

# Quark-Hadron Phase Transition in an Extended NJL Model with Scalar-Vector Eight-Point Interaction

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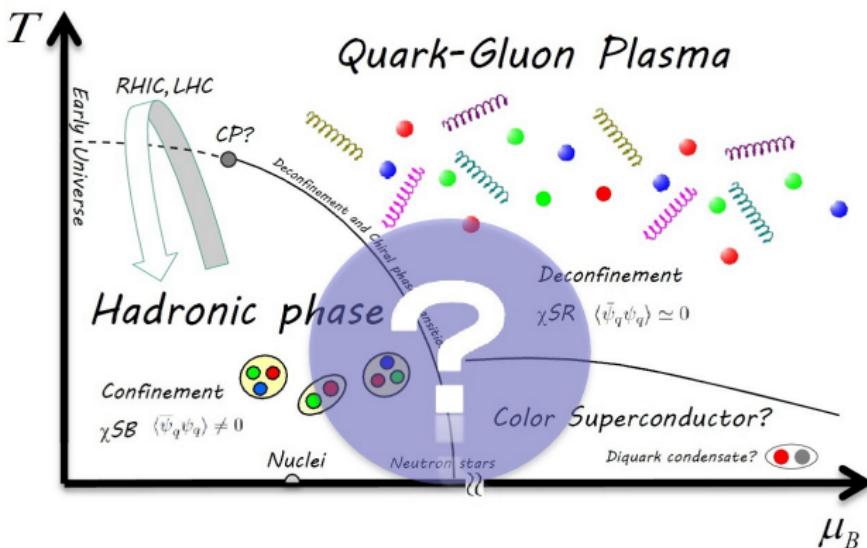
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arXiv:1207.1499

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# Schematic QCD phase diagram



## Finite density systems

- ▶ Difficulty in the Lattice-QCD calculations (**Sign problem**)
- ▶ Effective model approach (e.g. NJL model [Nambu et al., '61])
- ▶ Quark-Hadron phase transition (**Confinement problem**)

# Objective and Methods

## Objective and Methods

- ▶ To investigate the quark-hadron phase transition at finite temperature and density.
- ▶ We use the extended NJL model with the scalar-vector interaction.

## The extended NJL model

- ▶ Reproduction of nuclear saturation property
  - ▶ Introduction of  $G_v^N, G_{sv}^N$   
 $G_v^N$  : 4-point vector-type interaction  
 $G_{sv}^N$  : 8-point scalar-vector-type interaction [Koch et al., '87]
- ▶ Influence on the chiral phase transition with  $G_{sv}^q$ 
  - ▶ Introduction of  $G_{sv}^q$   
Strengthening attractive quark-antiquark interaction  
(Role in pushing the chiral restoration point to higher density side)  
→ Tuning parameter of the chiral symmetry restoration point
- ▶ Nucleon/Quark is treated as a fundamental fermion.
- ▶ Chiral symmetry is preserved.



# Objective and Methods

## Procedure for investigating

### Assumption

- ▶ Hadronic phase side  $\Rightarrow$  Symmetric nuclear matter
- ▶ Quark phase side  $\Rightarrow$  Free quark phase (No quark-pair correlations)

### Model application

- ▶ Symmetric nuclear matter
  - $\Rightarrow$  The extended 2-flavor NJL model with  $G_v^N$  and  $G_{sv}^N$
  - $\Rightarrow$  Nucleon field is treated as a fundamental field with  $N_c^N=1$ .
- ▶ Quark matter
  - $\Rightarrow$  The extended 2-flavor NJL model with  $G_{sv}^q$
  - $\Rightarrow$  Quark field is treated as a fundamental field with  $N_c^q=3$ .

### Phase determination

- ▶ Comparison of the pressure of nuclear matter with that of quark matter
  - $\Rightarrow$  The phase which has the largest pressure is physically realized.
  - $\Rightarrow$  Phase diagram (Paying attention to the order of phase transition)

# Outline

1 Introduction

2 Formalism

3 Numerical Results

4 Summary and Future Work

# Formalism

## Thermodynamics

(Extended NJL model + Mean field approximation)

# Formalism

- 
- ▶ Lagrangian density for nuclear and quark matters :

$$\begin{aligned}\mathcal{L}_i = & \bar{\psi}_i i\gamma^\mu \partial_\mu \psi_i + G_s^i [(\bar{\psi}_i \psi_i)^2 + (\bar{\psi}_i i\gamma_5 \boldsymbol{\tau} \psi_i)^2] \\ & - G_v^i (\bar{\psi}_i \gamma^\mu \psi_i)^2 - G_{sv}^i [(\bar{\psi}_i \psi_i)^2 + (\bar{\psi}_i i\gamma_5 \boldsymbol{\tau} \psi_i)^2] (\bar{\psi}_i \gamma^\mu \psi_i)^2\end{aligned}$$

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- ▶ For nuclear matter ( $i = N$ )  
⇒  $N_f^N = 2$ ,  $N_c^N = 1$ ,  $G_v^N \neq 0$ ,  $G_{sv}^N \neq 0$
- ▶ For quark matter ( $i = q$ )  
⇒  $N_f^q = 2$ ,  $N_c^q = 3$ ,  $G_v^q = 0$ ,  $G_{sv}^q \neq 0$

# Formalism

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► Mean field approximation :

$$\mathcal{L}_i^{MF} = \bar{\psi}_i(i\gamma^\mu\partial_\mu - m_i)\psi_i + \tilde{\mu}_i\bar{\psi}_i\gamma^0\psi_i + C_i$$

$$\mathcal{H}_i^{MF} = -i\bar{\psi}_i\boldsymbol{\gamma}\cdot\nabla\psi_i + m_i\bar{\psi}_i\psi_i + \tilde{\mu}_i\bar{\psi}_i\gamma^0\psi_i - C_i$$

with

$$C_i \equiv -G_s^i \langle\langle \bar{\psi}_i\psi_i \rangle\rangle^2 + G_v^i \langle\langle \bar{\psi}_i\gamma^0\psi_i \rangle\rangle^2 + 3G_{sv}^i \langle\langle \bar{\psi}_i\psi_i \rangle\rangle^2 \langle\langle \bar{\psi}_i\gamma^0\psi_i \rangle\rangle^2$$

$$m_i = -2 [G_s^i + 2G_{sv}^i \langle\langle \bar{\psi}_i\gamma^0\psi_i \rangle\rangle^2] \langle\langle \bar{\psi}_i\psi_i \rangle\rangle$$

$$\tilde{\mu}_i = 2 [G_v^i + 2G_{sv}^i \langle\langle \bar{\psi}_i\psi_i \rangle\rangle^2] \langle\langle \bar{\psi}_i\gamma^0\psi_i \rangle\rangle$$

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MFA :  $\bar{\psi}\Gamma\psi \rightarrow \langle\langle \bar{\psi}\Gamma\psi \rangle\rangle + (\bar{\psi}\Gamma\psi - \langle\langle \bar{\psi}\Gamma\psi \rangle\rangle)$

$\langle\langle \bar{\psi}_i\psi_i \rangle\rangle \neq 0$ ,  $\langle\langle \bar{\psi}_i\gamma^0\psi_i \rangle\rangle \neq 0$ , others = 0

( Fermion number density :  $\rho_i \equiv \langle\langle \psi_i^\dagger\psi_i \rangle\rangle = \langle\langle \bar{\psi}_i\gamma^0\psi_i \rangle\rangle$  )

$\langle\langle \dots \rangle\rangle$  : The finite-temperature expectation value which represents thermal average.

# Formalism

- ▶ Introduce the chemical potential  $\mu_i$  :

$$\begin{aligned}\mathcal{H}'_i &= \mathcal{H}_i^{MF} - \mu_i \psi_i^\dagger \psi_i \\ &= -i\bar{\psi}_i \boldsymbol{\gamma} \cdot \nabla \psi_i + m_i \bar{\psi}_i \psi_i - \mu_i^r \bar{\psi}_i \gamma^0 \psi_i - C_i\end{aligned}$$

- ⇒ The effective chemical potential  $\mu_i^r$  :

$$\begin{aligned}\mu_i^r &= \mu_i - \tilde{\mu}_i \\ &= \mu_i - 2 [G_v^i + G_{sv}^i \langle\langle \bar{\psi}_i \psi_i \rangle\rangle^2] \langle\langle \bar{\psi}_i \gamma^0 \psi_i \rangle\rangle\end{aligned}$$

# Formalism

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- ▶ Gap equation (Self-consistent equation for  $m_i$ ) :

$$m_i = -2G_s^i \left[ 1 - \frac{G_{sv}^i}{G_s^i} \langle\langle \bar{\psi}_i \gamma^0 \psi_i \rangle\rangle^2 \right] \langle\langle \bar{\psi}_i \psi_i \rangle\rangle$$


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where

$$\langle\langle \bar{\psi}_i \psi_i \rangle\rangle = \nu_i \int \frac{d^3 p}{(2\pi)^3} \frac{m_i}{\sqrt{\mathbf{p}^2 + m_i^2}} (n_+^i - n_-^i)$$

$$\langle\langle \bar{\psi}_i \gamma^0 \psi_i \rangle\rangle = \nu_i \int \frac{d^3 p}{(2\pi)^3} (n_+^i + n_-^i - 1)$$

with

$$\nu_i = 2N_f^i N_c^i \quad \text{Degeneracy factor}$$

$$n_{\pm}^i = \left[ e^{\beta(\pm\sqrt{\mathbf{p}^2 + m_i^2} - \mu_i^r)} + 1 \right]^{-1} \quad \text{Fermion number distribution function}$$

$$\beta = 1/T \quad \text{Temperature}$$

$$\mu_i^r = \mu_i - 2 \left[ G_v^i + G_{sv}^i \langle\langle \bar{\psi}_i \psi_i \rangle\rangle^2 \right] \langle\langle \bar{\psi}_i \gamma^0 \psi_i \rangle\rangle \quad \text{Effective chemical potential}$$

# Formalism

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► Pressure of nuclear and quark matters :

$$p_i(T, \mu_i) = - \left[ \langle\langle \mathcal{H}_i^{MF} \rangle\rangle_{(T, \mu_i)} - \langle\langle \mathcal{H}_i^{MF} \rangle\rangle_{(T=0, \mu_i=m_i(T=0))} \right] + \mu_i \langle\langle \mathcal{N}_i \rangle\rangle + T \langle\langle S_i \rangle\rangle$$


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where

$$\begin{aligned} \langle\langle \mathcal{H}_i^{MF} \rangle\rangle &= \langle\langle \bar{\psi}_i (\boldsymbol{\gamma} \cdot \mathbf{p}) \psi_i \rangle\rangle - G_s^i \langle\langle \bar{\psi}_i \psi_i \rangle\rangle^2 \\ &\quad + G_v^i \langle\langle \bar{\psi}_i \gamma^0 \psi_i \rangle\rangle^2 + G_{sv}^i \langle\langle \bar{\psi}_i \psi_i \rangle\rangle^2 \langle\langle \bar{\psi}_i \gamma^0 \psi_i \rangle\rangle^2 \end{aligned}$$

$$\langle\langle \mathcal{N}_i \rangle\rangle = \langle\langle \bar{\psi}_i \gamma^0 \psi_i \rangle\rangle = \rho_i$$

$$\begin{aligned} \langle\langle S_i \rangle\rangle &= -\nu_i \int \frac{d^3 p}{(2\pi)^3} \left[ n_+^i \ln n_+^i + (1 - n_+^i) \ln(1 - n_+^i) \right. \\ &\quad \left. + n_-^i \ln n_-^i + (1 - n_-^i) \ln(1 - n_-^i) \right] \end{aligned}$$

and

$$\langle\langle \bar{\psi}_i (\boldsymbol{\gamma} \cdot \mathbf{p}) \psi_i \rangle\rangle = \nu_i \int \frac{d^3 p}{(2\pi)^3} \frac{\mathbf{p}^2}{\sqrt{\mathbf{p}^2 + m_i^2}} (n_+^i - n_-^i)$$

# Numerical Results

Parameter set, Results at finite temperature and density

# Parameter set

## ► The parameter set for nuclear matter( $i = N$ ) and quark matter ( $i = q$ )

[Y. Tsue, J. da Providênci, C Providênci and M. Yamamura, Prog. Theor. Phys. 123, (2010), 1013]

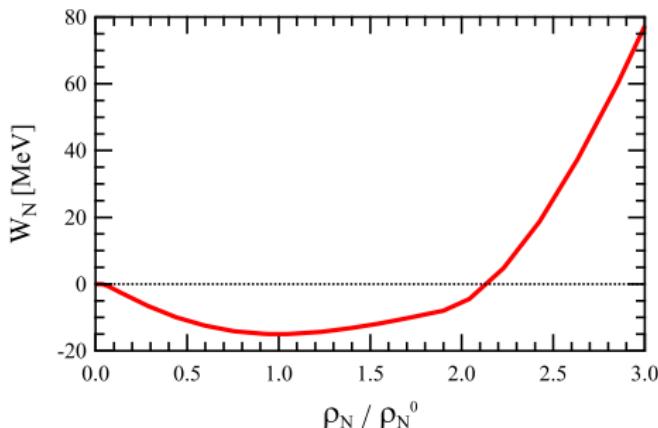
$\Lambda_N$	377.8 MeV	$\Lambda_q$	653.961 MeV
$G_s^N \Lambda_N^2$	19.2596	$G_s^q \Lambda_q^2$	2.13922
$G_v^N \Lambda_N^2$	-1069.89	$G_v^q \Lambda_q^2$	0
$G_{sv}^N \Lambda_N^8$	17.9824	$G_{sv}^q \Lambda_q^8$	free*)

\*)  $G_{sv}^q = 0$ ,  $G_{sv}^q \Lambda_q^8 = -68.4$ ,  $G_{sv}^q \Lambda_q^8 = -81.9$  [T.-G. Lee, et al., arXiv:1207.1499]

# Reproduction of nuclear saturation property

- Energy density per single nucleon  $W_N$

vs Normal nuclear density  $\rho_N / \rho_N^0$



- Incompressibility of nuclear matter

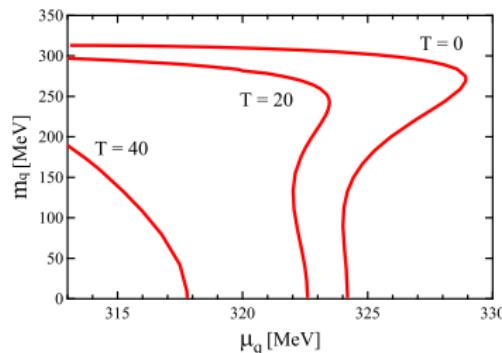
$$K = 9\rho_N^0 \frac{\partial^2 W_N(\rho_N)}{\partial \rho_N^2} \Big|_{\rho_N=\rho_N^0} \simeq 260 \text{ MeV}$$

# Numerical Results

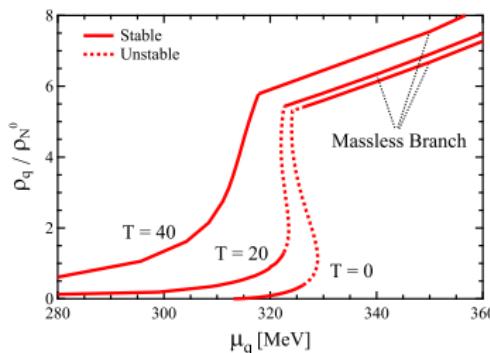
with  $G_{sv}^q \Lambda_q^8 = -68.4$

# Numerical Results with $G_{sv}^q \Lambda_q^8 = -68.4$

- ▶ Dynamical quark mass  $m_q$   
vs Quark chemical potential  $\mu_q$



- ▶ Quark number density  $\rho_q$   
vs Quark chemical potential  $\mu_q$



- ▶ Unphysical regions which have unstable solutions

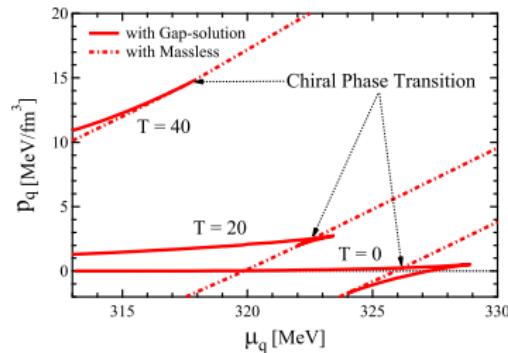
Comparison of pressure

⇒ Determine the physically realized solution (stable solution)

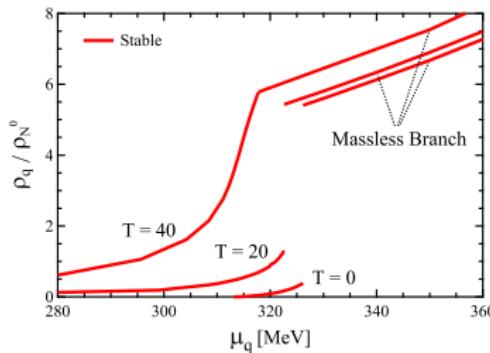
- ▶ The solution with largest pressure = The physically realized solution

# Numerical Results with $G_{sv}^q \Lambda_q^8 = -68.4$

► Pressure of quark matter  $p_q$   
vs Quark chemical potential  $\mu_q$



► Quark number density  $\rho_q$   
vs Quark chemical potential  $\mu_q$

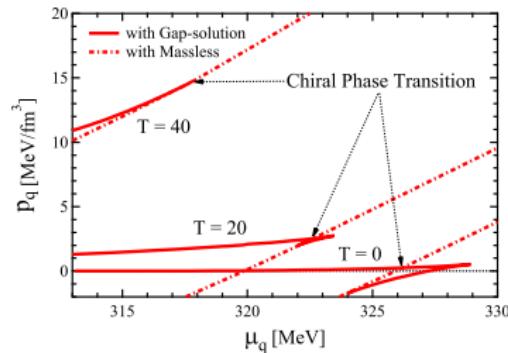


►  $T = 0$  MeV

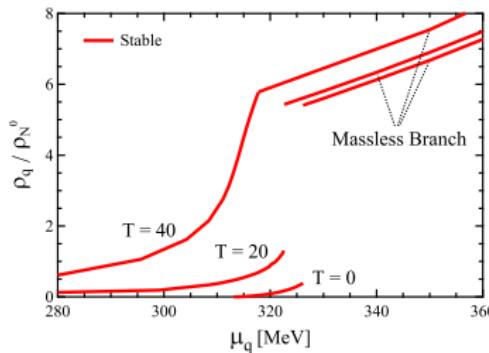
- $\mu_q^{\text{chiral}} \approx 326$  MeV : Chiral phase transition
- $\rho_q^{\text{coex}} = 0.38\rho_N^0 \sim 5.41\rho_N^0$  : 1<sup>st</sup>-order phase transition  
( $\rho_B = 0.13\rho_N^0 \sim 1.80\rho_N^0$ )

# Numerical Results with $G_{sv}^q \Lambda_q^8 = -68.4$

► Pressure of quark matter  $p_q$   
vs Quark chemical potential  $\mu_q$



► Quark number density  $\rho_q$   
vs Quark chemical potential  $\mu_q$

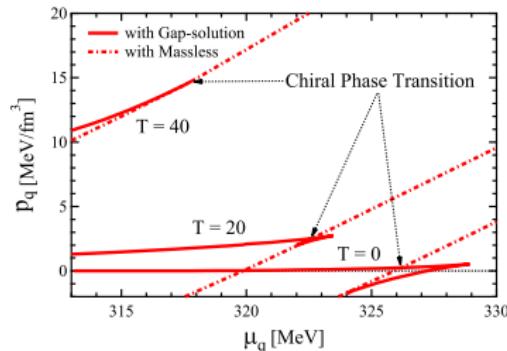


►  $T = 20$  MeV

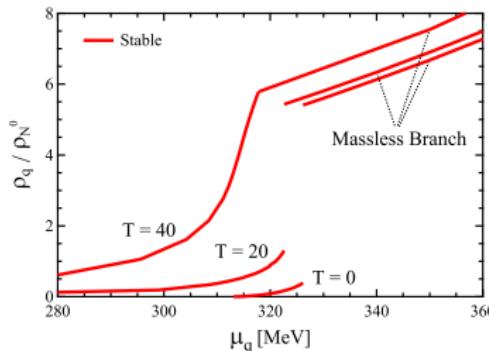
- $\mu_q^{\text{chiral}} \approx 323$  MeV : Chiral phase transition
- $\rho_q^{\text{coex}} = 1.30\rho_N^0 \sim 5.41\rho_N^0$  : 1<sup>st</sup>-order phase transition  
( $\rho_B = 0.43\rho_N^0 \sim 1.80\rho_N^0$ )

# Numerical Results with $G_{sv}^q \Lambda_q^8 = -68.4$

► Pressure of quark matter  $p_q$   
vs Quark chemical potential  $\mu_q$



► Quark number density  $\rho_q$   
vs Quark chemical potential  $\mu_q$



►  $T = 40$  MeV

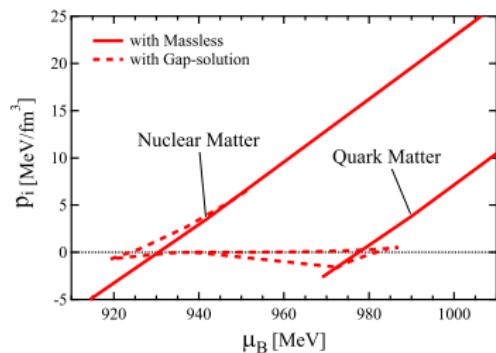
- $\mu_q^{\text{chiral}} \approx 318$  MeV : Chiral phase transition
- $\rho_q^{\text{chiral}} \sim 5.78 \rho_N^0$  : 2<sup>nd</sup>-order phase transition  
( $\rho_B \sim 1.93 \rho_N^0$ )

# Quark-Hadron phase transition

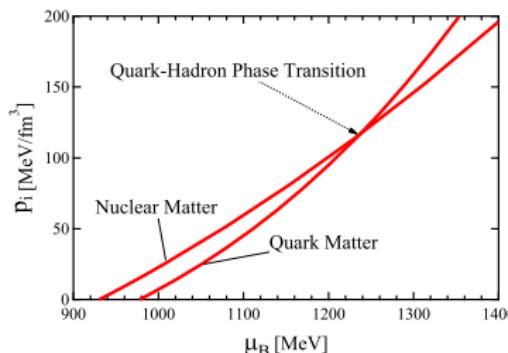
Numerical Results with  $G_{sv}^q \Lambda_q^8 = -68.4$

# Numerical Results with $G_{sv}^q \Lambda_q^8 = -68.4$

► Pressure  $p_i$  [Lower- $\mu_B$  side]  
vs Baryon number density  $\mu_B$



► Pressure  $p_i$  [Higher- $\mu_B$  side]  
vs Baryon number density  $\mu_B$



► The condition for thermodynamic equilibrium  
between the hadron and quark phases\*) :

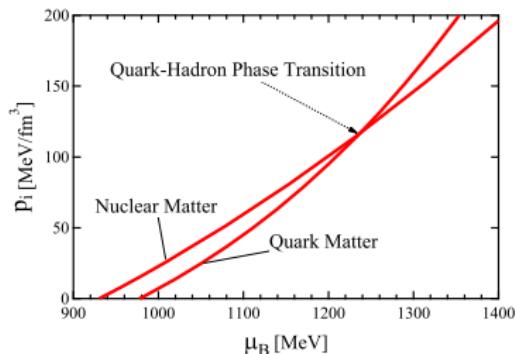
$$p_N(T, \mu_N) = p_q(T, 3\mu_q)$$

\*) The condition for chemical equilibrium (conservation of baryon number) :  $\mu_N(T) = 3\mu_q(T) \Leftrightarrow \mu_B(T)$

# Numerical Results with $G_{sv}^q \Lambda_q^8 = -68.4$

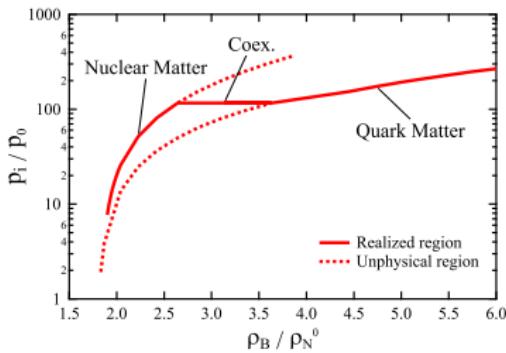
► Pressure  $p_i$

vs Baryon number density  $\mu_B$



► Pressure  $p_i/p_0$

vs Baryon number density  $\mu_B/\rho_N^0$



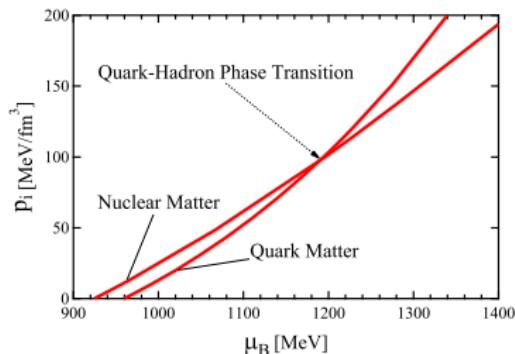
►  $T = 0$  MeV

- $\mu_B^{QH} \approx 1236$  MeV : Quark-Hadron phase transition
- $\rho_B^{\text{coex}} = 2.64\rho_N^0 \sim 3.63\rho_N^0$  : 1<sup>st</sup>-order phase transition  
( $\rho_N = 2.64\rho_N^0 \sim \rho_q = 10.9\rho_N^0$ )

# Numerical Results with $G_{sv}^q \Lambda_q^8 = -68.4$

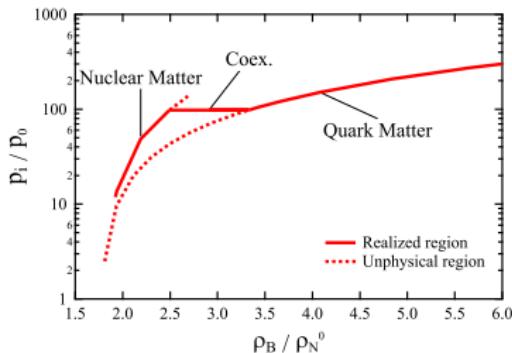
► Pressure  $p_i$

vs Baryon number density  $\mu_B$



► Pressure  $p_i/p_0$

vs Baryon number density  $\mu_B/\rho_N^0$



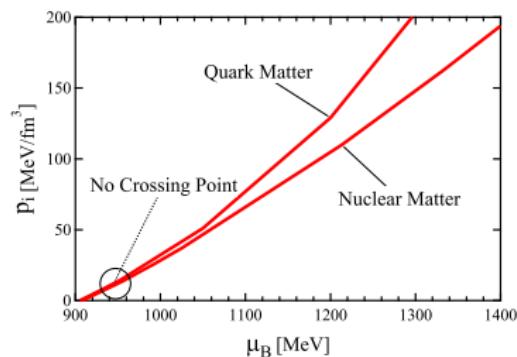
►  $T = 20$  MeV

- $\mu_B^{QH} \approx 1190$  MeV : Quark-Hadron phase transition
- $\rho_B^{\text{coex}} = 2.49\rho_N^0 \sim 9.99\rho_N^0$  : 1<sup>st</sup>-order phase transition  
( $\rho_N = 2.49\rho_N^0 \sim \rho_q = 3.33\rho_N^0$ )

# Numerical Results with $G_{sv}^q \Lambda_q^8 = -68.4$

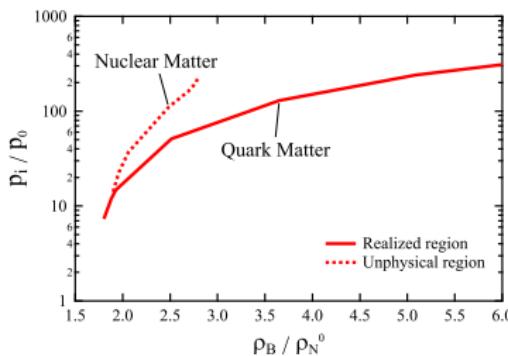
► Pressure  $p_i$

vs Baryon number density  $\mu_B$



► Pressure  $p_i/p_0$

vs Baryon number density  $\mu_B/\rho_N^0$



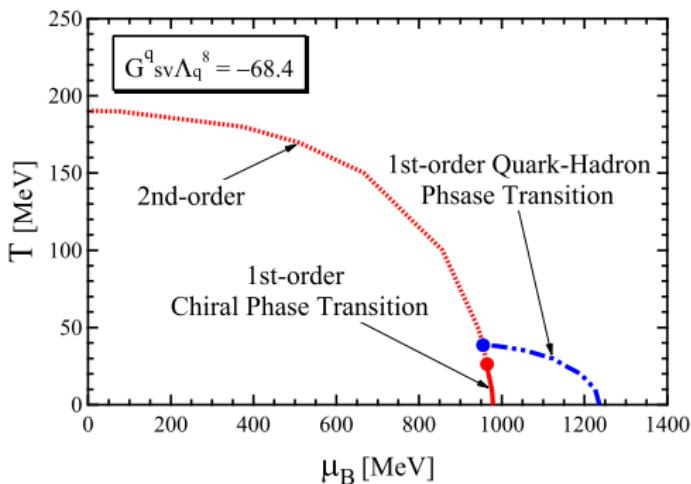
►  $T = 40 \text{ MeV}$

- There is no crossing point.

$\Rightarrow$  1<sup>st</sup>-order quark-hadron phase transition disappears.

# Phase diagram ( $\mu_B, T$ )

- Phase diagram with  $G_{sv}^q \Lambda_q^8 = -68.4$

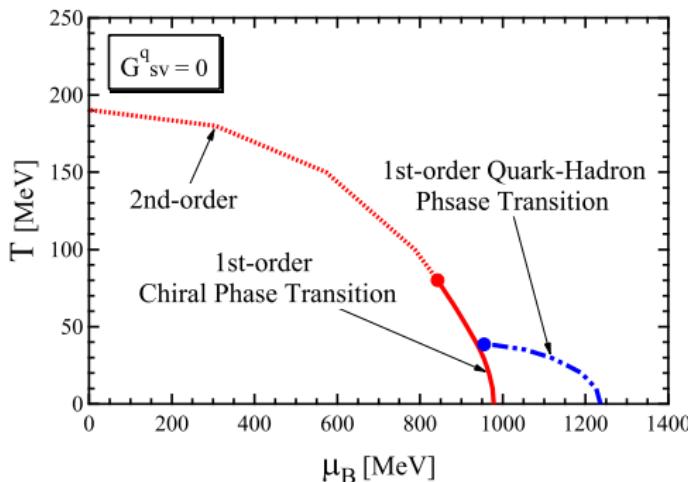


- ▷ Low- $T$  : 1<sup>st</sup>-order chiral phase transition → 1<sup>st</sup>-order quark-hadron phase transition
- ▷ High- $T$  : 2<sup>nd</sup>-order chiral phase transition
- ▷ Moderate- $\mu_B$  : Chiral restoration + Nucleonic/Hadronic excitation  
( Chiral symmetric nuclear phase ⇒ Quarkyonic-like phase ? )

# $G_{sv}^q$ -dependence of phase diagram

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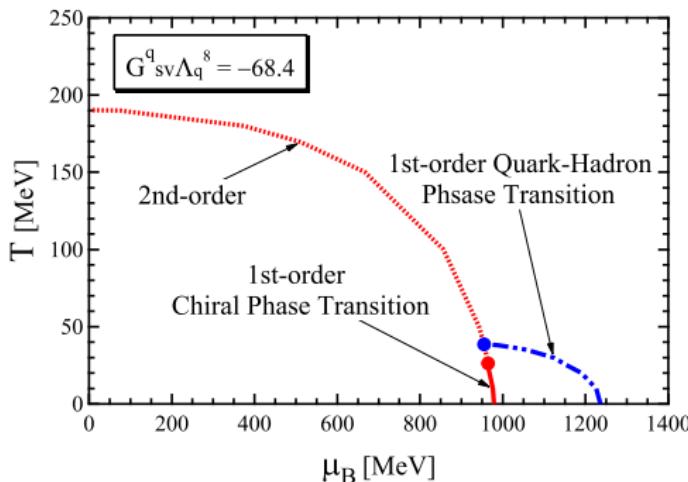
- Phase diagram without scalar-vector interaction



- 1<sup>st</sup>-order chiral phase transition :  $(\mu_B, T) \simeq (978, 0) \rightarrow (842, 80)$  MeV
- 2<sup>nd</sup>-order chiral phase transition :  $(\mu_B, T) \simeq (842, 80) \rightarrow (0, 190)$  MeV
- 1<sup>st</sup>-order quark-hadron transition :  $(\mu_B, T) \simeq (1236, 0) \rightarrow (955, 39)$  MeV

# $G_{sv}^q$ -dependence of phase diagram

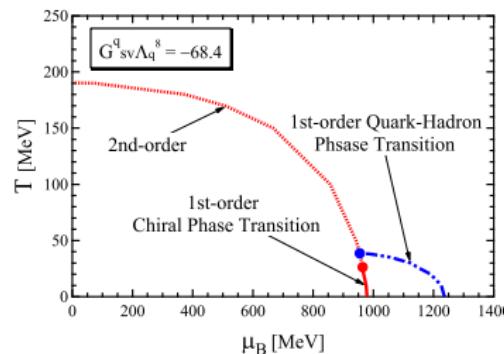
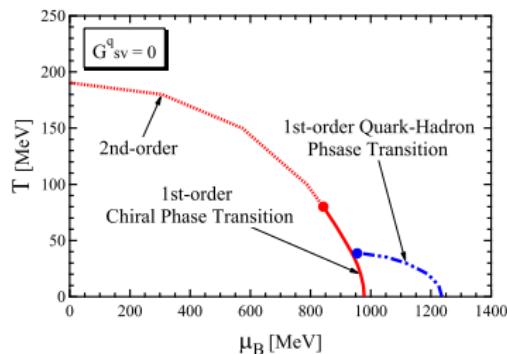
- ▶ Phase diagram with scalar-vector interaction



- ▷ **1<sup>st</sup>-order chiral phase transition** :  $(\mu_B, T) \simeq (979, 0) \rightarrow (964, 26)$  MeV
- ▷ **2<sup>nd</sup>-order chiral phase transition** :  $(\mu_B, T) \simeq (964, 26) \rightarrow (0, 190)$  MeV
- ▷ **1<sup>st</sup>-order quark-hadron transition** :  $(\mu_B, T) \simeq (1236, 0) \rightarrow (955, 39)$  MeV

# $G_{sv}^q$ -dependence of phase diagram

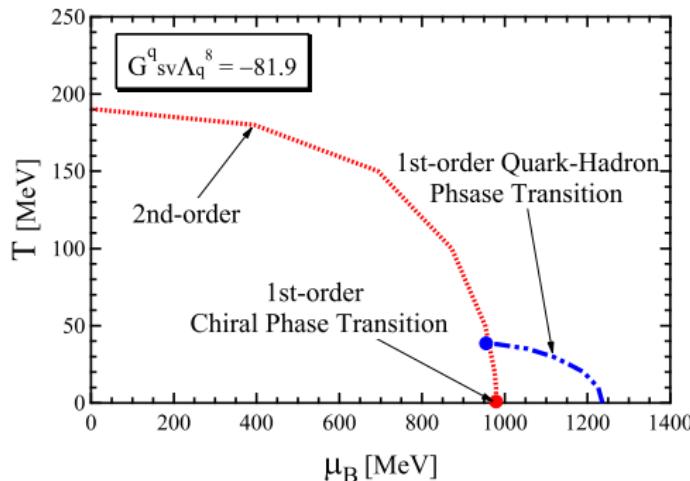
- Phase diagram without and with scalar-vector interaction



- **1<sup>st</sup>-order chiral phase transition** : Shrinks with increasing  $G_{sv}^q$ .
- **2<sup>nd</sup>-order chiral phase transition** : Tends to bloat outward.
- **1<sup>st</sup>-order quark-hadron transition** : There is no change.

# $G_{sv}^q$ -dependence of phase diagram

- Phase diagram with stronger scalar-vector interaction



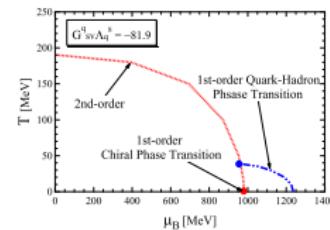
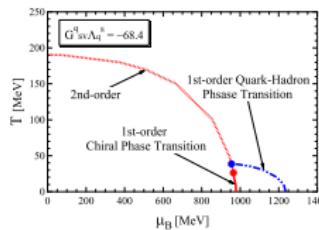
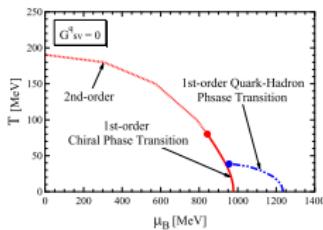
- 1<sup>st</sup>-order chiral phase transition :  $(\mu_B, T) \simeq (979, 0) \rightarrow (979, 1)$  MeV
- 2<sup>nd</sup>-order chiral phase transition :  $(\mu_B, T) \simeq (979, 1) \rightarrow (0, 190)$  MeV
- 1<sup>st</sup>-order quark-hadron transition :  $(\mu_B, T) \simeq (1236, 0) \rightarrow (955, 39)$  MeV

# Summary and Future Work

# Summary and Future Work

## Summary

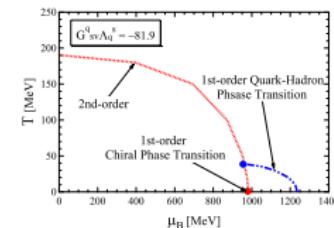
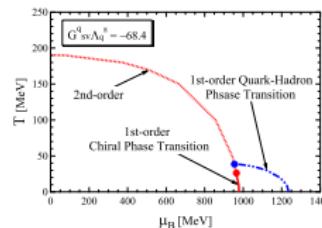
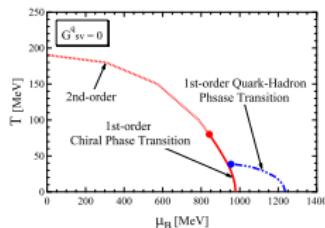
- ▷ The quark-hadron phase transition at finite temperature and density is investigated in an Extended NJL model.
- ▷ The extended NJL phase diagram with  $G_{sv}^q$  is obtained in the  $T$ - $\mu_B$  plane.
  - ▶ Phase diagram with  $G_{sv}^q \Lambda^q = -68.4$ 
    - ▶ 1<sup>st</sup>-order quark-hadron phase transition
    - ▶ 1<sup>st</sup> and 2<sup>nd</sup>-order chiral phase transition
    - ▶ The phase in which chiral symmetry is restored but the elementary excitation modes are nucleonic appears just before deconfinement.
  - ⇒ Quarkyonic phase? [McLerran et al.,'07]



# Summary and Future Work

## Summary

- ▶ Influence on the phase diagram with  $G_{sv}^q$ 
  - ▶ Does not affect to the 1<sup>st</sup>-order quark-hadron phase transition.  
⇒ The phase boundary is not changed. ( $G_{sv}^q$ -independence)
  - ▶ Affects the chiral phase transition. ( $G_{sv}^q$ -dependence)
    - ⇒ Critical line of 1<sup>st</sup>-order shrinks with increasing  $G_{sv}^q$
    - ⇒ Moves the tricritical point toward a larger  $\mu_B$  and a lower  $T$ .
    - ⇒ The endpoint of 1<sup>st</sup>-order quark-hadron phase transition is located on the critical line of chiral phase transition in the case of  $G_{sv}^q \Lambda_q = -68.4$ .



# Summary and Future Work

## Future Work

- ▶ Consideration of the color-superconducting phase
  - ⇒ Pairing interaction (Nuclear superfluidity in nuclear side, CSC in quark side)
- ▶ Consideration of the neutron star matter
  - ⇒ Phase transition between neutron matter and quark matter.  
(→ Physics of neutron star)

Thank you for your attention!

# Back up

# Parameter set

## ► Nuclear matter

- Model parameters :  $G_s^N, G_v^N, G_{sv}^N, \Lambda_N$       3-momentum cutoff :  $\Lambda_i$
- Conditions :

$$m_N(\rho_N=0) = 939 \text{ MeV}$$

$$\rho_N^0 = 0.17 \text{ fm}^{-3} \quad \text{Normal nuclear density}$$

$$m_N(\rho_N=\rho_N^0) = 0.6m_N(\rho_N=0) \text{ MeV}$$

$$W_N(\rho_N=\rho_N^0) = -15 \text{ MeV}$$

$$\text{Binding energy per single nucleon : } W_N(\rho_N) = \frac{\langle \mathcal{H}_i^{MF} \rangle(\rho_N) - \langle \mathcal{H}_i^{MF} \rangle(\rho_N=0)}{\rho_N} - m_N(\rho_N=0)$$


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## ► Quark matter

- Model parameters :  $G_s^q, \Lambda_q$

- Conditions :

$$m_q(\rho_q=0) = 313 \text{ MeV} \quad \text{Dynamical quark mass}$$

$$f_\pi^2 = 93 \text{ MeV} \quad \text{Pion decay constant}$$

- Free parameter :  $G_{sv}^q$

- Conditions :

$$m_q(\rho_q/3=\rho_N^0) = 0.625m_q(\rho_q=0) \text{ MeV} \rightarrow G_{sv}^q \Lambda_q^8 = -68.4$$

$$m_q(\rho_q/3=\rho_N^0) = 0.63m_q(\rho_q=0) \text{ MeV} \rightarrow G_{sv}^q \Lambda_q^8 = -81.9$$

# $G_{sv}^q$ -independence of the quark-hadron phase transition

## $G_{sv}^q$ -independence

- ▶ There is no influence on the 1<sup>st</sup>-order quark-hadron phase transition.
  - ▶ 1<sup>st</sup>-order quark-hadron phase transition occurs after the chiral restoration.  
 $\Rightarrow m_q = 0, \langle\langle \bar{\psi}_q \psi_q \rangle\rangle = 0$
  - ▶  $G_{sv}^q$ -independence of pressure  $p_q$

$$p_q(T, \mu_q) = - \left[ \langle\langle \mathcal{H}_q^{MF} \rangle\rangle_{(T, \mu_q)} - \langle\langle \mathcal{H}_q^{MF} \rangle\rangle_{(T=0, \rho_q=0)} \right] + \mu_q \langle\langle \mathcal{N}_q \rangle\rangle + T \langle\langle S_q \rangle\rangle$$

$$\begin{aligned} \langle\langle \mathcal{H}_q^{MF} \rangle\rangle &= \langle\langle \bar{\psi}_q (\boldsymbol{\gamma} \cdot \mathbf{p}) \psi_q \rangle\rangle - G_s^q \langle\langle \bar{\psi}_q \psi_q \rangle\rangle^2 \\ &\quad + G_v^q \langle\langle \bar{\psi}_q \gamma^0 \psi_q \rangle\rangle^2 + G_{sv}^q \langle\langle \bar{\psi}_q \psi_q \rangle\rangle^2 \langle\langle \bar{\psi}_q \gamma^0 \psi_q \rangle\rangle^2 \end{aligned}$$

$$\mu_q^r = \mu_q - 2 \left[ G_v^q + G_{sv}^q \langle\langle \bar{\psi}_q \psi_q \rangle\rangle^2 \right] \langle\langle \bar{\psi}_q \gamma^0 \psi_q \rangle\rangle$$

$\Rightarrow \langle\langle \mathcal{H}_i^{MF} \rangle\rangle$  and  $\mu_q^r$  do not depend on  $G_{sv}^q$  due to  $\langle\langle \bar{\psi}_q \psi_i \rangle\rangle = 0$ .

# Thermodynamic potential density

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- Thermodynamic potential density :

$$\omega_i = \langle\langle \mathcal{H}_i^{MF} \rangle\rangle - \mu_i \langle\langle \mathcal{N}_i \rangle\rangle - T \langle\langle S_i \rangle\rangle$$


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where

$$\begin{aligned}\langle\langle \mathcal{H}_i^{MF} \rangle\rangle &= \langle\langle \bar{\psi}_i (\boldsymbol{\gamma} \cdot \mathbf{p}) \psi_i \rangle\rangle - G_s^i \langle\langle \bar{\psi}_i \psi_i \rangle\rangle^2 \\ &\quad + G_v^i \langle\langle \bar{\psi}_i \gamma^0 \psi_i \rangle\rangle^2 + G_{sv}^i \langle\langle \bar{\psi}_i \psi_i \rangle\rangle^2 \langle\langle \bar{\psi}_i \gamma^0 \psi_i \rangle\rangle^2 \\ \langle\langle \mathcal{N}_i \rangle\rangle &= \langle\langle \bar{\psi}_i \gamma^0 \psi_i \rangle\rangle = \rho_i \\ \langle\langle S_i \rangle\rangle &= -\nu_i \int \frac{d^3 \mathbf{p}}{(2\pi)^3} \left[ n_+^i \ln n_+^i + (1 - n_+^i) \ln(1 - n_+^i) \right. \\ &\quad \left. + n_-^i \ln n_-^i + (1 - n_-^i) \ln(1 - n_-^i) \right]\end{aligned}$$

and

$$\langle\langle \bar{\psi}_i (\boldsymbol{\gamma} \cdot \mathbf{p}) \psi_i \rangle\rangle = \nu_i \int \frac{d^3 \mathbf{p}}{(2\pi)^3} \frac{\mathbf{p}^2}{\sqrt{\mathbf{p}^2 + m_i^2}} (n_+^i - n_-^i)$$

Minimize  $\omega_i$  w.r.t  $m_i$  and  $n_i^\pm \Rightarrow$  Gap eq. and Fermion number distribution func.

# Pressure

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- ▶ Pressure of nuclear and quark matters :

$$p_i(T, \mu_i) = - \left[ \langle\!\langle \mathcal{H}_i^{MF} \rangle\!\rangle_{(T, \mu_i)} - \langle\!\langle \mathcal{H}_i^{MF} \rangle\!\rangle_{(T=0, \mu_i=m_i(T=0))} \right] + \mu_i \langle\!\langle \mathcal{N}_i \rangle\!\rangle + T \langle\!\langle S_i \rangle\!\rangle$$


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where

$$\langle\!\langle \mathcal{H}_i^{MF} \rangle\!\rangle_{(T=0, \mu_i=m_i(T=0))} = \langle \bar{\psi}_i (\boldsymbol{\gamma} \cdot \mathbf{p}) \psi_i \rangle - G_s \langle \bar{\psi}_i \psi_i \rangle^2$$

$\langle \dots \rangle$  : Zero-temperature expectation value

$$n_+^i(T=0) = \theta(\mu_i^r - \sqrt{\mathbf{p}^2 + m_i^2}) \quad \text{Heaviside step function}$$

$$= \begin{cases} 1 & (\mathbf{p} < \sqrt{\mu_i^{r2} - m_i^2} \equiv \mathbf{p}_F^i) \quad \mathbf{p}_F^i : \text{Fermi momentum} \\ 0 & (\mathbf{p} > \sqrt{\mu_i^{r2} - m_i^2} \equiv \mathbf{p}_F^i) \end{cases}$$

$$n_-^i(T=0) = 1$$