Poster No.	Surname(s)	Given name(s)	Affiliation	Title	Abstract
1	Afrasiar	Mir	Indian Institute of Technology Kanpur	Entanglement Negativity Islands in Communicating Black Holes	We obtain the holographic conformal field theories du we analyze the mixed state separate models. The first boundaries describing com dimensional eternal JT blac described by \$CFT_2\$s. Fo obtained earlier in the com
2	Arai	Hayato	Nagoya University	Pseudo standard entanglement structure cannot be distinguished from standard entanglement structure	An experimental verification entangled state, but it does entanglement structure is local systems. Therefore, the is valid. To examine this iss under natural assumptions states. Surprisingly, there es structure. In our setting, ar obtained pseudo standard
3	Choudhury	Sayantan	The Thanu Padmanabhan Centre for Cosmology and Science Popularization (CCSP), SGT University, Gurugram, Haryana, India	Wormhole without averaging from O(N)q-1 tensor model	The SYK model has a worm couplings are fixed the con "half-wormholes". In this p model with O(N)q — 1 gaug the SYK model. We will res charges or non-trivial cobo especially at short distance wormholes and make the t our "half-wormholes". We leading order contributions non-trivial saddles from "h perturbative effects.
4	Čubrović	Mihailo	Institute of Physics Belgrade	Classical integrability and quantum chaos of open string dynamics	We consider an open string dynamics is always integra nonzero and roughly satura and energy levels, which sh system, where chaos is stre
5	-	-	-	-	Cancelled
6	Fujita	Mitsutoshi	Sun Yat-Sen University	Holographic entanglement entropy of the double Wick rotated BTZ black hole	In this paper, we analyze the hole (3 dimensional Kerr-An theory has negative energy which is obtained by a cont
7	Li	Yue-Zhou	McGill University	Causality constraints on corrections of Einstein gravity	Causality and unitarity as o modifications of Einstein g by using dispersive sum rul >0 limit, and prove the scal CFTs.

thic entanglement negativity for bipartite mixed states at a finite temperature in baths described by dual to configurations involving two communicating black holes in braneworld geometries. In this context ate entanglement structure characterized by the information transfer between the black holes for two st model involves communicating black holes in a Karch-Randall braneworld and \$BCFT_2\$s with two ommon bath systems for the radiation flux. The second model corresponds to a configuration of two lack holes in a braneworld geometry involving two Planck branes coupled through shared bath systems For both the models our results reproduce analogue of the Page curves for the entanglement negativity bottext of random matrix theory and from geometric evaporation in JT black hole configurations.

tion of the maximally entangled state ensures that the constructed state is close to the maximally bes not guarantee that the state is exactly the same as the maximally entangled state. Further, the is not uniquely determined in general probabilistic theories even if we impose reasonable postulate about , the existence of the maximally entangled state depends on whether the standard entanglement structure issue, we introduce pseudo standard entanglement structure as a structure of quantum composite system ns based on the existence of projective measurements and the existence of approximations of all standard e exist infinitely many pseudo standard entanglement structures different from the standard entanglement any maximally entangled state can be arbitrarily approximated by an entangled state that belongs to our rd entanglement structure. That is, experimental verification does not exclude the possibility of our rd entanglement structure that is different from the standard entanglement structure.

mhole-like solution after averaging over the fermionic couplings in the nearly AdS2 space. Even when the ontribution of these wormholes continues to exist and new saddle points appear which are interpreted as spaper, we will study the fate of these wormholes in a model without quenched disorder namely a tensor uge symmetry whose correlation function and thermodynamics in the large N limit are the same as that of estate the factorization problem linked with the wormhole threaded Wilson operator, in terms of global pordism classes associated with disconnected wormholes. Therefore for the partition function to factorize ices, there must exist certain topological defects which break the global symmetry associated with e theory devoid of global symmetries. We will interpret these wormholes with added topological defects as /e will also comment on the late time behavior of the spectral form factor, particularly its leading and subons coming from higher genus wormholes in the gravitational sector. Finally we will show how, the other "half-wormhole" dominate and give rise to unusual thermodynamics in the bulk sector due to non-

ing in AdS space in various configurations corresponding to light and heavy quarks in CFT and show that its rable in sufficiently symmetric configurations. Despite this fact, its finite-time Lyapunov exponent its urates the chaos bound. We explain this by considering the quantum dynamics, i.e. scattering amplitudes show clear presence of quantum chaos. The existence of contraints makes the open string an unusual tronger (for most systems it is weaker) in the quantum regime.

the holographic covariant entanglement entropy in the double Wick rotated version of a rotating BTZ black -AdS solution), where the periodicity of Euclidean time and spatial direction are changed. The dual field rgy in the Lorentzian signature. The holographic entanglement entropy agrees with its CFT counterpart, onformal transformation of the correlation functions of twisted operators.

s old principles, together with modern techniques, can put strong constraints on EFT. We think of gravity as low-energy gravity EFT and demonstrate causality and unitarity can constrain the modifications rules of 2->2 amplitudes. Our bounds imply that gravitational interactions must shut off uniformly in the Gcaling with higher-spin mass. In D=5, we also provide numerical bounds on central charges in holographic

8	Matsuno	Ken	Osaka Metropolitan University	Hawking radiation of scalar particles from four- dimensional Einstein-Gauss-Bonnet black holes based on a generalized uncertainty principle	We study Hawking radiation particles. We consider phe measurable length. We de particle, Gauss-Bonnet con obtain some known Hawk show that generalized und to thermodynamic stable hole.
9	Moreno	Javier	Pontificia Universidad Católica de Valparaíso	Shape-dependence of entanglement entropy in three dimensional CFTs	The entanglement entropy For a disk region, this term general theories. We argu theories. The proof makes place an osculating circle v contribution in general CF Mutual Information mode
10	Mori	Takato	КЕК	Holographic local operator quenches in BCFTs	We present a gravity dual excitation in a three-dime EOW brane gets deformed from the gravity dual and EOW brane in an elegant v
11	Osawa	Yuki	Nagoya University	Particle Creation and Entanglement Structure in the Dispersive Model	The theory of Hawking rac consider a dispersive wave This kind of modification in Hawking radiation is modi spectrum of the radiation will discuss the behavior a
12	Yadav	Gopal	Department of Physics, Indian Institute of Technology Roorkee	Page Curves of Reissner-Nordstrøm Black Hole in HD Gravity	My talk is based on arXiv:2 presence of higher derivat Maxwell term, in four dim is increases linearly with t entropy of the Hawking ra of the black hole and we of Gauss-Bonnet coupling ind curve of the Reissner-Nord
13	Yamashika	Shion	Department of physics, Chuo university	Entanglement dynamics of bosons in a 1D optical lattice	Motivated by the experim bosons in an optical lattic quenched into the Mott-ir Wigner transformation, w compare it with the quasi

ation from four-dimensional charged Einstein-Gauss-Bonnet black hole by tunneling of charged scalar ohenomenological quantum gravity effects predicted by generalized uncertainty principle with minimal derive corrections of Hawking temperature to general relativity, which are related to energy of emitted coupling constant, charge of black hole and existence of minimal length in the black hole geometry. We wking temperatures in four-dimensional black hole spacetimes by taking limits in modified temperature. We uncertainty principle may slow down the increase of Hawking temperature due to radiation, which may lead le remnant of the order of Planck mass after evaporation of four-dimensional Einstein-Gauss-Bonnet black

opy corresponding to a smooth region in general three-dimensional CFTs contains a constant universal term. arm coincides with the free energy on a three sphere and provides a renormalization group monotone for gue this finite contribution is globally minimized by disks with respect to arbitrary regions and for general area use of the strong subadditivity of entanglement entropy and the geometric fact that one can always e within a given smooth entangling region. In addition, we provide accurate approximations of the finite CFTs in the case of elliptic regions as well as numerically for more general shapes in the so-called "Extensive del", verifying the general bound.

al of local operator quench in a two-dimensional CFT with conformal boundaries. This is given by a massive nensional AdS space with the end of the world brane (EOW brane). Due to the gravitational backreaction, the ned in a nontrivial way. We show that the energy-momentum tensor and entanglement entropy computed nd from the BCFT in the large c limit match perfectly. Interestingly, this comparison avoids the folding of the nt way. This talk is based on JHEP05(2022)060 [arXiv:2203.03851].

radiation has a problem called Trans-Planckian problem. One of the approaches to this problem is to ave with a cutoff of the wave number by adding the higher order derivative term into the wave equation. In induces additional modes called Planckian modes and many researches showed that the behavior of odified but the thermality of the radiation still remains in the low energy regimes. We analyzed the power on and the partner structure of modes for the steplike geometry with and without the horizon. In this talk, I r and the thermality of the radiation from the viewpoint of the partner structure of the modes.

v:2204.11882. I will explain the Page curves computation of an eternal Reissner-Nordstrøm black hole in the vative terms which are $(cal O)(R^2)$ terms plus Maxwell term and Einstein-Gauss-Bonnet gravity plus mensions. In both the cases entanglement entropy of the Hawking radiation in the absence of island surface time. After including contribution from the island surface, we found that after the Page time entanglement radiation in both the cases reaches a constant value which is the twice of the Bekenstein-Hawking entropy e obtained the Page curves. Further we found that Page curves will appear at later or earlier time when the increases or decreases. As a consistency check, in the limit of vanishing GB coupling we obtain the Page ordstr?m black hole obtained in arXiv:2101.06867.

imental developments of ultracold atoms, we theoretically investigate the entanglement dynamics of tice. Specifically, we calculate the time-evolution of the 2nd-order Rényi entropy (RE) when the system is t-insulating state from a product state. Developing the effective theory based in the generalized Jordanwe derive the analytic expression for the time-evolution of the RE. We reveal the entire dynamics of RE and asi-particle picture for entanglement dynamics proposed by Calabrese and Cardy.