

Equilibration of Scalar Fields in an Expanding System

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Aug 22nd, 2012.

arXiv:1206.4743

Relativistic Heavy Ion Collision at RHIC and LHC

$$\sqrt{s_{NN}} = 0.2 \text{ TeV}$$

Au+Au (RHIC)

$$\sqrt{s_{NN}} = 2.76 \text{ TeV}$$

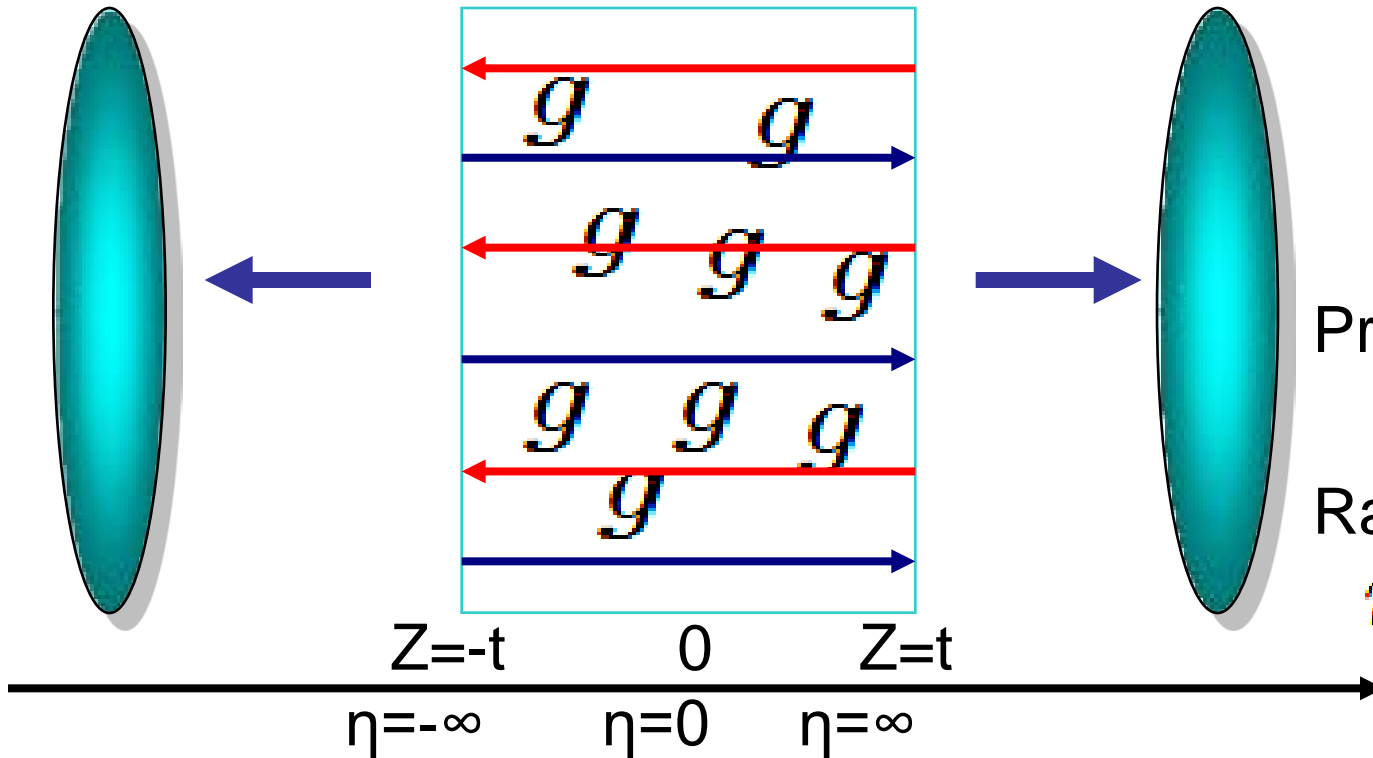
Pb+Pb (LHC)

Proper time

$$\tau = \sqrt{t^2 - z^2}$$

Rapidity

$$\eta = \tanh^{-1} \frac{z}{t}$$



Time 0 fm/c

1fm/c

10fm/c

15fm/c

Black box

Before collision

Glasma

Hydrodynamics
QGP

hydro+
hadron gas

CGC

Nonequilibrium dynamics of gluons

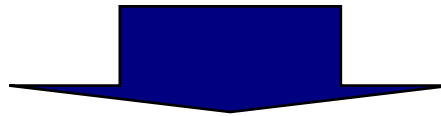
Formation of **Quark-Gluon Plasma** (QGP)

Success of nearly ideal hydrodynamics after thermalization.
Early Thermalization of gluons (0.6-1fm/c)! (RHIC and LHC)
Kolb and Heinz (2002), Hirano et al. (2010)

Comparative to formation time of partons ($1/Q_s \sim 0.2\text{fm}/c$)
Semi-Classical Boltzmann eq. should not be applied,
since 2-3fm/c is predicted for **$gg \rightarrow gg$, $gg \rightarrow ggg$** (**Boltzmann**).

Decoherence: Muller, Schafer (2006)

Baier, Mueller, Schiff, Son (2001 and 2011)



New method is needed.

Quantum nonequilibrium processes based on field theory

Application of **Kadanoff-Baym eq.**
to early thermalization of gluons.

Purpose of this talk

Introduction of time evolution equation for classical field and Kadanoff-Baym equation for quantum fluctuation in $O(N)$ scalar model in an expanding metric.

To show particle production and equilibration in Numerical Analyses.

Comparison of expanding and nonexpanding system.

Rest of this

- **Time Evolution Equation I, II**
- **Initial condition**
- **Comparison of expanding and nonexpanding system**
- **Summary and Remaining Problems**

Time Evolution Equation I

Action of scalar O(N) model

$$S = \int d^4x \sqrt{-g} \left[\frac{g^{\mu\nu}}{2} \partial_\mu \varphi_a \partial_\nu \varphi_a - \frac{m^2}{2} \varphi_a \varphi_a - \frac{\lambda (\varphi_a \varphi_a)^2}{4!N} \right]$$

$$a=1, \dots, N$$

$$g = \text{diag}(1, -\tau^2, -1, \dots)$$

Interaction term

Equation of motion of classical field

$$\phi_a(t, x) = \langle \hat{\varphi}_a(t, x) \rangle$$

$$\left[\partial_\tau^2 + \frac{1}{\tau} \partial_\tau + m^2 + \frac{\lambda}{6N} \phi^2(\tau) \right] \phi_a(\tau) = 0$$

Or effect of fluctuations

Damping of classical field for an expanding system

Time Evolution Equation II

- Kadanoff-Baym equation: Quantum evolution equation of 2-point Green's function (fluctuations) **statistical (distribution)** and **spectral** functions

$$F_{ab}(x, y) = \frac{1}{2} \langle \{ \tilde{\phi}_a(x), \tilde{\phi}_b(y) \} \rangle$$

$$\rho_{ab}(x, y) = \langle [\tilde{\phi}_a(x), \tilde{\phi}_b(y)] \rangle$$

$$F(\tau, \tau', p) \approx \frac{1}{2m_{\perp} \sqrt{\tau \tau'}} \cos m_{\perp}(\tau - \tau') (2n_p + 1)$$

$\tau, \tau' \rightarrow \infty$ **Boson**

$$\rho(\tau, \tau', p) \approx \frac{1}{m_{\perp} \sqrt{\tau \tau'}} \sin m_{\perp}(\tau - \tau')$$

$m_{\perp} = \sqrt{m^2 + p_{\perp}^2}$

$$(-G_0^{-1} + \Sigma_{\text{loc}}) F(x, y) = \int_0^{y^0} dz \Sigma_F(x, z) \rho(z, y) - \int_0^{x^0} dz \Sigma_{\rho}(x, z) F(z, y)$$

$$(-G_0^{-1} + \Sigma_{\text{loc}}) \rho(x, y) = \int_{x^0}^{y^0} dz \Sigma_{\rho}(x, z) \rho(z, y) \quad \text{Memory integral}$$

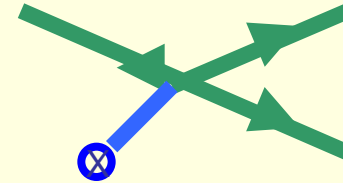
$$G_0^{-1} \equiv -\frac{\partial^2}{\partial \tau^2} - \frac{1}{\tau} \frac{\partial}{\partial \tau} + \frac{1}{\tau^2} \frac{\partial^2}{\partial \eta^2} + \nabla_{\perp}^2 + m^2 \quad \Sigma = \text{Self-energies}$$

Self-energies: local Σ_{loc} mass shift, nonlocal real Σ_F and imaginary part Σ_{ρ}

Merit

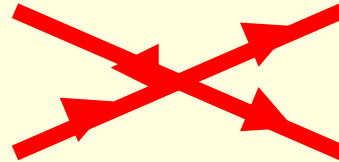
- **Field-Particle Conversion: Particle production from classical field.**

(Parametric resonance) +



- **Collision of particles → Bose-Einstein distribution**

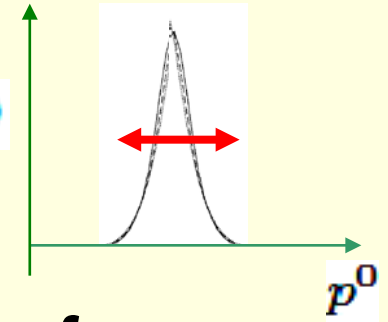
binary



Finite decay width

- **Off-shell effect: Memory effects and finite spectral width**

$\rho(p^0, p)$



binary collisions (2-to-2) → Rapid Change of distribution functions (Lindner and Muller 2006) +
3-to-1 → entropy production + **chemical equilibrium.**

Demerit

Numerical simulation needs much memory of computers.

Initial condition

Initial condition: Classical field with **vacuum**
quantum fluctuations (Color Glass Condensate ?)

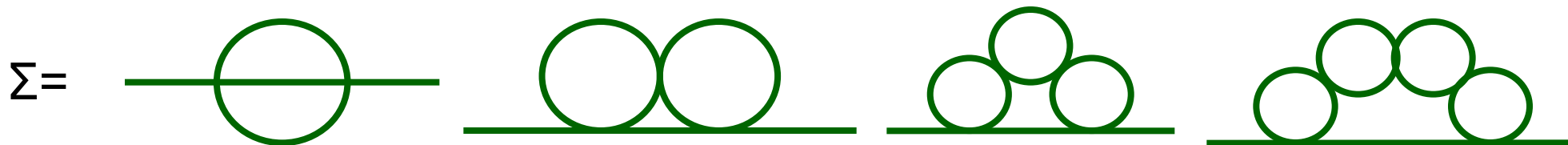
$$\phi_a(\tau) = \phi(\tau)\delta_{a1} \quad \phi(\tau_0) = \sqrt{\frac{6N}{\lambda}} \sigma_0 \quad \begin{array}{l} \lambda = 10 \\ N = 4 \end{array} \quad m/\sigma_0 = 0.1$$

$$F_{ab} = \text{diag}(F_{\parallel}, F_{\perp}, \dots, F_{\perp}) \quad \text{vacuum}$$

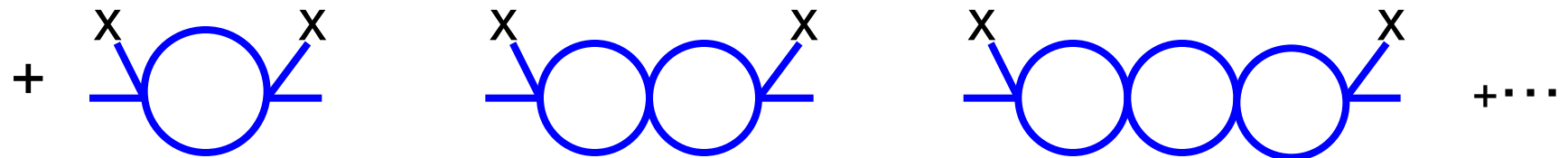
We assume homogeneity in space.

Collision term

Our results (collaboration with Y. Hatta): Quantum collision term.



Normal collision term.



Source induced amplification.

Summation of Next-to-Leading Order of $1/N$ expansion. This approach covers all evolution of F from $O(1)$ to $O(\lambda^{-1})$

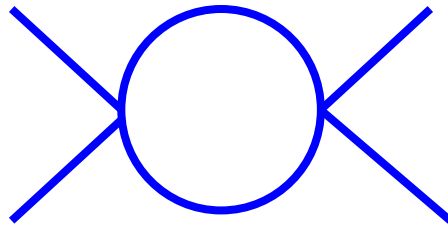
Classical Statistical Approximation

For example

Berges et al. (2012)

Gelis et al. (2012)

$\Pi_F =$



~~$= FF - (1/4)\rho\rho$~~

$N_{eq} \sim T/p - 1/2$

$F \sim (1/\tau)n_p, \rho \sim 1/\tau$

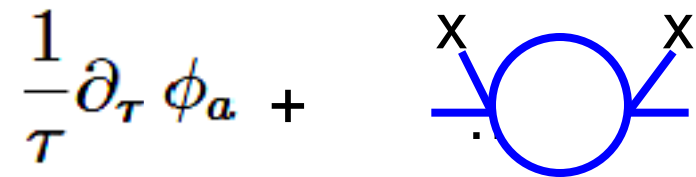
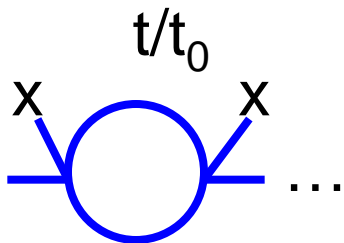
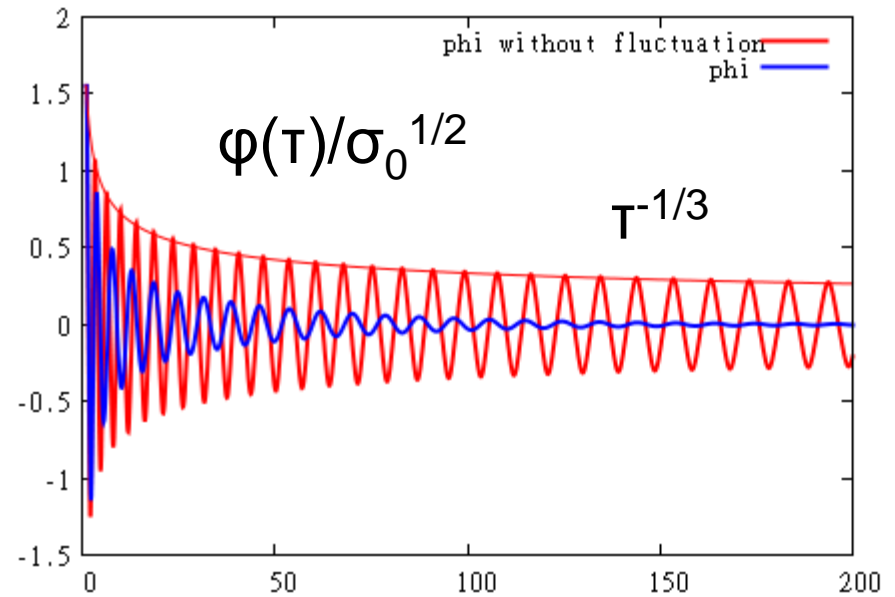
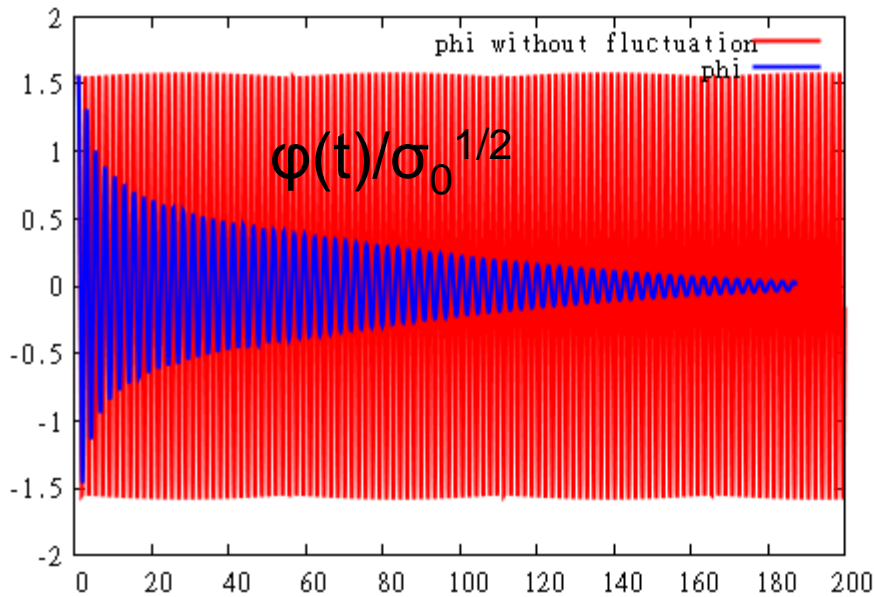
$n_p \sim 1/\lambda, \text{at late time}$

This approximation is good in the case of dense system $FF \gg \rho\rho$ (Weakly coupled). But if $FF \sim \rho\rho$ (Normal coupling), the difference appears.

Comparison of Nonexpanding and Expanding systems (2+1 dim)

Evolution of classical field

Nonexpanding Expanding



The above term + source induced amplification

Evolution of Green's functions F

Quantum evolution

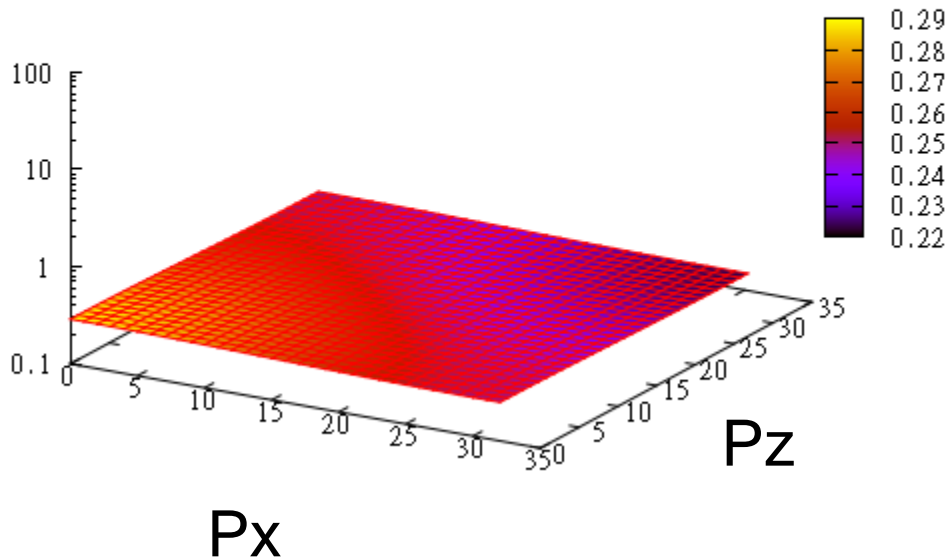
Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

t/t0=1 ———



Evolution of Green's functions F

Quantum evolution

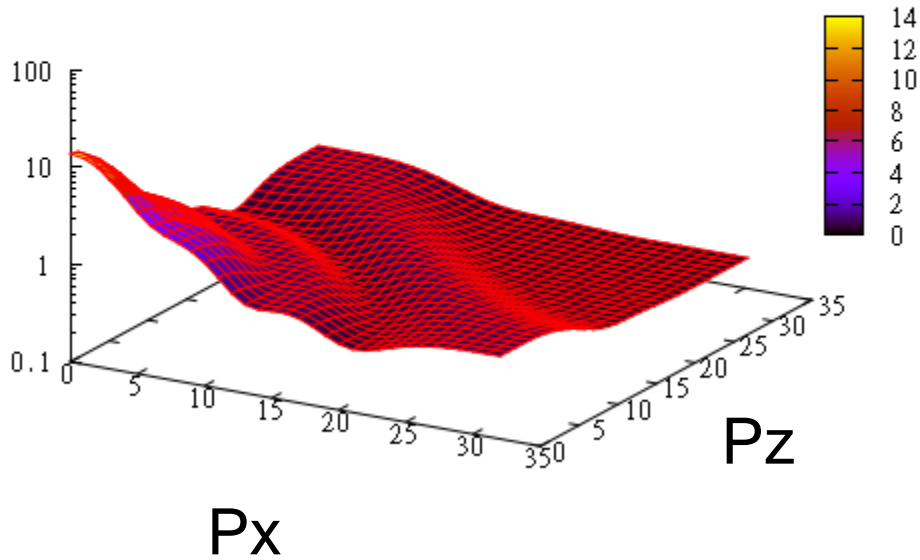
Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

t/t0=5



Evolution of Green's functions F

Quantum evolution

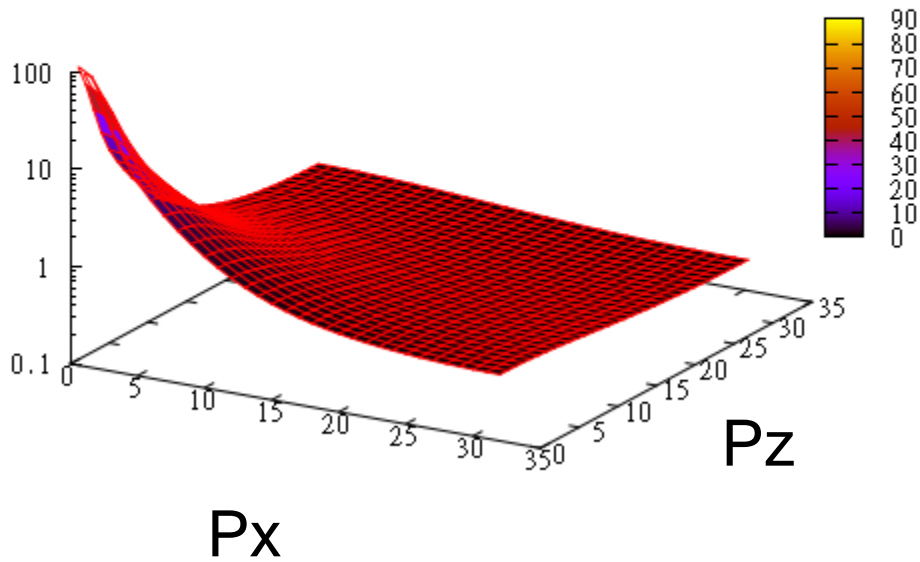
Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

t/t0=20



Evolution of Green's functions F

Quantum evolution

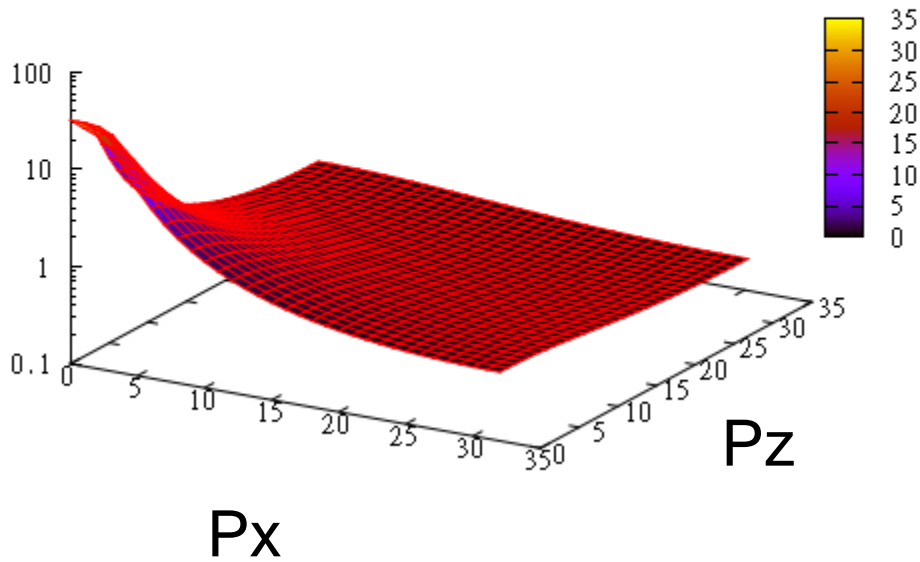
Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

t/t0=40



Evolution of Green's functions F

Quantum evolution

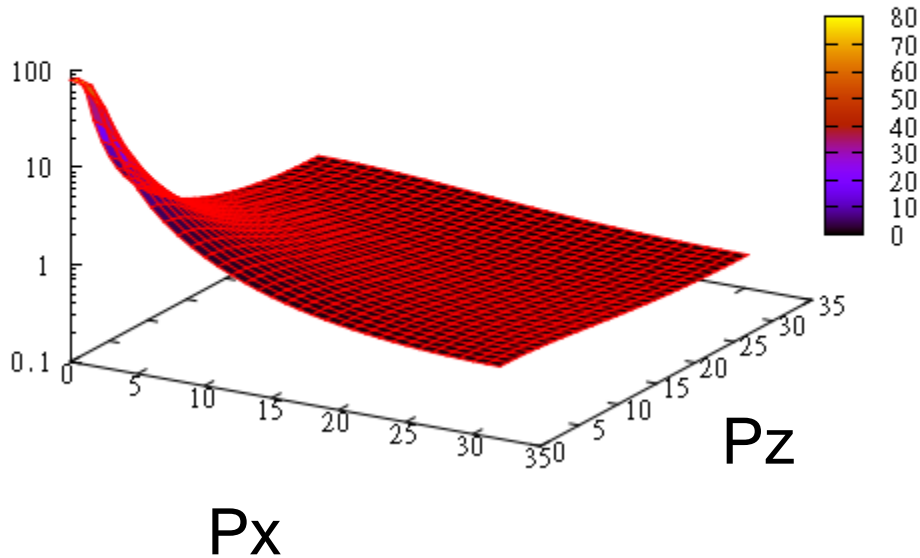
Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

t/t0=60



Evolution of Green's functions F

Quantum evolution

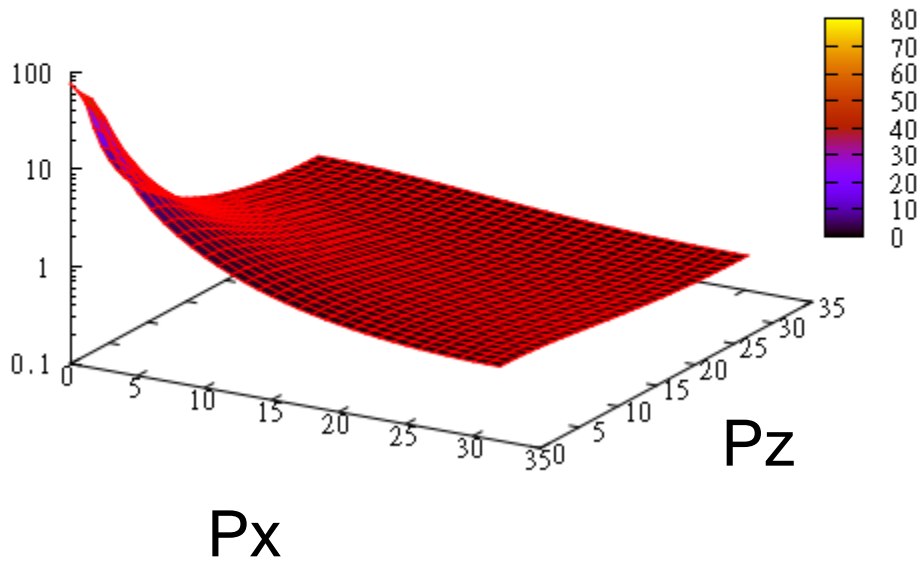
Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

t/t0=80



Evolution of Green's functions F

Quantum evolution

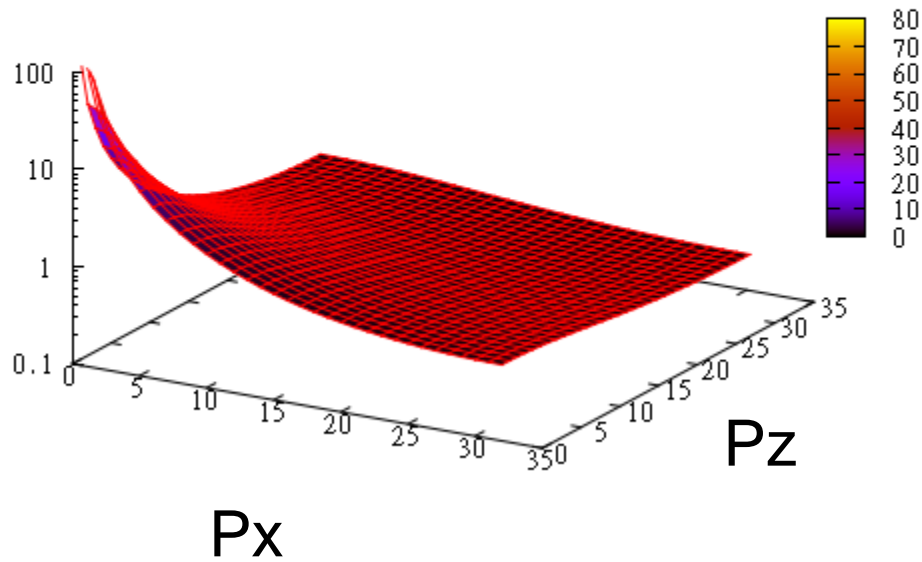
Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

t/t0=100



Evolution of Green's functions F

Quantum evolution

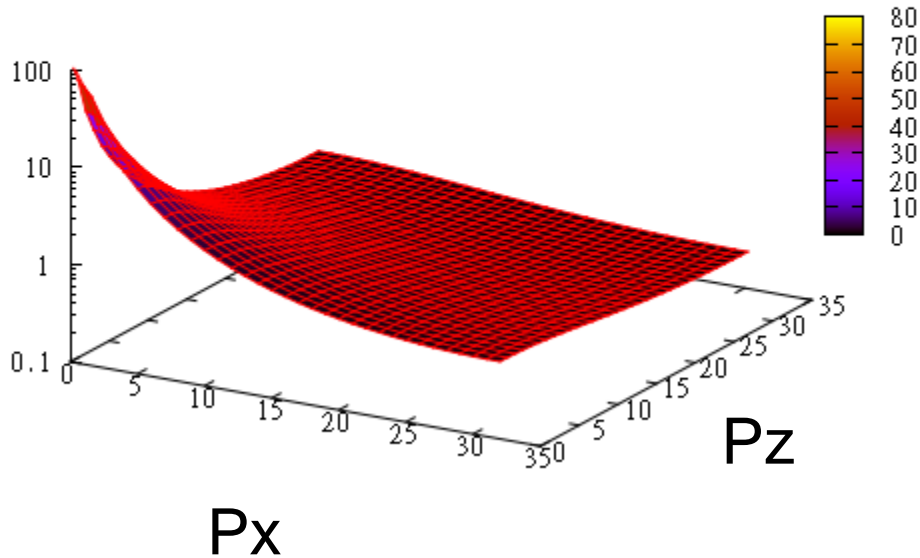
Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

$t/t_0=120$ ———



Evolution of Green's functions F

Quantum evolution

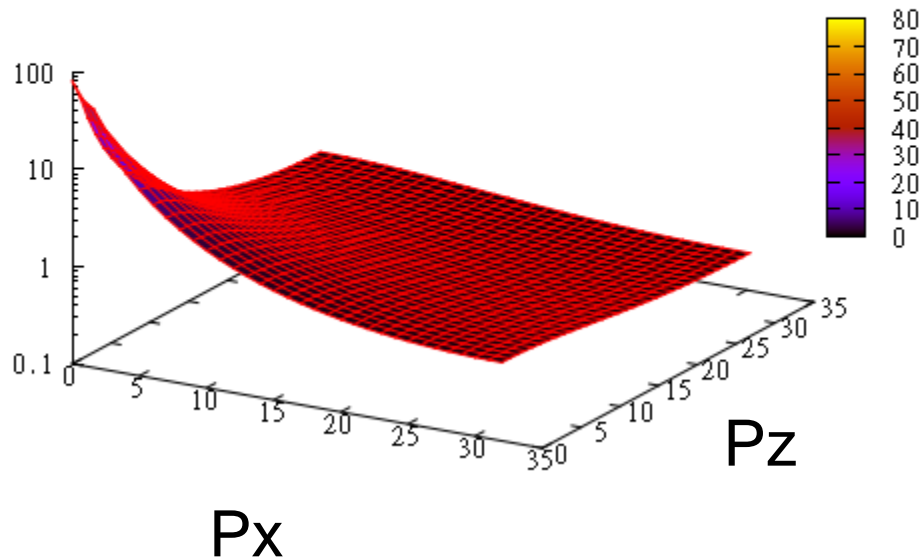
Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

$t/t_0=140$ ———



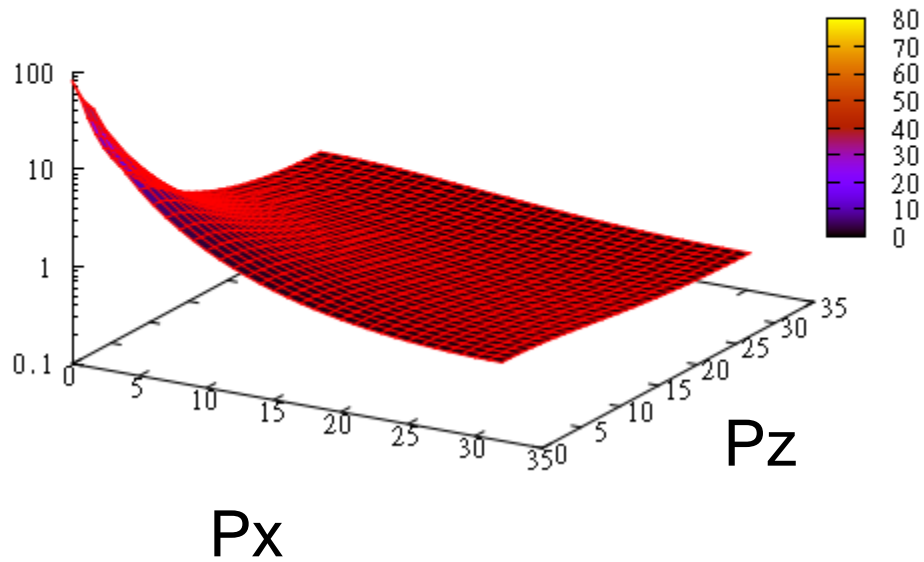
Evolution of Green's functions F

Quantum evolution

Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

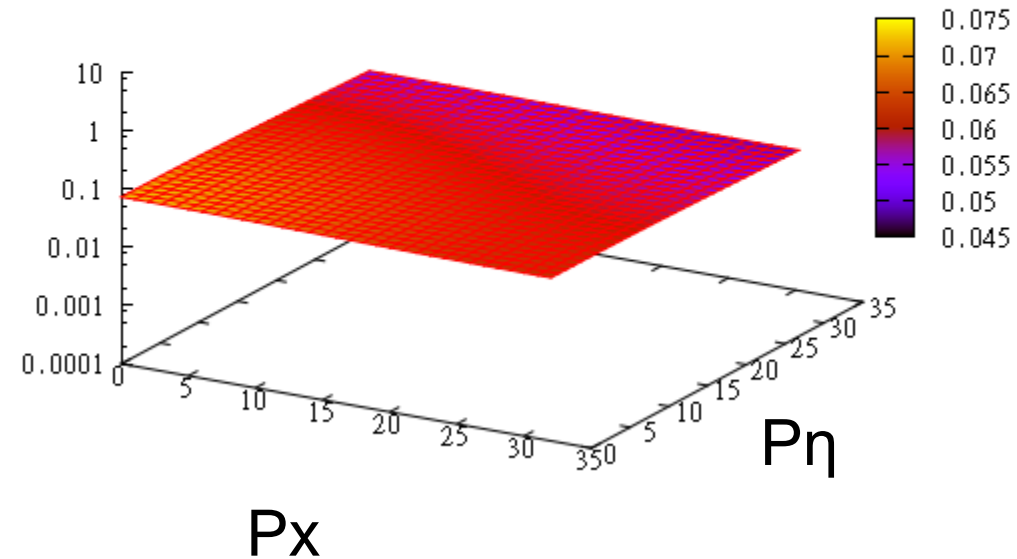
$t/t_0=140$



Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

$\tau/\tau_0=1$



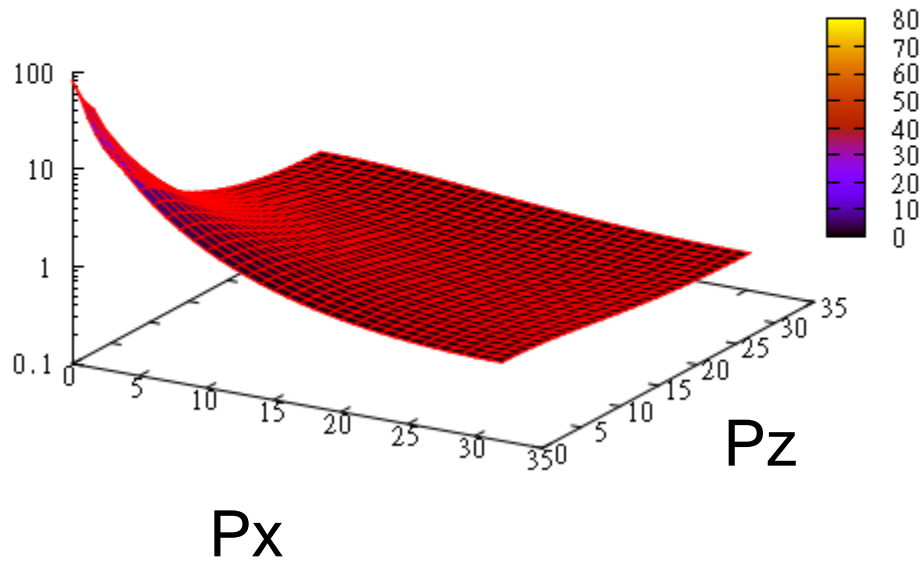
Evolution of Green's functions F

Quantum evolution

Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

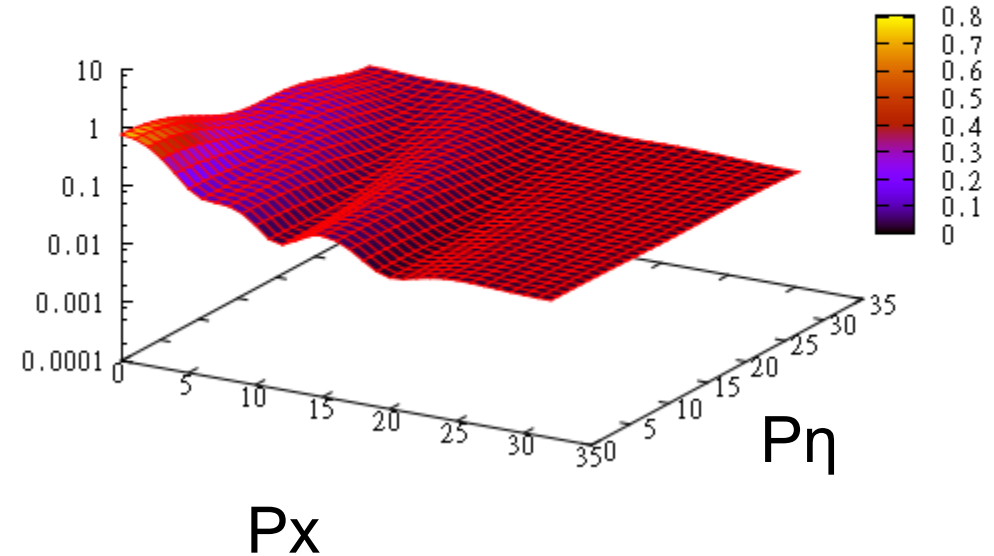
$t/t_0=140$



Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

$\tau/\tau_0=5$



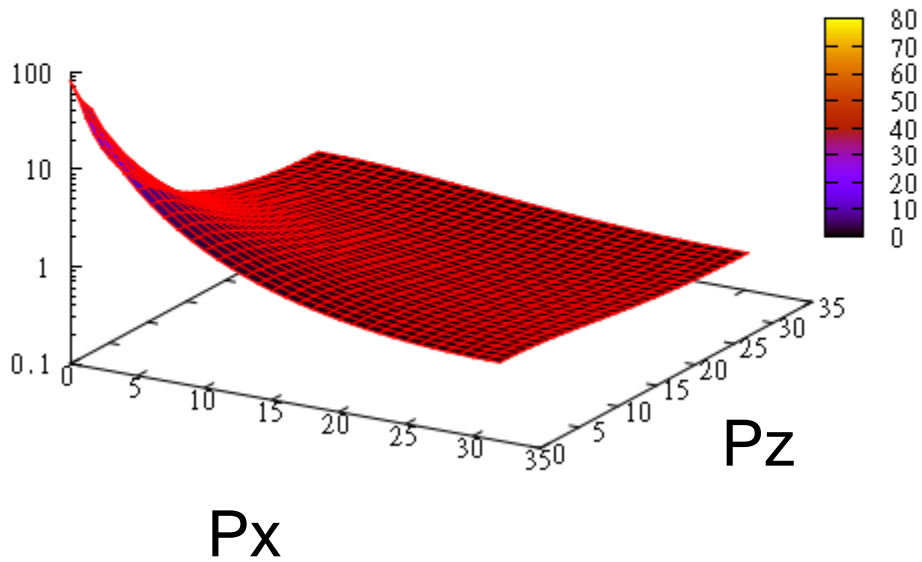
Evolution of Green's functions F

Quantum evolution

Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

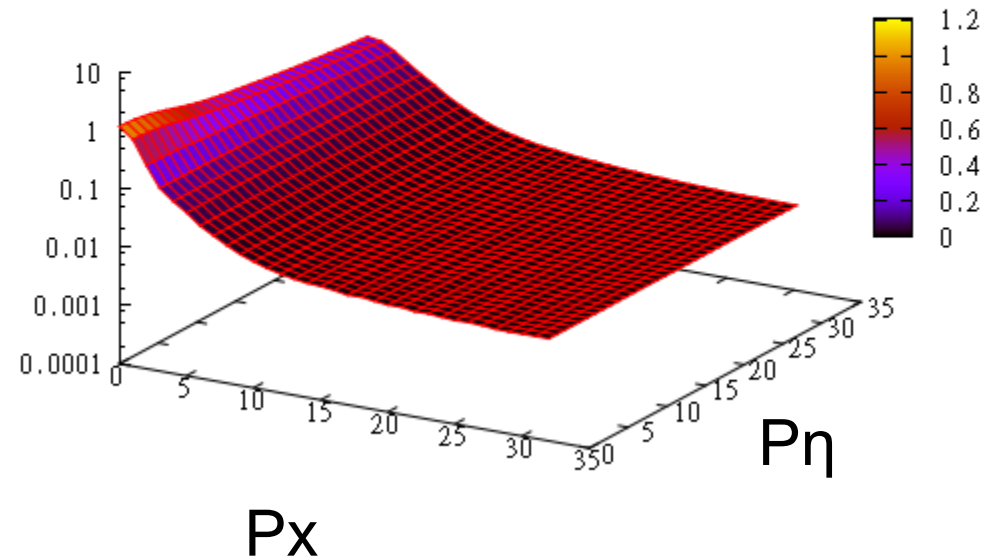
$t/t_0=140$ —



Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

$\tau/\tau_0=20$ —



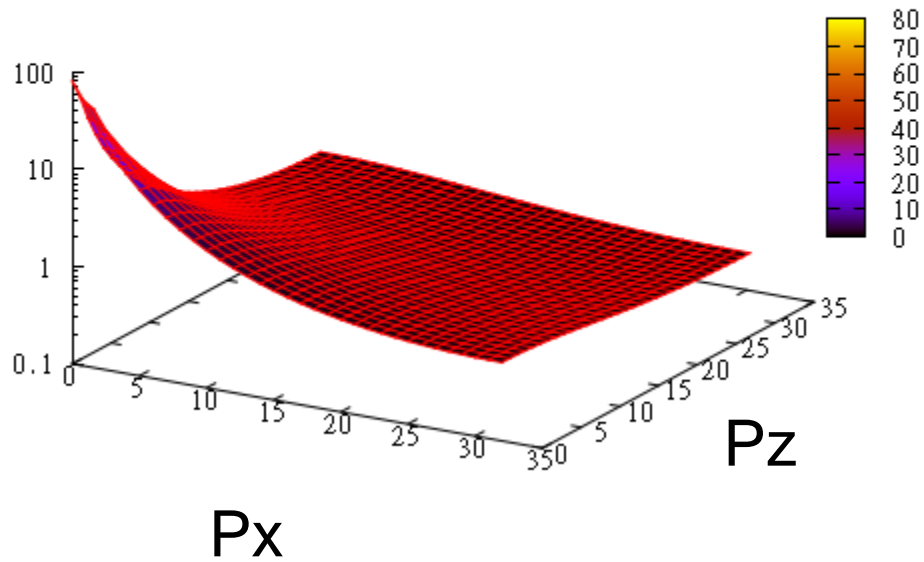
Evolution of Green's functions F

Quantum evolution

Nonexpanding

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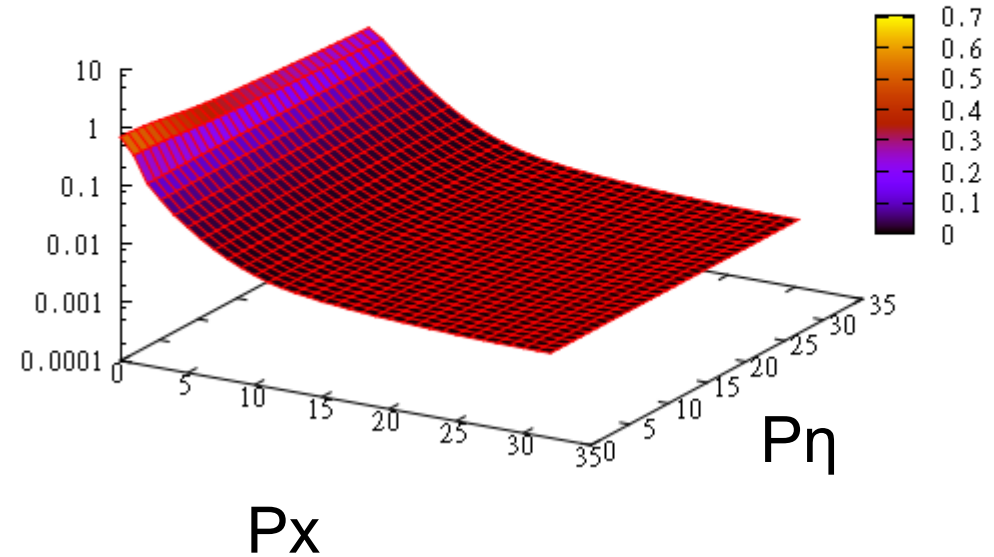
$t/t_0=140$



Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

$\tau/\tau_0=40$



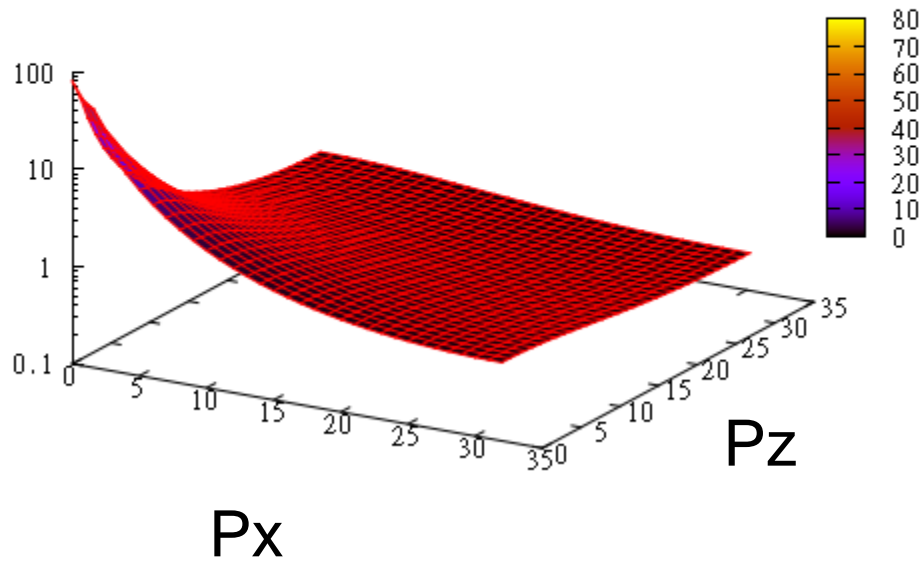
Evolution of Green's functions F

Quantum evolution

Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

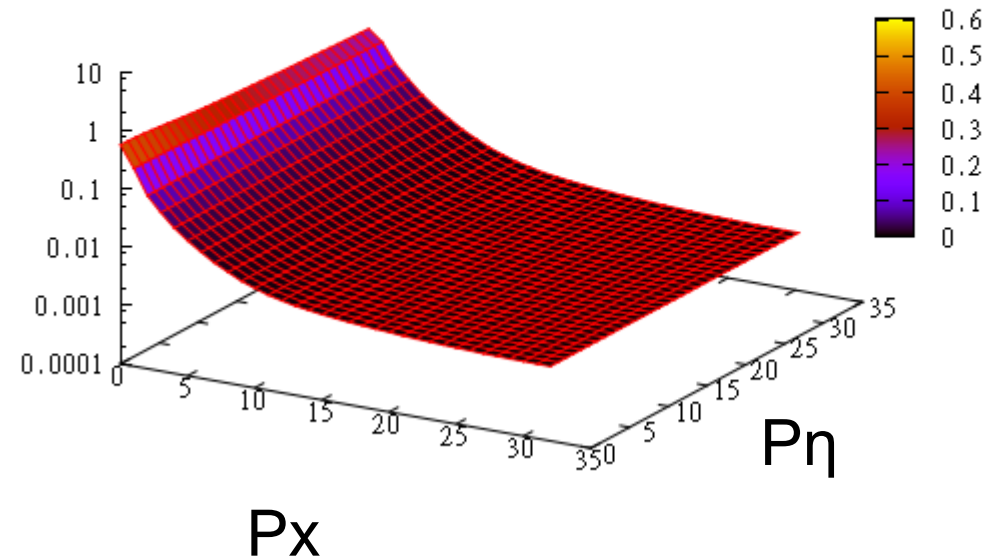
$t/t_0=140$



Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

$\tau/\tau_0=60$



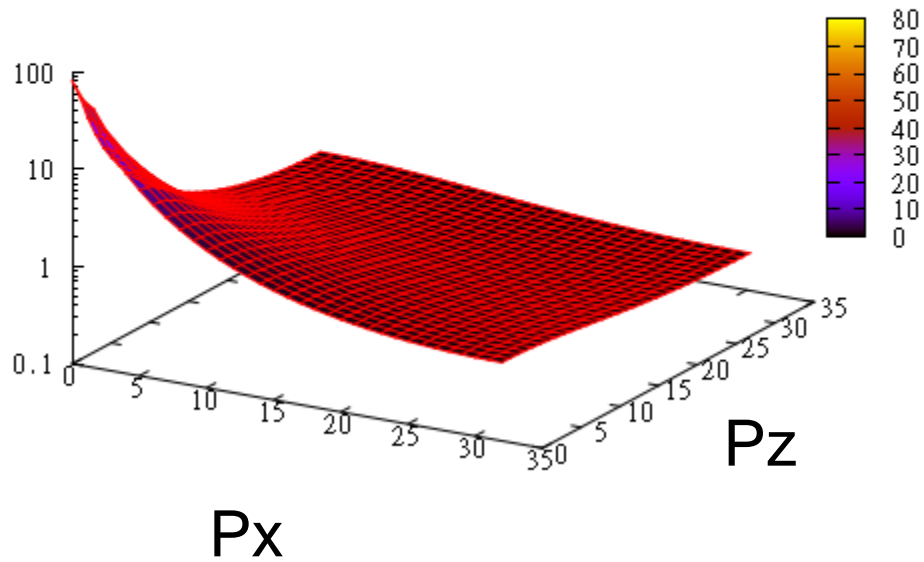
Evolution of Green's functions F

Quantum evolution

Nonexpanding

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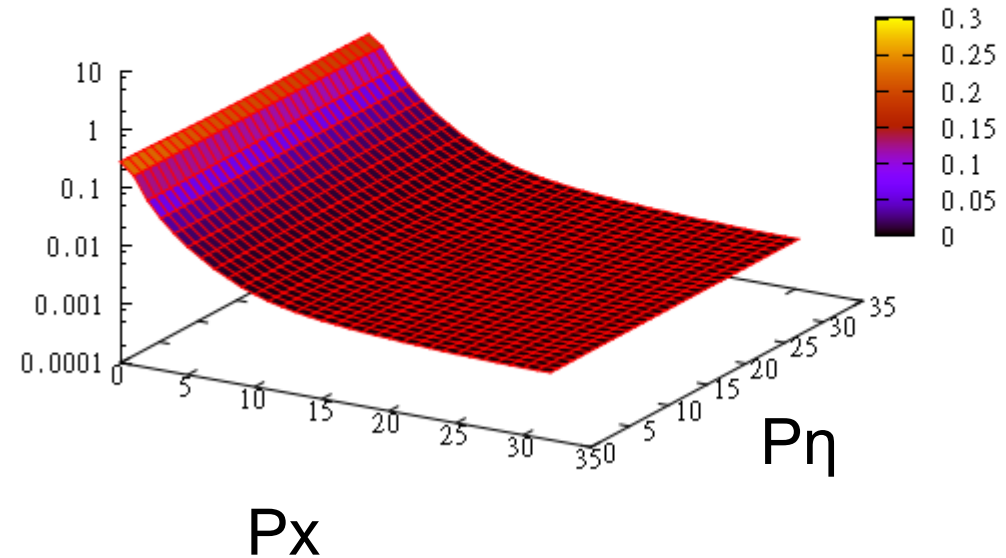
$t/t_0=140$



Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

$\tau/\tau_0=80$



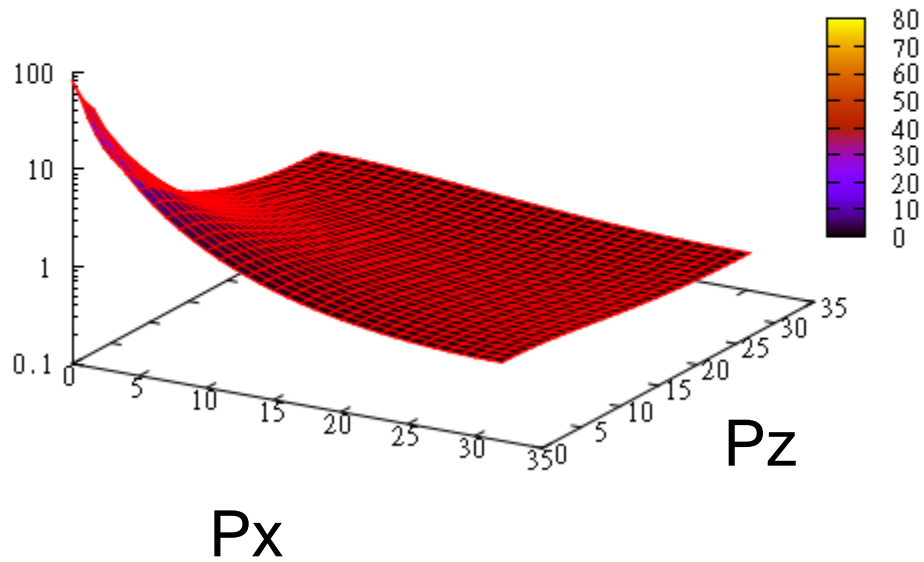
Evolution of Green's functions F

Quantum evolution

Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

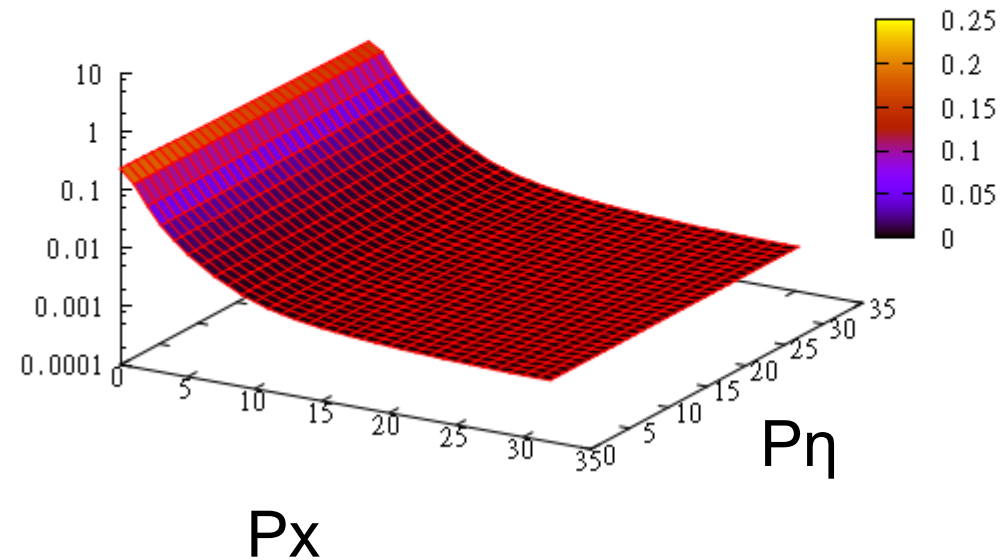
$t/t_0=140$



Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

$\tau/\tau_0=100$



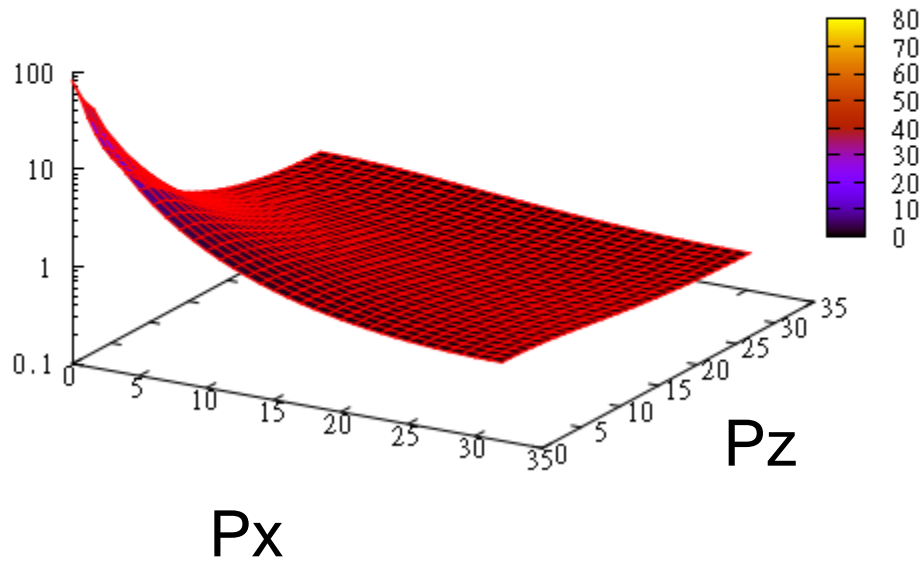
Evolution of Green's functions F

Quantum evolution

Nonexpanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

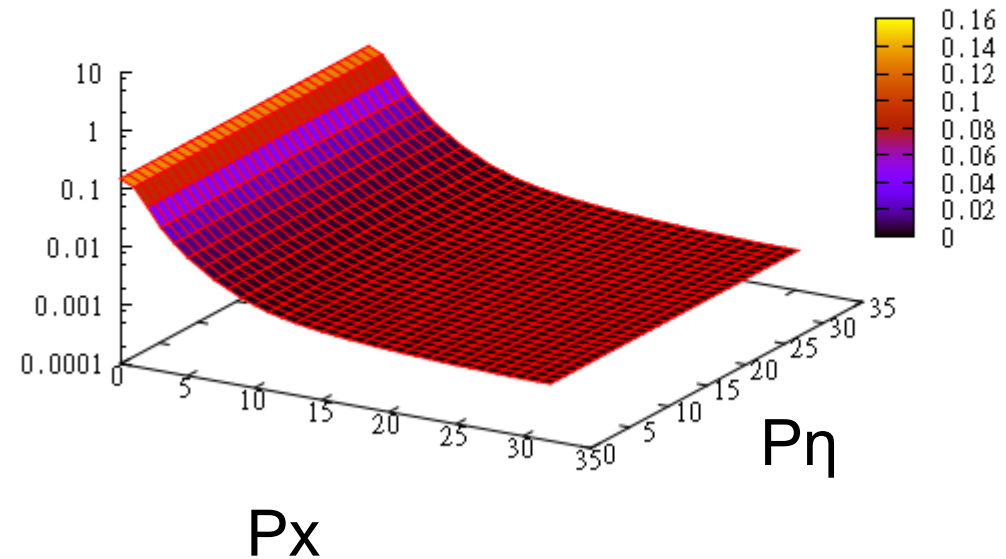
t/t0=140



Expanding

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

tau/tau0=120



Evolution of Green's functions F

Quantum evolution

Nonexpanding

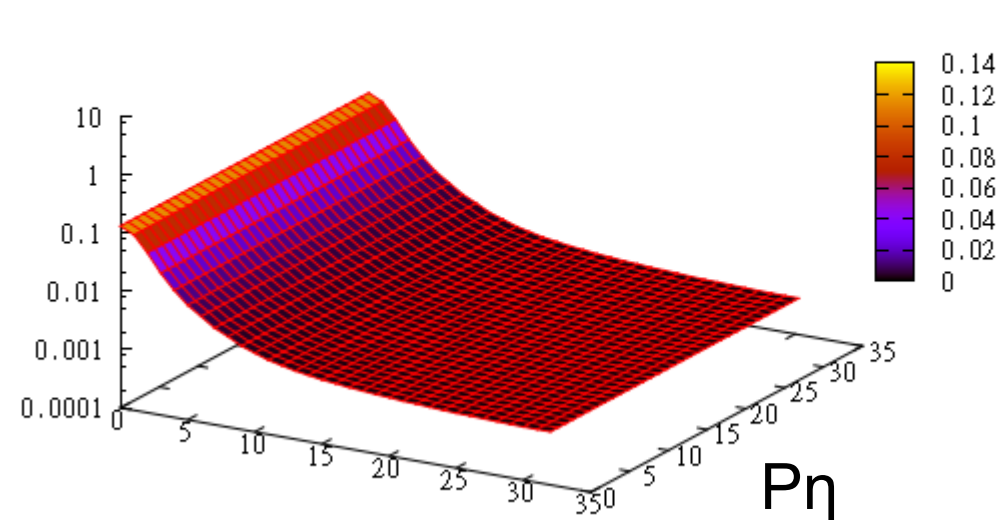
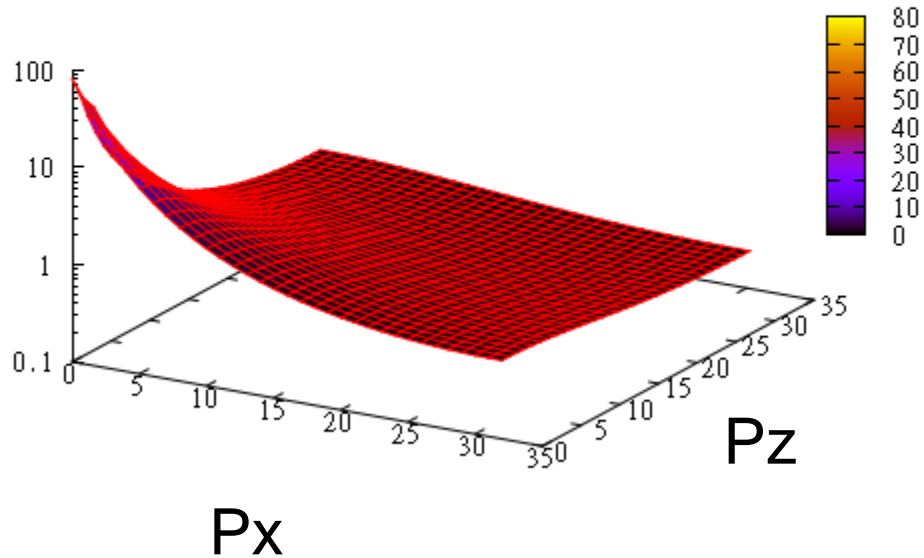
Expanding

$$\sigma_0 F_{\parallel}(t, t, p_T, p_z)$$

$$F_{\parallel}(\tau, \tau, p_T, p_{\eta})$$

t/t0=140

tau/tau0=140



$$G_0^{-1} \equiv -\frac{\partial^2}{\partial \tau^2} - \frac{1}{\tau} \frac{\partial}{\partial \tau} + \frac{1}{\tau^2} \frac{\partial^2}{\partial \eta^2} + \nabla_{\perp}^2 + m^2$$

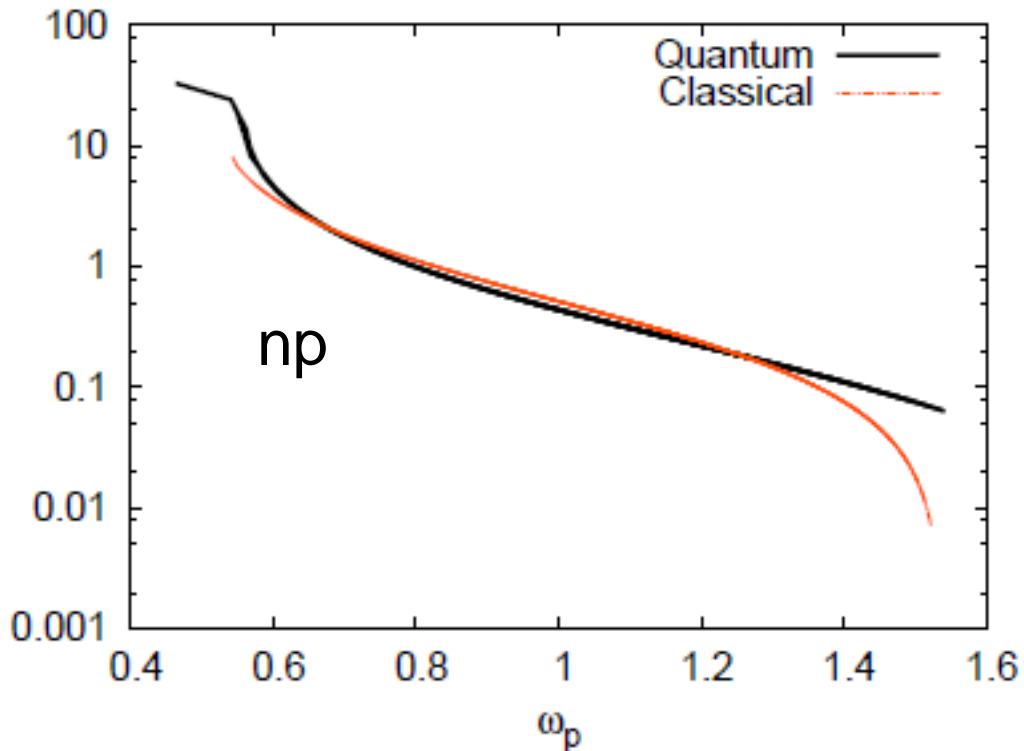
Pz = Pη/τ

\downarrow
 p_{η}^2/τ^2

Late-time distribution function n_p

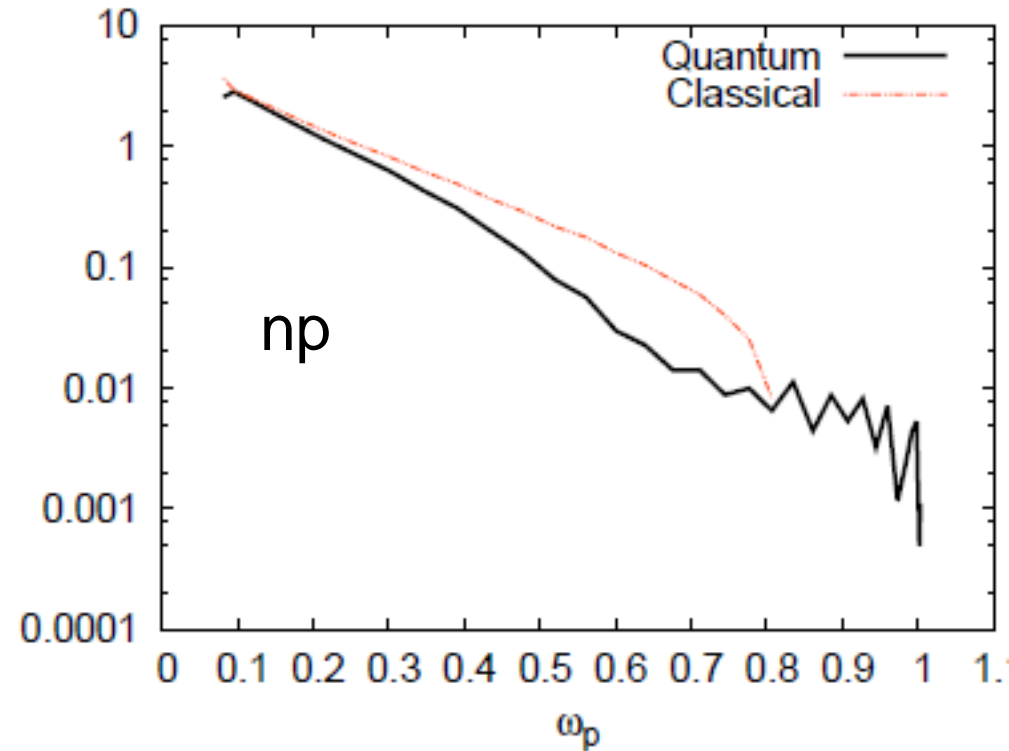
Nonexpanding

$t/t_0=150$



Expanding

$\tau/\tau_0=150$



Classical statistical approximation: $N_{eq} \sim T/p - 1/2$

Summary

- **We have considered the Kadanoff-Baym approach to thermalization of $O(N)$ scalar fields from initial background classical field with longitudinal expansion in $2+1$ dimensions.**
- **Field-particle conversion occurs when we include effects of fluctuation. Then classical field damps rapidly due to expansion of the system compared with nonexpanding system.**
- **In both nonexpanding and expanding system, late-time Boltzmann tails are realized in Quantum evolution.**
- **In classical statistical approximation, the late-time distribution is not Bose-Einstein type (Normal coupling)**

Remaining problems

- **Application to non-Abelian gauge theories in expanding system.**
- **Initial condition in an expanding system (Color Glass Condensate with vacuum fluctuations).**
- **Renormalization procedure in an expanding system.**
- **Tuning of program codes.**

F

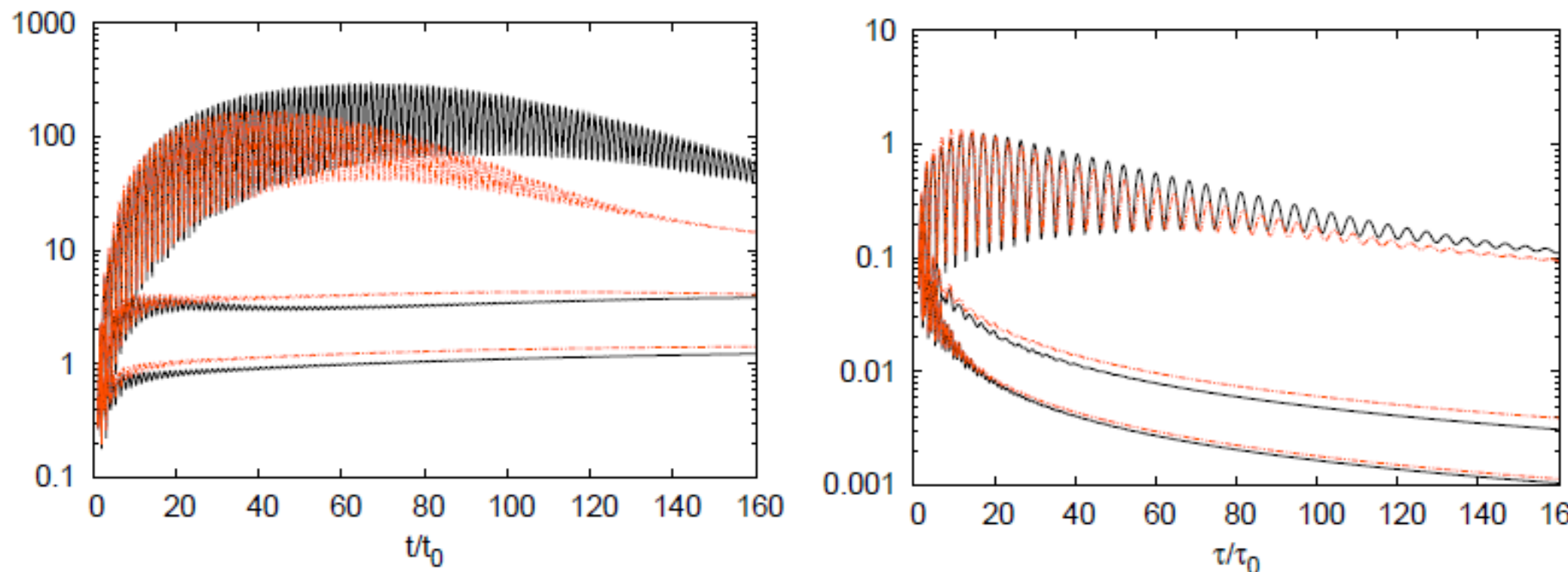
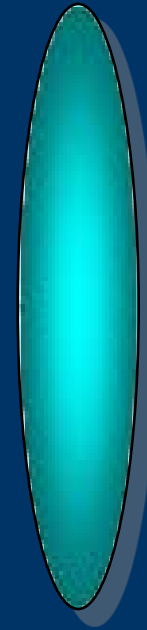


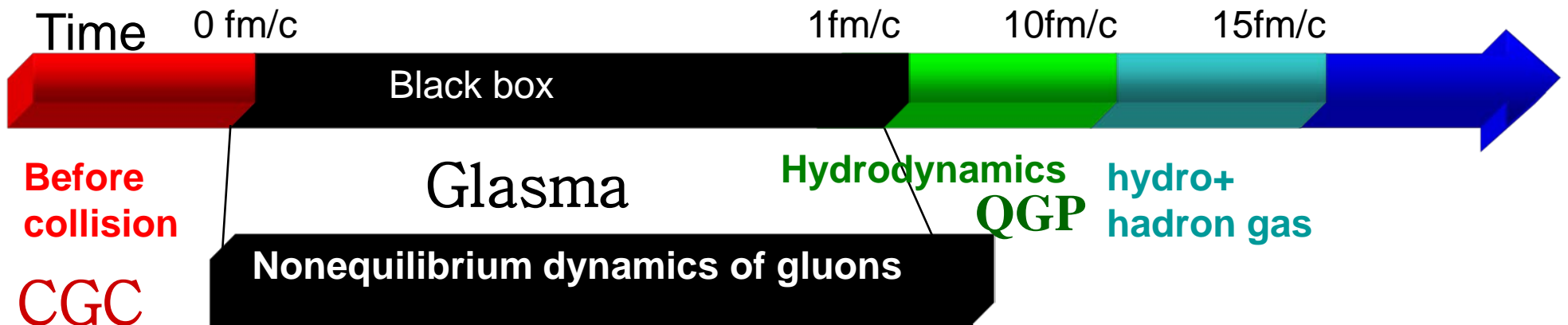
Fig. 4. Time evolution of the statistical function at strong coupling $\lambda = 10$ for three different values of n_T : $n_T = 0$, $n_T = 8$, $n_T = 16$ (from top to bottom).

Relativistic Heavy Ion Collision at RHIC and LHC

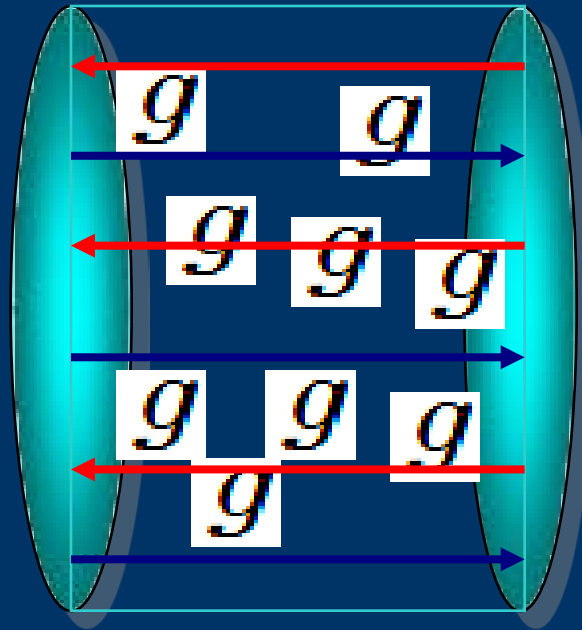


$\sqrt{s_{NN}}=0.2 \text{ TeV}$
Au+Au (RHIC)
 $\sqrt{s_{NN}}=2.76 \text{ TeV}$
Pb+Pb (LHC)

Quark-Gluon Plasma



Relativistic Heavy Ion Collision at RHIC and LHC



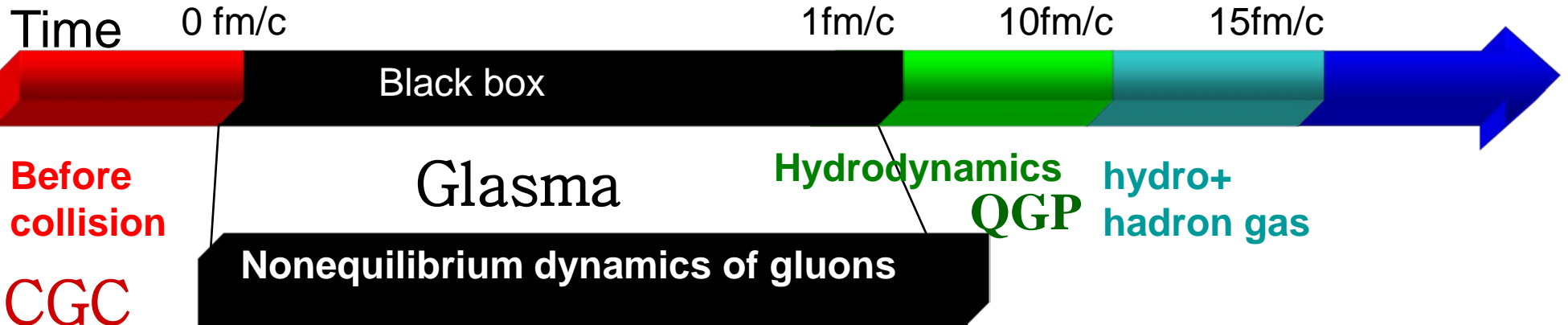
$$\sqrt{s_{NN}}=0.2 \text{ TeV}$$

Au+Au (RHIC)

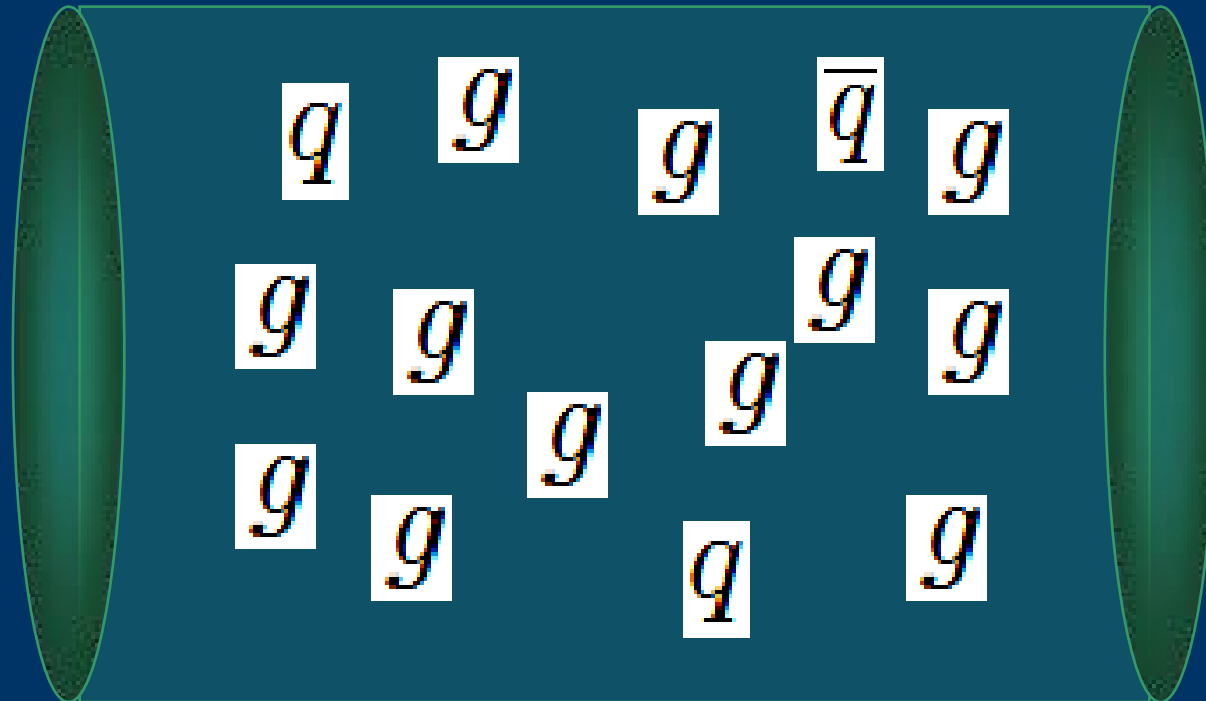
$$\sqrt{s_{NN}}=2.76 \text{ TeV}$$

Pb+Pb (LHC)

Quark-
Gluon
Plasma



Relativistic Heavy Ion Collision at RHIC and LHC



$\sqrt{s_{NN}}=0.2$ TeV
Au+Au (RHIC)
 $\sqrt{s_{NN}}=2.76$ TeV
Pb+Pb (LHC)

Quark-
Gluon
Plasma

Time 0 fm/c 1fm/c 10fm/c 15fm/c

Black box

Before
collision

Glasma

Hydrodynamics
QGP

hydro+
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CGC

Nonequilibrium dynamics of gluons

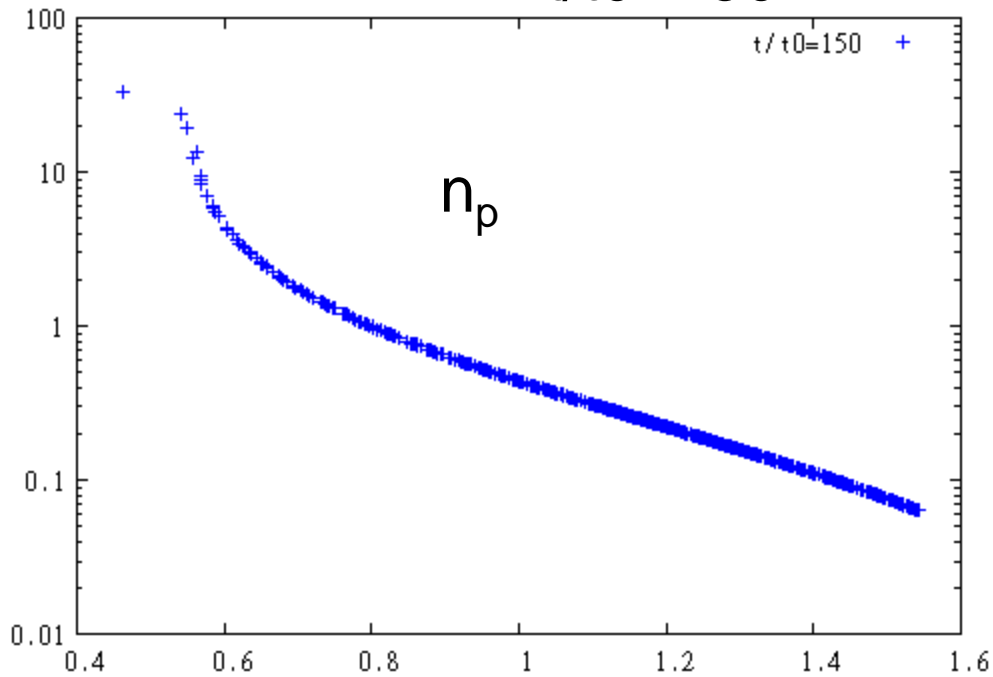
Late-time distribution function (Quantum evolution)

Nonexpanding

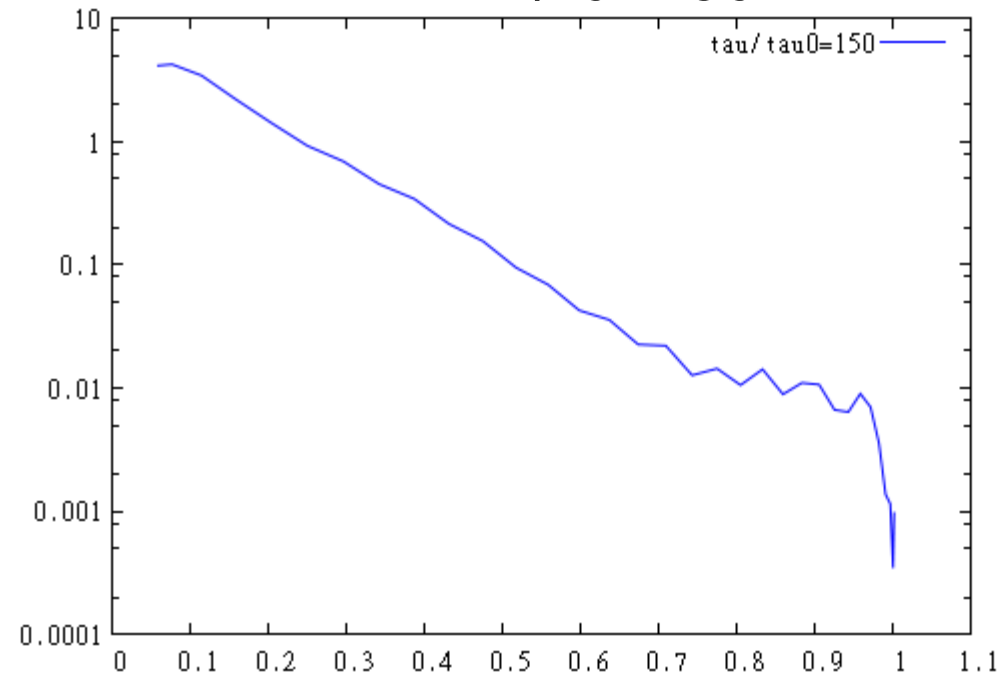
Expanding

$t/t_0=150$

$\tau/\tau_0=150$



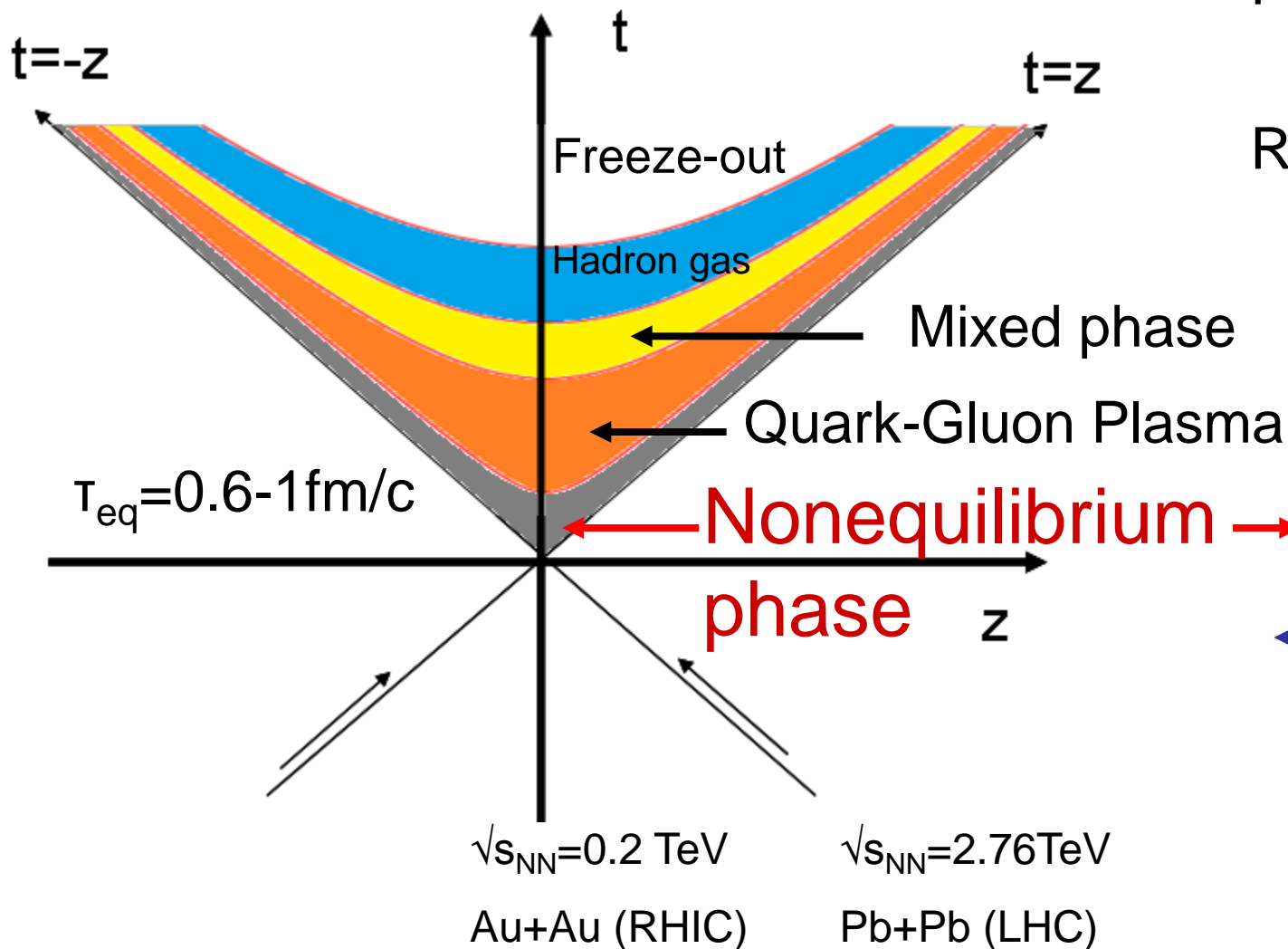
ϵ_p



ω_p

$$n_p(\tau) \equiv \tau \omega_p F(\tau, \tau, p) / C - \frac{1}{2}$$

Relativistic Heavy Ion Collision at RHIC and LHC

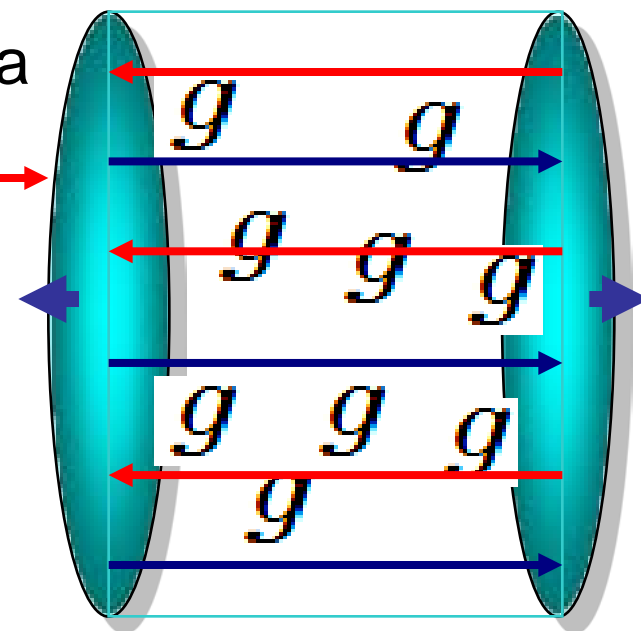


Proper time

$$\tau = \sqrt{t^2 - z^2}$$


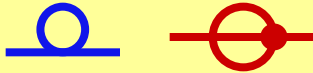

Rapidity

$$\eta = \tanh^{-1} \frac{z}{t}$$






Time irreversibility

Symmetric phase $\langle \Phi \rangle = 0$

	$\lambda\Phi^4$	$O(N)$	$SU(N)$
Exact 2PI (no truncation)	✗	✗	✗
Truncation	NLO of λ  Δ	NLO of $1/N$  Δ	LO of g^2  $\Delta(\text{TAG})$
LO of Gradient expansion H-theorem	\bigcirc	\bigcirc	$\bigcirc(\text{TAG})$

Numerical Simulation for KB eq.

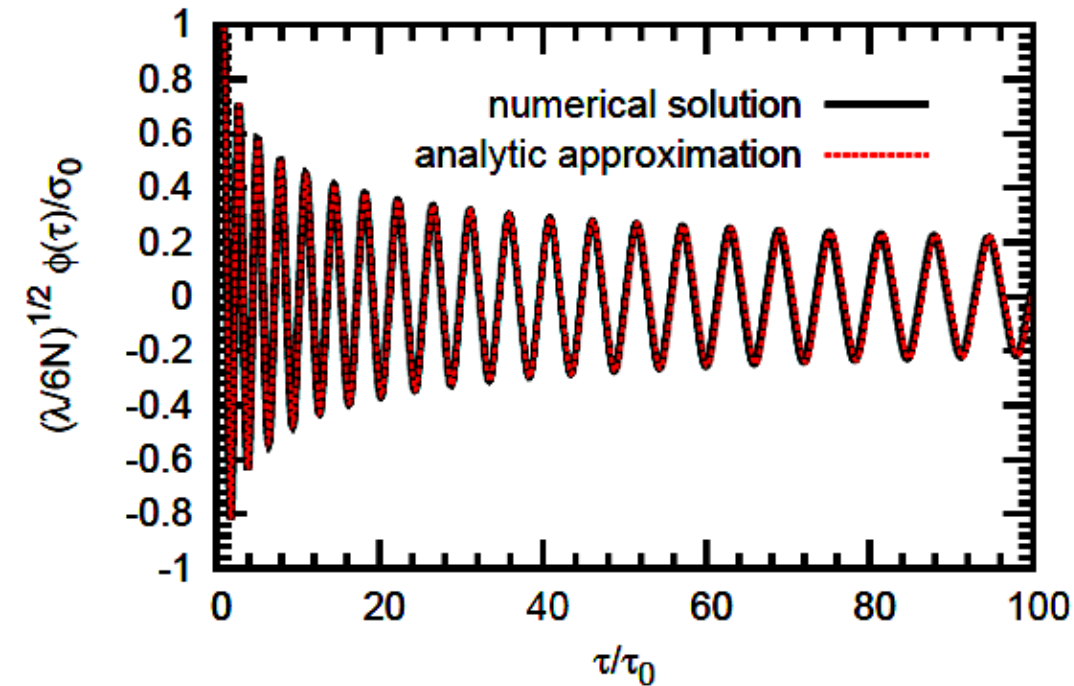
Symmetric phase $\langle \Phi \rangle = 0$

	$\lambda\Phi^4$	$O(N)$	$SU(N)$
Truncation	NLO of λ 	NLO of $1/N$ 	LO of g^2 
Others' Code	1+1 dim 2+1 dim 3+1 dim	1+1 dim 3+1 dim	?
Our Code	1+1 dim 2+1 dim 3+1 dim	1+1 dim 2+1 dim 3+1 dim	2+1 dim 3+1 dim

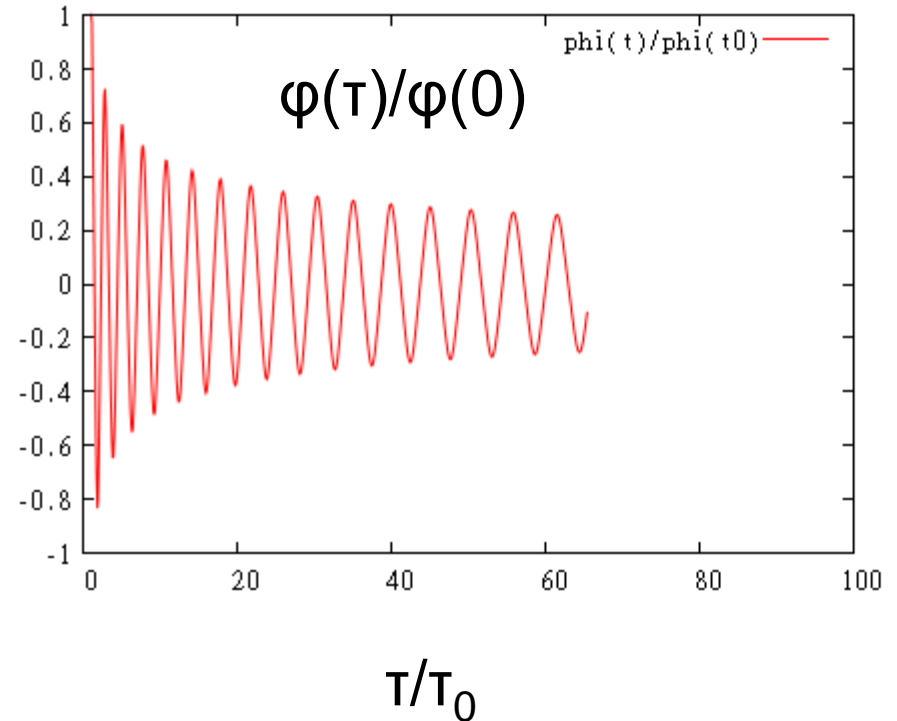
Evolution of classical field and fluctuation

Reproduction of J. Berges, K. Boguslavski, S. Schlichting, hep-ph 1201.3582.

Case without collision term



$$\varphi \sim T^{-1/3}$$



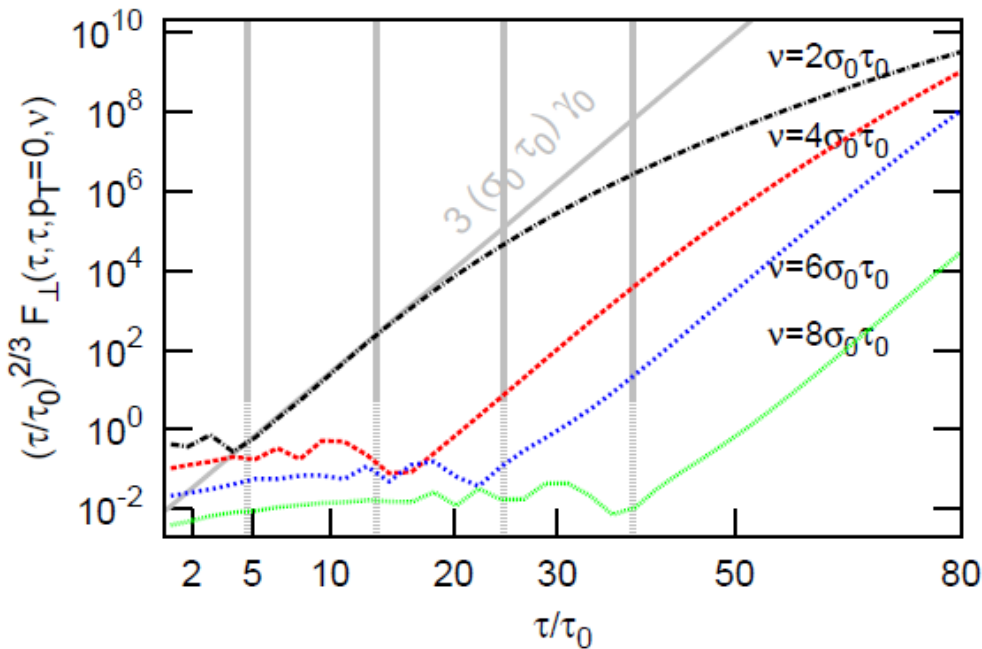
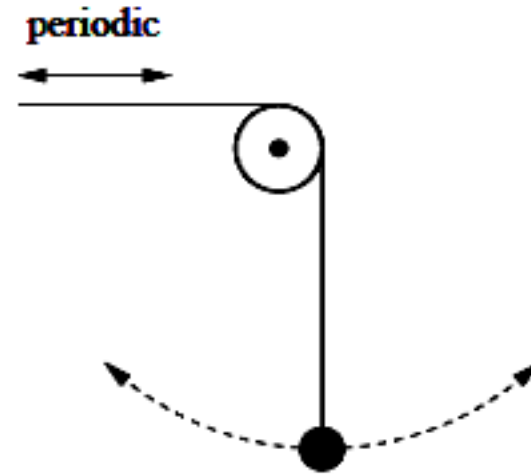
Parametric Resonance instability

Fluctuation $\ddot{y} + \omega^2(t)y = 0$

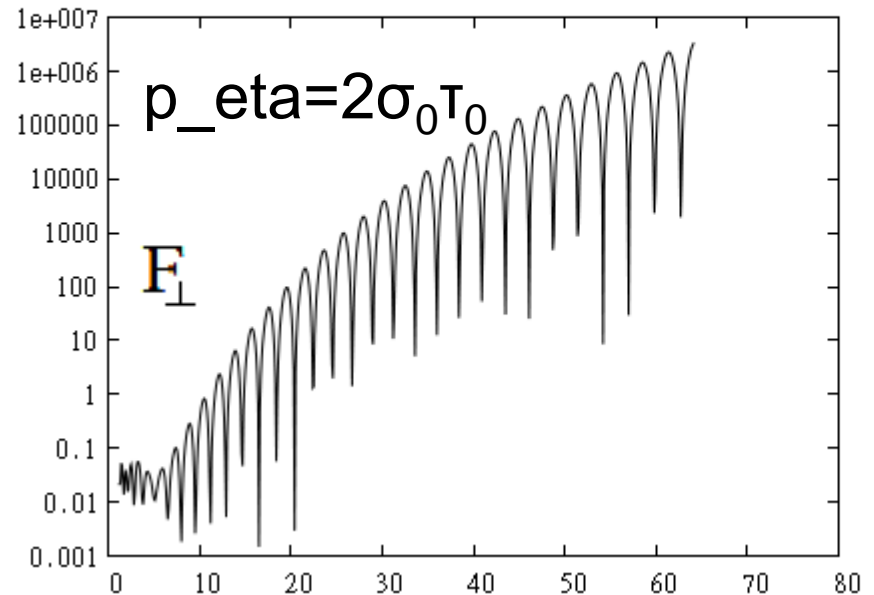
periodic $\omega(t + T) = \omega(t)$

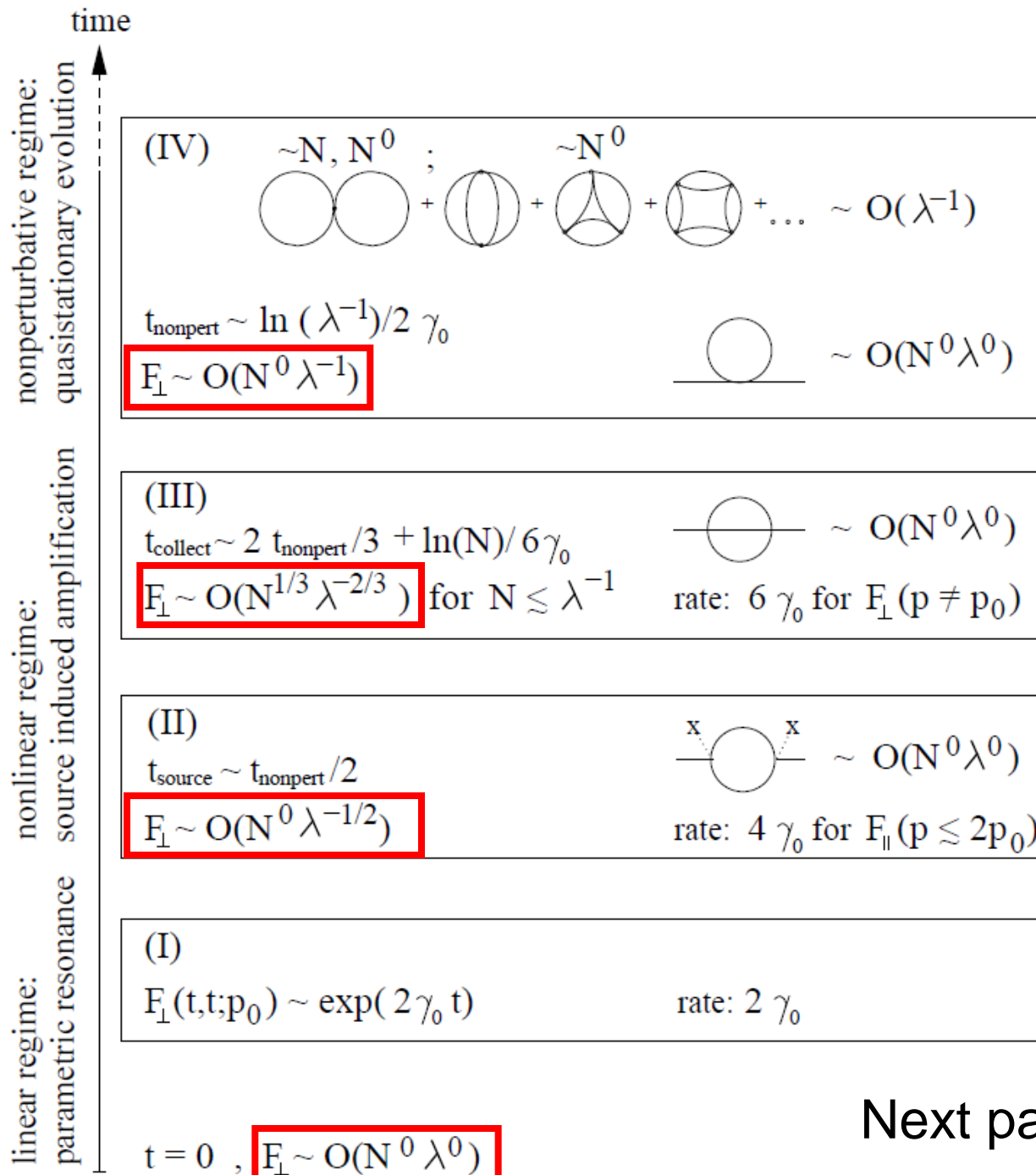
$y(t) = c^{t/T} \Pi(t) \quad c > 1$

$\omega^2(t) \sim \phi^2(t) + \dots$ Flat



Curved $\exp(\gamma_0 \tau^{2/3})$





Berges 2004.

(Flat metric)

Next page: Our results