



De Sitter Horizons and Holography

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Based on work in collaboration with D. Anninos and D. Hofman

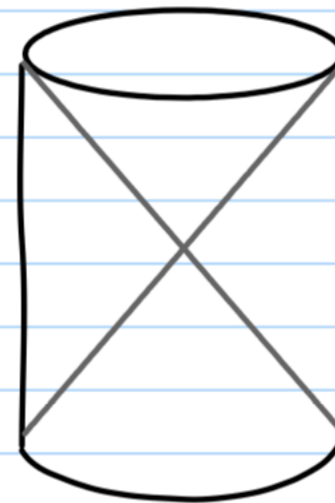
AdS, dS: similar, but different



(from Eva's talk)

More generally, we need a complete framework for the $V > 0$ landscape.

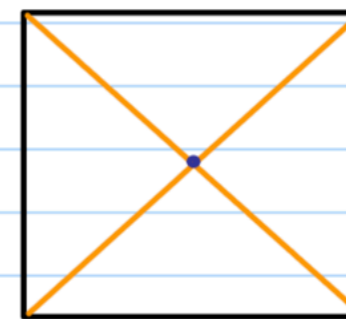
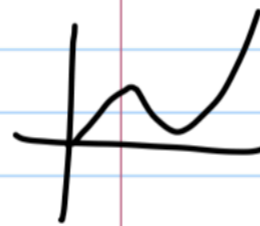
$\Lambda < 0$



AdS:

observables are
conformal field
theory (CFT)
correlation functions
**Timelike boundary at
infinity pins fields.**

$\Lambda > 0$

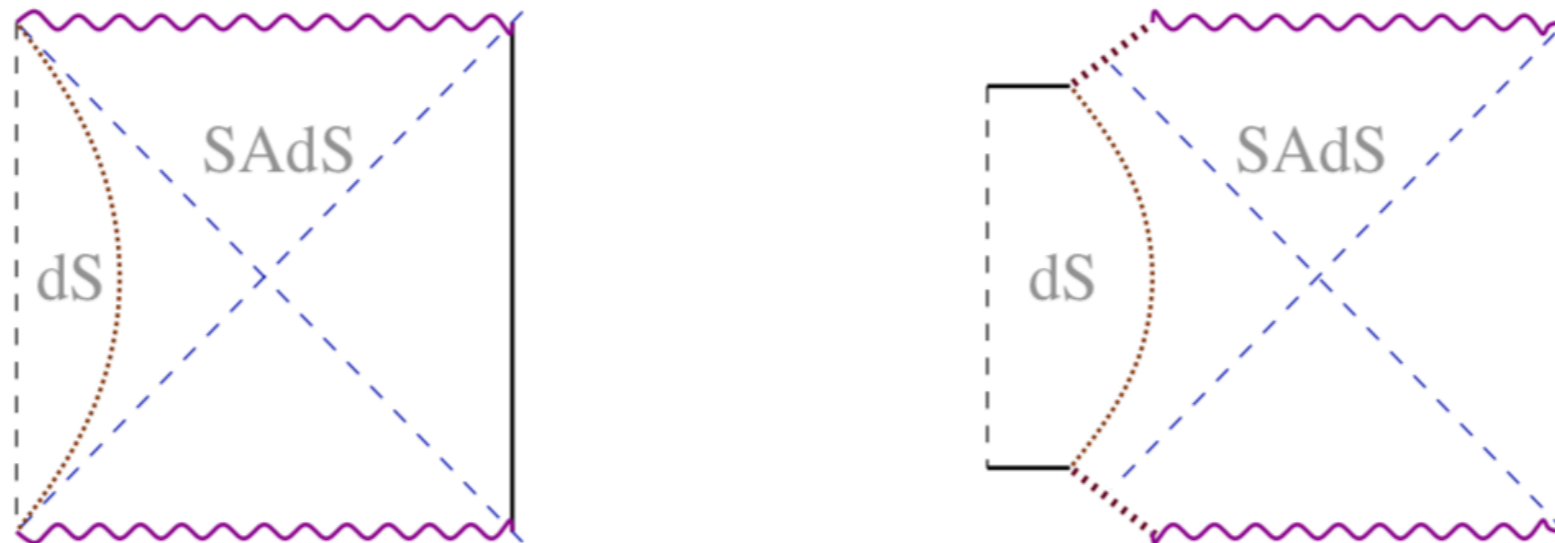


dS

**no boundary
analogous to the
AdS one. (Finite,
fluctuating space.)**

Embedding dS into AdS

- A natural idea would be to try to embed a patch of de Sitter into an AdS geometry.
- This was fostered by [Freivogel, Hubeny, Maloney, Myers, Rangamani, Shenker, 2005].
- But, EoMs + Energy conditions required the dS patch to hide behind the horizon of an AdS Schwarzschild solution.



- But this is for $d > 2$. What about (A)dS₂?

A few words on dS_2

- The global metric is described by

$$ds_{global}^2 = \frac{-d\tau^2 + d\varphi^2}{\cos^2 \tau}, \quad \tau \in [-\pi/2, \pi/2]$$

- A single observer only has access to the static patch geometry:

$$ds_{static}^2 = \frac{-dt^2 + d\rho^2}{\cosh^2 \rho}, \quad \rho \in [-\infty, \infty]$$

- The patch is always finite.
- The static patch has two horizons.
- It appears as a solution in higher dimensions, $dS_2 \times S^{d-2}$, in the so-called Nariai limit.

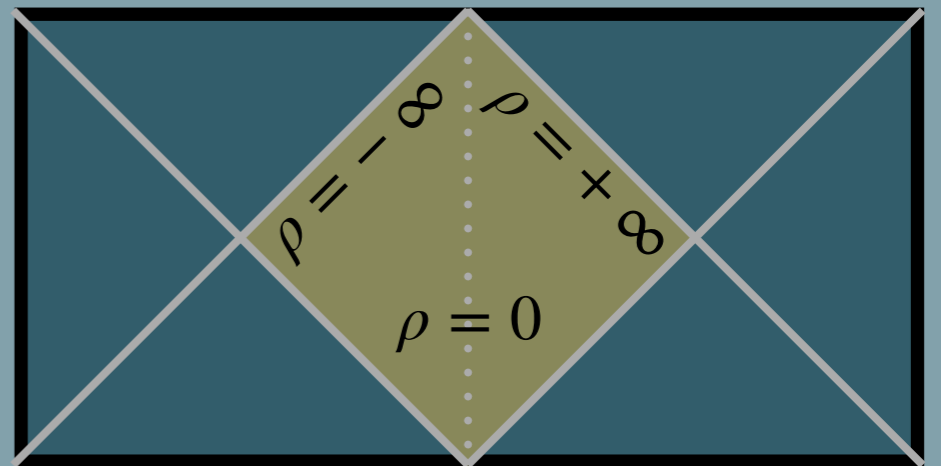
[Maldacena, Turiaci, Yang, '19]
[Cotler, Jensen, Maloney, '19]

\mathcal{F}^+



\mathcal{F}^-

\mathcal{I}^+



\mathcal{I}^-

An action for the centaurs

- In 2d, the Einstein action is topological

$$S_{top} = -\frac{\phi_0}{2} \int d^2x \sqrt{g} R - \phi_0 \int_{\partial\mathcal{M}} du \sqrt{h} K$$

- We will consider the following (Euclidean) gravity-dilation action

$$S_E = -\frac{1}{2\kappa} \int d^2x \sqrt{g} (\phi R + V(\phi)) - \frac{1}{\kappa} \int_{\partial\mathcal{M}} du \sqrt{h} \phi K$$

- It is important that the full dilaton remains positive, so that

$$\phi_{tot} = (\phi_0 + \phi/\kappa) \gg 0$$

- The equations of motion can be simply recasted as

$$R = -V'(\phi) \quad , \quad T_{\mu\nu}^\phi = 0 \quad .$$

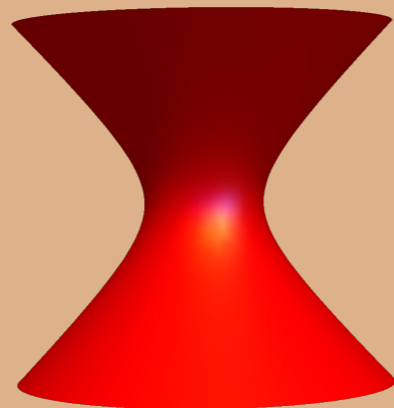
An action for the centaurs II

- The equations of motion can be simple recasted as

$$R = -V'(\phi) \quad , \quad T_{\mu\nu}^{\phi} = 0 .$$

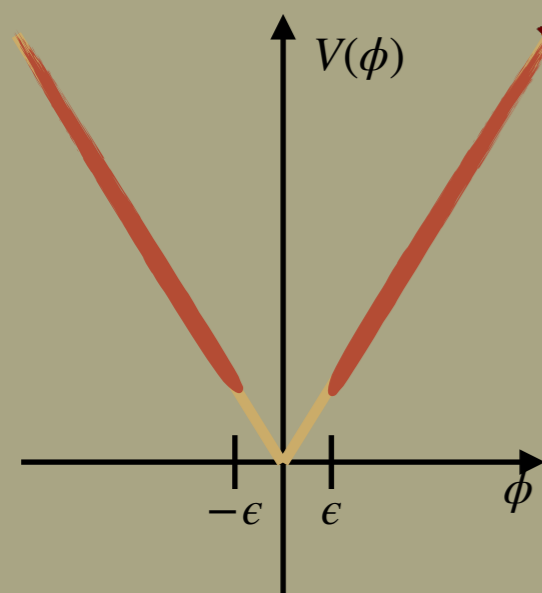
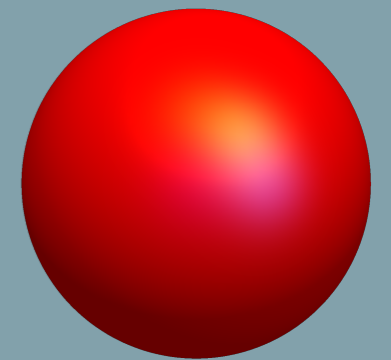
- If the potential is $+2\phi$, then the solution is (n)AdS₂:

$$\begin{cases} ds^2 = \frac{dt^2 + d\rho^2}{\cos^2 \rho} \\ \phi = \phi_h \tan \rho \end{cases}$$



- If the potential is -2ϕ , then the solution is (n)dS₂:

$$\begin{cases} ds^2 = \frac{dt^2 + d\rho^2}{\cosh^2 \rho} \\ \phi = \phi_h \tanh \rho \end{cases}$$

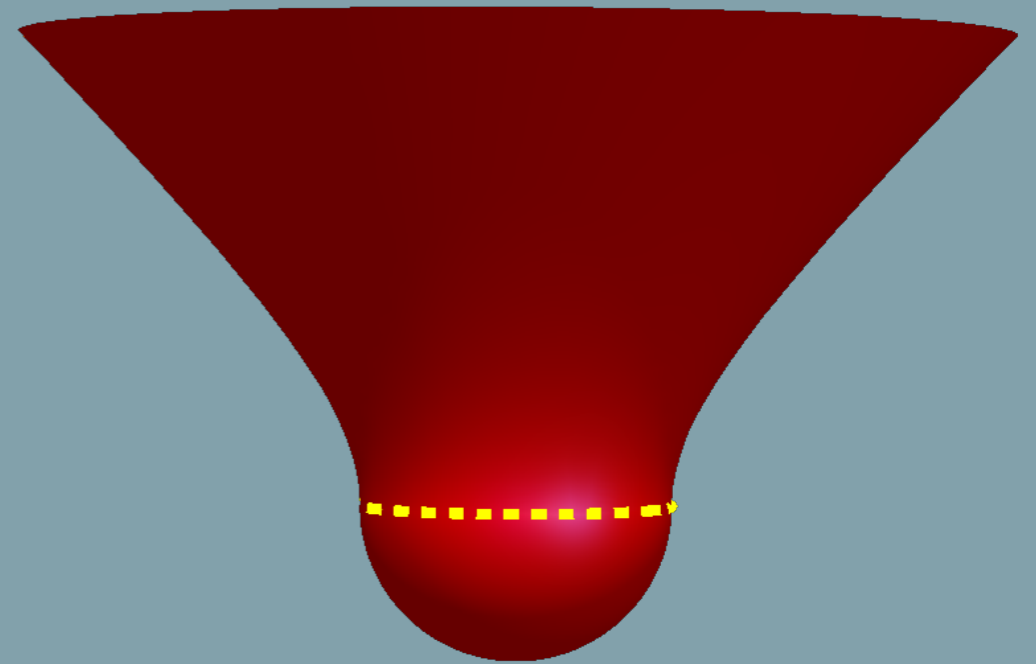
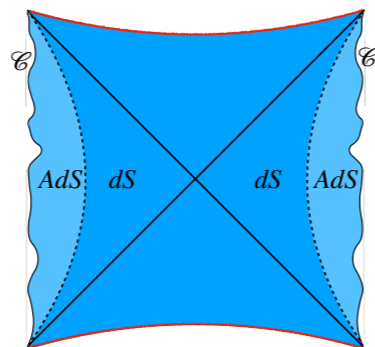


$$= \begin{cases} \cos^{-2} \rho (d\rho^2 + d\tau^2) \quad , \quad \rho \in (\pi/2, 0) \quad , \\ \cosh^{-2} \rho (d\rho^2 + d\tau^2) \quad , \quad \rho \in (-\infty, 0) \quad . \end{cases}$$

$$\phi = \begin{cases} -\phi_h \tan \rho \quad , \quad \rho \in (\pi/2, 0) \quad , \\ -\phi_h \tanh \rho \quad , \quad \rho \in (-\infty, 0) \quad . \end{cases}$$

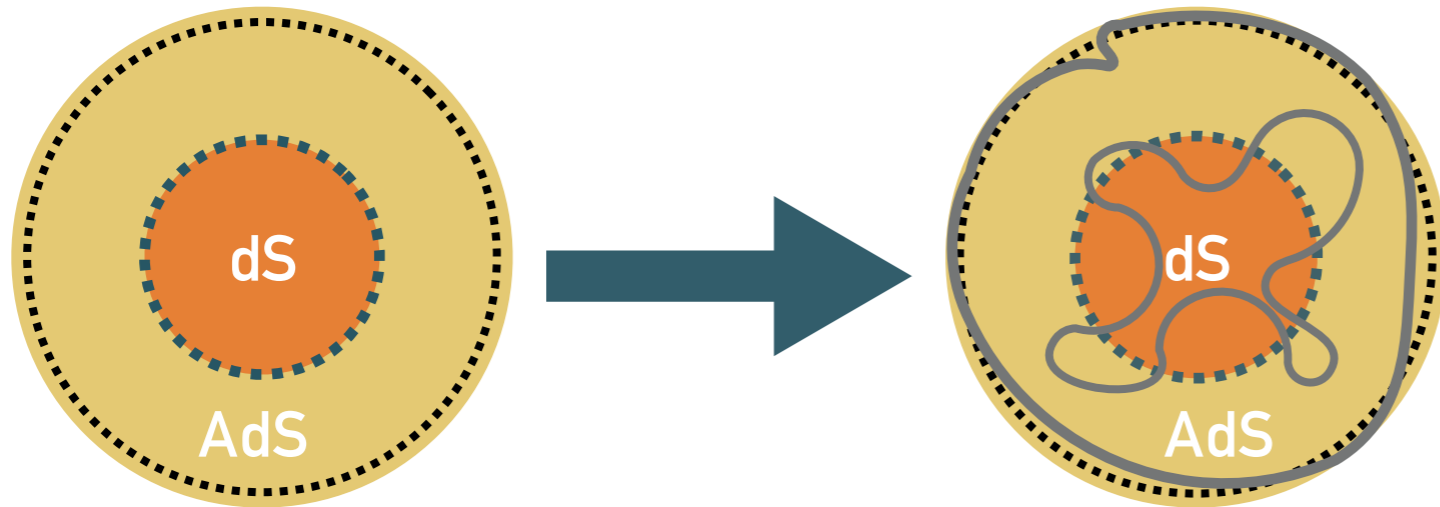
Properties of the centaur

- It is a smooth, differentiable geometry that interpolates between AdS and dS.
- It has the conformal boundary of AdS but deep in the infrared it goes to the static patch of dS.
- The dilaton grows monotonically, so there is no violation of the (higher dimensional) NEC.
- We can also build the Lorentzian centaur via an analogue Hartle-Hawking mechanism.

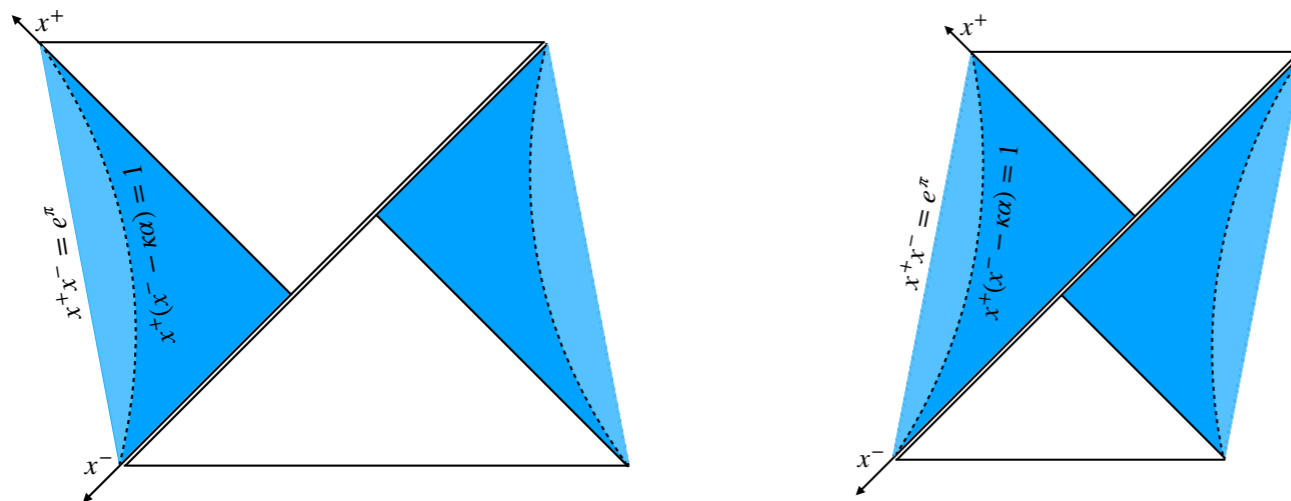


Some holographic experiments

- Perturbative Euclidean deformations

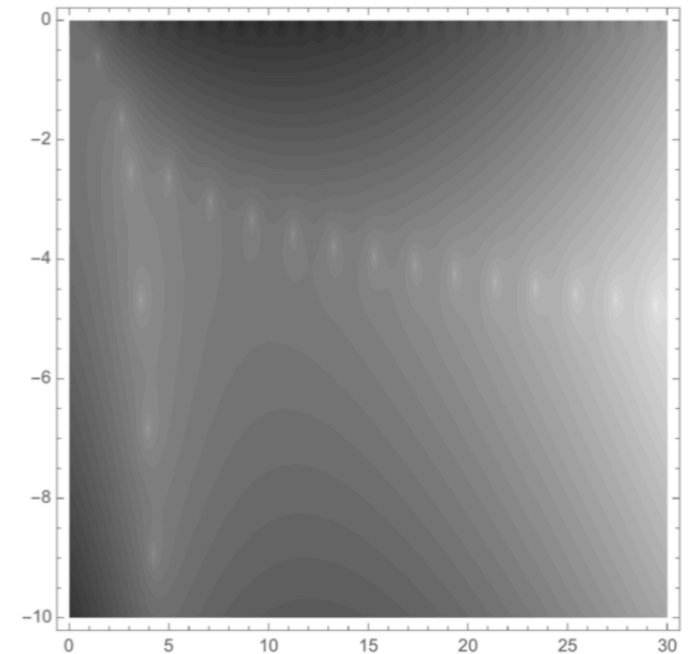


- Shockwave deformations



- **Boundary modes and chaos**

- Quasinormal modes



See 1703.04622 and 1811.08153 or ask me later!!

The boundary theory

- We still need to fix boundary conditions. The Dirichlet problem defines a curve close to the boundary $\mathcal{C} = \{\tau(u), \rho(u)\}$

$$h(u) = e^{2\omega(\rho(u), \tau(u))} \left((\partial_u \tau(u))^2 + (\partial_u \rho(u))^2 \right) , \quad \phi_b(u) = \phi(\rho(u), \tau(u))$$

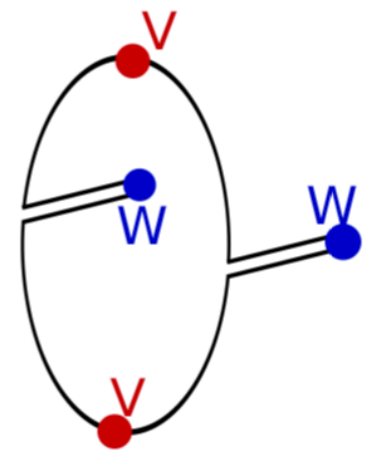
- Evaluating the action we obtain a boundary action

$$S_{bdy} = \frac{\tilde{\phi}_b}{\kappa} \int du \left(\frac{1}{2} (\partial_u \tau(u))^2 - Sch[\tau(u), u] \right) .$$

- The clock close to the boundary of the centaur geometries is that of the hyperbolic cylinder instead of the hyperbolic disk.
- Expanding around the saddle $\tau=u$, we obtain an effective action for the fluctuations $\tau=u+\delta\tau(u)$:

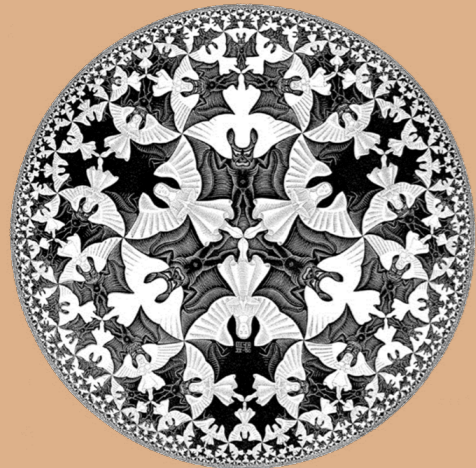
$$S_{fluct} = \frac{\tilde{\phi}_b}{2\kappa} \int du \left((\partial_u^2 \delta\tau(u))^2 + (\partial_u \delta\tau(u))^2 \right) .$$

Out-of-time ordered correlators



- Now we will add some matter perturbations and compute the out-of-time ordered correlator for these geometries:

$$F(t) \equiv \langle V(\pi/2)W(it)V(-\pi/2)W(-\pi + it) \rangle_c$$



The AdS₂ case is maximally chaotic!!

$$S_{fluct} = \frac{\tilde{\Phi}_b}{2\kappa} \int du \left((\partial_u^2 \delta\tau(u))^2 - (\partial_u \delta\tau(u))^2 \right)$$

compute, compute, compute ...

$$F(t) \propto \frac{\kappa\beta}{\tilde{\Phi}_b} \exp \frac{2\pi}{\beta} t \quad \lambda_L = \frac{2\pi}{\beta}$$



The centaurs have oscillatory OTOC!!

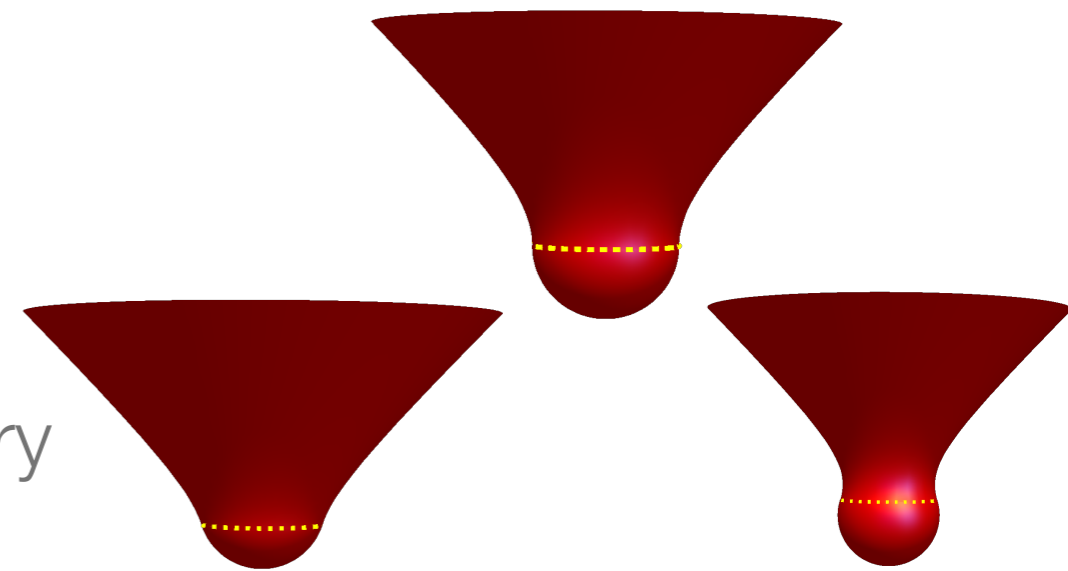
$$S_{fluct} = \frac{\tilde{\Phi}_b}{2\kappa} \int du \left((\partial_u^2 \delta\tau(u))^2 + (\partial_u \delta\tau(u))^2 \right)$$

compute, compute, compute, compute,
compute, compute, compute, compute,
compute, compute, compute, compute ...

$$F(t) \propto \frac{\kappa\beta}{\tilde{\Phi}_b} \cos \frac{2\pi}{\beta} t$$

De Sitter horizons and chaos

- We found that dS horizons do not exhibit a chaotic behaviour in their out-of-time-ordered correlator. This is quite striking since the naive argument is that the gravity result is universal for horizons.
- We consider this a distinctive feature of dS horizons rather than a “bug”.
- We can even generalise our construction:
- We obtain a generalised Schwarzian theory



$$S_{bdy} = \frac{\phi_b}{\kappa} \int du \left(\frac{\gamma}{2} (\partial_u \tau(u))^2 - Sch[\tau(u), u] \right)$$

Exponential, when $\gamma < 0$

Power law, when $\gamma = 0$

Oscillatory, when $\gamma > 0$

Questions and future directions

- Can we understand the γ -Schwarzian theory from a quantum mechanical point of view? Is there an SYK-like model that exhibits these features? [T. Anous, J. Sonner; T. Mertens, J. Turiaci; J. Yoon; K. Jensen]
- What about the $\kappa < 0$ solutions and the non-suppressed fluctuations? How can we make sense of this theory?
- Would it possible to get rid of the AdS part? [e.g.: poster by A. Rolph]
- Are there centaurs in higher dimensions?
- What is the relation between this construction and dS/CFT correspondence? Or the dS/dS correspondence?

Lots to explore...