

New Phases in QCD from Topological Solitons Stabilized by Background Fields

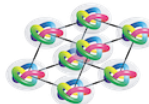
YITP Workshop 場の理論と弦理論 2024
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References (QCD) Collaborators

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Zebin Qiu, **Yuki Amari**, **Ryo Yokokura** (Keio U.), **Shin Sasaki** (Kitasato U.)

QCD in Strong Magnetic Field:

[1] M.Eto, K.Nishimura & MN, [2304.02940](#)

[2] M.Eto, K.Nishimura & MN, *JHEP*, [2311.01112](#)

[3] Y.Amari, M.Eto & MN, [2403.06778](#)

[4] Y.Amari, MN & R.Yokokura, [2406.14419](#)

[5] Z.Qiu & MN, *JHEP*, [2403.07433](#) **Vortex Skymion Phase**

[6] M.Eto & MN, *JHEP*, [2207.00211](#) quantum creation

[7] Z.Qiu & MN, *JHEP*, [2304.05089](#) quasi-crystal

[8] MN & S.Sasaki [2404.12066](#) supersymmetry

**Domain-wall
Skymion Phase**

QCD in Rapid Rotation:

[9] M.Eto, K.Nishimura & MN, *JHEP*, [2112.01381](#)

[10] M.Eto, K.Nishimura & MN, *JHEP*, [2310.17511](#)

[11] M.Eto, K.Nishimura & MN, *JHEP*, [2312.10927](#)

**← Domain-wall
Skymion Phase**

References (Cond-mat) Collaborators

Yuki Amari (Keio U.), R.Calum (Edge Hill U.),
S.B.Gudnason (Henan U.), Seungho Lee, Se Kwon Kim (KAIST),
Yutaka Akagi (U. Tokyo), Yasha Shnir (Dubna),
Toshiaki Fujimori (Dokkyo U.)

Chiral Magnets:

- [12] C.Ross & MN, *PRB* [2205.11417](#)
- [13] Y.Amari & MN, *JHEP* [2307.11113](#)
- [14] Y.Amari, C.Ross & MN, *PRB* [2311.05174](#)
- [15] S.B.Gudnason, Y.Amari & MN, [2406.19056](#)
- [16] S.Lee, T.Fujimori, MN & S.K.Kim, [2407.04007](#)

$\mathbb{C}P^2$ Magnets:

- [17] Y.Amari, Y.Akagi, S.B.Gudnason, MN & Y.Shnir, *PRB* [2204.01476](#)

Main messages

- (1) Unstable topological solitons can be stabilized by a **background (gauge) field**.
- (2) When the background is strong enough, the **soliton's energy** becomes **negative**.
→ **Inhomogeneous (Solitonic) Phases**
- (3) Examples in **QCD** & **Cond-mat**.
- (4) Maybe in Early Universe as well?

Main messages

(1) Unstable topological solitons can be stabilized by a **background (gauge) field**.

(2) When the background is strong enough, the **soliton's energy** becomes **negative**. **Not today**
→ **Inhomogeneous (Solitonic) Phases**

(3) Examples in **QCD** & **Cond-mat**.

(4) Maybe in Early Universe as well? **In future**

The stability of topological solitons $\vec{n} = (n_1, n_2, n_3), \vec{n}^2 = 1$

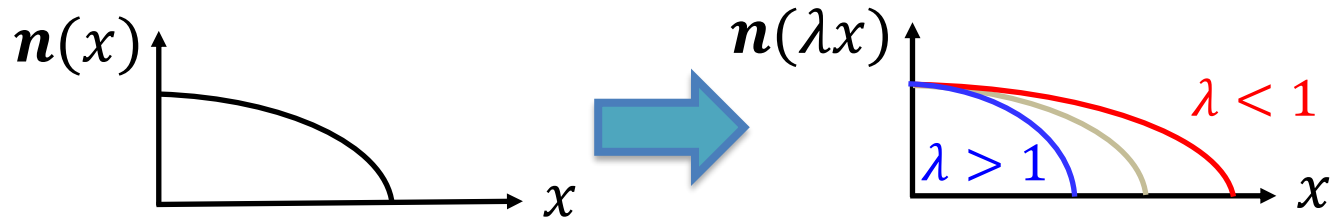
Criteria: The conventional Derrick's scaling argument.

e.g. O(3) model: $\mathcal{L} = \overset{E_2}{\partial_\mu \vec{n} \partial^\mu \vec{n}} - \overset{E_0}{V(n)} + \kappa \left(\overset{E_4}{\partial_\mu \vec{n} \times \partial_\nu \vec{n}} \right)^2$

$x \rightarrow x' = \lambda x \quad n(x) \rightarrow n(\lambda x)$ Skyrme term

$$\begin{aligned} E(\lambda) &= E_0(\lambda) + E_2(\lambda) + E_4(\lambda) \\ &= \lambda^{-d} E_0(\lambda = 1) + \lambda^{2-d} E_2(\lambda = 1) + \lambda^{4-d} E_4(\lambda = 1) \end{aligned}$$

$$E'(\lambda)|_{\lambda=1} = -dE_0 + (2-d)E_2 + (4-d)E_4 = 0$$



The stability of topological solitons $\vec{n} = (n_1, n_2, n_3), \vec{n}^2 = 1$

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e.g. O(3) model: $\mathcal{L} = \partial_\mu \overset{E_2}{\vec{n}} \partial^\mu \vec{n} - \overset{E_0}{V(n)} + \kappa (\partial_\mu \vec{n} \times \partial_\nu \vec{n})^2$
Spatial dim $d = 2$: $E'(\lambda)|_{\lambda=1} = -2E_0 + 2E_4 = 0$ Skyrme term

$E_0 = 0, E_4 = 0$: lump (scale invariant)

$E_0 \neq 0, E_4 \neq 0$: baby Skyrmion

$E_0 \neq 0, E_4 = 0$: **unstable to shrink**

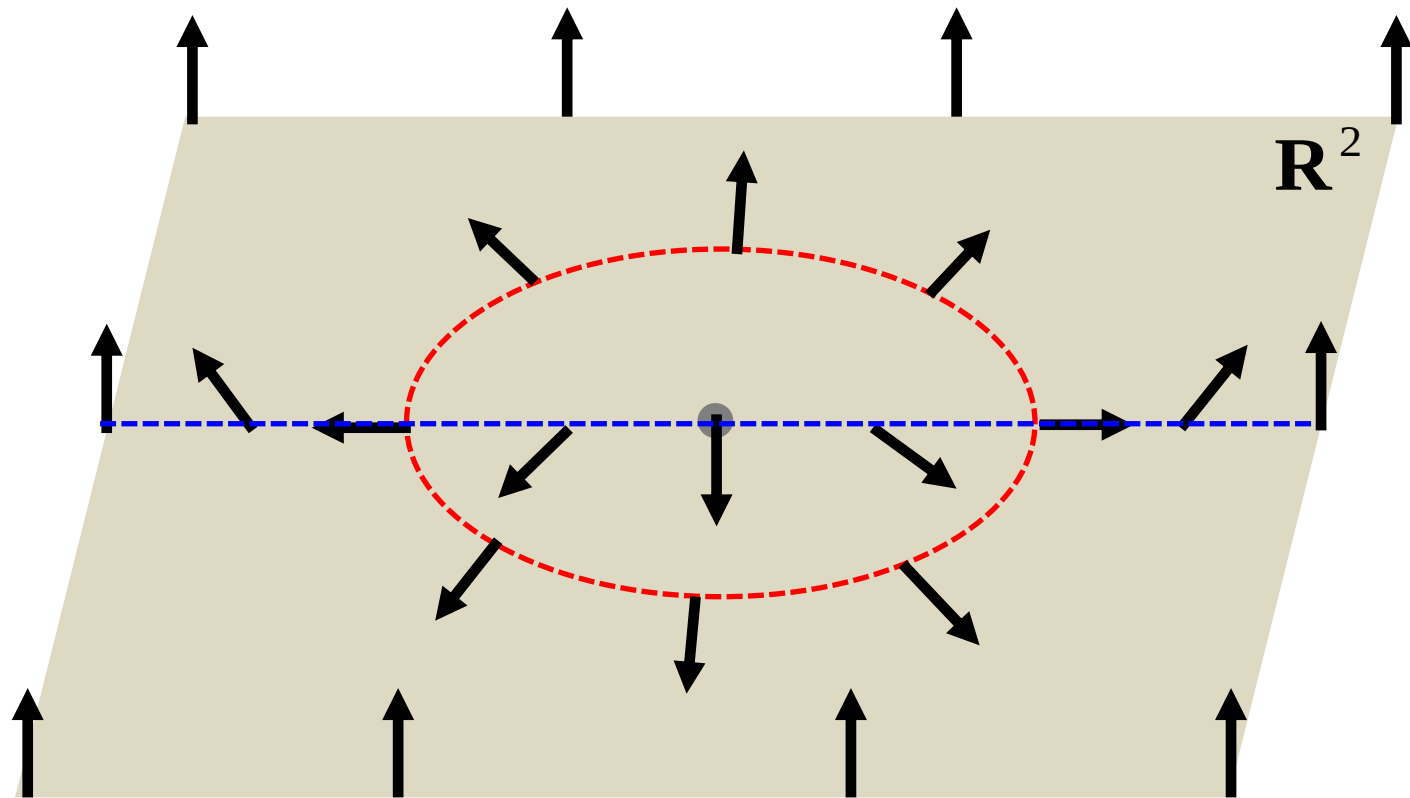
$d = 3$: $E'(\lambda)|_{\lambda=1} = -3E_0 - E_2 + E_4 = 0$

$E_2 \neq 0, E_4 \neq 0$: Skyrmion & Hopfion

$E_2 \neq 0, E_4 = 0$: **unstable to shrink**

2D (baby) Skyrmion (or lump)

$$\pi_2(S^2) \cong \mathbb{Z}$$



$$\pi_2(S^2) \cong \mathbb{Z}$$



[Amazon.co.jp: STK スイカ 吊り下げスイカネット 園芸ネット 果物袋 カボチャ メロン 丈夫 通気 栽培用 \(30cm\(普通、ロープ付き\), 10PCS\) : DIY・工具・ガーデン](https://www.amazon.co.jp/dp/B000APR08C)

The stability of topological solitons

Criteria: The conventional Derrick's scaling argument.

We should think it more carefully in the presence of a **background (gauge) field**.

Examples:

1. Condensed matter systems
(e.g. chiral magnets,
ultracold atomic gases with synthetic gauge field)
2. QCD (or field theory) under external fields
e.g. magnetic field / rotation (relevant to QGP, neutron stars)

1. Examples in Cond-mat

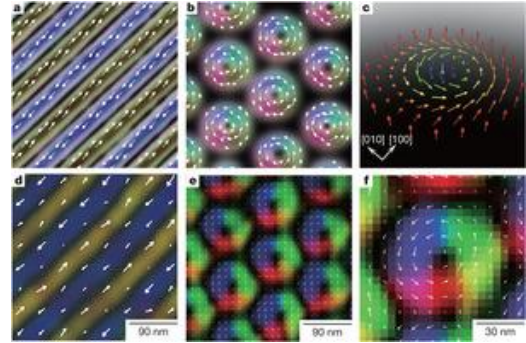
Magnetic Skyrmions in chiral magnets

Important for nanotechnology such as
low-energy consumption **magnetic memories**.

O(3) model +
Dzyaloshinsky-Moriya (DM) interaction

$$\mathcal{H}_{\text{DM}} = \kappa \mathbf{n} \cdot (\nabla \times \mathbf{n})$$

a thin film of $\text{Fe}_{0.5}\text{Co}_{0.5}\text{Si}$
X. Z. Yu et.al
Nature 465, 901 (2010)



Derrick's theorem

$$E_{1(\text{DM})} + 2E_{0(\text{pot})} = 0$$

< 0 > 0 : prefers shrinking

Negative energy prevents shrinking
: possible only for either of (anti-)Skyrmion

O(3) model + a background $SU(2)$ gauge field

B. J. Schroers, SciPost Phys. 7 (2019) 030 [1905.06285].

$$\begin{aligned}\mathcal{H} &= \frac{1}{2} (D_k \mathbf{n})^2 & A_\mu^a &= (A_0^a = 0, A_k^a) & \mu &= 0, 1, 2: \text{spacetime index} \\ & & & & a &= 1, 2, 3: SU(2) \text{ adj index} \\ &= \frac{1}{2} \partial_k \mathbf{n} \cdot \partial_k \mathbf{n} + \boxed{A_k \cdot (\mathbf{n} \times \partial_k \mathbf{n})} + \frac{1}{2} (A_k \times \mathbf{n})^2 \\ & & \mathcal{H}_{\text{DM}} & \text{DM interaction} & & \text{potential } V_{\text{tot}}\end{aligned}$$

$$SU(2) \text{ field strength } F_{12} = \kappa^2 \sigma_3$$

Cf. Magnetized D-brane

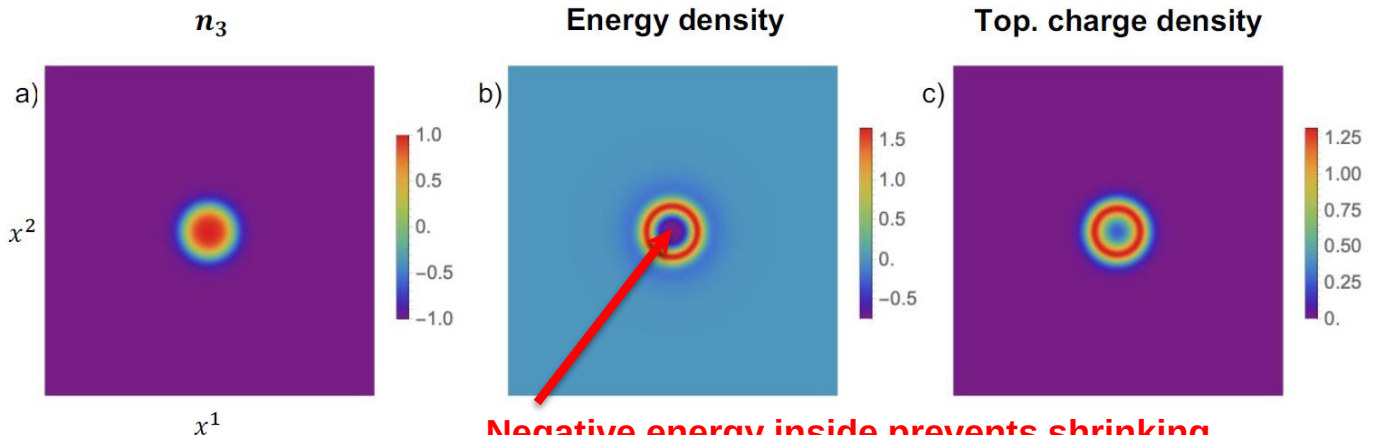
$$V(\vec{n}) = m^2(1 - n_3^2)$$

Skyrmion & anti-Skyrmion are stable (without Zeeman).

Skyrmion

@ the ferro g.s.

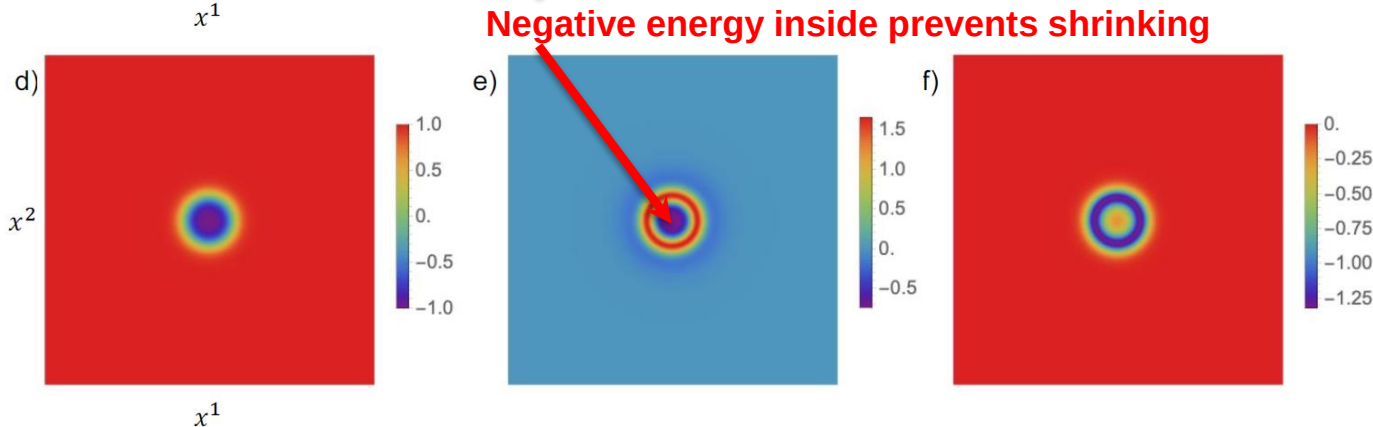
$n_3 = -1$
(Anti-Skyrmion
is unstable)



Anti-Skyrmion

@ the ferro g.s.

$n_3 = +1$
(Skyrmion
is unstable)



Ultracold atomic gas with a synthetic gauge field

Lin, Spiermal et.al, Nature 462 ('09)

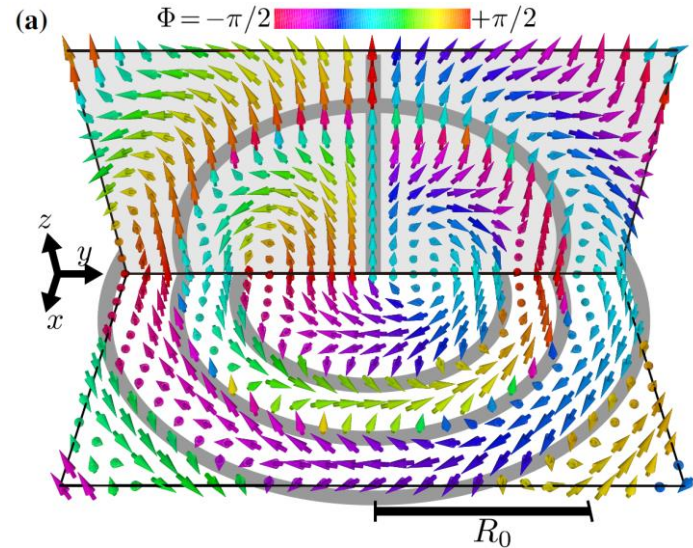
3D Skyrmions (without Skyrme term)

T. Kawakami, T. Mizushima,
MN, and K. Machida
PRL 109, 015301 (2012)

$$\mathcal{H}_0 = \int d^3r [D_{\mu\nu} \Psi_\nu(\mathbf{r})]^\dagger \cdot [D_{\nu\eta} \Psi_\eta(\mathbf{r})]$$

$$D_{\mu\nu} = (-i\nabla\sigma_0 + A)_{\mu\nu}$$

$$A = \kappa_\perp (\sigma_x \hat{x} + \sigma_y \hat{y}) + \kappa_z \sigma_z \hat{z}.$$



2. Examples in Sigma Model & QCD

$U(1)$ gauged $O(3)$ [or $\mathbb{C}P^1$] model

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + D_\mu \vec{n} \cdot D^\mu \vec{n} - V(n) \quad \vec{n} = \begin{pmatrix} n_1 \\ n_2 \\ n_3 \end{pmatrix}, \quad \vec{n}^2 = v^2$$

gauge tr. $n_1 + in_2 \rightarrow e^{-i\theta(x)} (n_1 + in_2)$, $n_3 \rightarrow n_3$

$$D_\mu n_1 = \partial_\mu n_1 + eA_\mu n_2, \quad D_\mu n_2 = \partial_\mu n_2 - eA_\mu n_1, \quad D_\mu n_3 = \partial_\mu n_3$$

Well-known model?

1. With a potential $V(n) = m^2 n_3^2$, half gauged-lumps (merons) are BPS and stable. B.J.Schroers('95)

1'. $V(n) = m^2 (n_3 - c)^2$, fractional gauged-lumps

2. Without potential, lumps are unstable to expand.

2. Without potential, lumps are unstable to expand.



Background gauge field (constant magnetic field)

$$\mathcal{A}_0 = 0, \quad \mathcal{A}_1 = -\frac{By}{2}, \quad \mathcal{A}_2 = \frac{Bx}{2} \quad F_{12} = B(= -B_3)$$

Y. Amari, M.Eto & MN,
[2403.06778](#)

We find

1. For non-dynamical gauge field,
 $k = 1$ stable & $k = -1$ unstable (*similar to magnetic Skyrmion*)
2. For dynamical gauge field, both are stable.
 $k = 1$ lighter & $k = -1$ heavier
3. For higher k , droplet behavior: Area $\sim k$

Non-dynamical gauge field

$$\begin{aligned} E[\vec{n}] &= \int d^2x |D_i \vec{n}|^2 \\ &= \int d^2x \left[\underbrace{|\partial_i \vec{n}|^2}_{E_2} + \underbrace{2e \mathcal{A}_i \varepsilon_{ab} \partial_i n_a n_b}_{E_1} + \underbrace{e^2 \mathcal{A}_i^2 (n_1^2 + n_2^2)}_{E_0} \right] \end{aligned}$$

**Derick's
Scaling**

$$\delta \equiv E_1 + 2E_0 = 0$$

< 0

> 0 : prefers shrinking

$$\vec{n}^{(\lambda)}(x) = \vec{n}(\lambda x), \quad \partial_i \vec{n}^{(\lambda)}(x) = \lambda \partial'_i \vec{n}(\lambda x)$$

$$\mathcal{A}_i^{(\lambda)}(x) = \lambda^{-1} \mathcal{A}_i(\lambda x)$$

$$e(\lambda) = E[\vec{n}^{(\lambda)}(x)] = E_2 + \lambda^{-2} E_1 + \lambda^{-4} E_0 \quad \frac{de(\lambda)}{d\lambda} \big|_{\lambda=0} = 0$$

Axisymmetric ansatz

$$n_1 = v \frac{x}{r} \sin \Theta(r), \quad n_2 = v \frac{y}{r} \sin \Theta(r), \quad n_3 = v \cos \Theta(r)$$

$$\textbf{EOM} \quad \Theta'' + \frac{\Theta'}{r} - \frac{(2 - eBr^2)^2}{4r^2} \sin \Theta \cos \Theta = 0$$

$$\mathbf{k} = +\mathbf{1}: \mathbf{n}_3(\infty) = -\mathbf{v}, \quad \mathbf{n}_3(\mathbf{0}) = +\mathbf{v}$$

$$\mathbf{k} = -\mathbf{1}: \mathbf{n}_3(\infty) = +\mathbf{v}, \quad \mathbf{n}_3(\mathbf{0}) = -\mathbf{v}$$

$$E_0 = \frac{\pi e^2 B^2 v^2}{2} \int_0^\infty dr r^3 \sin^2 \Theta > 0$$

$$\delta \equiv E_1 + 2E_0 = 0$$

$$E_1 = -2\pi e B v^2 \int_0^\infty dr r \sin^2 \Theta \quad \begin{cases} < 0 \text{ for } B > 0: \text{ OK} \\ > 0 \text{ for } B < 0: \end{cases}$$

Axisymmetric ansatz

Charge conjugation

$$n_1 = v \frac{x}{r} \sin \Theta(r), \quad n_2 = v \frac{-y}{r} \sin \Theta(r), \quad n_3 = v \cos \Theta(r)$$

$$\textbf{EOM} \quad \Theta'' + \frac{\Theta'}{r} - \frac{(2 + eBr^2)^2}{4r^2} \sin \Theta \cos \Theta = 0$$

$$k = -1: n_3(\infty) = -v, \quad n_3(0) = +v$$

$$k = +1: n_3(\infty) = +v, \quad n_3(0) = -v$$

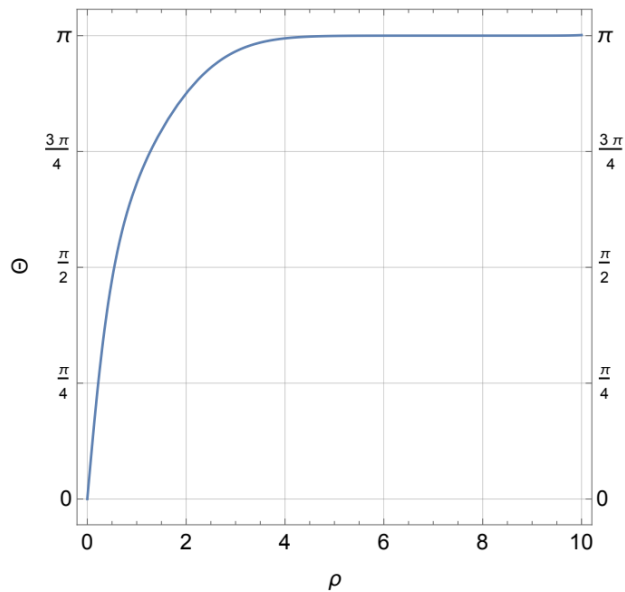
$$E_0 = \frac{\pi e^2 B^2 v^2}{2} \int_0^\infty dr r^3 \sin^2 \Theta > 0$$

$$\delta \equiv E_1 + 2E_0 = 0$$

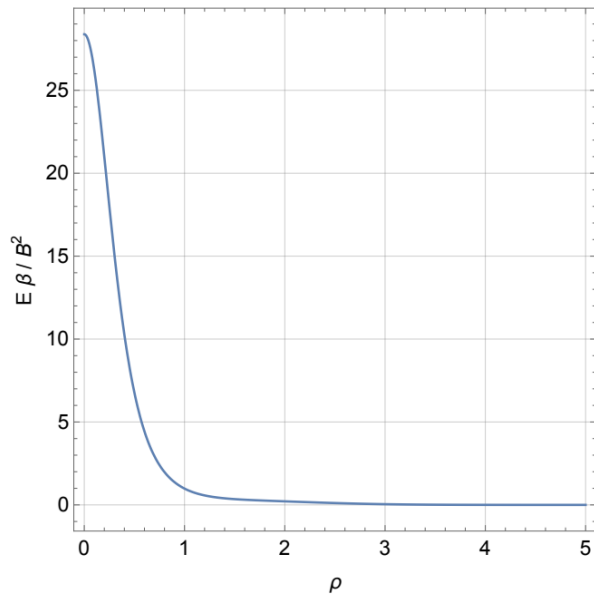
$$E_1 = +2\pi e B v^2 \int_0^\infty dr r \sin^2 \Theta \quad \begin{cases} < 0 \text{ for } B < 0: \text{ OK} \\ > 0 \text{ for } B > 0 \end{cases}$$

Stable solution

$$n_1 = v \frac{x}{r} \sin \Theta(r), \quad n_2 = v \frac{y}{r} \sin \Theta(r), \quad n_3 = v \cos \Theta(r)$$



Profile

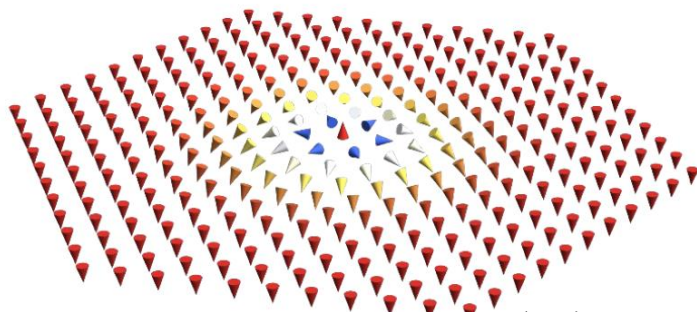


Energy density

Lump

$$k = 1$$

$$B > 0$$

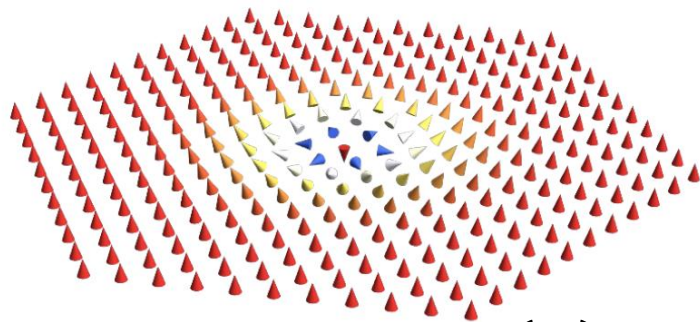


$$\mathbf{n}_3(\infty) = -\mathbf{v}$$

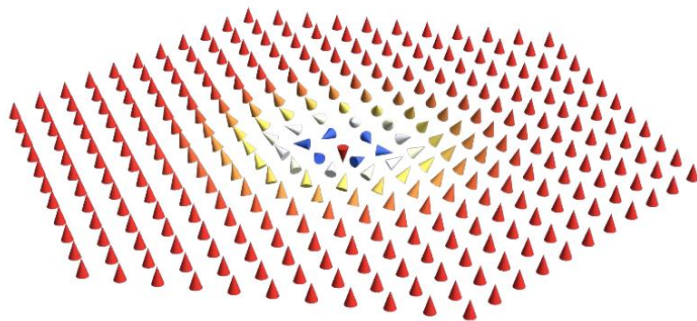
Anti-lump

$$k = -1$$

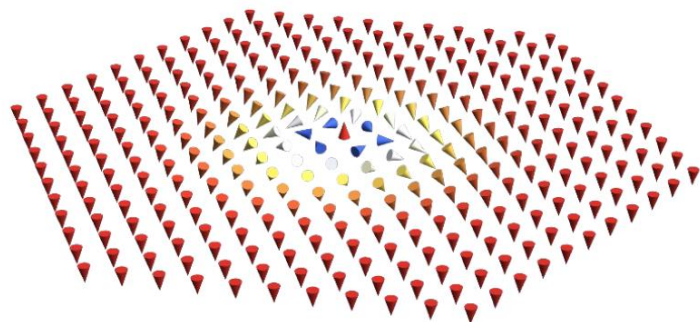
$$B < 0$$



$$\mathbf{n}_3(\infty) = +\mathbf{v}$$



$$\mathbf{n}_3(\infty) = +\mathbf{v}$$



$$\mathbf{n}_3(\infty) = -\mathbf{v}$$

Profiles

(1) Mass: $\frac{M}{2\pi v^2} = 3.62188$

$< 4 = \frac{M_{BPS}}{2\pi v^2}$: **90.5% of BPS lump mass**

(2) Size: $r_0 = \frac{0.544784}{\sqrt{eB}}$ **40% of** $\tilde{r}_0 = \sqrt{\frac{2}{eB}}$

(3) Magnetic flux inside a superconducting ring:

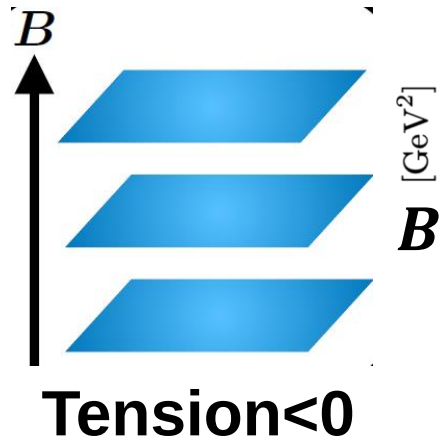
$\Phi = \pi r_0^2 B \simeq 0.15 \times \frac{2\pi}{e}$ **15% of** $\tilde{\Phi} = \frac{2\pi}{e}$

(4) As for stability, very similar to magnetic skyrmions

Why do we consider this model? → It appears in QCD !

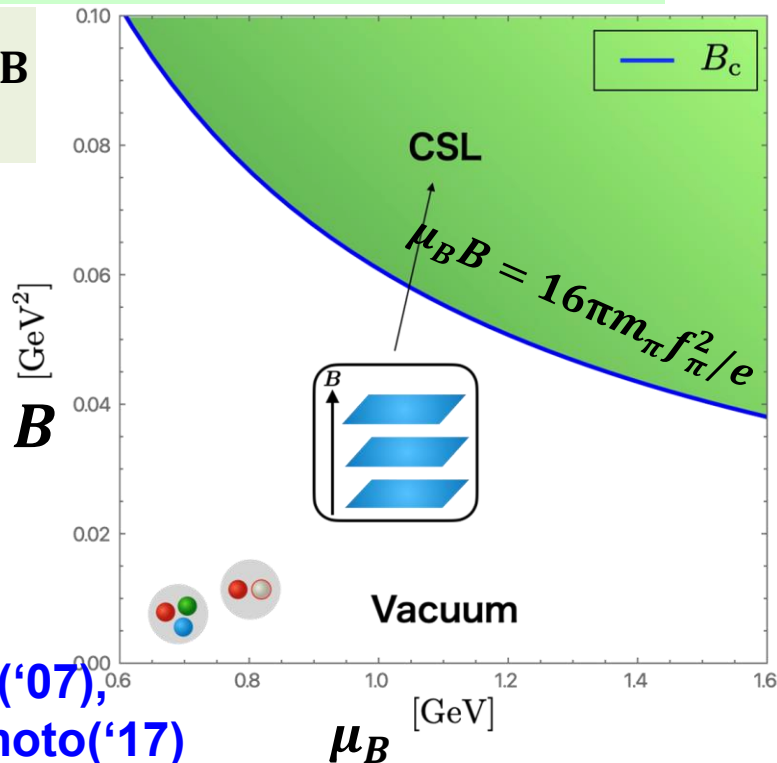
Chiral Soliton Lattice(CSL) phase

chemical pot. μ_B
magnetic field B



Son & Stephanov('07),
Brauner & Yamamoto('17)

Solitons carry baryon #

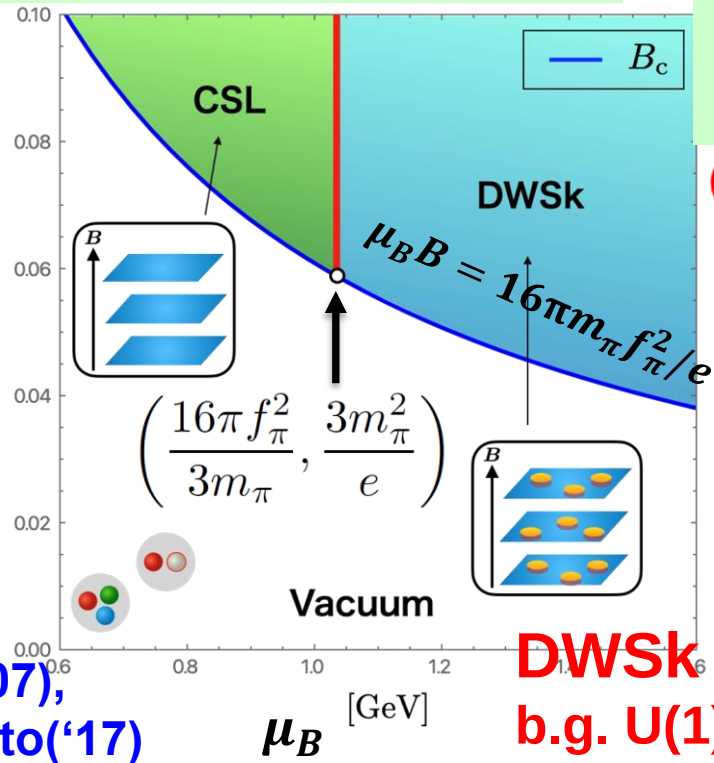
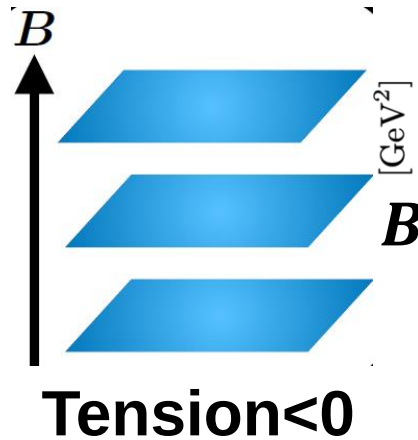


Why do we consider this model? → It appears in QCD !

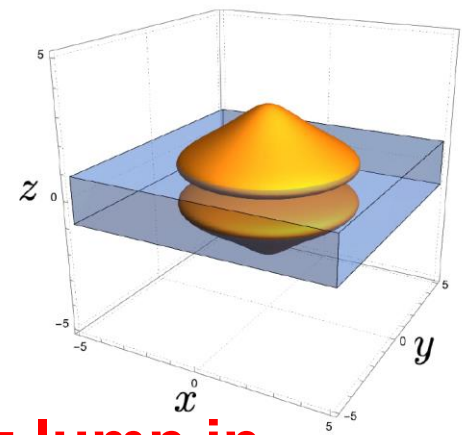
Chiral Soliton Lattice(CSL) phase

Domain-Wall Skymion Crystal (DW-SkX) phase

chemical pot. μ_B
magnetic field B



@ $\mu_B \geq 1.03$ GeV



Son & Stephanov('07),
Brauner & Yamamoto('17)

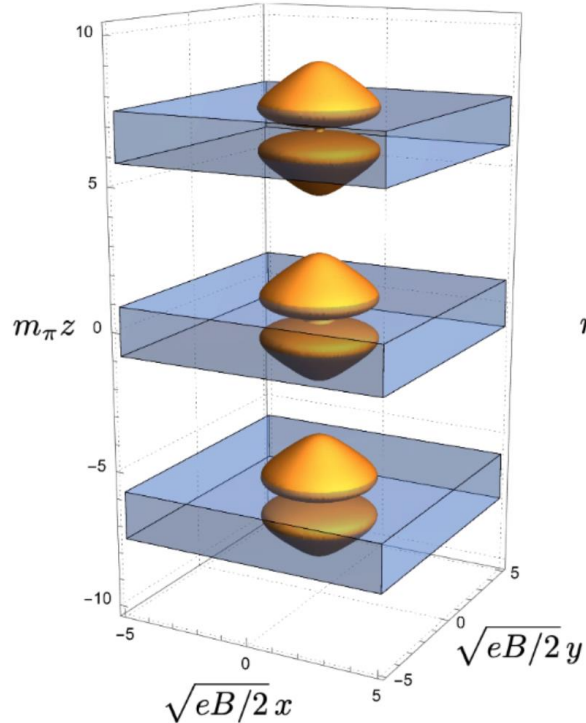
DWSk = lump in
b.g. U(1)-gauged O(3) model

Solitons carry baryon # M.Eto, K.Nishimura & MN, [2304.02940](#)

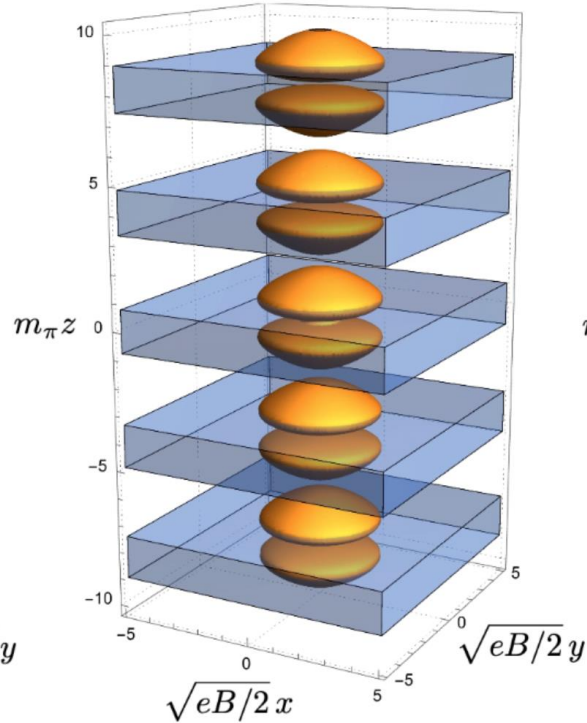
Macarons

small $\mu_B B$

$$\kappa = 0.9801$$



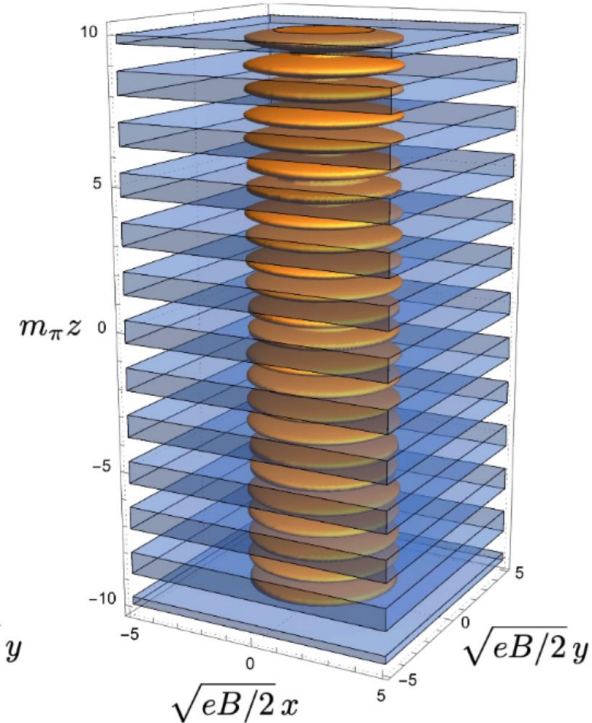
$$\kappa = 0.81$$

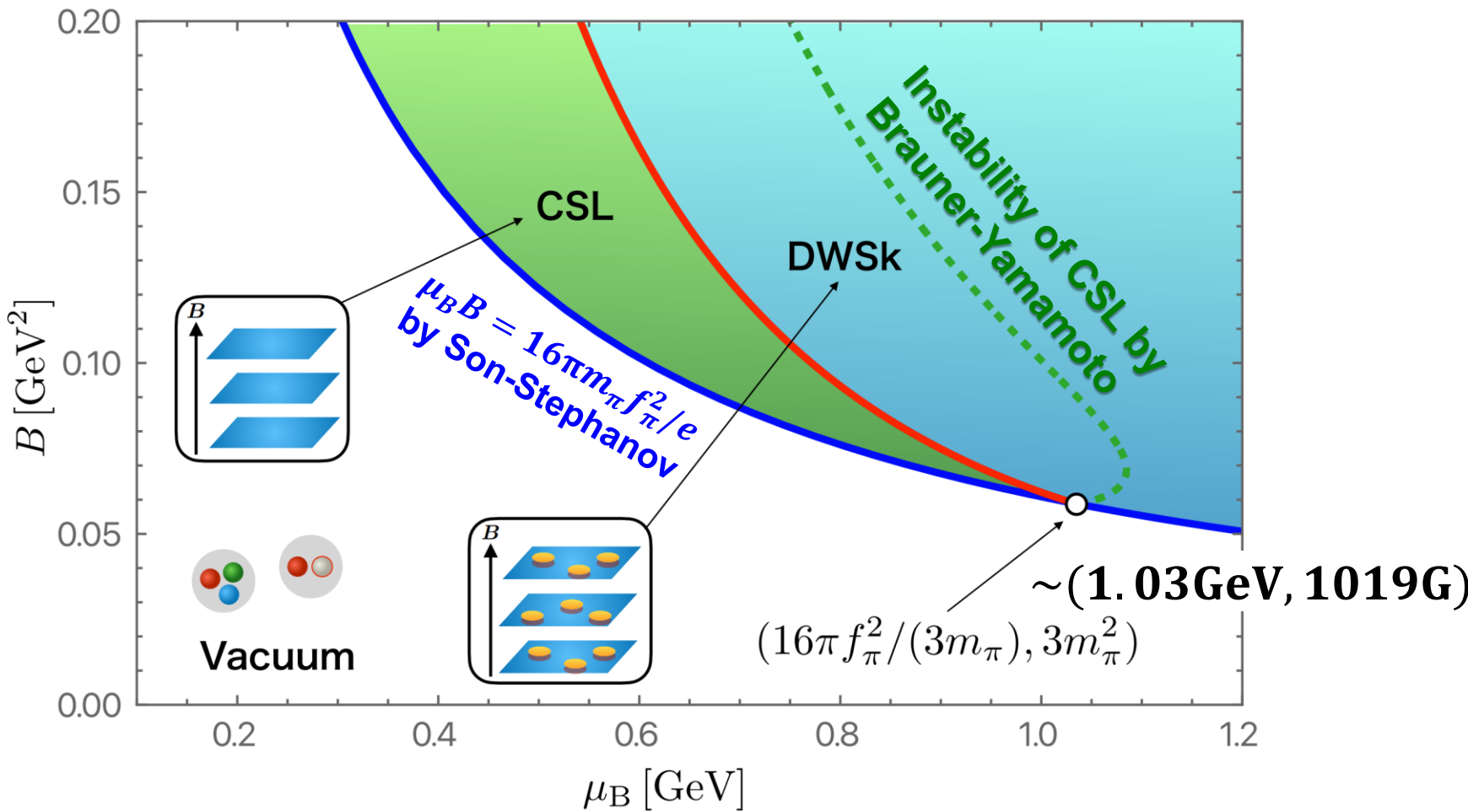


Pancake tower

large $\mu_B B$

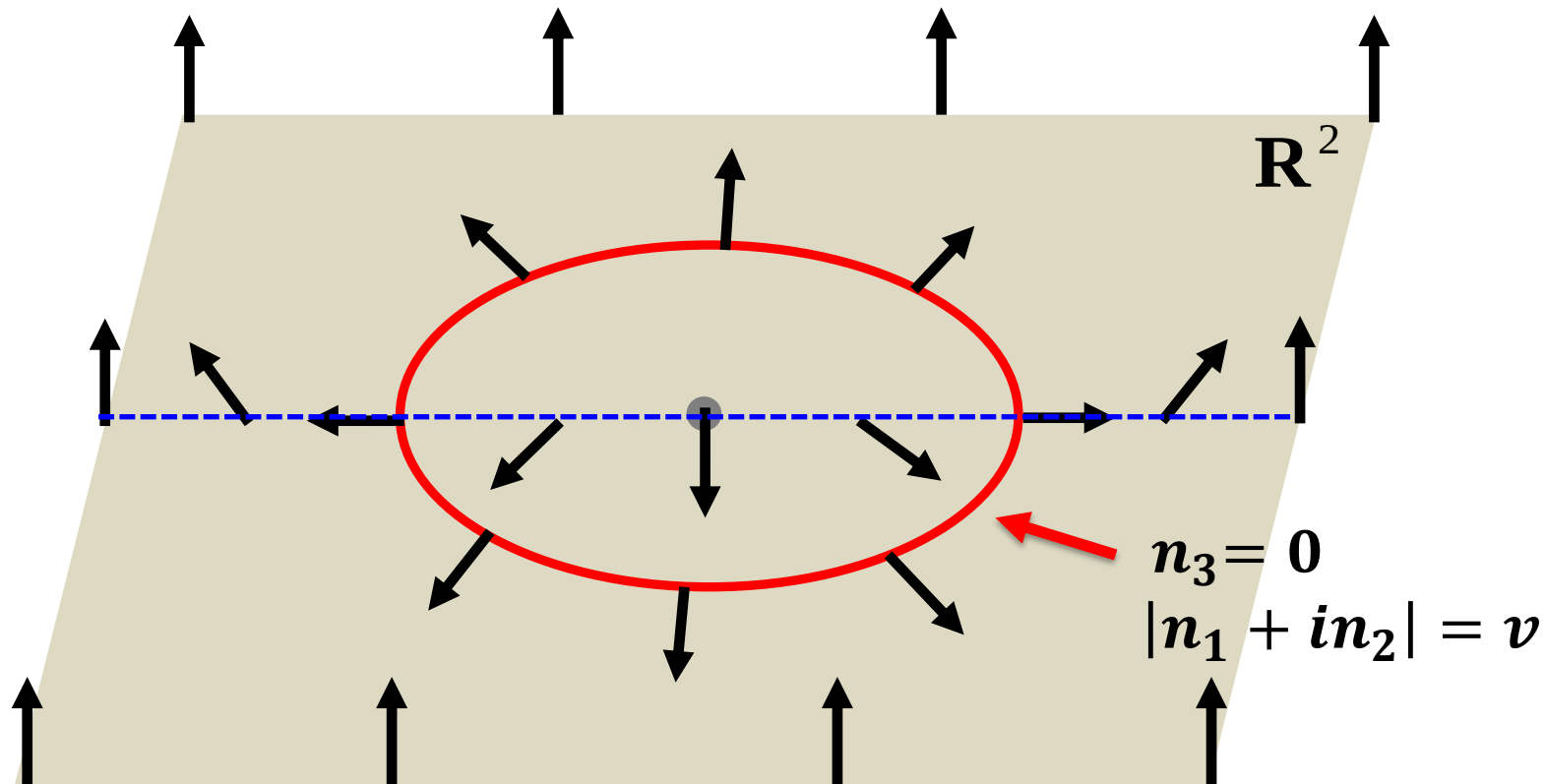
$$\kappa = 0.25$$



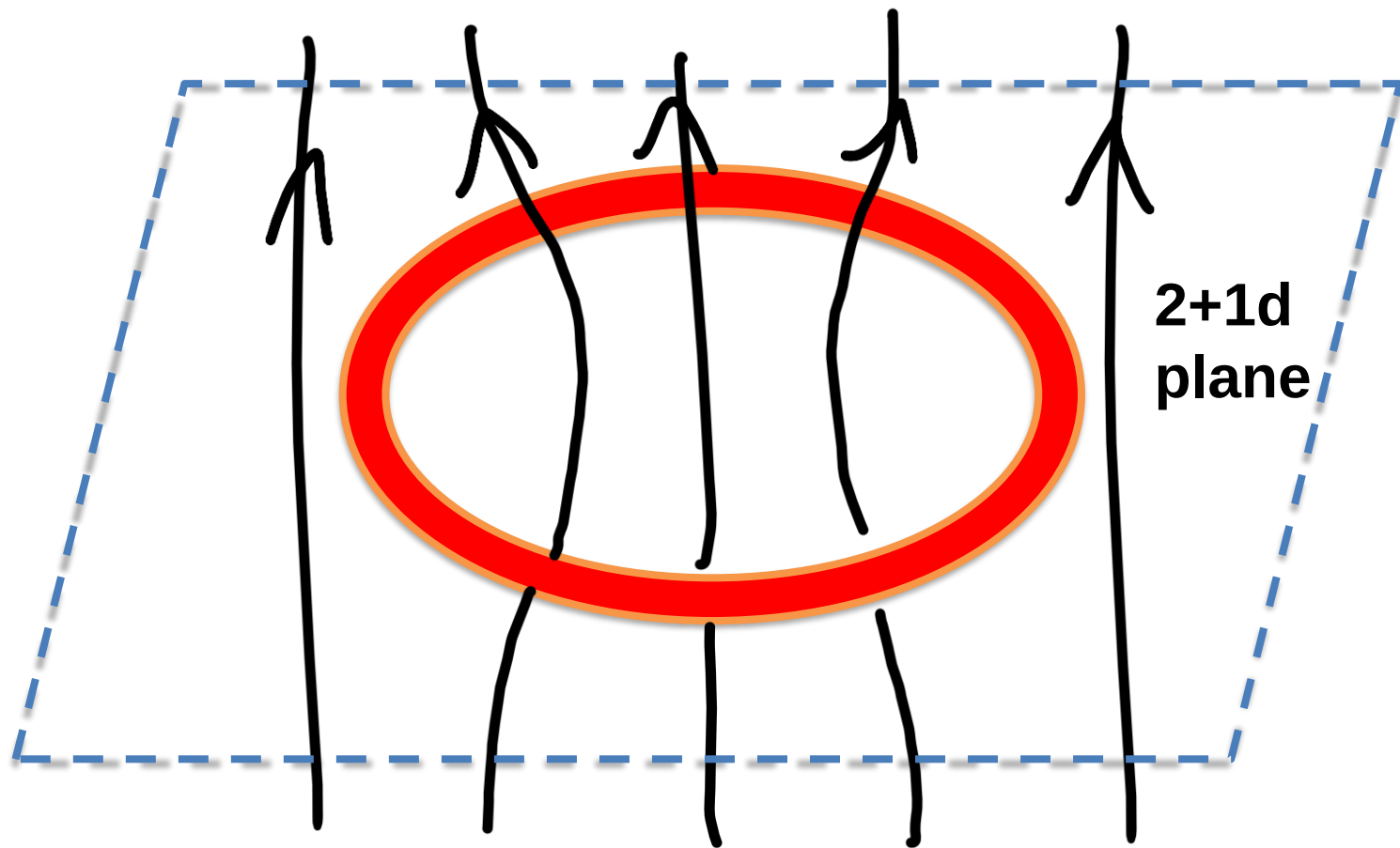


Lump = a superconducting ring

$$\pi_2(S^2) \cong \mathbb{Z}$$



Superconducting ring in magnetic field



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(2) When the background is strong enough, the **soliton's energy** becomes **negative**. Not today
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