Applications of inhomogeneous deformations in 2 d CFTs: from the revival of non-local correlation to curved spacetime
Masahiro Nozaki (KITS, UCAS \& iTHEMS, RIKEN)
This talk is based on collaboration with
Kanato Goto, Weibo Mao, Akihiro Miyata, Shinsei Ryu, Mao
Tian Tan, Kotaro Tamaoka, and Masataka Watanabe
Based on
arXiv:2112.14388, arXiv:2302.08009,
arXiv:23XX.XXXXX and arXiv:23XX.XXXX

 ITHEM.S

## Short summary of my talk

What we want to study:

1. The non-equilibrium process opposite to the quantum thermalization and scrambling.

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In other words, the process can endow the state with quantum properties such as non-local correlation. For example, the preparation for the vacuum state.

## Short summary of my talk

What we want to study:

1. The non-equilibrium process opposite to the quantum thermalization and scrambling.

In other words, the process can endow the state with quantum properties such as non-local correlation.

This is the main topic in this talk.

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What we want to study (if I have time):
2. Thermodynamics of QFT on the curved background.

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Two-dimensional conformal field theories (2d CFTs)

Short summary of my talk
What we want to study:
2. Thermodynamics of QFT on the curved background.

Phase transition induced by curvature of spacetime?
How about entanglement?

## Short summary of my talk

1.We explored the dynamical property of the system evolved with inhomogeneous Hamiltonians:

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H_{\text {Inho }}=\int_{0}^{L} d x f(x) h(x) \quad \begin{array}{ll}
\text { SSD Hamiltonian: } & f_{\text {SSD }}(x)=2 \sin ^{2}\left(\frac{\pi x}{L}\right) \\
\text { Mobius Hamiltonian: } & f_{\text {Möbius }}(x)=1-\tanh 2 \theta \cos \left(\frac{2 \pi x}{L}\right)
\end{array}
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For $f(x)=1$, this is undeformed Hamiltonian.

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1. Start from the boundary state (unentangled state), and the thermal state (mixed state).
2. Start from the thermofield double state or page state.

Vacuum entanglement or correlation approximately emerge or are recovered. (also quantum revival.) Non-local correlation emerges. State posses interesting properties.

## Short summary of my talk

2.We explored the dynamical property of the thermal state whose distribution is determined by the Mobius Hamiltonian:

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## Equivalent

## CFT Hamiltonian on the curved background.

$$
\begin{gathered}
d s^{2}=-f_{\text {Möbius }}^{2}(x, \theta) d t^{2}+d x^{2} \\
R(x, \theta)=\frac{8 \pi^{2} \tanh 2 \theta \cos \left(\frac{2 \pi x}{L}\right)}{L^{2}\left(\tanh 2 \theta \cos \left(\frac{2 \pi x}{L}\right)-1\right)}
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By varying $\theta$,the system in
2d holographic CFT may exhibit phase transition.

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By varying $\theta$, the system may
exhibit phase transition thank to the entanglement phase transition.

Note

All the theories considered in this talk are two-dimensional conformal field theories.

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This is closely relevant to quantum thermalization.

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Local observables in A depend on initial condition.

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\left|\Psi^{i}(t)\right\rangle=U(t)\left|\Psi_{0}^{i}\right\rangle
$$


$U(t)$


## Introduction

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$$
\left|\Psi_{\text {late }}^{i}\right\rangle=\lim _{t \rightarrow \infty} U(t)\left|\Psi_{0}^{i}\right\rangle
$$



$$
\operatorname{tr}_{\bar{A}}\left[\left|\Psi_{\text {late }}^{i}\right\rangle\left\langle\Psi_{\text {late }}^{i}\right|\right] \approx \operatorname{tr}_{\bar{A}} e^{-\beta H}
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Information on the initial states is locally
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## This is the process of

## (information) scrambling.

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Quantum thermalization (the thermalization of the subsystems ) occurs because the reduced density matrices are approximated by the thermal state with the effective temperature.

## Introduction

Scrambling is one of the cutting-edge research topics.

## It is believed that if the Hamiltonian has a strong scrambling ability, information scrambling (quantum thermalization) occurs.


niaden (lost).

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Scrambling is one of the cutting-edge research topics.

## It is believed that 2d holographic CFT, CFT having gravity dual has such strong scrambling ability.


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Non-local correlation
Let us consider the behavior of non-local correlation during the quantum thermalization,
The behavior of the reduced density: $\operatorname{tr}_{\bar{A}}\left(e^{-i H t}|\Psi\rangle\langle\Psi| e^{i H t}\right) \approx \operatorname{tr}_{\bar{A}} e^{-\epsilon H}$

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S_{A} \approx l_{A}, S_{B} \approx l_{B}, S_{A \cup B} \approx\left(l_{A}+l_{B}\right) \Rightarrow I_{A, B \underset{1 \gg}{ } \approx}^{\approx}
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## Quantum thermalization occurs.

$S_{A} \approx l_{A}, S_{B} \approx l_{B}, S_{A \cup B} \approx\left(l_{A}+\right.$
, where $1 \gg \epsilon:$ an effective inverse temperature

$$
I_{A, B}=S_{A}+S_{B}-S_{A \cup B}
$$

$$
\rho_{A \cup B} \underset{1 \gg \epsilon}{\approx} \rho_{A} \otimes \rho_{B}
$$

## Non-local correlation

Let us consider the behavior of non-local correlation during the quantum thermalization,

The behavi
Non-local correlation is destroyed by the dynamics.
Divide the Hilbert space into $A, B$, and the complement to $A$ and $B$

$$
\begin{aligned}
S_{A} \approx l_{A}, S_{B} \approx l_{B}, S_{A \cup B} \approx\left(l_{A}+l_{B}\right) & \Rightarrow I_{A, B} \underset{\approx}{\approx} 0 \\
& \Rightarrow \rho_{A \cup B} \approx \rho_{A} \otimes \rho_{B}
\end{aligned}
$$

Relation to our papers
Meaning of $\rho_{A \cup B} \approx \rho_{A} \otimes \rho_{B}$
No non-local correlation between $A$ and $B$
For example, $\left\langle\mathcal{O}\left(X_{1} \in A\right) \mathcal{O}\left(X_{2} \in B\right)\right\rangle \approx\left\langle\mathcal{O}\left(X_{1} \in A\right)\right\rangle \times\left\langle\mathcal{O}\left(X_{2} \in B\right)\right\rangle$ $1 \gg \epsilon$

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Thus, quantum properties (here, non-local correlation) may be destroyed during quantum thermalization

## Relation to our papers

Thus, quantum properties (here, non-local correlation) may be completely destroyed during quantum thermalization.
New research topic is to explore the non-equilibrium processes or quantum quench where the state has quantum properties even in the late times.
for example, Many-body-localization (MBL)
Quantum-many-body-scars
This may lead to the non-equilibrium phenomena beyond statical mechanics.

This may be applicable for the quantum computation.

## Why we consider this new topic?

1. These phenomena may be beyond the statical mechanics.
2. These phenomena lead to implementation of quantum computer (quantum computation=non-equilibrium process)
3. In AdS/CFT, this may lead to new finding about black holes

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2. These phenomena lead to implementation of quantum computer (quantum computation=non-equilibrium process)
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## Thermal state on the curved background (on-going)

- New thermodynamical properties related to curvature.
- Insight on the quantum matters near the black hole horizon.


## Contents

- Introduction
- Motivation
- Summary
- Results on this project
- Preliminary
- About summary 2 (Mao Tian's talk)
- About summary 4 (on-going)
- Discussion \& Future directions

Note

The parameter region considered in this talk is

$$
L \gg l_{\mathcal{V}}, t \gg \epsilon \gg 1,
$$

where these parameters are dimensionless and their unit is the lattice spacing.

Note

The parameter region considered in this talk is


Subsystem size Inverse temperature

Note

All the theories considered in this talk are two-dimensional conformal field theories.

## Motivation on these papers

Setup for the time dependent case:
Theories considered are 2d CFTs on spatial circle.
Start from
(Circumstance $=L$ )

$$
\begin{aligned}
& \rho=\frac{e^{-2 \epsilon H}}{\operatorname{tre} e^{-2 \epsilon H}},|\Psi\rangle=\frac{1}{\sqrt{\operatorname{tr} e^{-2 \epsilon H}} \sum_{a} e^{-\epsilon H}|a\rangle_{1} \otimes|a\rangle_{2},} \\
& \text { and }|\Psi\rangle=\frac{\left.e^{-\epsilon H} \mid \text { Bdy }\right\rangle}{\sqrt{\left.\langle\mathrm{Bdy}| e^{-2 \epsilon H} \mid \text { Bdy }\right\rangle}} .
\end{aligned}
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Motivation on these papers Thermal state
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and $|\Psi\rangle=\frac{e^{-\epsilon H}|\mathrm{Bdy}\rangle}{\sqrt{\langle\mathrm{Bdy}| e^{-2 \epsilon H}|\mathrm{Bdy}\rangle}}$.

Motivation on these papers
Thermofield double state Setup for the time dependent case:
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They don't have non-local correlation.
Therefore, we can check whether or not Inhomogeneous evolution endow these states with non-local correlations.

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and $|\Psi\rangle=\frac{e^{-1}|\mathrm{Bdy}\rangle}{\sqrt{\langle\mathrm{Bdy}| e^{-2 \epsilon H}|\mathrm{Bdy}\rangle}}$. $\begin{gathered}\text { This has strong non-local correlation between } \\ \text { Hilbert space one and two. We may be able to }\end{gathered}$ explore other property of this dynamics.

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and $|\Psi\rangle=\frac{\left.e^{-\epsilon H} \mid \text { Bdy }\right\rangle}{\sqrt{\left.\langle\text { Bdy }| e^{-2 \epsilon H} \mid \text { Bdy }\right\rangle}}$ This is easily preparable in the lab..

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Evolve the system with the sine-square deformed Hamiltonian and Mobius Hamiltonian.

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## Explain later

Evolve the system with the sine-square deformed Hamiltonian and Mobius Hamiltonian.

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- It is possible to study non-equilibrium process in the larger system than that numerically-computable!!
- We can study the dynamical property, independent of the finite-size effect, of the system.

Motivation on these papers

- In this setup, entanglement entropy, two point function and so on can be analytically computable.


## This is the reason we consider 2d CFTs.

- We can study the dynamical property, independent of the finite-size effect, of the system.


## Why we consider the inhomogeneous time evolution

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Entanglement entropy (EE) and local temperature may depend on the location of the subsystems.


## Why we consider the inhomogeneous time evolution

- The CFT Hamiltonians on the curved background.

Quasiparicles (excitations generated during the time evolution) may move with the velocity determined by, and distributes inhomogeneously.
Entanglement entropy (EE) and local temperature may depend on the location of the subsystems.

Mutual information may become non-zero.
Non-local correlation may emerge.

## For example

The relation between the thermofield double state and Bell state.

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Thermofield double state in high temperature limit is expected to be a product of Bell states.

$$
|\mathrm{TFD}\rangle \approx \Pi_{\tilde{x}}|\operatorname{Bell} ; \tilde{x}\rangle_{L}|\operatorname{Bell} ; \tilde{x}\rangle_{R}
$$

## For example

## The relation between the thermofield double state and Bell state.

Thermofield double state in high temperature limit is expected to be a product of Bell states.

$$
|\mathrm{TFD}\rangle \approx \Pi_{\tilde{x}}|\operatorname{Bell} ; \tilde{x}\rangle_{L}|\operatorname{Bell} ; \tilde{x}\rangle_{R}
$$

For example, for the infinite-temperature TFD in the spin system

In the infinite-temperature limit, the thermal state is given by the identity.

$$
\Rightarrow e^{0 \times H}=\mathbf{1}=\sum_{i_{1}=\uparrow, \downarrow} \cdots \sum_{i_{L}=\uparrow, \downarrow}\left|i_{1}, \cdots, i_{L}\right\rangle\left\langle i_{1}, \cdots, i_{L}\right|
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For example, for the infinite-temperature TFD in the spin system

$$
\text { On k-th site, } \mid \text { Bell, } k\rangle=\frac{1}{\sqrt{2}}\left(|\uparrow\rangle_{1}|\uparrow\rangle_{2}+|\downarrow\rangle_{1}|\downarrow\rangle_{2}\right)
$$

$e^{0 \times H}=\mathbf{1}=\sum_{i_{1}=\uparrow, \downarrow} \cdots \sum_{i_{L}=\uparrow, \downarrow}\left|i_{1}, \cdots, i_{L}\right\rangle\left\langle i_{1}, \cdots, i_{L}\right| \square$

## Motivation on this paper

The relation between the thermofield double state and Bell state.

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## For example

For $|\Psi(t)\rangle=e^{-i t H_{\text {inh }}^{1}}|\mathrm{TFD}\rangle$
During the inhomogeneous time evolution, the quasiparticles on the first Hilbert space may propagate with the velocity determined by the geometry.

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For $|\Psi(t)\rangle=e^{-i t H_{\text {inh }}^{1}} \mid$ TFD $\rangle$
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Undeformed
Inhomogeneous

$$
d s^{2}=-d t^{2}+d x^{2}
$$

$$
d s^{2}=-f^{2}(x) d t^{2}+d x^{2}
$$


SSD: $\quad f_{\text {sso }}(x)=2 \sin ^{2}\left(\frac{\pi x}{L}\right)$
Velocity: 1
Velocity: $\frac{d x}{d t}=f(x)$

## For example

For $|\Psi(t)\rangle=e^{-i t H_{\text {inh }}^{1}}|\mathrm{TFD}\rangle$
During the inhomogeneous time evolution, the quasiparticles on the first Hilbert space may propagate with the velocity determined by the geometry.


## For example

During the inhomogeneous time evolution, the distribution of quasiparticle on the first Hilbert space may inhomogeneously change with time.

We assume that the number of quasiparticles in the subsystem determines EE. $\Rightarrow$ EE may depend on the position.


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During the inhomogeneous time evolution, the distribution of quasi-particle on the first Hilbert space may inhomogeneously change with time.

EE may depend on the position.


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EE may depend on the position.

$$
S_{\alpha} \neq l_{\alpha}, \quad I_{A, B} \neq 0
$$



## For example

During the inhomogeneous time evolution, the distribution of quasi-particle on the first Hilbert space may inhomogeneously change with time.

Quantum property (non-local
EE may depend on the position. $\square$ correlation) may locally recover or emerges.


## For example

During the inhomogeneous time evolution, the distribution of quasi-particle on the first Hilbert space may inhomogeneously change with time.

Motivation: SSD/Mobius quenches may make the
E system have the temperature gradient(inhomogeneity of quasi-particle).
Quantum nature may emerge.


## Motivation on the thermodynamics

In the time dependent case, the entanglement entropy depends on the location of the subsystem.

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Even for the thermal state, the entanglement entropy depends on the location of the subsystem.

The local temperature may depend on the location.

Some interesting local phenomena (for example, entanglement phase transition) occur.

Summary 1: Preparation of nearly vacuum state by checking with EE
Evolution from the thermal state: $\rho=\frac{e^{-2 \epsilon H}}{\operatorname{tr} e^{-2 \epsilon H}}$


Solid lines are the time-dependences of EE in the twist operator formalism.

Dotted lines are those of EE in the quasi-particle picture.

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During SSD evolution, if subsystem does include the origin (a fixed point) the entanglement entropy is approximated by the thermal entropy.


Volume-law value $S_{A} / c=\frac{\pi X}{3 \epsilon}$


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Evolution from the boundary state: $|\Psi\rangle=\frac{e^{-\epsilon H}|\mathrm{Bdy}\rangle}{\sqrt{\langle\mathrm{Bdy}| e^{-2 \epsilon H}|\mathrm{Bdy}\rangle}}$



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Summary 1: Preparation of nearly vacuum state by checking with MI

We start from $\rho=\frac{e^{-2 \epsilon H}}{\text { tre } e^{-2 \epsilon H}}$ and $|\Psi\rangle=\frac{\left.e^{-H \mid} \mid \text { Bdy }\right\rangle}{\left.\sqrt{\mid \text { Bdy }} e^{-2 \cdot A} \mid \text { Bdy }\right\rangle}$, and then evolve the system with SSD Hamiltonian.

## The time dependence of mutual information (MI) show the mutual information approaches to the vacuum one for any subsystems. $I_{A, B}=S_{A}+S_{B}-S_{A \cup B}$ For example, free fermion

$$
I_{A, B} \approx 0 \longmapsto I_{A, B}^{\text {Vacuum }}
$$



Summary 1: Preparation of nearly vacuum state by checking with MI
 evolve the system with SSD Hamiltonian.

## The SSD time evolution operator endows the states with the vacuum non-local correlation.

$$
I_{A, B} \approx 0 \longmapsto I_{A, B}^{\text {Vacuum }}
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Summary 2: Revival of mutual information from the typical state
The setup considered:
The system in the pure state is unitarily evolved to the typical state with the strong scrambling Hamiltonian (2d holographic Hamiltonian.).

The entanglement entropy for this state follows the Page's curve:

$$
S_{A}=-\operatorname{tr}_{A} \rho_{A} \log \rho_{A} \approx \begin{cases}l_{A} \cdot \log d & \frac{L}{2}>l_{A}>0 \\ \left(L-l_{A}\right) \cdot \log d & L>l_{A}>\frac{L}{2}\end{cases}
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$L$ : system size
$l_{A}$ :subsystem size
$d$ : dimension of
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I_{A, B} \approx 0 \text { for } \frac{L}{2}>l_{A}, l_{B}, l_{A}+l_{B}>0
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$$
\rho_{A \cup B} \underset{1 \gg \epsilon}{\approx} \rho_{A} \otimes \rho_{B}
$$

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There are no non-local correlations of the typical state for the small subsystems $\frac{L}{2}>l_{A}, l_{B}, l_{A}+l_{B}>0$.

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We try to recover the non-local correlation from the thermofield double state by the SSD time evolution.

## Summary 2: Revival of mutual information from the typical

 stateInformation retrieval by using inhomogeneous quenches
(Non-local correlation)
We evolve the system with the 2 d uniform holographic Hamiltonian, $e^{-i H_{0}^{1} t_{0}} \otimes \mathbf{1}_{2}$.

In the large $t_{0}$-regime, the system may be approximated by a typical state.

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EE follows Page's curve.

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Then, we evolve it with the 2 d SSD holographic Hamiltonian, $e^{-i H_{\mathrm{SSD}}^{1} t_{1}} \otimes \mathbf{1}_{2}$.
The non-local correlation is recovered from the typical state: $I_{A, B} \approx \frac{2 c \pi l_{A}}{6 \epsilon}$

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Size of A
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proportional to the number of Bell
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## Summary 3: Genuine tripartite entanglement

The system considered is in:
$\left|\Psi\left(t_{0}, t_{1}\right)\right\rangle=\left(e^{-i t_{0} H_{0}^{1}} \otimes \mathbf{1}_{\mathcal{H}_{2}}\right)\left(e^{-i t_{1} H_{\text {SSD }}^{1}} \otimes \mathbf{1}_{\mathcal{H}_{2}}\right) \frac{1}{\sqrt{\operatorname{tr} e^{-2 \epsilon H_{0}}}} \sum_{a} e^{\frac{-\epsilon}{2}\left(H_{0}^{1}+H_{0}^{2}\right)}|a\rangle_{\mathcal{H}_{1}} \otimes|a\rangle_{\mathcal{H}_{2}}$
Let us divide $\mathcal{H}_{1}$ into $B_{1}, B_{2}$, and the complement to them. $A$ denotes the subsystem of $\mathcal{H}_{2}$.
$B_{1}=\left\{x \left\lvert\, L>L-Y_{1}>x>L-Y_{2}>\frac{L}{2}\right.\right\}$,

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Eigenstates of $H_{0}$
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where $\frac{L}{2}>Y_{1}>Y_{2}>0$.
In 2d Free fermion,


$$
I_{A, B_{i=1,2}} \geq 0, I_{B_{1}, B_{2}} \geq 0
$$

## Summary 3: Quasiparticle picture

The system considered is in:
$\left|\Psi\left(t_{0}, t_{1}\right)\right\rangle=\left(e^{-i t_{0} H_{0}^{1}} \otimes \mathbf{1}_{\mathcal{H}_{2}}\right)\left(e^{-i t_{1} H_{\mathrm{SSD}}^{1}} \otimes \mathbf{1}_{\mathcal{H}_{2}}\right) \frac{1}{\sqrt{\operatorname{tr} e^{-2 \epsilon H_{0}}}} \sum_{a} e^{\frac{-\epsilon}{2}\left(H_{0}^{1}+H_{0}^{2}\right)}|a\rangle_{\mathcal{H}_{1}} \otimes|a\rangle_{\mathcal{H}_{2}}$

- During the SSD time evolution, quasiparticles on $\mathcal{H}_{1}$ move to the fixed point and accumulate there.


Group of left and right-moving quasi-particles


## Summary 3: Quasiparticle picture

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$\left|\Psi\left(t_{0}, t_{1}\right)\right\rangle=\underline{\left(e^{-i t_{0} H_{0}^{1}} \otimes \mathbf{1}_{\mathcal{H}_{2}}\right)}\left(e^{-i t_{1} H_{\text {SSD }}^{1}} \otimes \mathbf{1}_{\mathcal{H}_{2}}\right) \frac{1}{\sqrt{\operatorname{tr} e^{-2 \epsilon H_{0}}}} \sum_{a} e^{\frac{-\epsilon}{2}\left(H_{0}^{1}+H_{0}^{2}\right)}|a\rangle_{\mathcal{H}_{1}} \otimes|a\rangle_{\mathcal{H}_{2}}$
During the SSD time evolution, quasiparticles on $\mathcal{H}_{1}$ move to the fixed point and accumulate there.


During the uniform time evolution, the groups of quasiparticles move left and right at the speed of light.


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The system consider $\left|\Psi\left(t_{0}, t_{1}\right)\right\rangle=\underline{\left(e^{-i t_{0} H_{0}^{1}} \otimes \mathbf{1}_{\mathcal{H}_{2}}\right)\left(e^{-i t_{1} H_{s}^{1}},\right.}$

During the SSD time quasiparticles on $\mathcal{H}_{1}$

Mutual information is given by the number of Bell pairs shared by two subsystems. In the time interval where the group of quasiparticles are in $B_{1}$ or $B_{2}$,

$$
I_{A, B_{i=1,2}} \geq 0
$$ fixed point and accumurate trore.

During the uniform time evolution, the groups of quasiparticles move left and right at the speed of light.


## Summary 3: Genuine tripartite entanglement

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$B$ does not include $x=0$ or $x=L / 2$.
where $\frac{L}{2}>Y_{1}>Y_{2}>0$.


$$
I_{A, B_{i=1,2}} \approx 0, I_{B_{1}, B_{2}} \approx 0
$$

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where $\frac{L}{2}>Y_{1}>Y_{2}>0$.
In 2d holographic CFT,
$I_{A, B_{i=1,2}} \approx 0, I_{B_{1}, B_{2}} \approx 0$

$\mathcal{H}_{1}$

In 2d holographic CFTs, the strong scrambling effect completely delocalize the quasiparticles in $\mathcal{H}_{1}$.

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How about the mutual information between A and $B_{1} \cup B_{2}$ ?

## Summary 3: Genuine tripartite entanglement

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$B$ does not include $x=0$ or $x=L / 2$, then mutual information is approximately zero.
where $\frac{L}{2}>Y_{1}>Y_{2}>0$.


There are time-regimes
where a non-local correlation shared by three parties exists.

$$
I_{A, B_{1} \cup B_{2}} \approx \begin{cases}0 & n L+Y_{2}>t_{0}>n L-Y_{2} \\ \frac{c \pi l_{A}}{3 \epsilon} & n L+Y_{1}>t_{0}>n L+Y_{2} \\ 0 & (n+1) L-Y_{1}>t_{0}>n L+Y_{1} \\ \frac{c \pi l_{A}}{3 \epsilon} & (n+1) L-Y_{2}>t_{0}>(n+1) L-Y_{1}\end{cases}
$$

## Summary 3: Genuine tripartite entanglement

Let us divide $\mathcal{H}_{1}$ into $B_{1}, B_{2}$, and the complement to them. $P_{C, B_{1}}=\frac{3 L}{4}$, and $P_{C, B_{1}}=\frac{L}{4}$. $A$ denotes the subsystem of $\mathcal{H}_{2}$. $B_{1}=\left\{x \left\lvert\, L>L-Y_{1}>x>L-Y_{2}>\frac{L}{2}\right.\right\}$,
$B_{2}=\left\{x \left\lvert\, \frac{L}{2}>Y_{1}>x>Y_{2}>0\right.\right\}$, where $\frac{L}{2}>Y_{1}>Y_{2}>0$.

$\mathcal{H}_{1}$ There are time-regimes
where a non-local correlation
shared by three parties exists.

$$
I_{A, B_{1} \cup B_{2}} \approx \begin{cases}0 & n L+Y_{2}>t_{0}>n L-Y_{2} \\ \frac{c \pi l_{A}}{3 \epsilon} & n L+Y_{1}>t_{0}>n L+Y_{2} \\ 0 & (n+1) L-Y_{1}>t_{0}>n L+Y_{1} \\ \frac{c \pi l_{A}}{3 \epsilon} & (n+1) L-Y_{2}>t_{0}>(n+1) L-Y_{1}\end{cases}
$$

Summary 3: Qu
The system consider $\left|\Psi\left(t_{0}, t_{1}\right)\right\rangle=\underline{\left(e^{-i t_{0} H_{0}^{1}} \otimes \mathbf{1}_{\mathcal{H}_{2}}\right)\left(e^{-i t_{1} H_{s}^{1}}\right.}$

In 2d free fermion and holographic CFT, there are the time intervals where the all quasiparticles on $\mathcal{H}_{1}$ are in $B_{1} \cup B_{2}$ In these time intervals,

$$
I_{A, B_{1} \cup B_{2}} \geq 0
$$

During the SSD time quasiparticles on $\mathcal{H}_{1}$ fixed point and accumulate there.

During the uniform time evolution, the groups of quasiparticles move left and right at the speed of light.


## Summary 3: Genuine tripartite entanglement

## Result 3: Tripartite entanglement

In 2d free fermion (no or weakly scrambling system.)

$$
I_{A, B_{i=1,2}} \geq 0, I_{B_{1}, B_{2}} \geq 0
$$

There are time-regimes
where a non-local correlation shared by three parties exist.

$$
I_{A, B_{1} \cup B_{2}} \approx\left\{\begin{array}{ll}
0 & n L+Y_{2}>t_{0}>n L-Y_{2} \\
\frac{c \pi l_{A}}{3 \epsilon} & n L+Y_{1}>t_{0}>n L+Y_{2} \\
0 & (n+1) L-Y_{1}>t_{0}>n L+Y_{1} \\
\frac{c \pi l_{A}}{3 \epsilon} & (n+1) L-Y_{2}>t_{0}>(n+1) L-Y_{1}
\end{array} .\right.
$$



## Summary 3: Genuine tripartite entanglement

## Result 3: Tripartite entanglement

In 2d holographic CFT (strong scrambling system.),

$$
I_{A, B_{i=1,2}} \approx 0, I_{B_{1}, B_{2}} \approx 0
$$

There are time-regimes
where a non-local correlation shared by three parties exist.

$$
I_{A, B_{1} \cup B_{2}} \approx\left\{\begin{array}{ll}
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\frac{c \pi l_{A}}{3 \epsilon} & (n+1) L-Y_{2}>t_{0}>(n+1) L-Y_{1}
\end{array} .\right.
$$



## Summary 3: Genuine tripartite entanglement

 Result 3: Tripartite entanglementIn 2d holographic CFT (strong scrambling system.)
Key property of this atypical state

- There are no correlations shared by the two parties.
- There are a correlation shared by the three parties.



## Summary 4: The thermodynamic on the curved spacetime (on-going)

We consider the thermodynamic property of the system in 2d holographic CFT on the curved background.
Our thermal state: $\rho=\frac{e^{-\beta H_{q-\text { Möbius }}}}{\operatorname{tr} e^{-\beta H_{q-\text { Möbius }}}}$


## Summary 4: The thermodynamic on the curved spacetime (on-going)

We consider the thermodynamic property of the system in 2d holographic CFT on the curved background.
Our thermal state: $\rho=\frac{e^{-\rho} \operatorname{tr}_{q-\text { Mobiil }}^{H_{q-\text { Mbius }}}}{\text { tre }}$
CFT Hamiltonian on the curved spacetime:

$$
\begin{aligned}
& H_{q-\text { Möbius }}=\int_{0}^{L} d x\left[1-\tanh 2 \theta\left(1-2 \sin ^{2}\left(\frac{q \pi x}{L}\right)\right)\right](T(x)+\bar{T}(x)) \\
& d s^{2}=-f^{2}(x, \theta) d t^{2}+d x^{2}, \quad f(x, \theta)=1-\tanh 2 \theta\left(1-2 \sin ^{2}\left(\frac{q \pi x}{L}\right)\right)
\end{aligned}
$$



## Summary 4: The thermodynamic on the curved spacetime (on-going)

We consider the thermodynamic property of the system in 2d

## Thermal entropy exhibits phase transition with respect to $\theta$.

CFT Hamiltonian on the curved spacetime:

$$
\begin{aligned}
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\end{aligned}
$$



Summary 4: The thermodynamic on the curved spacetime (on-going)

We consider the thermodynamic property of the system in 2d
Thermal entropy exhibits phase transition with respect to $\theta$.

## CFT Hamiltonian on the curved spacetime:



This may be induced by the entanglement phase transition (growth) induced by spacetime.

Details of this study

We will explain the details of

## Summary 2

and

## Summary 4.

## Mobius/SS deformation

The definition of Mobius and sine-square deformed Hamiltonians are

$$
H_{\text {Inho }}=\int_{0}^{L} d x f(x) h(x)
$$

where $\mathrm{h}(\mathrm{x})$ is Hamiltonian density of undeformed one: $H=\int_{0}^{L} d x h(x)$.
The envelop functions considered are

$$
f_{\text {Möbius }}(x)=1-\tanh 2 \theta \cos \left(\frac{2 \pi x}{L}\right), f_{\mathrm{SSD}}(x)=2 \sin ^{2}\left(\frac{\pi x}{L}\right) .
$$

## Mobius/SS deformation

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$$

For $\theta=0, H_{\text {Inho }}=H$. In SSD limit, $\theta \rightarrow \infty, H_{\text {Inho }} \rightarrow H_{\text {SSD }}$.

## The evolution of primary operator

The Mobius/SSD Hamiltonians considered are defined on the spatial circle with $L$, the circumstance.

The evolution of primary operators by these Hamiltonians is given by
where $\left(w_{X}, \bar{w}_{X}\right)=(i X,-i X) . h_{n}$ is the conformal dimension.

## The evolution of primary operator

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The evolution of primary operators by these Hamiltonians is given by
where $\left(w_{X}, \bar{w}_{X}\right)=(i X,-i X), h_{n}$ is the conformal dimension.

This simple transformation makes the computation of EE simpler as explained later.

## The evolution of primary operator

The Mobius/SSD Hamiltonians considered are defined on the spatial circle with $L$, the circumstance.

The evolution of primary operators by these Hamiltonians is given by

$$
e^{i H_{\text {Mbbiues/ss }} \mathrm{st}_{1}} \sigma_{n}\left(w_{X}, \bar{w}_{X}\right) e^{-i H_{\text {Mobiuius }} \text { ss } t_{1}}=\left|\frac{d w_{X}^{\mathrm{New}}}{d w_{x}}\right|^{h_{n}} \sigma_{n}\left(w_{X}^{\mathrm{New}}, \bar{w}_{X}^{\mathrm{New}}\right)
$$

where $\left(w_{X}, \bar{w}_{X}\right)=(i X,-i X) . h_{n}$ is the conformal diphension.

During the time evolution by the inhomogenous Hamiltonians, the operators move along the spatial circle.

Define the spatial position as $X_{X}^{\mathrm{New}}=\frac{w_{X}^{\mathrm{New}}-\bar{w}_{X}^{\mathrm{New}}}{2 i}$.

(a) The SSD time evolution

(b) The Möbius time evolution

During the SSD evolution, for $X=X_{f}^{1}=0, X=X_{f}^{2}=\frac{L}{2}, X_{X}^{\text {New }}$ doesn't move. We call them fixed points.

Define the spatial position as $X_{X}^{\text {New }}=\frac{w_{X}^{\mathrm{New}}-\bar{w}_{X}^{\mathrm{New}}}{2 i}$.

(a) The SSD time evolution

(b) The Möbius time evolution

During the Mobius evolution, the operators periodically move between $X_{f}^{1}$ and $X_{f}^{2}$.

Define the spatial position as $X_{X}^{\text {New }}=\frac{w_{X}^{\mathrm{New}}-\bar{w}_{X}^{\mathrm{New}}}{2 i}$.

(a) The SSD time evolution

(b) The Möbius time evolution

During the SSD evolution, the operators move to $X=X_{f}^{2}=\frac{L}{2}$.

## Preliminary

- Entanglement entropy (EE)

Definition: $S_{A}=\lim _{n \rightarrow 1} \frac{1}{1-n} \log \operatorname{tr}_{A} \rho_{A}^{n}=-\operatorname{tr}_{A} \rho_{A} \log \rho_{A}$

- In the twist operator formalism

$$
S_{A}=\lim _{n \rightarrow 1} \frac{1}{1-n} \log \left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle
$$

By computing two-point function, we can compute (Renyi) entanglement entropy.

## How to compute correlator

Suppose that $\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle$ is given by

$$
\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\sigma_{n}(X) \bar{\sigma}_{n}(Y) U \rho U^{\dagger}\right] .
$$

## How to compute correlator

Suppose that $\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle$ is given by

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$$

Schrödinger picture: $\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\sigma_{n}(X) \bar{\sigma}_{n}(Y) U \rho U^{\dagger}\right]$
Depends on the time evolution (Euclidean geometry is non-trivial)

## How to compute correlator

Suppose that $\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle$ is given by

$$
\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\sigma_{n}(X) \bar{\sigma}_{n}(Y) U \rho U^{\dagger}\right] .
$$

Two pictures: Schrödinger picture:

$$
\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\sigma_{n}(X) \bar{\sigma}_{n}(Y) U \rho U^{\dagger}\right]
$$

Heisenberg picture: $\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\hat{\sigma}_{n}(X) \hat{\bar{\sigma}}_{n}(Y) \rho\right]$
Here, $\hat{\sigma}_{n}(Y)=U^{\dagger} \sigma_{n}(Y) U, \hat{\bar{\sigma}}_{n}(Y)=U^{\dagger} \hat{\sigma}_{n}(Y) U \quad$ Time-independent.

## How to compute correlator

Suppose that $\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle$ is given by

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$$

Two pictures: Schrödinger picture:

$$
\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\sigma_{n}(X) \bar{\sigma}_{n}(Y) U \rho U^{\dagger}\right]
$$

Heisenberg picture: $\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\hat{\sigma}_{n}(X) \hat{\bar{\sigma}}_{n}(Y) \rho\right]$
Here, $\hat{\sigma}_{n}(Y)=U^{\dagger} \sigma_{n}(Y) U, \hat{\bar{\sigma}}_{n}(Y)=U^{\dagger} \hat{\sigma}_{n}(Y) U \quad$ Time-independent. If the operator in Heisenberg picture is simple, then the calculation is simple.

## How to compute correlator

Suppose that $\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle$ is given by

For $U=e^{-i H_{\text {Möbius/SSD }} t_{1}}$,
$e^{i H_{\text {Möbius } / \mathrm{SSD}^{2} t_{1}}} \sigma_{n}(X) e^{-i H_{\text {Möbius } / \mathrm{SSD}} t_{1}}=\left|\frac{d w_{X}^{\mathrm{New}}}{d w_{x}}\right|^{2 h_{n}} \sigma_{n}\left(w_{X}^{\text {New }}, \bar{w}_{X}^{\mathrm{New}}\right)$
where the conformal dimension is $h_{n}=\frac{c\left(n^{2}-1\right)}{24 n}$.

Heisenberg picture: $\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\hat{\sigma}_{n}(X) \hat{\bar{\sigma}}_{n}(Y) \rho\right]$
Here, $\hat{\sigma}_{n}(Y)=U^{\dagger} \sigma_{n}(Y) U, \hat{\bar{\sigma}}_{n}(Y)=U^{\dagger} \hat{\bar{\sigma}}_{n}(Y) U$

Time-independent.
(Euclidean geometry may be simple.)

## In AdS/CFT correspondence

In Schrödinger picture, the dual geometry evolves with time, while the locations of operators don't.

$$
\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\sigma_{n}(X) \bar{\sigma}_{n}(Y) U \rho U^{\dagger}\right]
$$

In Heisenberg picture, the dual geometry is static, while the locations of operators evolves with time.

$$
\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\hat{\sigma}_{n}(X) \hat{\bar{\sigma}}_{n}(Y) \rho\right]
$$

## In AdS/CFT correspondence

In Schrödinger picture, the dual geometry evolves with time, while the locations of operators don't.

$$
\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\sigma_{n}(X) \bar{\sigma}_{n}(Y) U \rho U^{\dagger}\right] \text { dual geometry may be complicated }
$$

In Heisenberg picture, the dual geometry is static, while the locations of operators evolves with time.

$$
\left\langle\sigma_{n}(X) \bar{\sigma}_{n}(Y)\right\rangle=\operatorname{tr}\left[\hat{\sigma}_{n}(X) \hat{\bar{\sigma}}_{n}(Y) \rho\right] \text { dual geometry may be simple }
$$

Results in Summary 2

Entanglement entropy in the twist operator formalism
As a consequence,
$S_{B}=-\frac{c}{12} \log \left[\prod_{i=1,2}\left|\frac{d w_{\gamma_{1}}^{\mathrm{New}}}{d w_{Y_{i}}}\right|^{2}\right]+\lim _{n \rightarrow 1} \frac{1}{1-n} \log \left\langle\sigma_{n}\left(w_{\gamma_{1}}^{\mathrm{Nev}}, \bar{w}_{\gamma_{1}}^{\mathrm{New}}\right) \bar{\sigma}_{n}\left(w_{\gamma_{2}}^{\mathrm{New}}, \bar{w}_{\gamma_{2}}^{\mathrm{Now}}\right)\right\rangle$ where B is the subsystem of the $\mathcal{H}_{1}$.

Entanglement entropy in the twist operator formalism
As a consequence,
$S_{B}=-\frac{c}{12} \log \left[\prod_{i=1,2}\left|\frac{d w_{Y_{i}}^{\mathrm{New}}}{d w_{Y_{i}}}\right|^{2}\right]+\lim _{n \rightarrow 1} \frac{1}{1-n} \log \left\langle\sigma_{n}\left(w_{Y_{1}}^{\mathrm{New}}, \bar{w}_{Y_{1}}^{\mathrm{New}}\right) \bar{\sigma}_{n}\left(w_{Y_{2}}^{\mathrm{New}}, \bar{w}_{Y_{2}}^{\mathrm{New}}\right)\right\rangle$ where B is the subsystem of the $\mathcal{H}_{1}$.

The piece depending on the detail of CFT.

Note

The parameter region considered in this talk is

$$
L \gg l_{\mathcal{V}}, t \gg \epsilon \gg 1,
$$

where these parameters are dimensionless and their unit is the lattice spacing.

Entanglement entropy in the twist operator formalism

## As a consequence,

$S_{B}=-\frac{c}{12} \log \left[\prod_{i=1,2}\left|\frac{d w_{Y_{i}}^{\mathrm{New}}}{d w_{Y_{i}}}\right|^{2}\right]+\lim _{n \rightarrow 1} \frac{1}{1-n} \log \left\langle\sigma_{n}\left(w_{Y_{1}}^{\mathrm{New}}, \bar{w}_{Y_{1}}^{\mathrm{New}}\right) \bar{\sigma}_{n}\left(w_{Y_{2}}^{\mathrm{New}}, \bar{w}_{Y_{2}}^{\mathrm{New}}\right)\right\rangle$ where B is the subsystem of the $\mathcal{H}_{1}$.

Geodesic length associated with B on the BTZ-black-hole geometry.

Vacuum entanglement entropy

$$
S_{B}=-\frac{c}{12} \log \left[\prod_{i=1,2}\left|\frac{d w_{Y_{i}}^{\mathrm{New}}}{d w_{Y_{i}}}\right|^{2}\right]+\lim _{n \rightarrow 1} \frac{1}{1-n} \log \left\langle\sigma_{n}\left(w_{Y_{1}}^{\mathrm{New}}, \bar{w}_{Y_{1}}^{\mathrm{New}}\right) \bar{\sigma}_{n}\left(w_{Y_{2}}^{\mathrm{New}}, \bar{w}_{Y_{2}}^{\mathrm{New}}\right)\right\rangle_{2 \epsilon}
$$



Vacuum entanglement entropy

## Geodesic length



Vacuum entanglement entropy


$$
\mathcal{O}(1) \ll \mathcal{O}(1 / \epsilon)
$$

Define the spatial position as $X_{X}^{\mathrm{New}}=\frac{w_{X}^{\mathrm{New}}-\bar{w}_{X}^{\mathrm{New}}}{2 i}$.

(a) The SSD time evolution

(b) The Möbius time evolution

During the SSD evolution, the operators move to $X=X_{f}^{2}=\frac{L}{2}$.

Define the spatial position as $X_{X}^{\mathrm{New}}=\frac{w_{X}^{\mathrm{New}}-\bar{w}_{X}^{\mathrm{New}}}{2 i}$.

(a) The SSD time evolution

Under the SSD evolution, the or

# The (effective) size of B monotonically decrease with times. 

## Vacuum entanglement entropy

$$
S_{B}=-\frac{c}{12} \log \left[\prod_{i=1,2}\left|\frac{d w_{Y_{i}}^{\mathrm{New}}}{d w_{Y_{i}}}\right|^{2}\right]+\lim _{n \rightarrow 1} \frac{1}{1-n} \log \left\langle\sigma_{n}\left(w_{Y_{1}}^{\mathrm{New}}, \bar{w}_{Y_{1}}^{\mathrm{New}}\right) \bar{\sigma}_{n}\left(w_{Y_{2}}^{\mathrm{New}}, \bar{w}_{Y_{2}}^{\mathrm{New}}\right)\right\rangle_{2 \epsilon}
$$



## EE decreases



The area of minimal surface decreases with time. ${ }^{X_{f}^{1}}$

## Vacuum entanglement entropy

$$
S_{B}=-\frac{c}{12} \log \left[\prod_{i=1,2}\left|\frac{d w_{Y_{i}}^{\mathrm{New}}}{d w_{Y_{i}}}\right|^{2}\right]+\lim _{n \rightarrow 1} \frac{1}{1-n} \log \left\langle\sigma_{n}\left(w_{Y_{1}}^{\mathrm{New}}, \bar{w}_{Y_{1}}^{\mathrm{New}}\right) \bar{\sigma}_{n}\left(w_{Y_{2}}^{\mathrm{New}}, \bar{w}_{Y_{2}}^{\mathrm{New}}\right)\right\rangle_{2 \epsilon}
$$

$$
\mathcal{O}(1)+\mathcal{O}(1) \approx S_{A}^{\mathrm{Vacuum}}=\frac{c}{3} \log \left[\frac{L}{\pi} \sin \left[\frac{\pi\left(Y_{1}-Y_{2}\right)}{L}\right]\right]
$$



The area of
minimal surface decreases with time.


## Vacuum entanglement entropy

$$
S_{B}=-\frac{c}{12} \log \left[\prod_{i=1,2}\left|\frac{d w_{Y_{i}}^{\mathrm{New}}}{d w_{Y_{i}}}\right|^{2}\right]+\lim _{n \rightarrow 1} \frac{1}{1-n} \log \left\langle\sigma_{n}\left(w_{Y_{1}}^{\mathrm{New}}, \bar{w}_{Y_{1}}^{\mathrm{New}}\right) \bar{\sigma}_{n}\left(w_{Y_{2}}^{\mathrm{New}}, \bar{w}_{Y_{2}}^{\mathrm{New}}\right)\right\rangle_{2 \epsilon}
$$

$$
\mathcal{O}(1)+\mathcal{O}(1) \approx S_{A}^{\text {Vacuum }}=\frac{c}{3} \log \left[\frac{L}{\pi} \sin \left[\frac{\pi\left(Y_{1}-Y_{2}\right)}{L}\right]\right]
$$


decreases with time.

## Mutual information of thermofield double state

As a simple example,
$\left|\phi\left(t_{1}\right)\right\rangle=\left(e^{-i t_{1} H_{\text {SSD }}^{1}} \otimes \mathbf{1}_{2}\right)|\mathrm{TFD}\rangle$
let us report the time-dependence of $I_{A, B}$.

Let $A$ and $B$ denote the subsystems of $\mathcal{H}_{2}$ and $\mathcal{H}_{1}$, respectively.
B includes $X_{f}^{1}$.
Here, $l_{\mathcal{V}=A, B}, P_{C, \mathcal{V}=A, B}$ denote the subsystem sizes and centers of $A$ and $B$, respectively.



## Mutual information of thermofield double state

As a simple example,
$\left|\phi\left(t_{1}\right)\right\rangle=\left(e^{-i t_{1} H_{\text {SSD }}^{1}} \otimes \mathbf{1}_{2}\right)|\mathrm{TFD}\rangle$
let us report the time-dependence of $I_{A, B}$.


## I Under the SSD-time evolution, mutual information between $A$ and $B$ returns to the initial value.

subsystem sizes and centers of $A$ and $B$, respectively.


## Why does mutual information revive?



## Why does mutual information revive?



## Why does mutual information revive?



## Why does mutual information revive?



## Mutual information of thermofield double state

We start from thermofield double state, then evolve the system with the 2d holographic Hamiltonian,

$$
\left|\Psi\left(t_{0}\right)\right\rangle=\left(e^{-i H_{0}^{1} t_{0}} \otimes \mathbf{1}_{2}\right)|\mathrm{TFD}\rangle
$$

For the time-regime, $t_{0} \gg \mathcal{O}(L)$, $I_{A, B}$ should be completely destroyed.


## Mutual information of thermofield double state

For the time-regime, $t_{0} \gg \mathcal{O}(L)$,
$I_{A, B}$ should be completely destroyed.
Then, we evolve this state with the SSD Hamiltonian,

$$
\left|\Psi\left(t_{1}, t_{0}\right)\right\rangle=\left(e^{-i H_{\mathrm{SSD}}^{1} t_{1}} \otimes \mathbf{1}_{2}\right)\left|\Psi\left(t_{0}\right)\right\rangle
$$



## Mutual information of thermofield double state

For the time-regime, $t_{0} \gg \mathcal{O}(L)$,
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$$

Correlation revives.


## Mutual information of thermofield double state

For the time-regime, $t_{0} \gg \mathcal{O}(L)$,
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$$

Correlation revives.


Results in Summary 4

## The thermodynamic on the curved spacetime


Hamiltonian: $H_{\tau-\text { Msouis }}=\int_{0}^{L} d x\left[1-\tanh 2 \theta\left(1-2 \sin ^{2}\left(\frac{q \pi x x}{L}\right)\right)\right](T(x)+\bar{T}(x))$
( $q$ is an integer.)

$$
=\int_{0}^{L} d x \sqrt{-\operatorname{det} g}(T(x)+\bar{T}(x))
$$

Geometry: $d s^{2}=-f^{2}(x, \theta) d t^{2}+d x^{2}, f(x, \theta)=1-\tanh 2 \theta\left(1-2 \sin ^{2}\left(\frac{q \pi x}{L}\right)\right)$.

Curvature: $\quad R(\theta, x)=-\frac{2 \partial_{x}^{2} f(\theta, x)}{f(\theta, x)}=\frac{8 \pi^{2} q^{2} \tanh (2 \theta) \cos \left(\frac{2 \pi x x}{L}\right)}{L^{2}\left(\tanh (2 \theta) \cos \left(\frac{2 \pi x a}{L}\right)-1\right)}$



## The thermodynamic on the curved spacetime


Hamiltonian: $H_{q-\text { Mäbius }}=\int_{0}^{L} d x\left[1-\tanh 2 \theta\left(1-2 \sin ^{2}\left(\frac{q \pi x}{L}\right)\right)\right](T(x)+\bar{T}(x))$
( $q$ is an integer.)

$$
=\int_{0}^{L} d x \sqrt{-\operatorname{det} g}(T(x)+\bar{T}(x))
$$

Geometry: $d s^{2}=-f^{2}(x, \theta) d t^{2}+d x^{2}, f(x, \theta)=1-\tanh 2 \theta\left(1-2 \sin ^{2}\left(\frac{q \pi x}{L}\right)\right)$
Curvature around $x=0, \mathrm{~L} / 4, \mathrm{~L} / 2,3 \mathrm{~L} / 4$
Curvat negatively grows with $\theta$.

## The thermodynamic on the curved spacetime


We assume $L / \beta<1$ (low temp.).
For $\theta=0$ (flat), entropy is $S \approx \mathcal{O}(1)$.
The moduli parameter of this torus is $\tau=\frac{L \cosh 2 \theta}{\beta}$


## The thermodynamic on the curved spacetime


We assume $L / \beta<1$ (low temp.).
For $\theta=0$ (flat), entropy is $S \approx \mathcal{O}(1)$.
The moduli parameter of this torus is $\tau=\frac{L \cosh 2 \theta}{\beta}$.
Therefore, if $\theta$ increases, then
$S \approx \begin{cases}\mathcal{O}(1) & L \cosh 2 \theta / \beta<1 \\ \frac{c \pi L \cosh 2 \theta}{6 \beta} \underset{\theta \gg 1}{\propto} \frac{L e^{2 \theta}}{\epsilon} & L \cosh 2 \theta / \beta>1\end{cases}$


## The thermodynamic on the curved spacetime


We assume $L / \beta<1$ (low temp.).
For $\theta=0$ (flat), entropy is $S \approx \mathcal{O}(1)$.
The moduli parameter of this torus is $\tau=\frac{L \cosh 2 \theta}{\beta}$.
Therefore, if $\theta$ increases, then

$$
S \approx \begin{cases}\mathcal{O}(1) & L \cosh 2 \theta / \beta<1 \\ \frac{c \pi L \cosh 2 \theta}{6 \beta} \underset{\theta \gg 1}{\propto} \frac{L e^{2 \theta}}{\epsilon} & L \cosh 2 \theta / \beta>1\end{cases}
$$



## The thermodynamic on the curved spacetime

Our thermal state: $\rho=\frac{e^{-\beta H_{-}}}{\text {tre } e^{-\beta H_{- \text {Shine }}}}$ We assume $L / \beta<1$ (low temp.).
For $\theta=0$ (flat), entropy is $S \approx \mathcal{O}(1)$.
The moduli parameter of this torus is $\tau=\frac{L \cosh 2 \theta}{\beta}$.

$$
\mathrm{q}=4
$$ Therefore, if $\theta$ increases, then

$$
S \approx \begin{cases}\mathcal{O}(1) & L \cosh 2 \theta / \beta<1 \\ \frac{c \pi L \cosh 2 \theta}{6 \beta} \underset{\theta \gg 1}{\propto} \frac{L e^{2 \theta}}{\epsilon} & L \cosh 2 \theta / \beta>1\end{cases}
$$

## The thermodynamic on the curved spacetime

Our thermal state: $\rho=\frac{e^{-\beta H_{-}} \operatorname{Hen}_{- \text {Nabiues }}}{\text { tre }}$
We assume $L / \beta<1$ (low temp.).
Here, A doesn't include $x=0, \mathrm{~L} / 4, \mathrm{~L} / 2,3 \mathrm{~L} / 4$.
$B$ includes $x=L / 2$.
In the large $\theta$ limit, the behavior of entanglement entropy is

$$
\begin{aligned}
& S_{A} \approx \frac{c}{3} \log \left[\frac{L}{4 \pi} \sin \left(\frac{4 \pi l_{A}}{L}\right)\right] \\
& S_{B} \approx \frac{c \cdot C_{\text {cof. }} L e^{2 \theta}}{\beta}
\end{aligned}
$$

$$
\text { Independent of } \theta \text {. }
$$

Similar to Vacuum EE

$$
\text { on the interval of } \mathrm{L} / 4
$$



## The thermodynamic on the curved spacetime


We assume $L / \beta<1$ (low temp.).
Here, A doesn't include $x=0, L / 4, L / 2,3 \mathrm{~L} / 4$.
$B$ includes $x=L / 2$.
In the large $\theta$ limit, the behavior of
 entanglement entropy is

$$
\begin{aligned}
& S_{A} \approx \frac{c}{3} \log \left[\frac{L}{4 \pi} \sin \left(\frac{4 \pi l_{A}}{L}\right)\right] \Rightarrow \begin{array}{l}
\text { In dependent of } \theta . \\
\text { Similar to Vacuum EE } \\
\text { on the interval of } \mathrm{L} / 4 .
\end{array} \\
& S_{B} \approx \frac{c \cdot C_{\text {cof. }} . L e^{2 \theta}}{\beta} \Rightarrow \text { Exponential growth with } \theta .
\end{aligned}
$$



## The thermodynamic on the curved spacetime


We assume $L / \beta<1$ (low temp.).
When the subsystem includes the region
3L/4. where curvature is negatively large, entanglement entropy grows with $\theta$.

entanglement entropy is
$\begin{aligned} S_{A} & \approx \frac{c}{3} \log \left[\frac{L}{4 \pi} \sin \left(\frac{4 \pi l_{A}}{L}\right)\right] \Rightarrow \begin{array}{l}\text { In dependent of } \theta . \\ \text { Similar to Vacuum EE } \\ \text { on the interval of } \mathrm{L} / 4 .\end{array} \\ S_{B} \approx \frac{c \cdot C_{\mathrm{cof} .} . L e^{2 \theta}}{\beta} & \Rightarrow \text { Exponential growth with } .\end{aligned}$


## The thermodynamic on the curved spacetime


We assume $L / \beta<1$ (low temp.).

When the subsystem includes the region where curvature is negatively large, entanglement entropy grows with $\theta$.

Phase transition may be induced by the entanglement phase transition in the region where curvature is negatively large.
entanglement entropy is

$$
\begin{array}{ll}
S_{A} \approx \frac{c}{3} \log \left[\frac{L}{4 \pi} \sin \left(\frac{4 \pi l_{A}}{L}\right)\right] \quad \begin{array}{l}
\text { In dependent of } \theta . \\
\text { Similar to Vacuum EE } \\
\text { on the interval of L/4. }
\end{array} \\
S_{B} \approx \frac{c \cdot C_{\text {cof. }} L e^{2 \theta}}{\beta} & \Rightarrow \text { Exponential growth wi }
\end{array}
$$



## The thermodynamic on the curved spacetime


We assume $L / \beta<1$ (low temp.).
When the subsystem includes the region where curvature is negatively large, entanglement entropy grows with $\theta$.

Phase transition may be induced by the entanglement phase transition in the region where curvature is negatively large.
entanglement entropy is
Spacetime (maybe, curvature? or geometry?) generates entanglement or induces entanglement phase transition.

## Summaries

- We studied the time dependence and thermodynamic properties of entanglement entropy and mutual information in two-dimensional inhomogeneous conformal field.
- During the time evolution from the TFD, mutual information can revive(non-local correlation retrieval).
- In holographic CFT on the curved spacetime, the phase transition (entanglement phase transition) related to curvature may occurs.


## Future directions

- Quantum many body-scars
- Measurement-induced phase transition
- ETH and thermalization on the curved spacetime
- Quantum simulation
- Cosmology
- Experiments


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