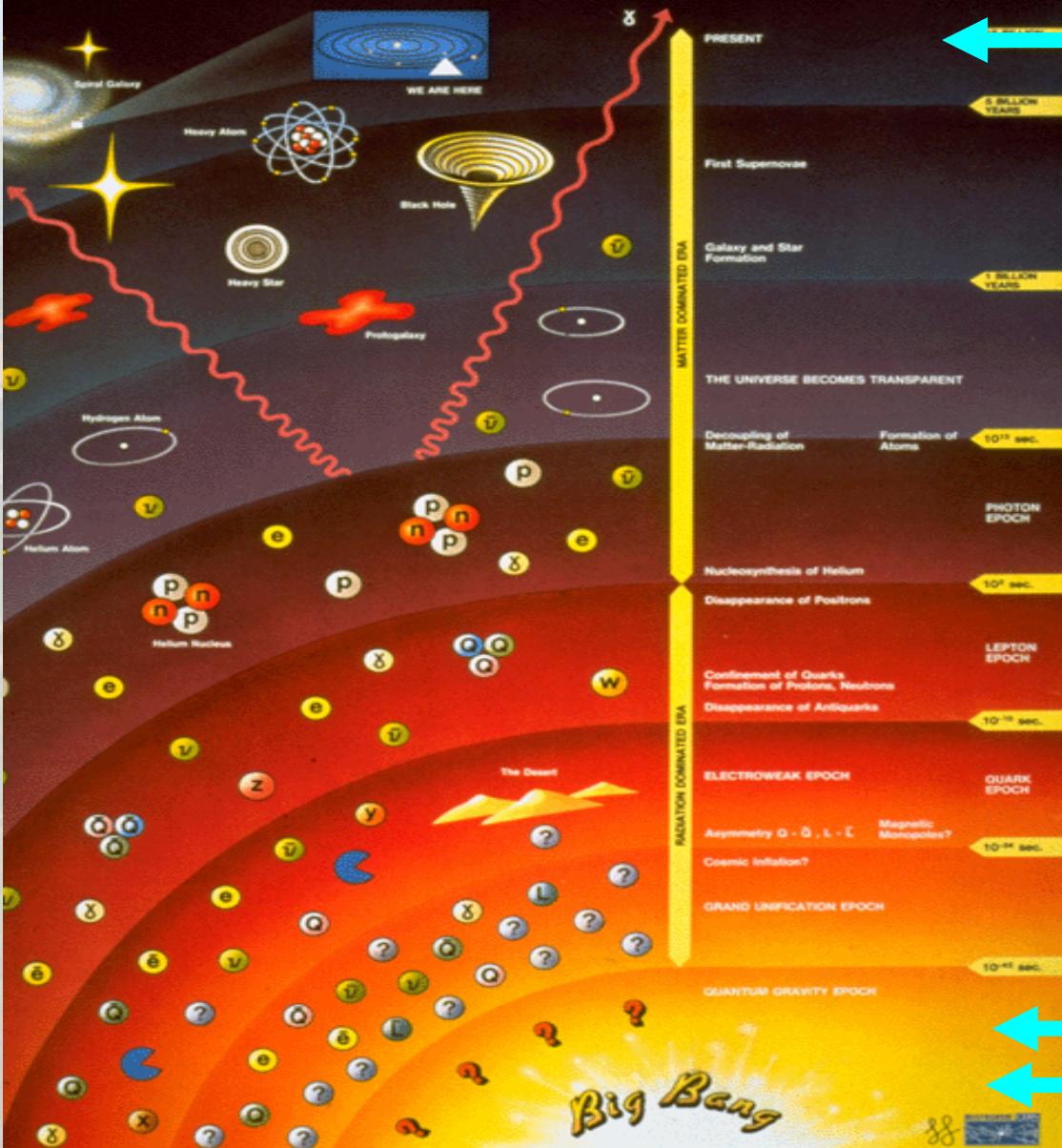
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 Kobayashi

 Maskawa

# History of the Universe



**T=3 deg (= -270°C)**  
**Present Universe is dominated by matter (w/o anti-matter)**

**Universe expanded & T decreased**

**There must have been asymmetry between particles & anti-particles at some stage**

**Temperature T=10<sup>32</sup>deg**  
**At the beginning of universe, #(particle) = #(anti-particles)**  
**→ There must be equal amount of matter & anti-matter**

**Universe was born by Big Bang**

**CP symmetry**  
(Invariance under  
CP transformation)

$$\mathbf{CP} = \mathbf{C} \times \mathbf{P}$$

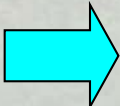
**C:** Charge conjugation

**P:** Parity transformation

If CP symmetry is broken, then there can be difference between the speeds of the following reactions:

Heavy particle  $\rightarrow$  Light particle + . . .

Heavy anti-particle  $\rightarrow$  Light anti-particle + . . .

 If **CP symmetry** is broken, then we may be able to explain **matter-anti-matter asymmetry** of the Universe by cosmology + particle theory!

# Summary (3): elementary particles

- There are 3 generation of particles & anti-particles
- In our Universe, we have only matter (made of particles) but have no anti-matter (made of anti-particles)

↳ Matter-anti-matter asymmetry is a mystery at present

particles				
	1st	2nd	3rd	
quarks				
	up	charm	top	
	down	strange	bottom	
	leptons			
		electron ν	mu ν	tau ν
electron		muon	tauon	

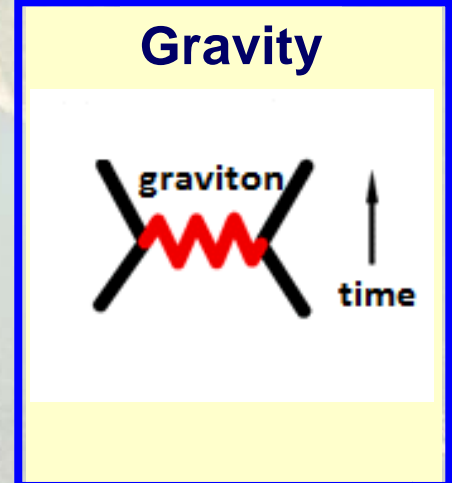
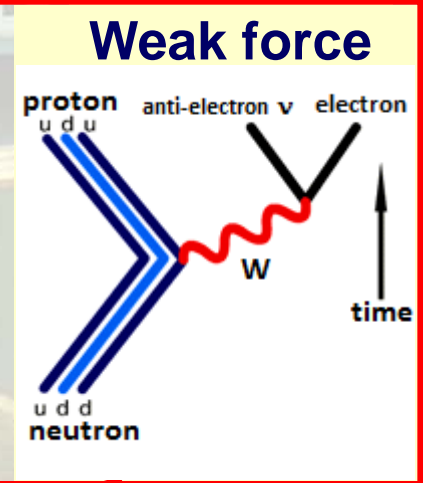
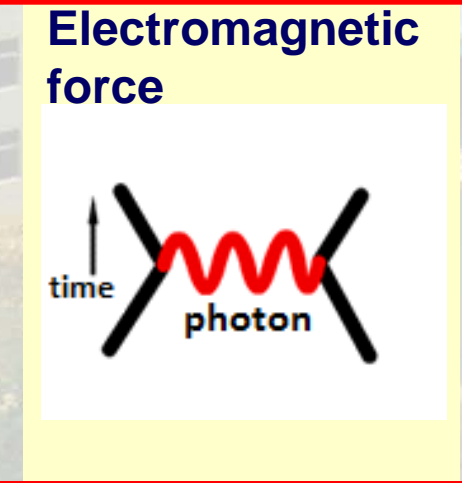
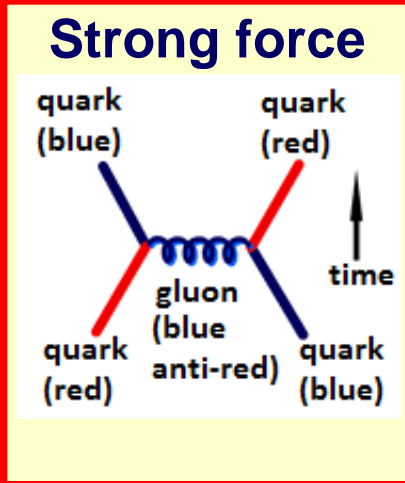
anti-particles				
	1st	2nd	3rd	
anti-quarks				
	anti-up	anti-charm	anti-top	
	anti-down	anti-strange	anti-bottom	
	anti-leptons			
		anti-electron ν	anti-mu ν	anti-tau ν
positron		anti-muon	anti-tauon	



**(2) Interactions of elementary particles (the Standard Model)**

# Interactions of Elementary particles

Interactions (force)	Strong force	Electromagnetic force	Weak force	Gravity
Force mediating particles	Gluon	Photon	W,Z boson	Graviton
Strength of force	1	$10^{-2}$	$10^{-5}$	$10^{-40}$



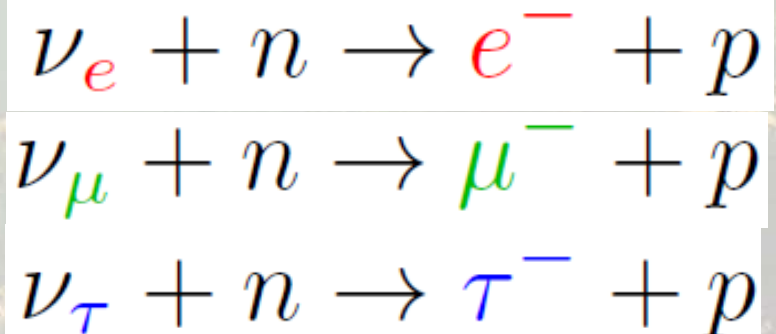
So-called **Standard Model** describes 3 interactions (Strong, Electromagnetic, Weak forces)

Gravity among particle is so weak that it is **ignored**

**(NB) Neutrinos** feel only weak force → It is extremely difficult to observe them.

particles		Strong force	Electro magnetic force	Weak force	Gravity
quark	u	✓	✓	✓	✓
	d	✓	✓	✓	✓
leptons	e	×	✓	✓	✓
	$\nu_e$	×	×	✓	✓

**Flavor** of neutrino is inferred by observing the charged lepton.





# Part II. Physics beyond the Standard Model

## (3) Neutrino oscillation (Physics beyond the Standard Model)

$\nu$  oscillation: **quantum mechanical interference**

- Neutrinos are **massless** in the Standard Model, while they are massive in the theory beyond the Standard Model.

Theory	Neutrino mass	Flavor vs <b>Mass</b> eigenstate
Standard Model	0	the same
Beyond Standard Model	≠0	different

Flavor eigenstate

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Mass eigenstate

# $\nu$ oscillation in vacuum

If  $\nu$  of two different flavor eigenstates  $\nu_\mu, \nu_\tau$  are related to two  $\nu$  mass eigenstates  $\nu_1, \nu_2$  (mass  $m_1, m_2$ ) by a 2x2 matrix

$$\begin{pmatrix} \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$\theta$ : mixing angle

then probability of transforming from  $\nu_\mu$  to  $\nu_\tau$  while propagating for distance  $L$  is given by

$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2 2\theta \sin^2 \left( 1.27 \frac{(\Delta m^2 c^4 / \text{eV}^2) (L / \text{km})}{(E / \text{GeV})} \right)$$

$\Delta m^2 \equiv m_2^2 - m_1^2$   
: mass squared difference

Probability in natural units  
( $\hbar=c=1$ )

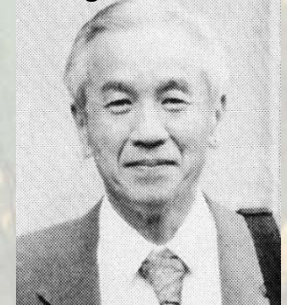
1962 Maki-Nakagawa-Sakata

[www2.yukawa.kyoto-u.ac.jp/~sg/](http://www2.yukawa.kyoto-u.ac.jp/~sg/)



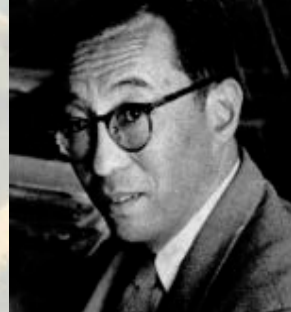
Maki

Publ. Committee Sci.  
Work of Prof.  
Nakagawa



Nakagawa

Publ. Committee Sci.  
Work of Prof. Sakata



Sakata

Probability has an **oscillatory** behavior with respect to L

$$\sin^2 2\theta \sin^2 \left( 1.27 \frac{(\Delta m^2 / \text{eV}^2) (L / \text{km})}{E / \text{GeV}} \right)$$

P=maximum  
→ Argument= $\pi/2$

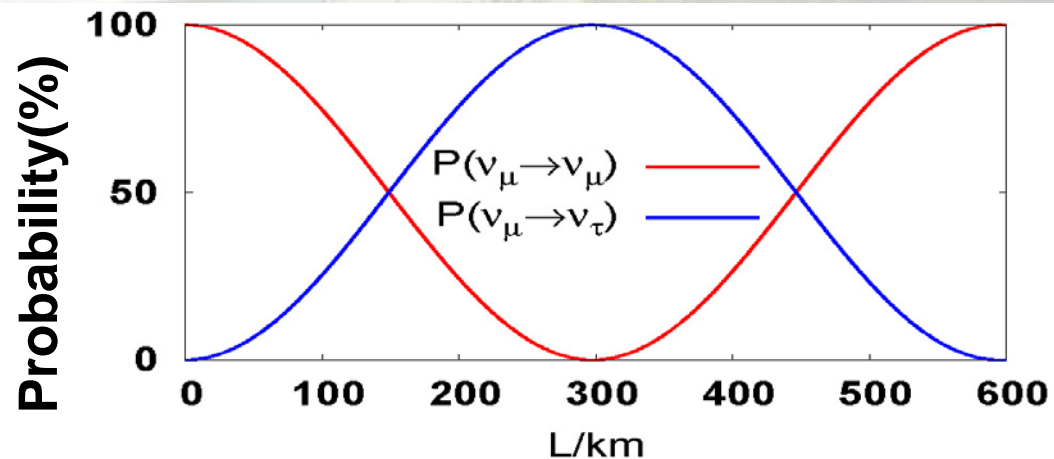
$\Delta m^2 = 3 \times 10^{-3} \text{eV}^2 \rightarrow E = 0.6 \text{GeV}, L = 300 \text{km}$  (accelerator)

$\Delta m^2 = 3 \times 10^{-3} \text{eV}^2 \rightarrow E = 4 \text{MeV}, L = 2 \text{km}$  (short L reactor)

$\Delta m^2 = 8 \times 10^{-5} \text{eV}^2 \rightarrow E = 4 \text{MeV}, L = 60 \text{km}$  (long L reactor)

$\Delta m^2 \equiv m_2^2 - m_1^2$  is mass squared difference and not mass itself

Experiments give us info on  $\theta$  and  $\Delta m^2$



• In the presence of  $\nu$  mass & mixing, flavor transition occurs.

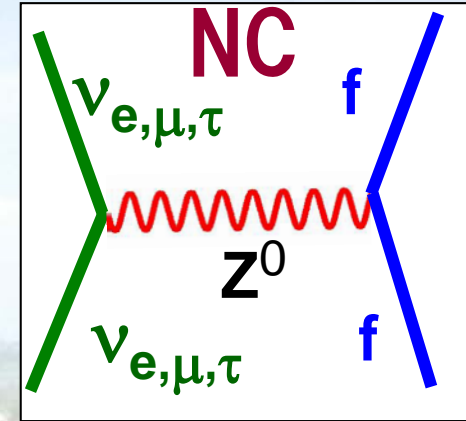
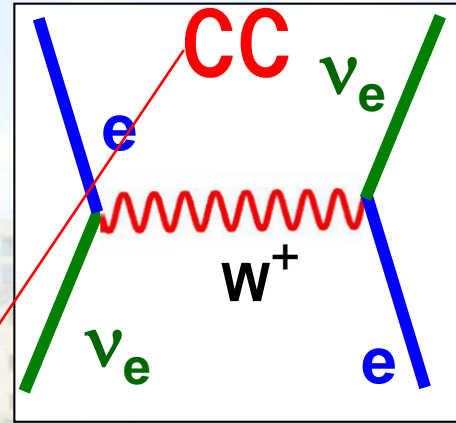
• Macroscopic distance is required to see flavor transitions.

# $\nu$ oscillation in vacuum (MSW effect)

$f = e, \mu, \tau$

$$L_{\text{eff}} = \sqrt{2}G_F \bar{\nu}_e \nu_e \bar{e} e$$

$$\rightarrow \sqrt{2}G_F \bar{\nu}_e \nu_e n_e(x)$$



$$i \frac{d}{dx} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \left[ U \begin{pmatrix} E_1 & 0 \\ 0 & E_2 \end{pmatrix} U^{-1} + \begin{pmatrix} A & 0 \\ 0 & 0 \end{pmatrix} \right] \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix}$$

$$E_j = \sqrt{\vec{p}^2 + m_j^2}$$

$$A \equiv \sqrt{2}G_F n_e(x)$$

If the density  $N_e$  is constant, we have a similar formula:

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\tilde{\theta} \sin^2 \left( \frac{\Delta \tilde{E} L}{2} \right)$$

$$\tan 2\tilde{\theta} \equiv \frac{\Delta E \sin 2\theta}{\Delta E \cos 2\theta - A}$$

$$\Delta \tilde{E} \equiv \left[ (\Delta E \cos 2\theta - A)^2 + (\Delta E \sin 2\theta)^2 \right]^{1/2}$$

$$\Delta E \equiv \Delta m^2 / 2E$$



**(4) Atmospheric  $\nu$  +  
Accelerator  $\nu$**

# Atmospheric $\nu$

● So-called primary cosmic rays are falling onto ground, and they collide w/ nucleons in the atmosphere, and produce 2ndary cosmic rays.

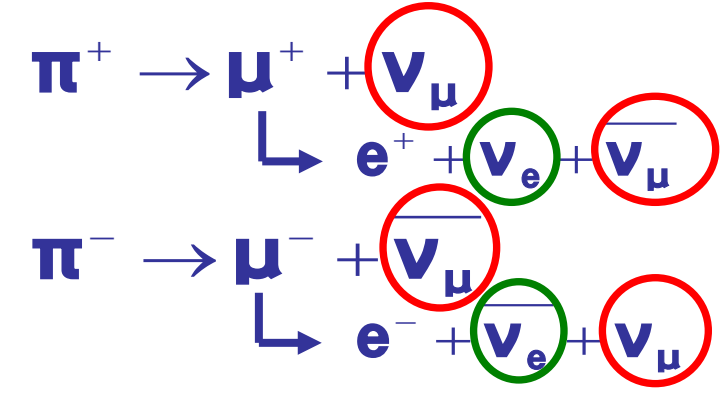
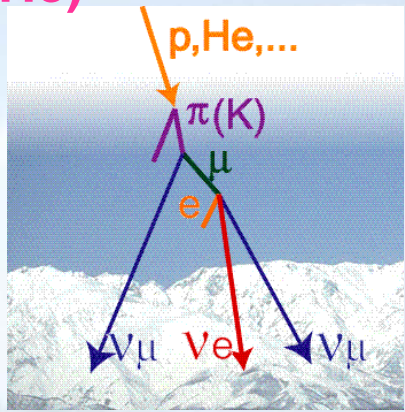
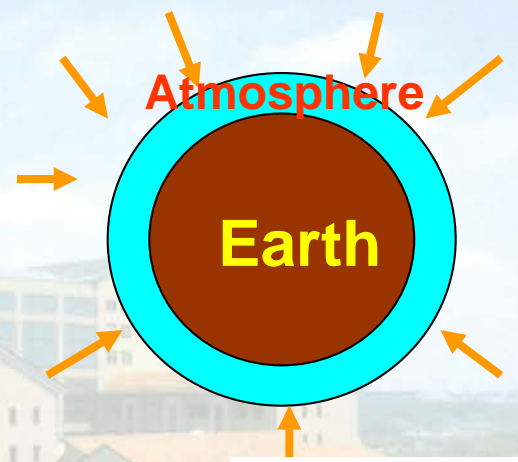
● Almost all the particles become  $\pi^\pm$  mesons, which decay into  $\mu^\pm$  and then  $\mu^\pm$  decay into electrons and positrons.

● If we ignore the difference between  $\nu$  and  $\bar{\nu}$ , then

$$(\nu_\mu + \bar{\nu}_\mu) : (\nu_e + \bar{\nu}_e) = 2:1$$

Is predicted.

Primary cosmic rays (p,He)



However, the observation was

$$(\nu_\mu + \bar{\nu}_\mu) : (\nu_e + \bar{\nu}_e) = 1.3:1$$

which disagrees w/ prediction.

# Cause of Atmospheric $\nu$ anomaly:

Because of  $\nu_{\mu} \Leftrightarrow \nu_{\tau}$  oscillation,

$\nu_{\mu}$  decreases (SK cannot observe  $\nu_{\tau}$ )

Experimental value of

$$(\overline{\nu_{\mu}} + \overline{\nu_{\tau}}) : (\overline{\nu_e} + \overline{\nu_{\mu}})$$

depends on L & E and

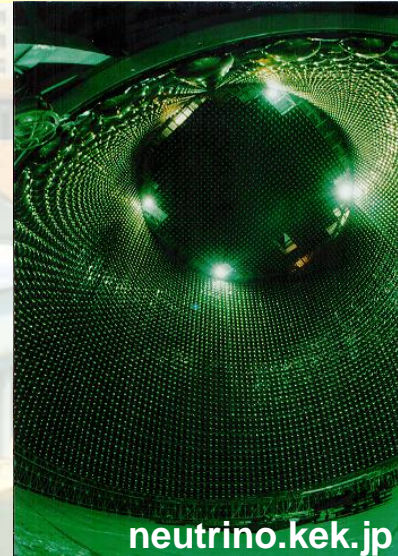
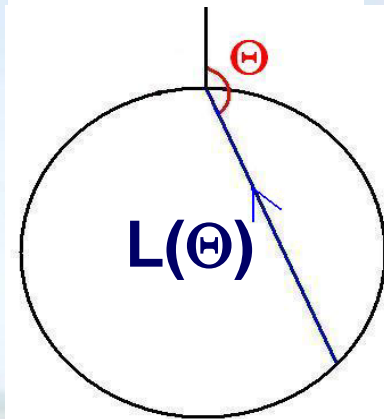
**Superkamiokande** proved that

it is consistent with the

formula

$$P(\nu_{\mu} \rightarrow \nu_{\tau}) = \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

zenith angle



[www2.kek.jp](http://www2.kek.jp)

[nobelprize.org](http://nobelprize.org)



Totsuka



Kajita



# Accelerator $\nu$

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

## Experiments in the past

- K2K (JP, KEK $\rightarrow$ SK, 1999-2004)

L=250km, E $\sim$ 1.3GeV

$$\nu_\mu \rightarrow \nu_\mu$$

- MINOS (US, FNAL $\rightarrow$ Soudan, MN, 2005-2012)

L=735km, E $\sim$ 4GeV

$$\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu + \nu_\mu \rightarrow \nu_e$$

- OPERA (CH, CERN $\rightarrow$ Gransasso, IT, 2010-2018)

L=730km, E $\sim$ 17GeV

$$\nu_\mu \rightarrow \nu_\tau$$

## Experiments in operation

$$\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu + \bar{\nu}_\mu \rightarrow \bar{\nu}_e$$

- T2K(JP, JPARC  $\rightarrow$  SK, 2009-) L=295km, E $\sim$ 0.6GeV

- MINOS+(US, FNAL  $\rightarrow$  Soudan, MN, 2013-)L=735km, E $\sim$ 4GeV

- Nova(US, FNAL  $\rightarrow$  Ash River, MN, 2014-), L=810km, E $\sim$ 2GeV



All the results are consistent with the atmospheric  $\nu$  experiments

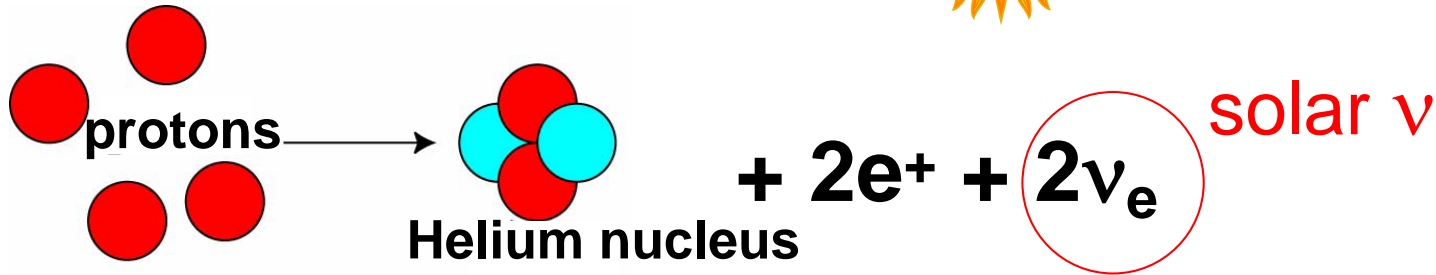


**(5) Solar  $\nu$  + Long baseline reactor  $\nu$**

# Solar $\nu$ deficit



## The fusion reaction



produces electron neutrinos: They are called solar  $\nu$

● Solar  $\nu$  were detected since 1960's by Davis at Homestek, SD. Observed flux was less than  $\frac{1}{2}$  of theoretical prediction: Solar  $\nu$  problem

It turned out that flux of  $\nu$  is reduced due to conversions  $\nu_e \rightarrow \nu_\mu$ ,  $\nu_e \rightarrow \nu_\tau$

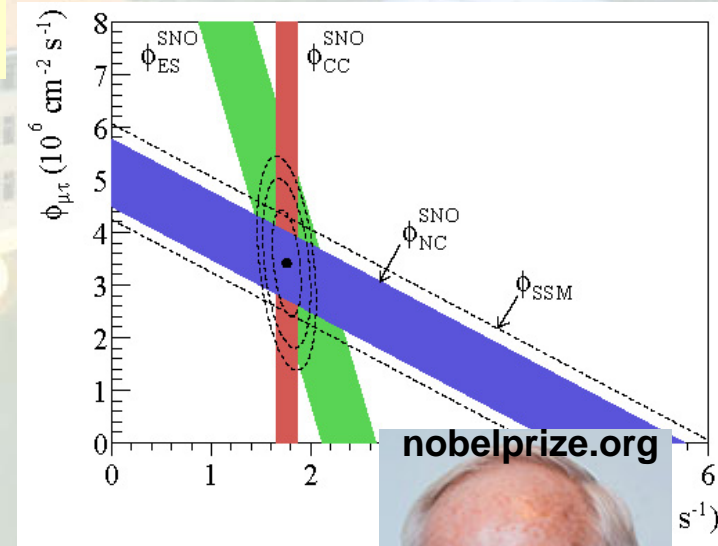
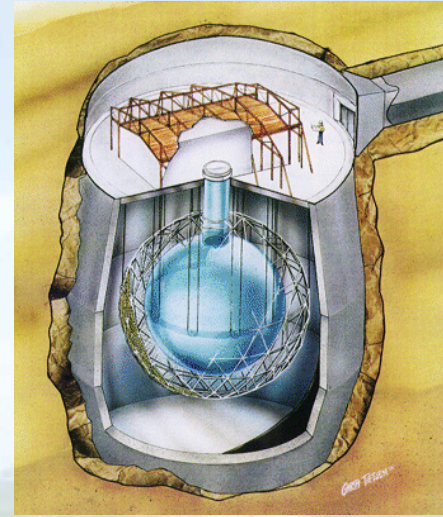


# SNO (Sudbury Neutrino Observatory, 1999-2006)

- Detector w/ heavy water (1 kt) D<sub>2</sub>O, d=(pn), neutron
- Underground laboratory (~2km) (To reduce Background)
- Direct proof for solar ν deficit  
SNO can detect the both reactions:



x=e,μ,τ



From the data of these 2 reactions, it was concluded that  $\nu_e + \nu_\mu + \nu_\tau$  agrees w/ theory, but  $\nu_e$  is less than theory



McDonald

# KamLAND (JP, 2002-, long baseline reactor $\nu$ )

$$\bar{\nu}_e \rightarrow \bar{\nu}_e$$

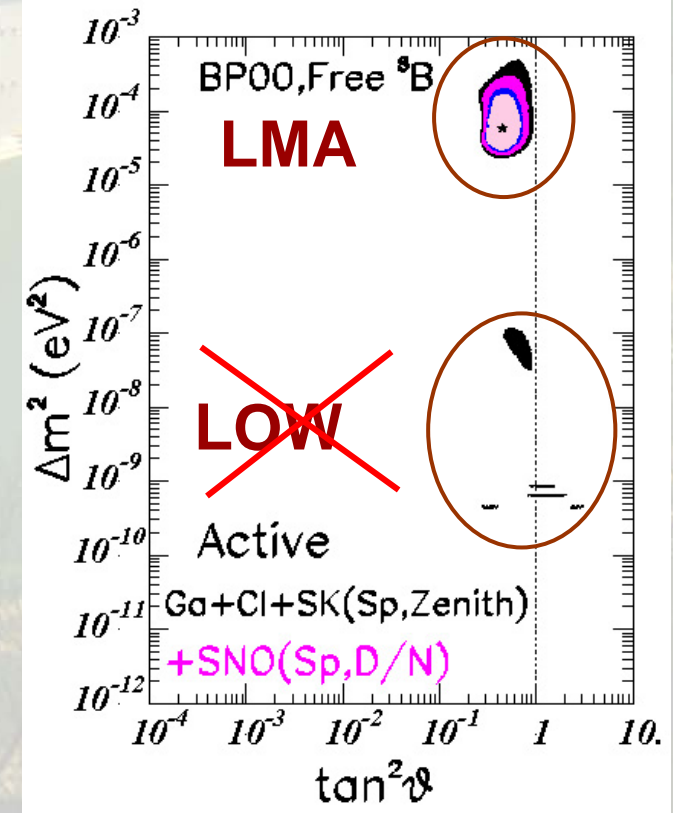
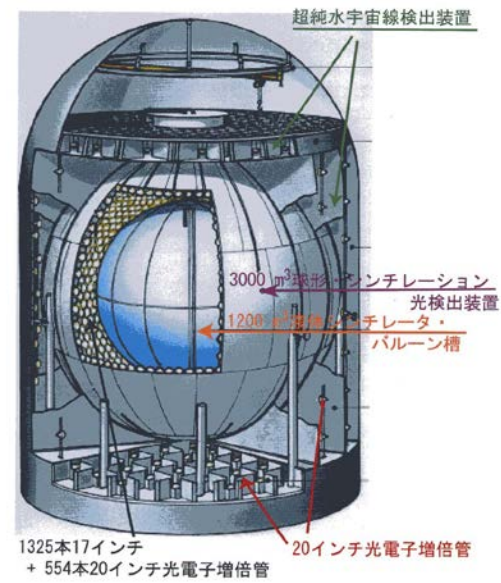
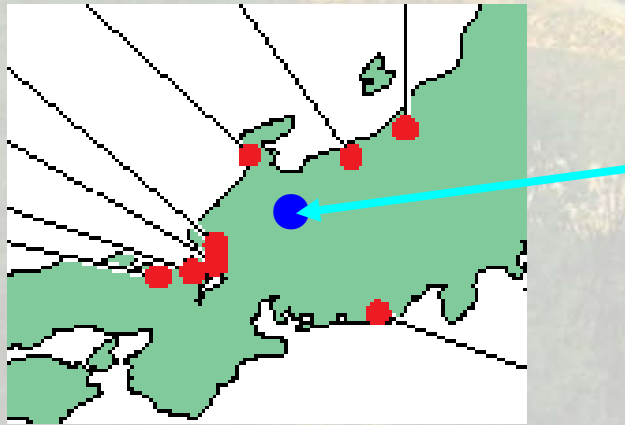
$$L \sim 200\text{km}, E \sim 4\text{MeV}$$



Atsuto Suzuki

- Detector w/ liquid scintillator
- Detected  $\bar{\nu}_e$  from various nuclear power plants (average distance 200km)
- Observed deficit of reactor neutrinos for the 1<sup>st</sup> time

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$



An aerial photograph of a modern university campus. The central focus is a large, multi-story building with a prominent curved section and a covered walkway. To the right of this building is a large, irregularly shaped pond. The campus is surrounded by greenery and other buildings in the background under a blue sky with scattered clouds.

# **(6) Short baseline reactor $\nu$**

# Reactor $\bar{\nu}$ (short baseline)

$$\bar{\nu}_e \rightarrow \bar{\nu}_e$$

$L \sim 2\text{km}, E \sim 4\text{MeV}$

Double CHOOZ (Fr) (2016/3)

$$\sin^2 2\theta = 0.111 \pm 0.018$$

Daya Bay (Cn) (2015/5)

$$\sin^2 2\theta = 0.084 \pm 0.005$$

Reno (Kr) (2015/12)

$$\sin^2 2\theta = 0.082 \pm 0.011$$

$\theta=0$  is excluded at  $168\sigma$

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

An aerial photograph of a university campus. In the foreground, there is a large, irregularly shaped pond with a light-colored path or walkway around its edge. Several multi-story buildings with light-colored facades and red-tiled roofs are visible in the middle ground. The background shows more campus buildings and a clear blue sky with some light clouds.

# **(7) 3 flavor neutrino oscillation**



# 3 flavor mixing framework (in the real world)

$$\begin{pmatrix} \mathbf{v}_e \\ \mathbf{v}_\mu \\ \mathbf{v}_\tau \end{pmatrix} = \mathbf{U} \begin{pmatrix} \mathbf{v}_1 \\ \mathbf{v}_2 \\ \mathbf{v}_3 \end{pmatrix}$$

Mixing matrix is 3x3

$$\mathbf{U} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \mathbf{C}_{23} & \mathbf{S}_{23} \\ 0 & -\mathbf{S}_{23} & \mathbf{C}_{23} \end{pmatrix} \begin{pmatrix} \mathbf{C}_{13} & 0 & \mathbf{S}_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\mathbf{S}_{13} e^{i\delta} & 0 & \mathbf{C}_{13} \end{pmatrix} \begin{pmatrix} \mathbf{C}_{12} & \mathbf{S}_{12} & 0 \\ -\mathbf{S}_{12} & \mathbf{C}_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

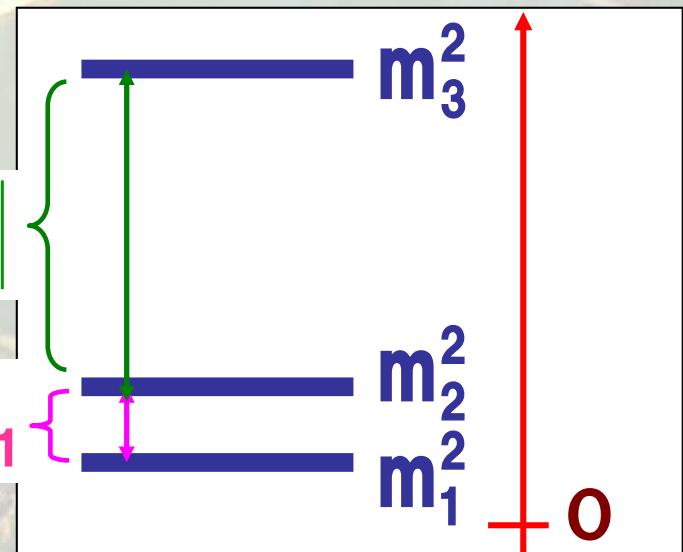
There are 2 independent mass squared differences

$\theta_{12}, \theta_{23}, \theta_{13}$ : mixing angle

$\delta$ : CP violating phase

$$|\Delta m_{32}^2|$$

$$\Delta m_{21}^2$$



# Features of 3 flavor mixing framework

## (1) Mass hierarchy

$$\Delta m_{21}^2 \ll |\Delta m_{32}^2| \cong |\Delta m_{32}^2|$$

→ Oscillation probabilities are simplified

## (2) Small $\theta_{13}$

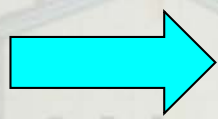
$$\sin^2 2\theta_{13} = 0.08$$

→ In the 0-th approximation, we can work with  $\theta_{13} = 0$

→ We have further simplification.

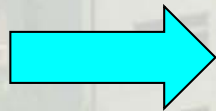
# Determination of 3 $\nu$ oscillation parameters

(i) Solar  $\nu$  deficit + Long baseline reactor  $\nu$  deficit (KamLAND)



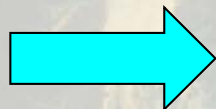
$$\Delta m_{21}^2 \sim 8 \times 10^{-5} \text{eV}^2 \quad \sin^2 2\theta_{12} \sim 0.8$$

(ii) Atmospheric  $\nu$  anomaly + Accelerator  $\nu$  oscillation (K2K, MINOS, OPERA, T2K, Nova)



$$|\Delta m_{32}^2| \sim 3 \times 10^{-3} \text{eV}^2 \quad \sin^2 2\theta_{23} \sim 1$$

(iii) Short baseline reactor  $\nu$  deficit (Double CHOOZ, Daya Bay, RENO) + Accelerator  $\nu$  appearance (T2K, MINOS, Nova)

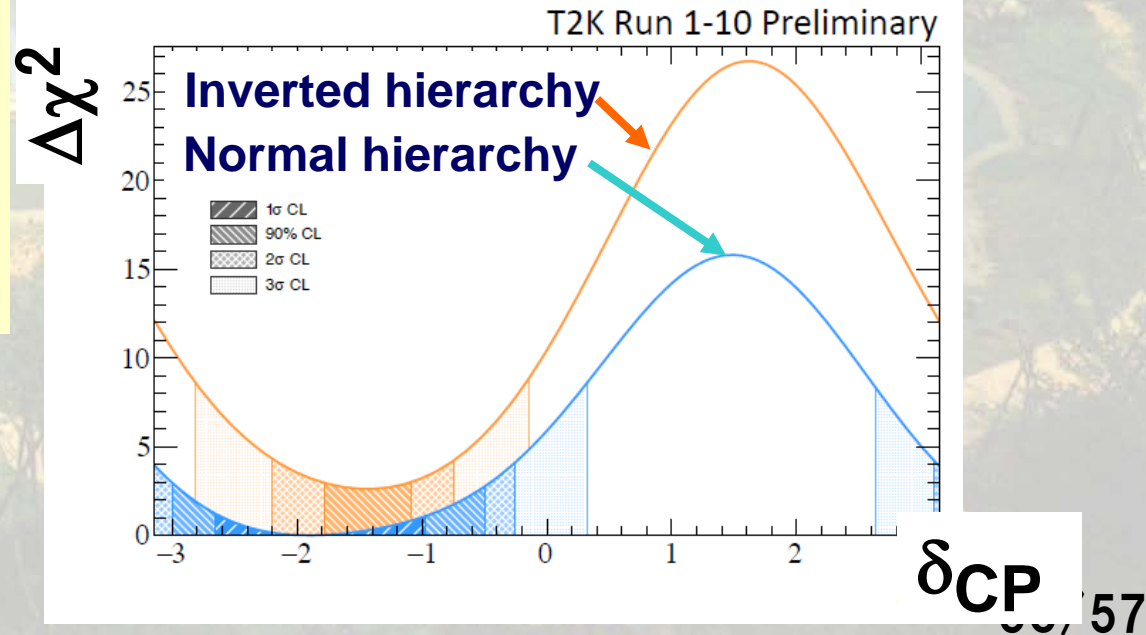
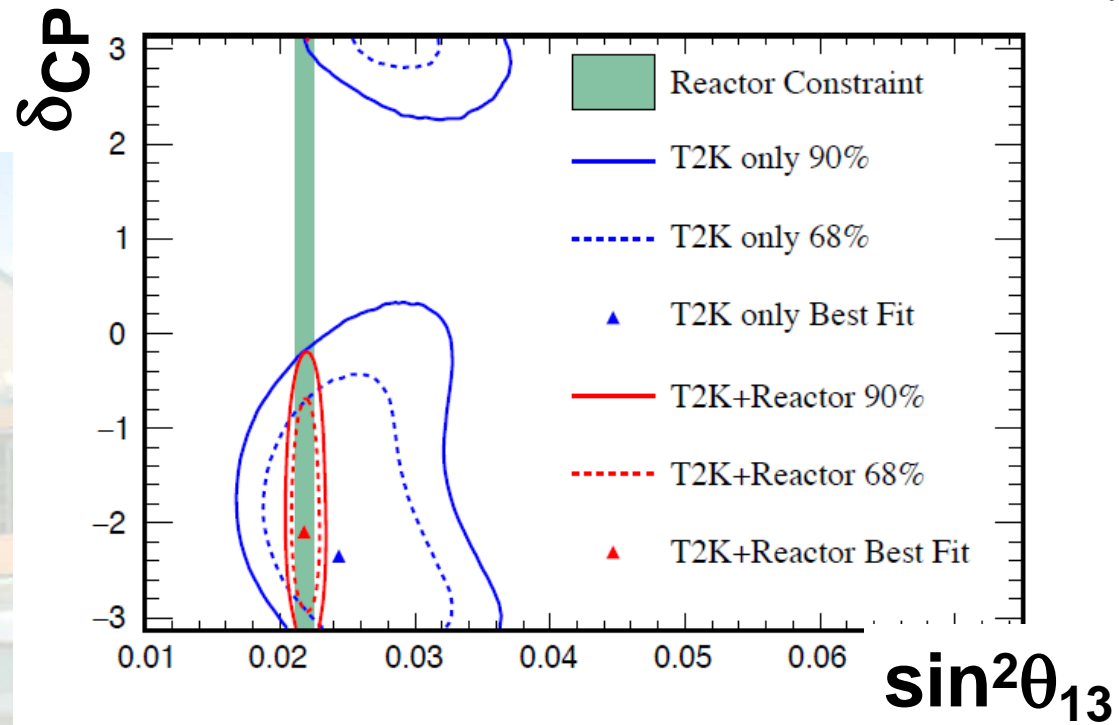


$$\sin^2 2\theta_{13} = 0.08$$

# Recent T2K results

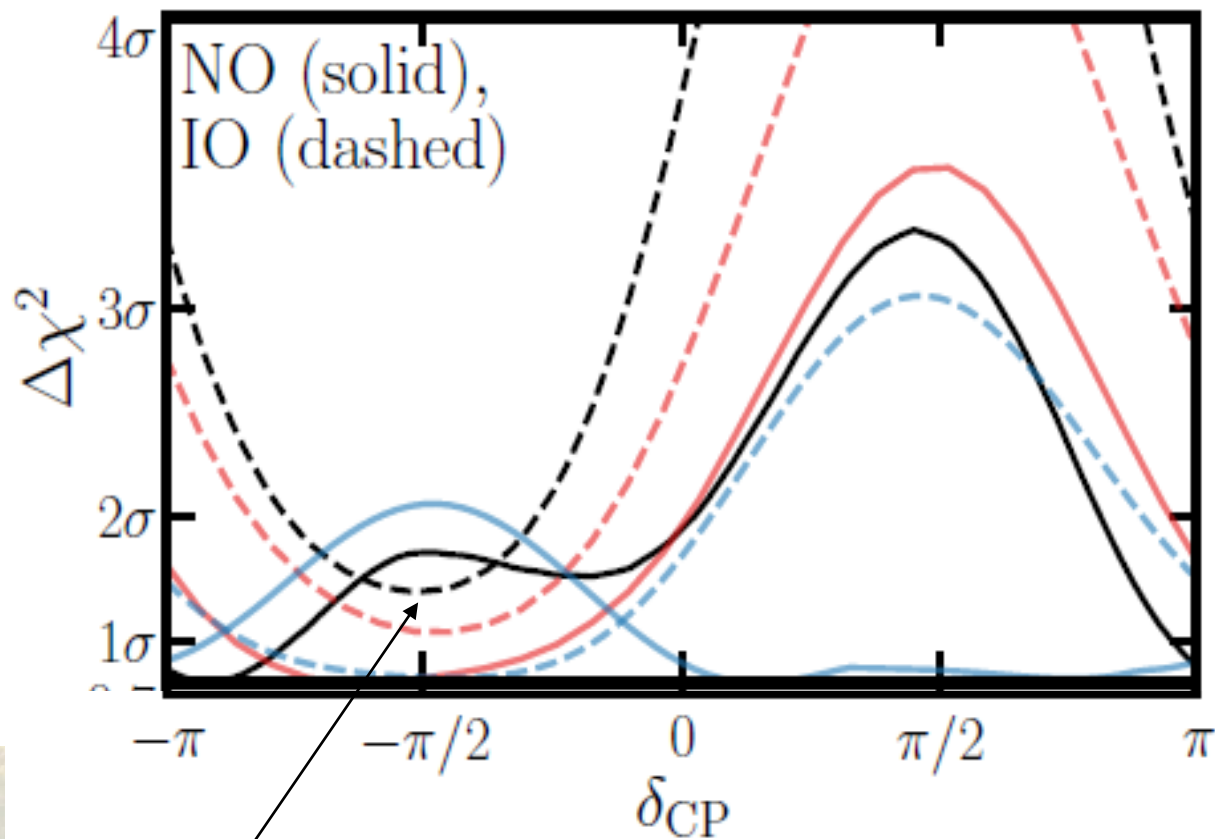
Dunne@Neutrino2020

Normal hierarchy &  $\delta_{CP} \sim -\pi/2$  seems to be favored, but we need more data to conclude



# Recent status: Tension between T2K and Nova?

Kelly et al,  
arXiv:2007.08526v1 [hep-ph]



Black lines: a joint fit of  
T2K/NOvA/SK18  
Blue: NOvA alone  
Red: T2K alone

Joint fit may indicate preference  
for Inverted Hierarchy

# Present status of 3 flavor mixing framework

$$N_\nu=3 : \nu_{\text{atm}} + \nu_{\text{solar}} + \nu_{\text{reactor}} + \nu_{\text{accelerator}}$$

## Mixing matrix

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \cong \begin{pmatrix} c_{12} & s_{12} & \epsilon \\ -s_{12}/\sqrt{2} & c_{12}/\sqrt{2} & 1/\sqrt{2} \\ s_{12}/\sqrt{2} & -c_{12}/\sqrt{2} & 1/\sqrt{2} \end{pmatrix}$$

## Mixing angles & mass squared differences

$$\theta_{12} \cong \pi/6, \quad \theta_{23} \cong \pi/4$$

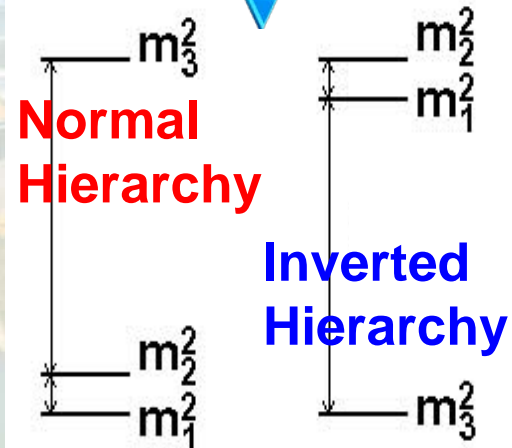
$$\theta_{13} \cong \pi/20$$

$$\Delta m_{21}^2 \cong 8 \times 10^{-5} \text{ eV}^2$$

$$|\Delta m_{32}^2| \cong 2.5 \times 10^{-3} \text{ eV}^2$$

● Less known parameters :  
 $\delta_{\text{CP}}$ ,  $\text{sign}(\Delta m_{31}^2)$

Both hierarchy patterns are allowed



$$\Delta m_{32}^2 > 0$$

$$\Delta m_{32}^2 < 0$$

An aerial photograph of a university campus. In the foreground, a large, irregularly shaped pond is surrounded by a paved walkway and greenery. Behind the pond, several multi-story buildings with light-colored facades and red-tiled roofs are visible. The sky is blue with scattered white clouds. A yellow rectangular box is overlaid on the image, containing the text '(8) Future plans' in bold black font.

## **(8) Future plans**

## Future plans

Next task is to measure  $\text{sign}(\Delta m_{31}^2)$  and  $\delta_{\text{CP}}$  precisely

## Proposed experiments

$$\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{\mu} + \overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$$

- T2HK(JP, JPARC-->HK) L=295km, E~0.6GeV
- DUNE (US, FNAL-->Homestake, SD), E~2GeV, L=1300km

These experiments are expected to measure  $\text{sign}(\Delta m_{31}^2)$  and  $\delta_{\text{CP}}$



# Future plan: T2HK

## ● Phase 2

1.66MW  $\nu$  beam  $\Rightarrow$  Hyperkamiokande  
(300 times K2K)  $\Rightarrow$  (20 times SK)

## ● Extension of T2K

## ● Measurement of CP phase $\delta_{CP}$

t2k-experiment.org

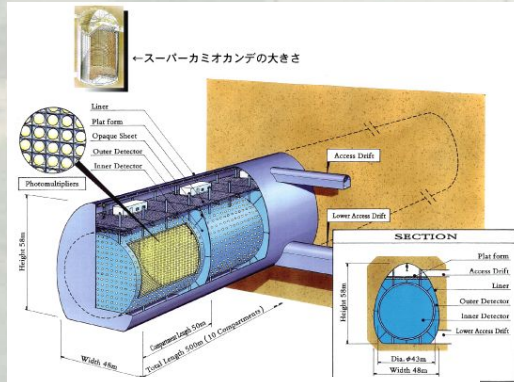
www.he.scphys.kyoto-u.ac.jp



Kobayashi



Nakaya



Hyper-kamiokande



J-PARC Main Ring  
(KEK-JAEA, Tokai)



# Future plan: DUNE

2.3MW  $\nu$  beam @ Fermilab  
 $\Rightarrow$  40-kt Liquid Argon  
detector @ Sanford  
Underground RF

$E \sim 2\text{GeV}$ ,  $L \sim 1300\text{km}$



## Deep Underground Neutrino Experiment

Sanford Underground  
Research Facility  
Lead, South Dakota

Fermilab  
Batavia, Illinois

20 miles  
800 miles

naturalsciences.ch www.hep.phy.cam.ac.uk



A. Rubbia



Thomson

An aerial photograph of a university campus. In the foreground, a large, dark blue lake is surrounded by green trees and a winding path. In the middle ground, several modern, multi-story buildings with light-colored facades and red-tiled roofs are visible. The background shows a clear blue sky with some light clouds and distant hills.

## **(9) Beyond the standard scenario**

# Nonstandard scenarios

High precision measurements of  $\nu$  oscillation in future experiments can be used also to probe physics beyond SM by looking at deviation from SM+massive  $\nu$

T2HK, DUNE,  $\nu_{\text{atm}}$  @Hyperkamiokande

# New Physics discussed in this talk

Scenario beyond SM+m <sub><math>\nu</math></sub>	Experimental indication ?	Phenomenological constraints on the magnitude of the effects
(1) Light sterile $\nu$	Maybe	O(10%)
(2) Non Standard Interaction	Maybe	e- $\tau$ : O(100%) Others: O(1%)

**Neither sterile  $\nu$  nor Non Standard Interaction is required from theory.**

**→ They are introduced phenomenologically.**

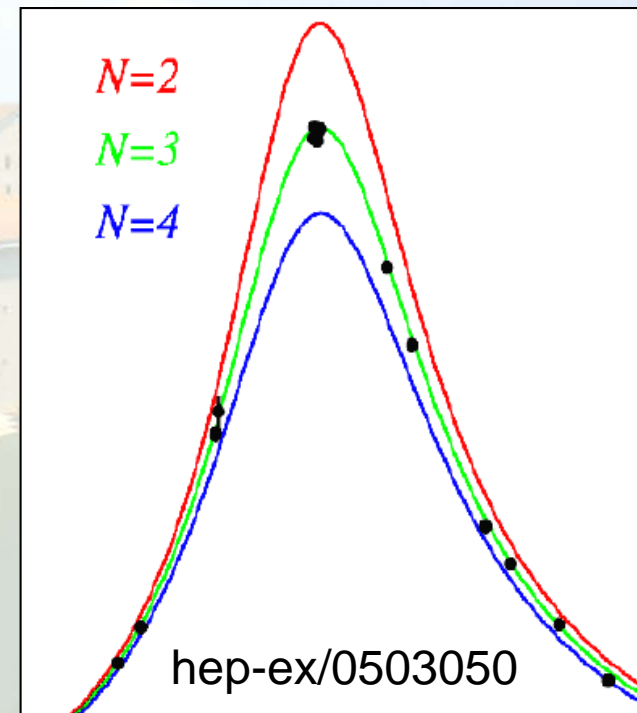
# (1) Light sterile neutrinos ( $\nu_s$ )

## ● Motivation for $\nu_s$

A) 4<sup>th</sup> neutrino mass eigenstate has been phenomenologically motivated by the following affirmative results:

- LSND anomaly ( $E \sim 50\text{MeV}$ ,  $L \sim 30\text{m}$ )
- Reactor anomaly ( $E \sim 4\text{MeV}$ ,  $L < 10\text{m}$ )
- Gallium anomaly ( $E < 1\text{MeV}$ ,  $L < 5\text{m}$ )

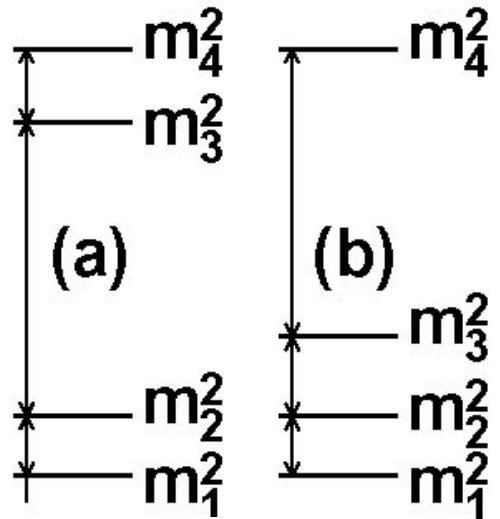
B) From LEP result,  $\#(\nu \text{ coupled to } Z)=3$



A)  $\rightarrow \Delta m^2 \sim O(1)\text{eV}^2 \gg \Delta m^2(\text{atm}) \gg \Delta m^2(\text{solar})$

B)  $\rightarrow$  4<sup>th</sup>  $\nu$  flavor eigenstate has to be sterile (i.e., it has no weak interaction)

# Mass pattern for sterile neutrinos ( $\nu_s$ )



(2+2)-scheme

(3+1)-scheme

(a): (2+2)-scheme is completely excluded by

$\nu_{\text{solar}}$  &  $\nu_{\text{atm}}$

(b): (3+1)-scheme has tension between  $\nu_{\mu} \rightarrow \nu_{\mu} + \nu_e \rightarrow \nu_e$  &

$\nu_{\mu} \rightarrow \nu_e$

$$\Delta m_{41}^2 = \Delta m_{\text{LSND}}^2 \cong 1 \text{ eV}^2$$

$$\Delta m_{21}^2 = \Delta m_{\text{solar}}^2, \Delta m_{32}^2 = \Delta m_{\text{atm}}^2$$

## (2+2)-scheme

$$\eta_s \equiv |\mathbf{U}_{s1}|^2 + |\mathbf{U}_{s2}|^2 \rightarrow 0$$

$$\mathbf{v}_{\text{atm}} : \mathbf{v}_\mu \rightarrow \mathbf{v}_s \text{ (100\%)}$$

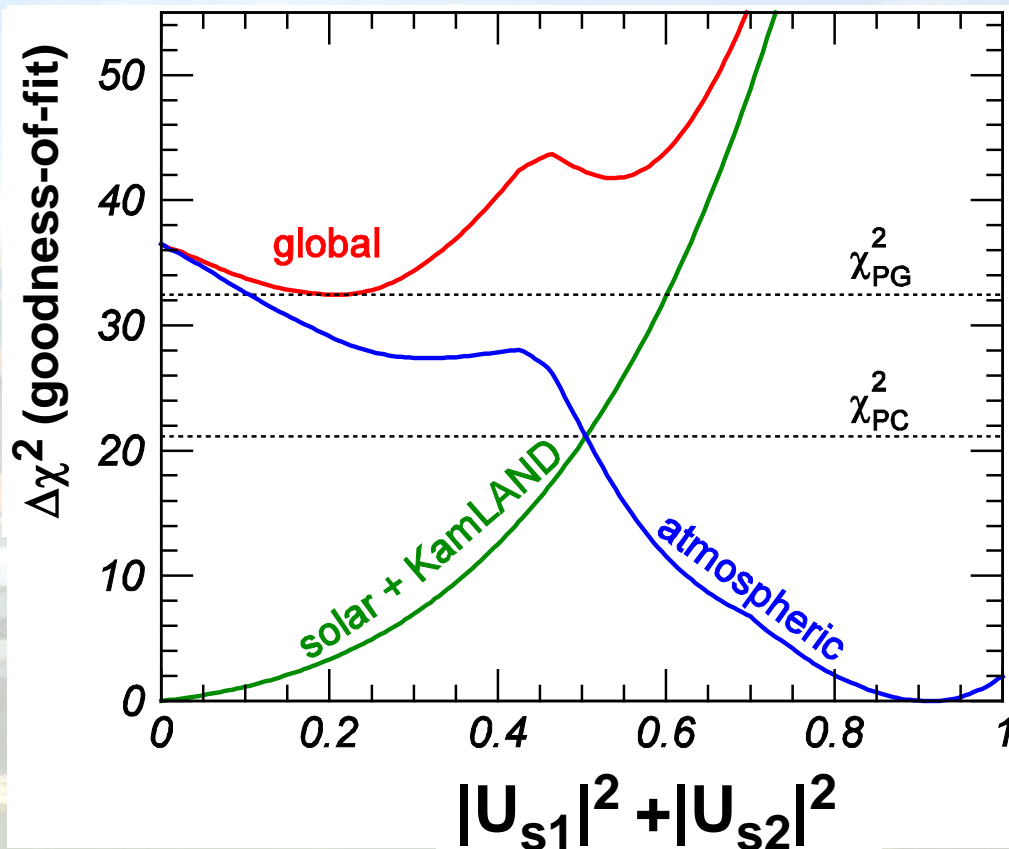
Strongly disfavored  
by SK  $\nu_{\text{atm}}$  data

$$\eta_s \equiv |\mathbf{U}_{s1}|^2 + |\mathbf{U}_{s2}|^2 \rightarrow 1$$

$$\mathbf{v}_{\text{sol}} : \mathbf{v}_e \rightarrow \mathbf{v}_s \text{ (100\%)}$$

Strongly disfavored  
by SNO  $\nu_{\text{sol}}$  data

For any value of  $|\mathbf{U}_{s1}|^2 + |\mathbf{U}_{s2}|^2$ , fit to sol+atm data is bad.



PC: parameter consistency test  
PG: parameter goodness-of-fit test



# (3+1)-scheme

## Bugey (reactor): negative

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - 4|U_{e4}|^2(1 - |U_{e4}|^2) \sin^2(\Delta m_{41}^2 L/4E)$$

$$\sin^2 2\theta_{\text{Bugey}} > 4|U_{e4}|^2(1 - |U_{e4}|^2) \cong 4|U_{e4}|^2$$

## CDHSW (accelerator): negative

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - 4|U_{\mu 4}|^2(1 - |U_{\mu 4}|^2) \sin^2(\Delta m_{41}^2 L/4E)$$

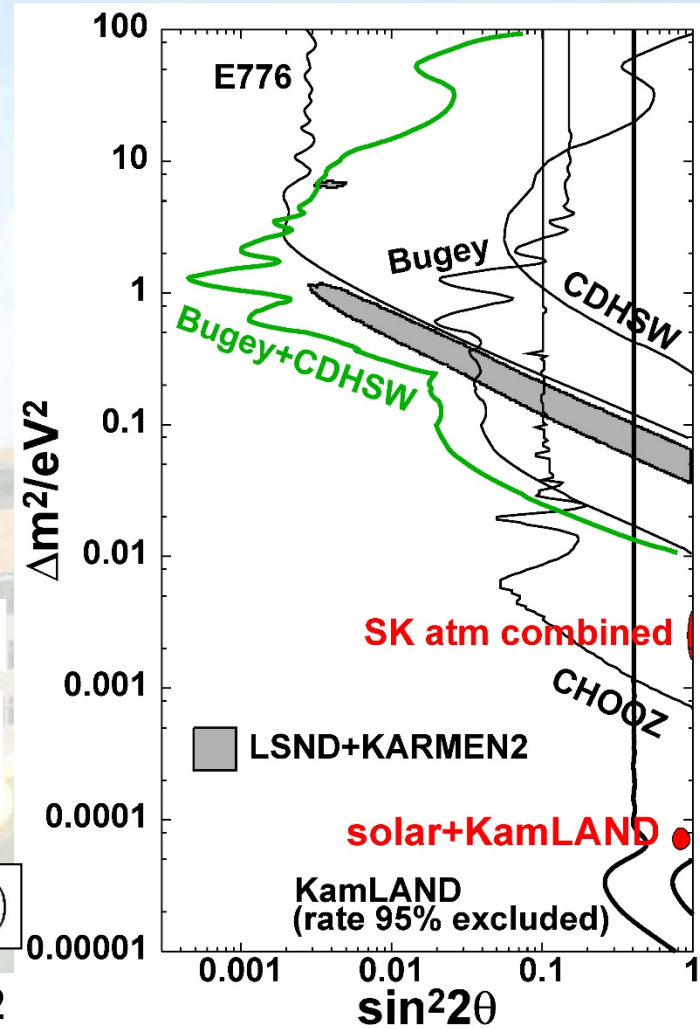
$$\sin^2 2\theta_{\text{CDHSW}} > 4|U_{\mu 4}|^2(1 - |U_{\mu 4}|^2) \cong 4|U_{\mu 4}|^2$$

## LSND (accelerator): affirmative

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = 4|U_{e4}|^2|U_{\mu 4}|^2 \sin^2(\Delta m_{41}^2 L/4E)$$

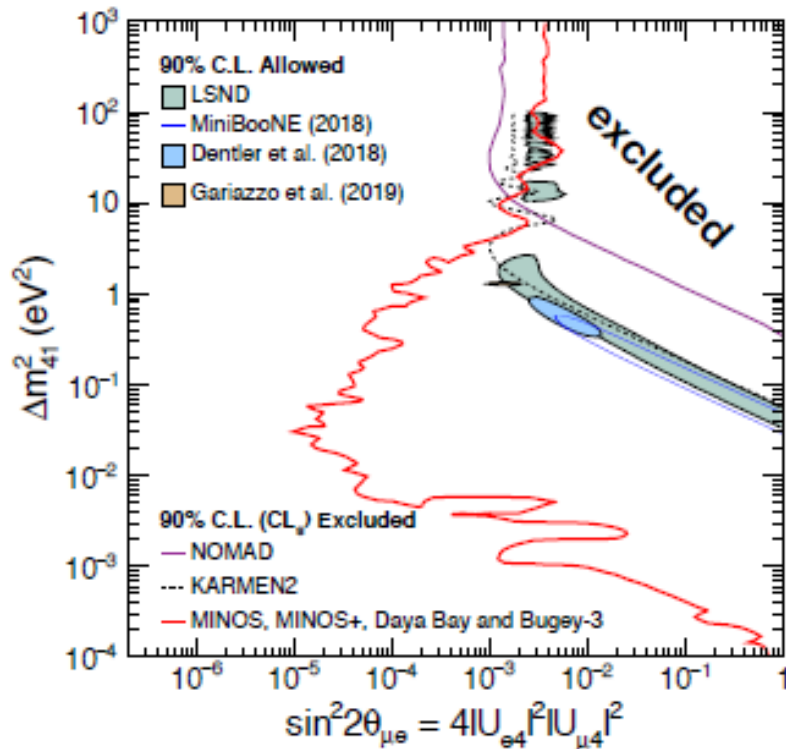
$$\sin^2 2\theta_{\text{LSND}} = 4|U_{e4}|^2|U_{\mu 4}|^2$$

$$\sin^2 2\theta_{\text{LSND}}(\Delta m^2) < \frac{1}{4} \sin^2 2\theta_{\text{Bugey}}(\Delta m^2) \sin^2 2\theta_{\text{CDHSW}}(\Delta m^2)$$

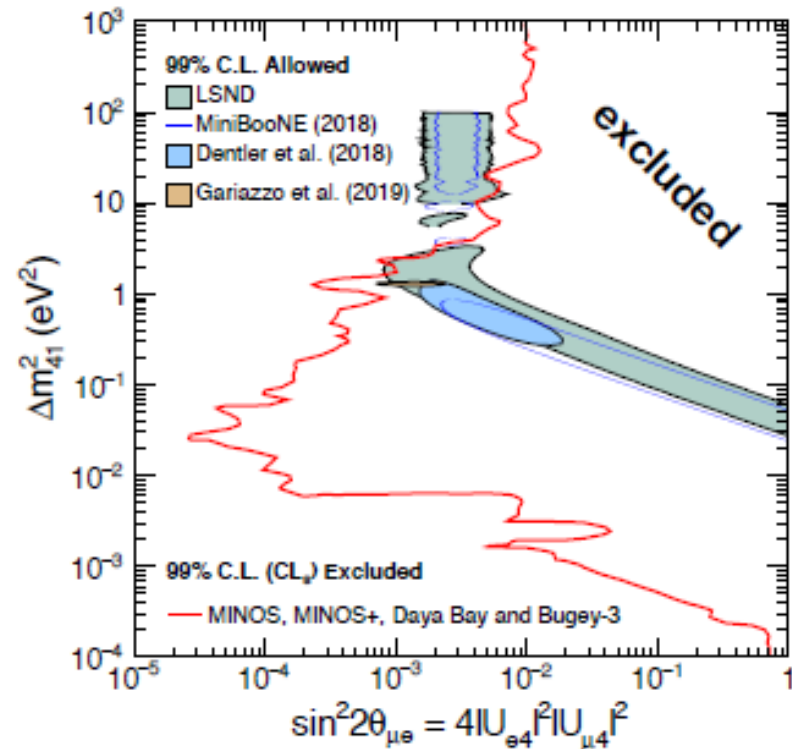


must be satisfied but there is no overlap between the left side of **Bugey+CDHSW** and the inside of **LSND** (Okada-OY Int.J.Mod.Phys.A12:3669,1997)

90% C.L.



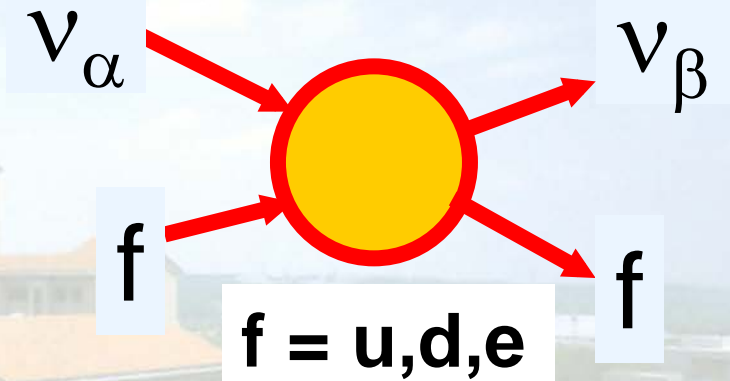
99% C.L.



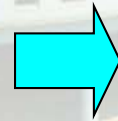
arXiv:2002.00301 (accepted by PRL)

## (2) Nonstandard Interactions

$$\mathcal{L}_{eff} = G_{NP}^{\alpha\beta} \bar{\nu}_\alpha \gamma^\mu \nu_\beta \bar{f} \gamma_\mu f'$$



Matter potential is modified  
by **Physics Beyond the  
Standard Model**



$\nu$  experiments can give  
constraints or hints on  
**Physics BSM**

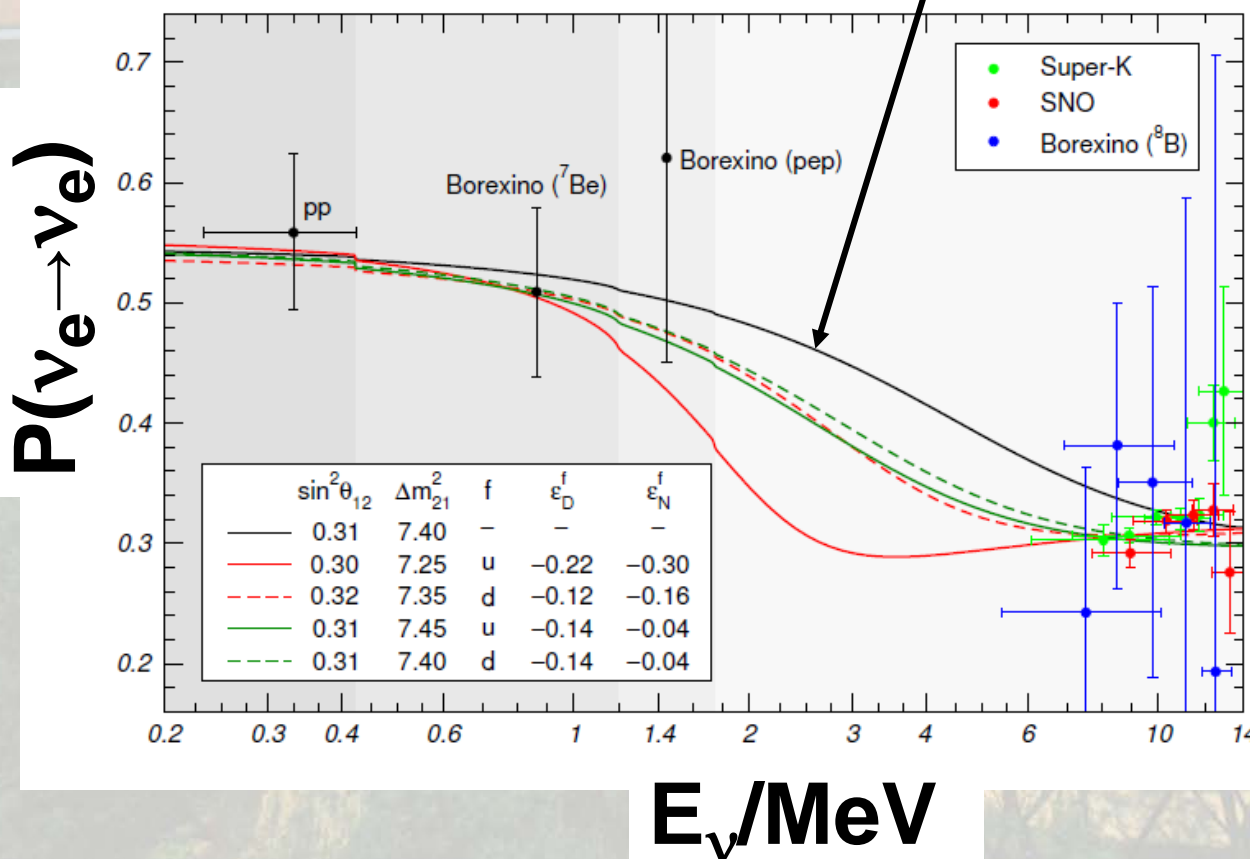
$$\mathbf{SM} \quad \mathcal{A}_0 \equiv A \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \quad \rightarrow \quad \mathbf{NP} \quad \mathcal{A} \equiv A \begin{pmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{pmatrix}$$

$$A \equiv \sqrt{2} G_F N_e \quad N_e \equiv \text{electron density}$$

# Motivation for Non Standard Interactions

Tension between solar  $\nu$  & KamLAND data comes from little observation of **upturn** by SK & SNO

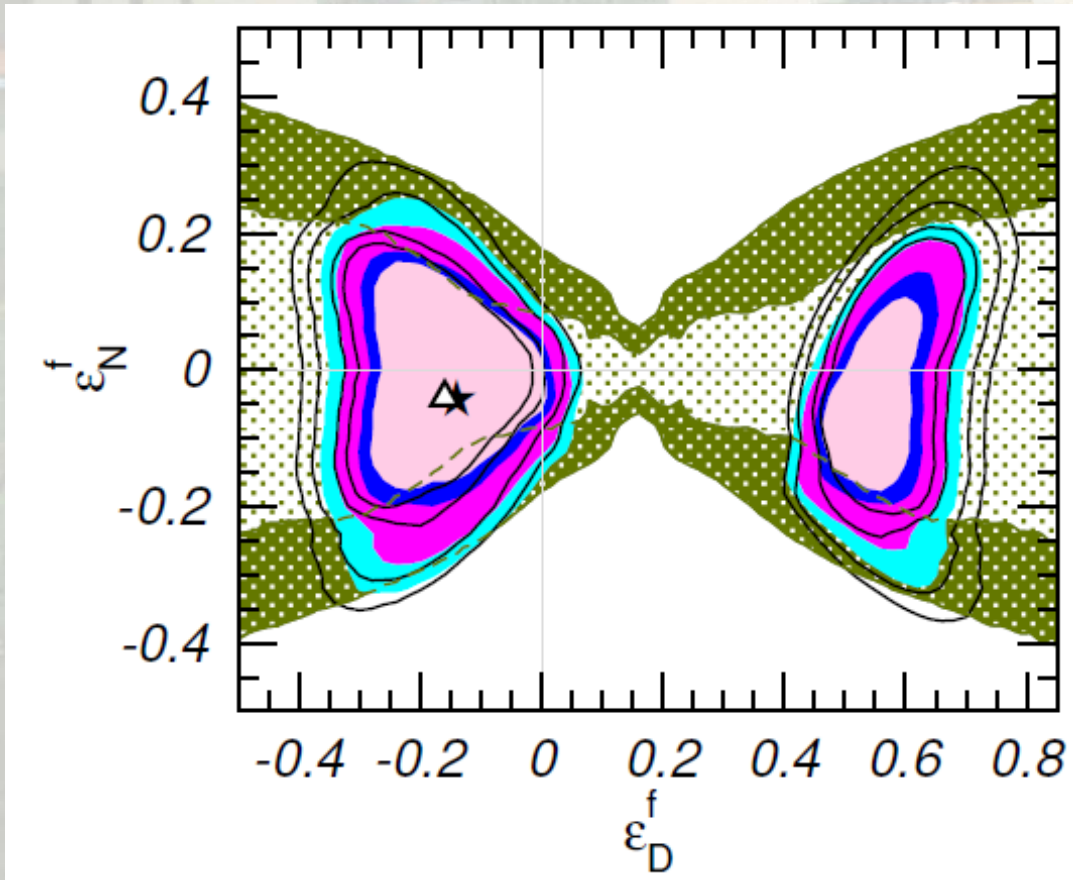
Standard scenario w/  $\Delta m^2_{21}$  by KamLAND



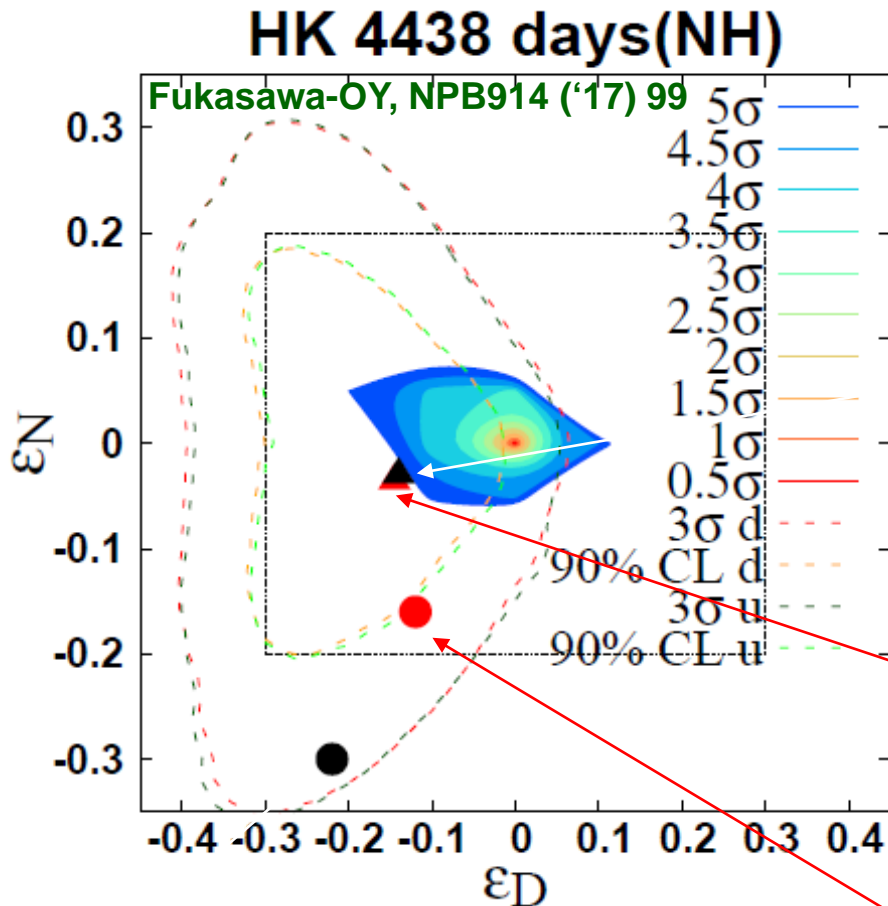
Gonzalez-Garcia,  
Maltoni, JHEP  
1309 (2013) 152

The tension between solar  $\nu$  & KamLAND data may be resolved by Non Standard Interaction.

Gonzalez-Garcia, Maltoni, JHEP 1309 (2013) 152



# Sensitivity of future experiment HK $\nu_{\text{atm}}$ to NSI



HK  $\nu_{\text{atm}}$  has sensitivity to some region of the  $\nu_{\text{solar}}$  anomaly

$$(\epsilon_D^u, \epsilon_N^u) = (-0.140, -0.030)$$

Best fit point of global analysis for  $f=u$ : significance:5 $\sigma$

$$(\epsilon_D^d, \epsilon_N^d) = (-0.145, -0.036)$$

Best fit point of global analysis for  $f=d$ : significance:5 $\sigma$

$$(\epsilon_D^d, \epsilon_N^d) = (-0.12, -0.16)$$

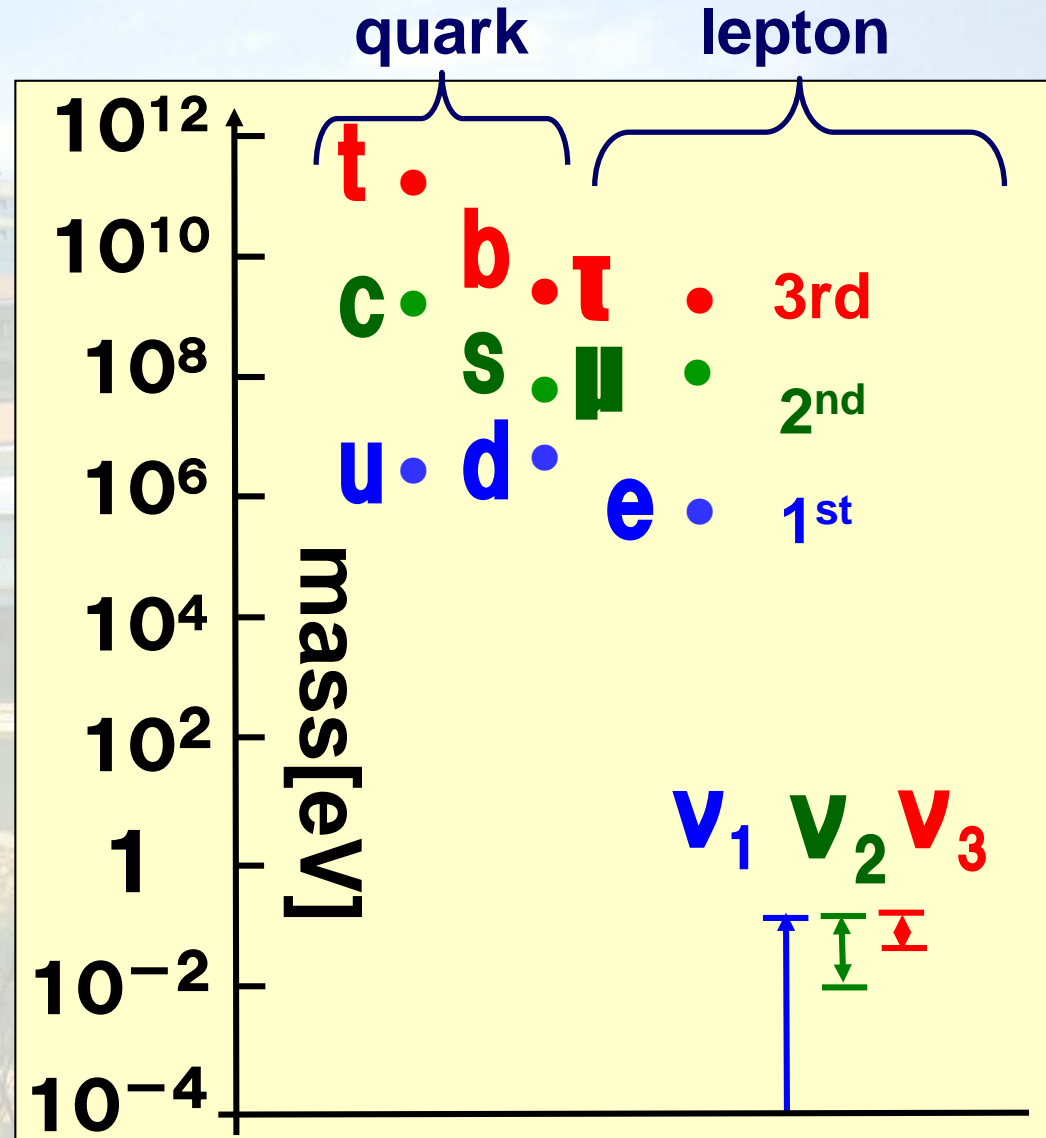
Best fit point of solar & KamLAND for  $f=d$ : significance:11 $\sigma$

$$(\epsilon_D^u, \epsilon_N^u) = (-0.22, -0.30)$$

Best fit point of solar & KamLAND for  $f=u$ : significance:38 $\sigma$

# Implication of discovery of $\nu$ mass

- $\nu$  mass is evidence of physics **Beyond the Standard Model** → It gives us a clue for BSM
- $\nu$  mass is much smaller than that of other quarks & leptons → New mystery for hierarchy
- $\delta_{CP}$  stands for difference between  $\nu$  &  $\bar{\nu}$  → It is expected to give us a clue on matter-anti-matter asymmetry of our Universe



# Physics beyond the Standard Model & $\nu$

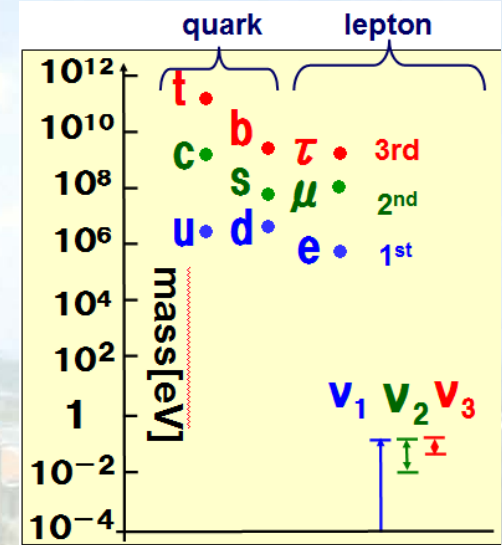
● Small  $\nu$  mass --> Hint for physics at high energy scale?

Seesaw mechanism

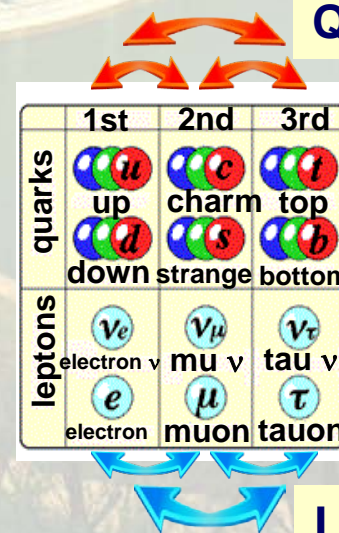
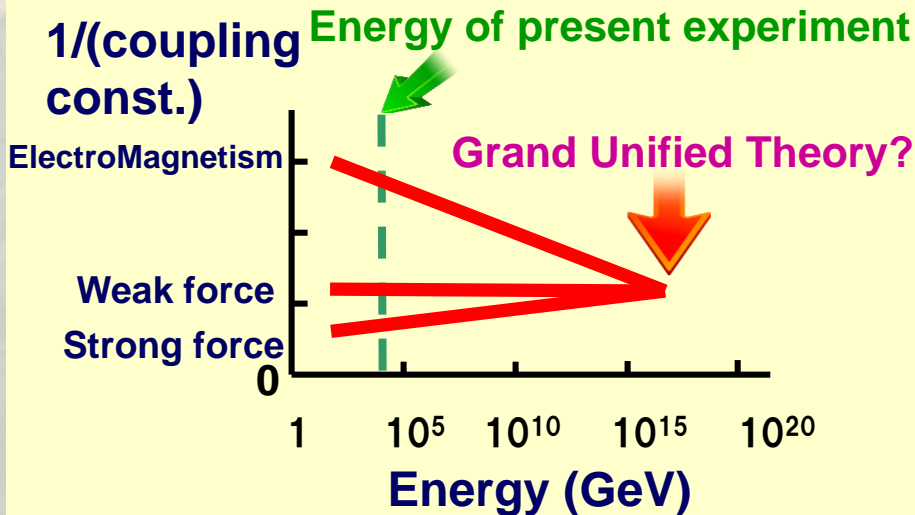
If  $1\text{ GeV} \sim m \ll M$

$$\begin{pmatrix} 0 & m \\ m & M \end{pmatrix} \rightarrow \begin{pmatrix} -m^2/M & 0 \\ 0 & M \end{pmatrix}$$

$$m_\nu = m^2/M < 1 \text{ eV} \rightarrow M > 10^9 \text{ GeV}$$



● Relation between quark & lepton mixings-->Symmetry at high energy?



If Grand Unified Theory is chosen by Nature, the mixings must coincide



# Summary

- From various  $\nu$  oscillation experiments, 3 mixing angles and 2 mass squared difference have been determined. Undetermined parameters are  $\delta$  &  $\text{sign}(\Delta m^2_{31})$ .
- Future experiments are planned to determine  $\delta$  &  $\text{sign}(\Delta m^2_{31})$ .
- New physics can be investigated at  $\nu$  oscillation experiments by looking for deviation from the standard scenario.

An aerial photograph of a university campus. In the foreground, a large, multi-story building with a light-colored facade and a red-tiled roof is visible. The building has a prominent entrance with a curved portico supported by columns. To the right of the building is a large, irregularly shaped pond with a dark, still surface. The pond is surrounded by a landscaped area with trees and a winding path. In the background, several other buildings of varying heights and styles are visible, along with a clear blue sky with scattered white clouds. The overall scene is a well-maintained and modern university campus.

**Backup slides**

## In natural units

$$\hbar = c = 1 \quad (1)$$

In the units (1), every quantity can be expressed in terms of power of mass or power of length. In the units (1), we have

$$\begin{aligned} 1 &= \hbar c = 0.197 \text{ GeV} \cdot \text{fm} \\ &= 0.197 \times 10^{9-15} \text{ eV} \cdot \text{m} = 0.197 \times 10^{-6} \text{ eV} \cdot \text{m} \\ &= 0.197 \times 10^{9-18} \text{ eV} \cdot \text{km} = 0.197 \times 10^{-9} \text{ eV} \cdot \text{km} \end{aligned}$$

Thus the argument of sine factor can be calculated as

$$\begin{aligned} \frac{\Delta m^2 L}{4E} &= \frac{\Delta m^2 L}{4E\hbar c} \\ &= \frac{(\Delta m^2/\text{eV}^2) \text{eV}^2 (L/\text{km}) \text{km}}{4 \times (E/\text{GeV}) \text{GeV} \times 0.197 \times 10^{-9} \text{eV} \cdot \text{km}} \\ &= \frac{(\Delta m^2/\text{eV}^2) (L/\text{km})}{4 \times 0.197 (E/\text{GeV})} \\ &= 1.269 \frac{(\Delta m^2/\text{eV}^2) (L/\text{km})}{(E/\text{GeV})} \end{aligned}$$

# $\nu$ oscillation

## 2 flavor case in vacuum

1 component of Dirac eq. for mass eigenstate (w/ common  $\vec{p}$ )

$$\begin{cases} i \frac{d}{dx} \mathbf{v}_1(x) = E_1 \mathbf{v}_1(x) \\ i \frac{d}{dx} \mathbf{v}_2(x) = E_2 \mathbf{v}_2(x) \end{cases}$$

$$E_j = \sqrt{\vec{p}^2 + m_j^2}$$

Mixing angle

Flavor eigenstates

$$\begin{pmatrix} \mathbf{v}_\mu \\ \mathbf{v}_\tau \end{pmatrix} = \mathbf{U} \begin{pmatrix} \mathbf{v}_1 \\ \mathbf{v}_2 \end{pmatrix}$$

$$U \equiv \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix}$$

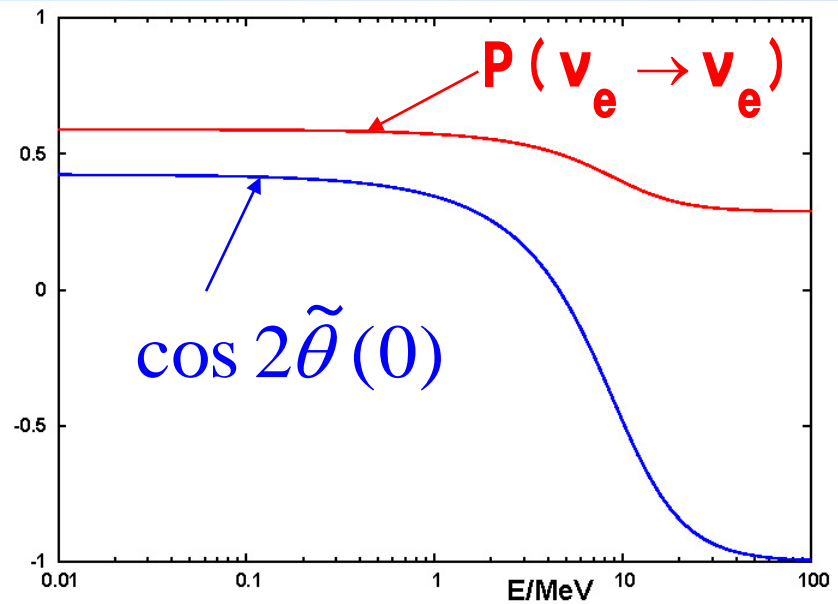
$$P(\mathbf{v}_\mu \rightarrow \mathbf{v}_\tau) = \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

$$\Delta E = E_2 - E_1 \approx \frac{m_2^2 - m_1^2}{2E} = \frac{\Delta m^2}{2E}$$

• In the presence of  $\nu$  mass & mixing, flavor transition occurs.

• The probability of flavor transition has an **oscillatory** behavior with respect to  $L$

Probability for solar  $\nu$  can be obtained by an adiabatic approximation and the limit  $L \rightarrow \infty$ . It is expressed in terms of the initial and final mixing angles, and depends on  $E_\nu$  through the initial mixing angle.



$$\begin{aligned}
 P(\nu_e \rightarrow \nu_e; \theta_\odot; A)_{N_\nu=2} &\simeq c_\odot^2 \tilde{c}^2(t_1) + s_\odot^2 \tilde{s}^2(t_1) \\
 &= \frac{1}{2} \left[ 1 + \cos 2\theta_\odot \cos 2\tilde{\theta}(t_1) \right] \\
 &= \frac{1}{2} \left( 1 + \cos 2\theta_\odot \frac{\Delta E \cos 2\theta_\odot - A}{\Delta \tilde{E}(t_1)} \right)
 \end{aligned}$$

Final mixing angle (in vacuum)

$$A \equiv \sqrt{2G_F n_e(x)}$$

Initial mixing angle (in matter)

Expression in the case of adiabatically varying  $N_e$

$$\Delta E \equiv E_2 - E_1 \simeq \Delta m^2 / 2E$$

$$\tan 2\tilde{\theta}(t_1) \equiv \frac{\Delta E \sin 2\theta_\odot}{\Delta E \cos 2\theta_\odot - A(t_1)}$$

$$\Delta \tilde{E}(t_1) \equiv \left\{ [\Delta E \cos 2\theta_\odot - A(t_1)]^2 + (\Delta E \sin 2\theta_\odot)^2 \right\}^{1/2}$$

**solar  $\nu$  experiments**  
**Ga:Gallex-GNO, SAGE**

**Cl:Homestake**

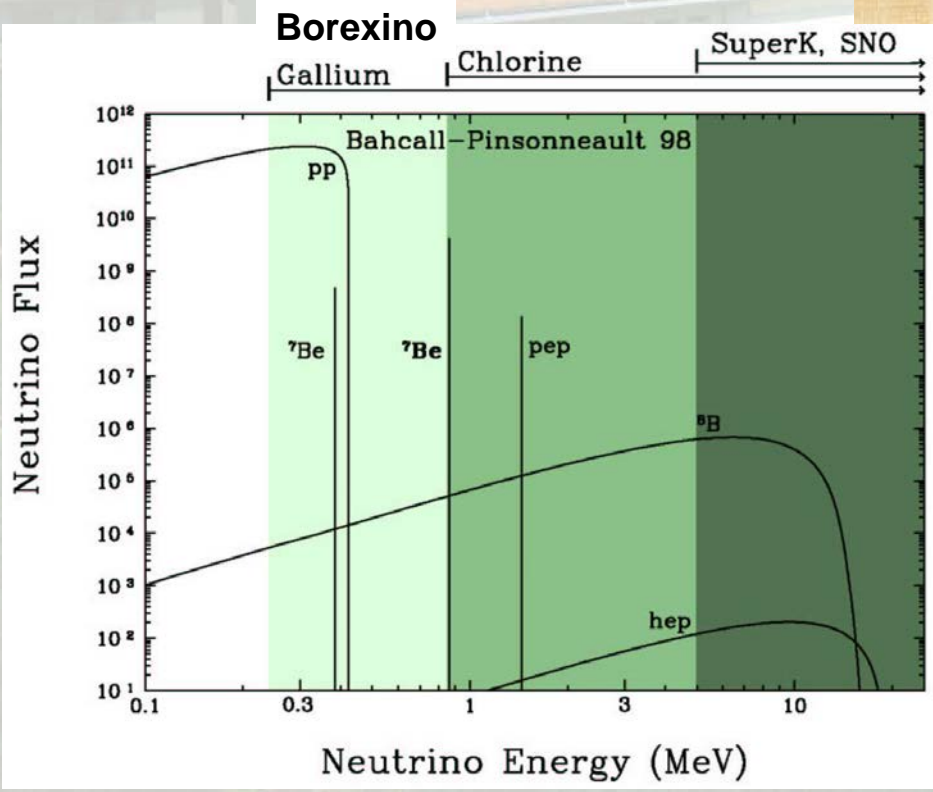
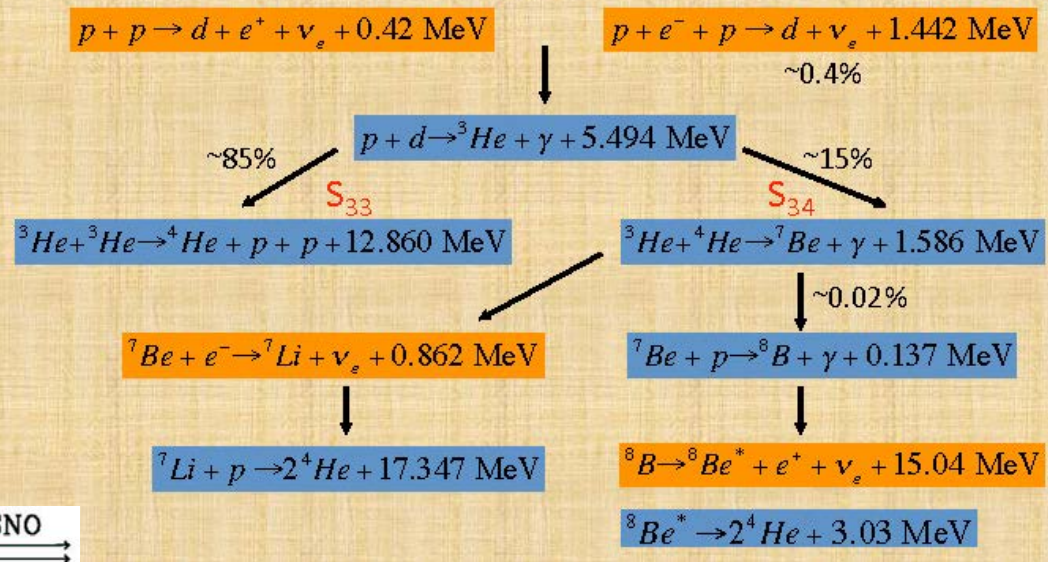
**H<sub>2</sub>O:Kam, SK**

**D<sub>2</sub>O:SNO**

**(CH<sub>2</sub>)<sub>n</sub>:Borexino**

**in operation**

**Solar Neutrinos Sources: pp chain**



**From various solar  $\nu$  experiments with different threshold energies, info on  $\Delta m^2$  and  $\sin^2 2\theta$  can be obtained**

# Recent status: Tension between T2K and Nova?

Kelly et al,  
arXiv:2007.08526v1 [hep-ph]

Black lines: a joint fit  
of T2K/NOvA/SK18

Blue: NOvA alone

Red: T2K alone

Joint fit may  
indicate preference  
for Inverted  
Hierarchy

