



W-boson mass with singlet scalars

Mainly based on 2204.04770 to appear in PLB, in collaboration with **Kodai Sakurai** (Tohoku -> Warsaw), Fuminobu Takahashi.

@23th July, 北陸合宿

殷文 (東北大学)

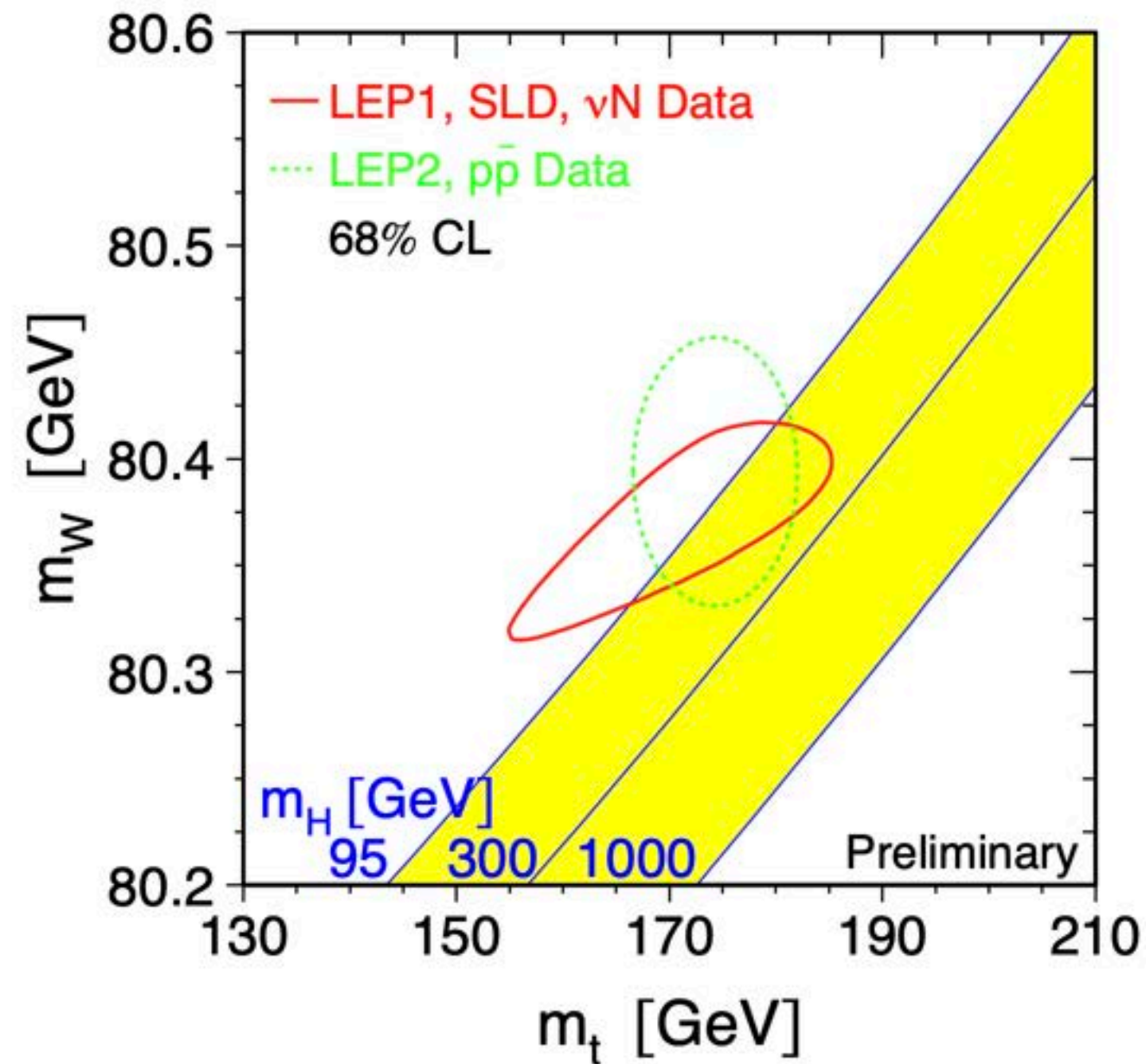
Contents

- 1. Introduction
- 2. W -boson mass with singlet extensions
- 3. Conclusions

1. Introduction

An example of the success of the SM

EW fit -> Higgs boson discovery

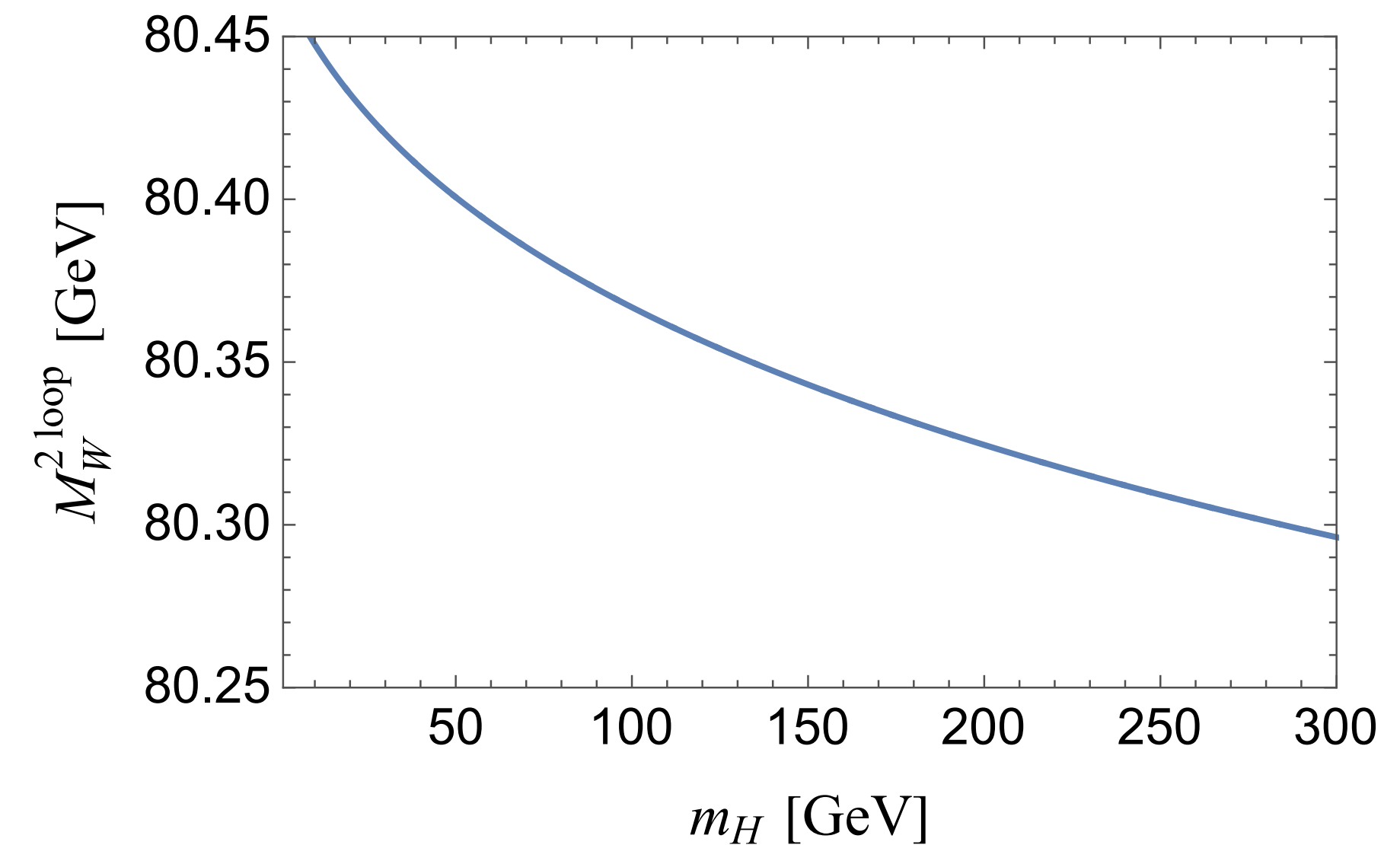
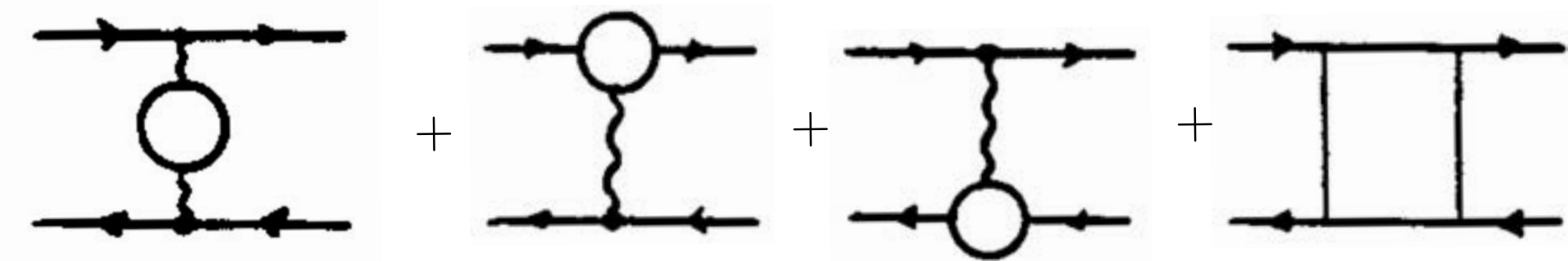


LEP EW WG summer 99 conference
<http://lepewwg.web.cern.ch/lepewwg/plots/summer1999/>

Matching the Fermi theory and SM prediction of muon life-time we get

$$M_W = \frac{M_Z}{2} \left(1 + \sqrt{1 - \frac{4\pi\alpha_{em}}{\sqrt{2}G_F M_Z^2 (1 - \Delta r)}} \right),$$

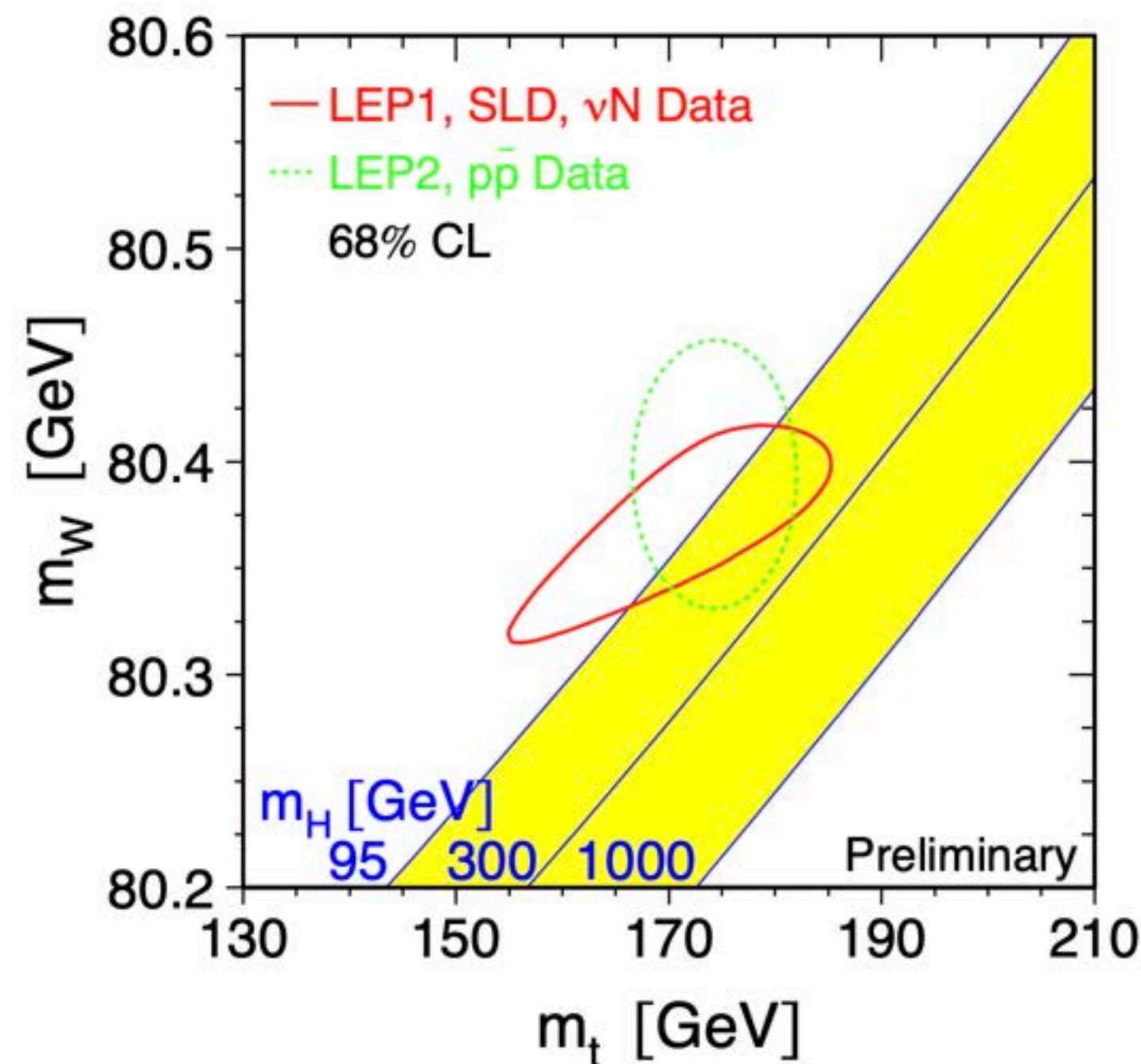
Δr includes the effects



この図や式はKodai Sakurai's slideから持ってきました🙄

An example of the success of the SM

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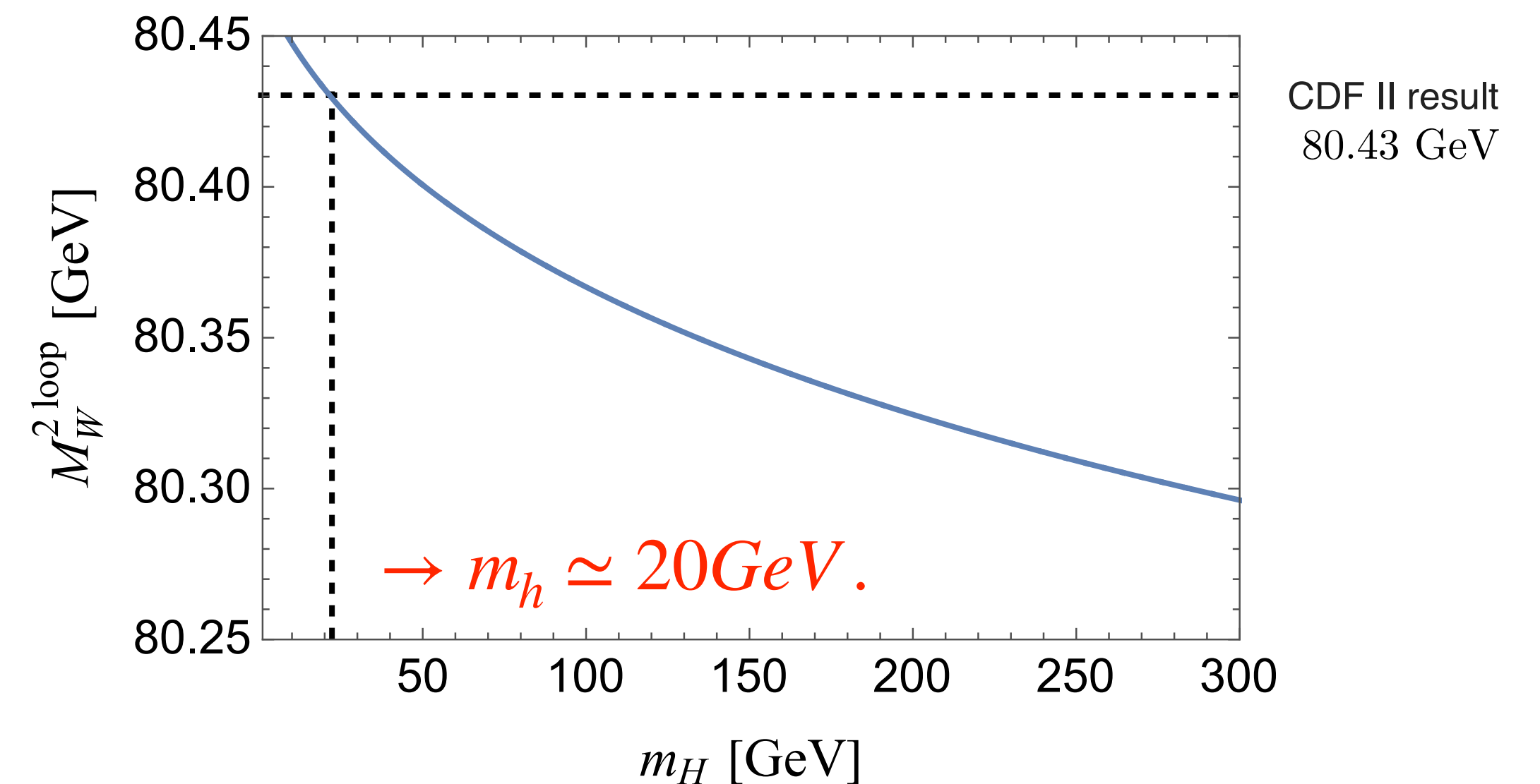
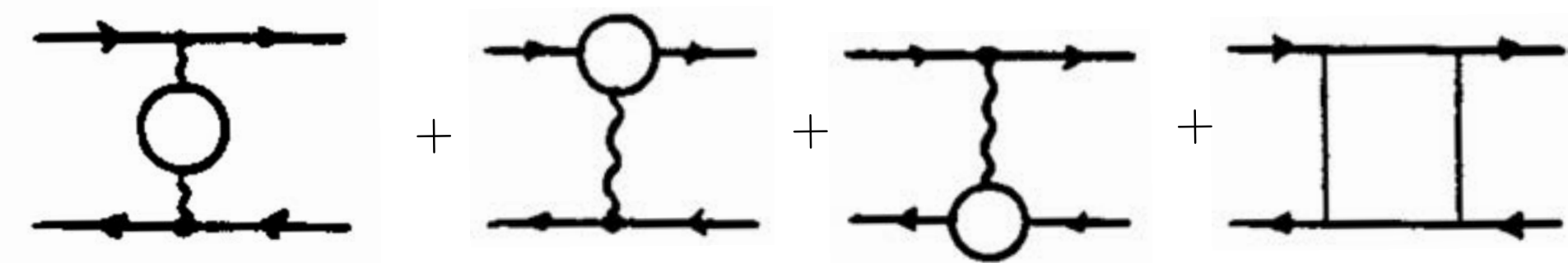
$80433.4 \pm 9.4 \text{ MeV}$ (CDFII collaboration 2022)

まぐれだった! ? 🤖

Matching the Fermi theory and SM prediction of muon life-time we get

$$M_W = \frac{M_Z}{2} \left(1 + \sqrt{1 - \frac{4\pi\alpha_{em}}{\sqrt{2}G_F M_Z^2 (1 - \Delta r)}} \right),$$

Δr includes the effects



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W-boson mass shiftの意味することは？

Success of SM

VS

待ちに待った
BSM 但し
SM success
は一部諦める。

$80376 \pm 23 \text{ MeV}$ (D0 II)

$80370 \pm 19 \text{ MeV}$ (Atlas)

$\sim 2-3\sigma$

$80433.4 \pm 9.4 \text{ MeV}$

(CDFII collaboration. 2022)

一つのBSM解としてFundamental constantの時空変化の可能性は今日は話さない。

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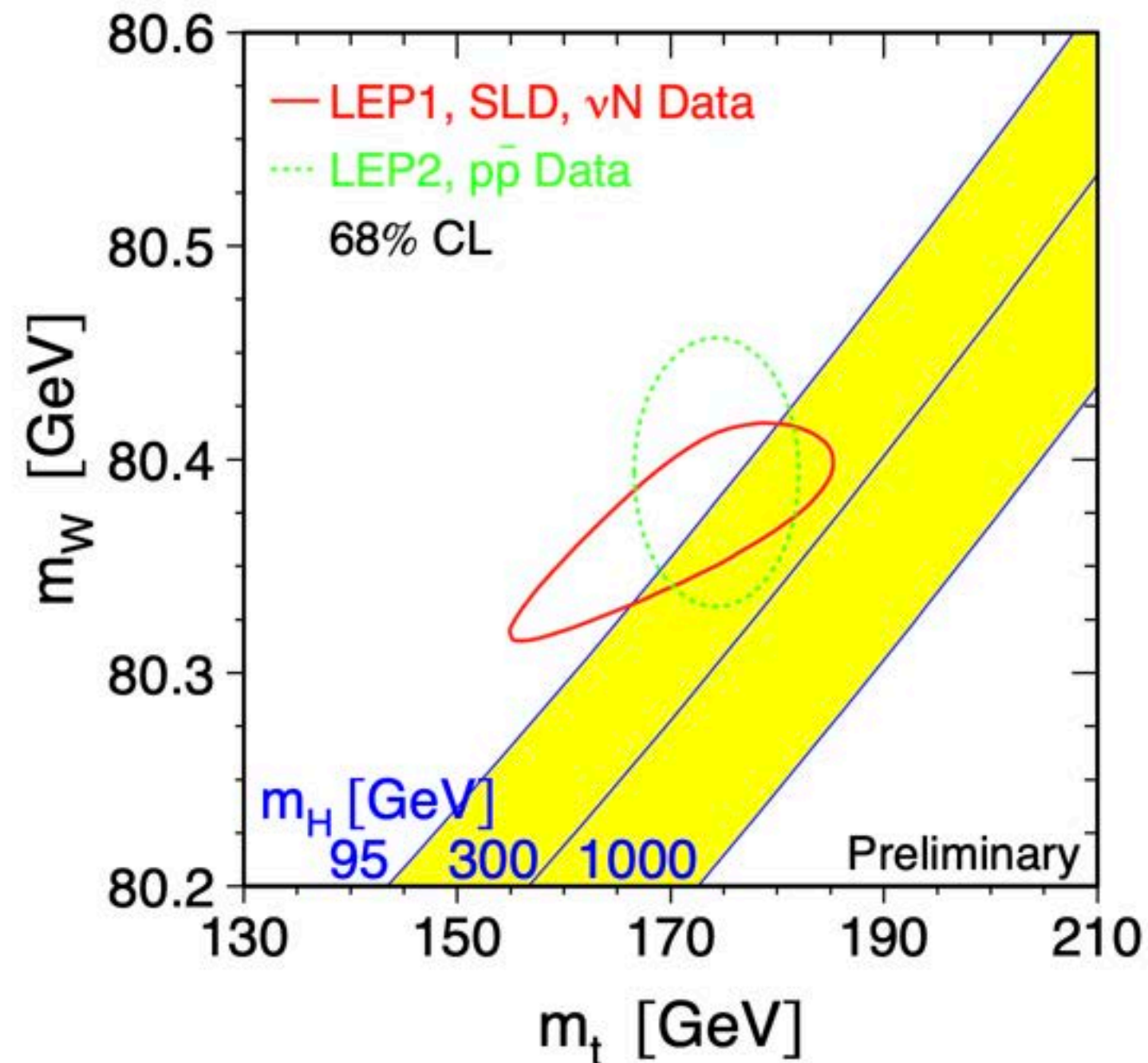
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(CDFII collaboration. 2022)

一つのBSM解としてFundamental constantの時空変化の可能性は今日は話さない。

A few examples of the success of the SM

EW fit -> Higgs boson discovery



LEP EW WG summer 99 conference

<http://lepewwg.web.cern.ch/lepewwg/plots/summer1999/>

$80433.4 \pm 9.4 \text{ MeV}$ (CDFII collaboration 2022)

$\rightarrow m_h \simeq 20 \text{ GeV}$. まぐれだと認めよう。

Prediction

Suppressed flavor violation!

Suppressed CP violation!

Proton stability!

They are predicted from particle contents
+ renormalizability.

I would like to consider BSM with this structure.

まぐれと認めたくない。

BSMs with similar properties.

Singlet scalar (dark sector) extensions

$$\delta V = V(|H|^2, s)$$

[Silveira, 1985; Burgess et al 0011335](#)

s can be thought to be CP-even if renormalizable, i.e. the potential is accidentally CP-conserving.

Flavor violation, proton decay are suppressed as in the SM.

Massive dark gauge boson (SM is not charged) also belong to this category since we need a dark higgs.

This talk

Non-trivially charged particles
*Real adjoint Higgs field

*Field with very large representation etc

Others, including the case that BSMs are heavy.

Singlet scalar extensions can connect light dark sector:

G_{dark} symmetry extension. $\Phi \supset s$

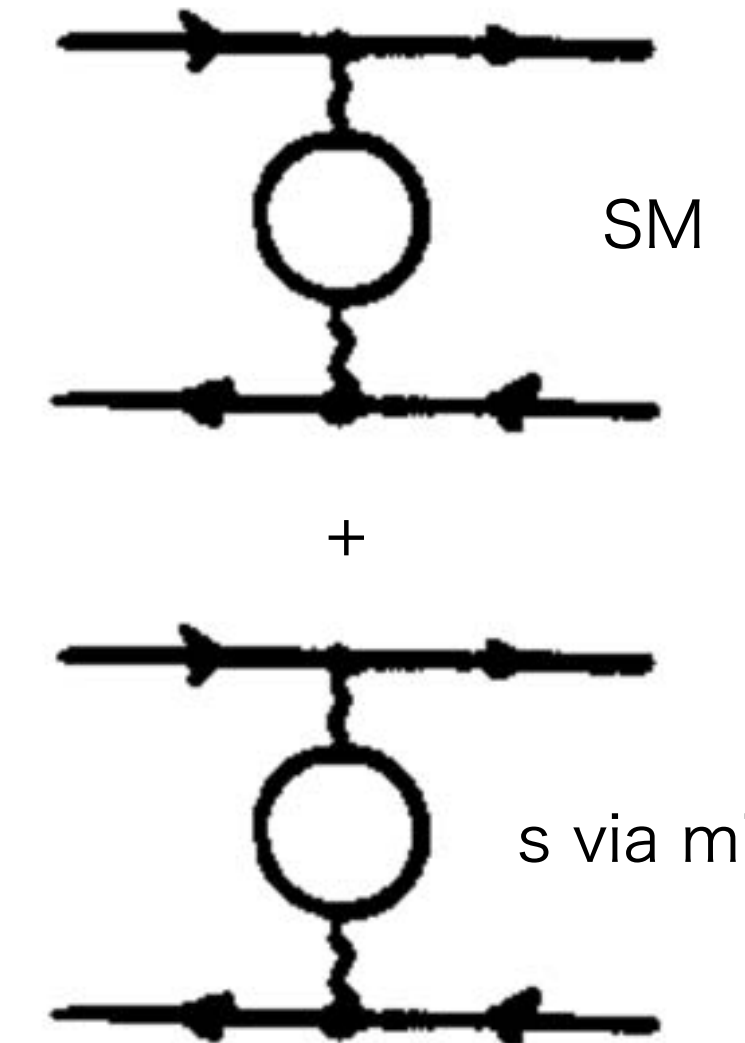
See CxSM Barger et al, 0811.0393

$$V = -m_{\Phi}^2 |\Phi|^2 + \lambda |\Phi|^4 + \lambda_P |\Phi|^2 |H|^2 + \lambda_H |H|^4 - \mu_H^2 |H|^2.$$

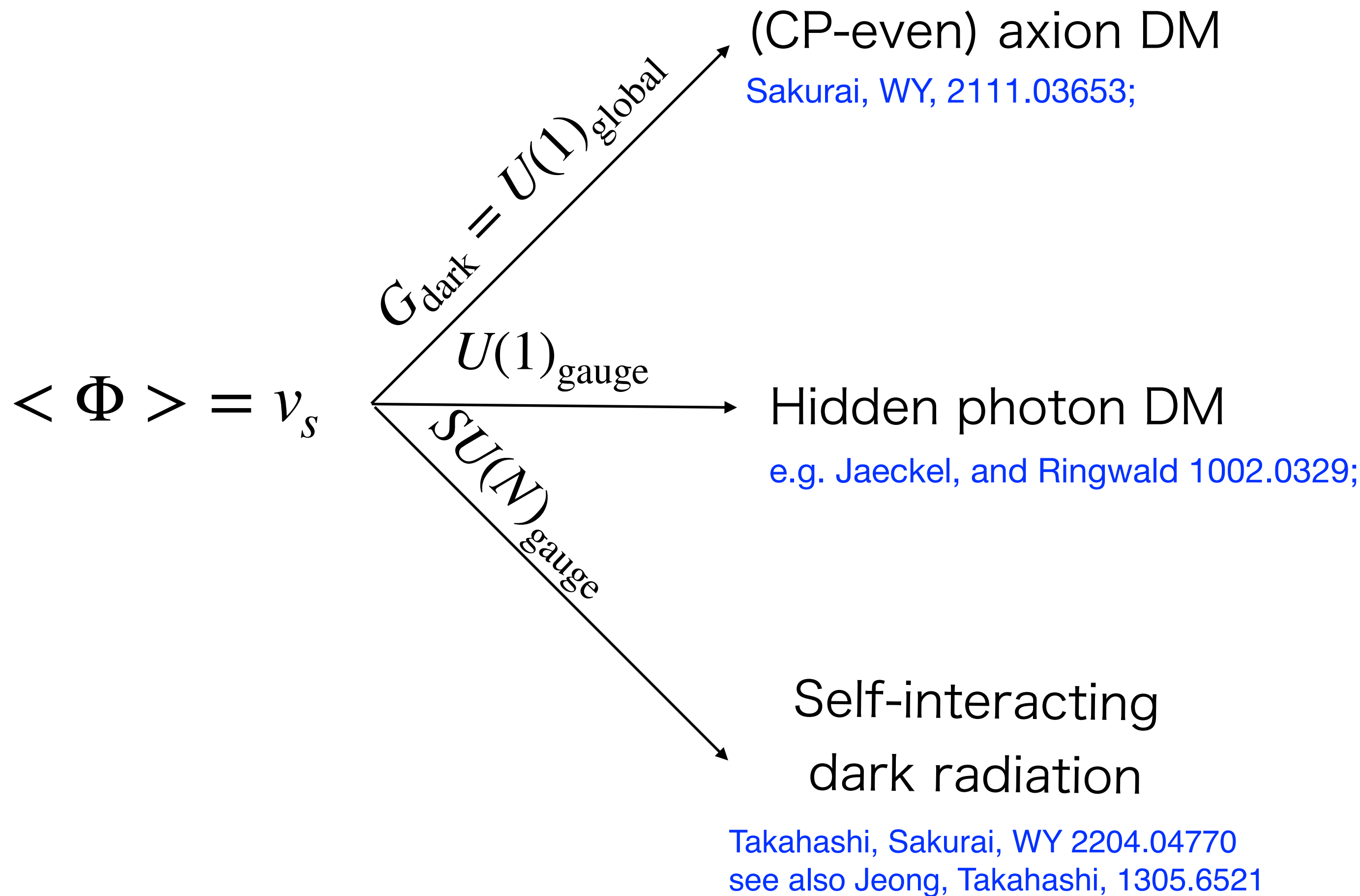
いずれもh-s mixingが予言。

$$h \text{ --- } \times \text{ --- } s$$

$\propto \lambda_P v_{EW} v_s$



$\Delta r \supset$



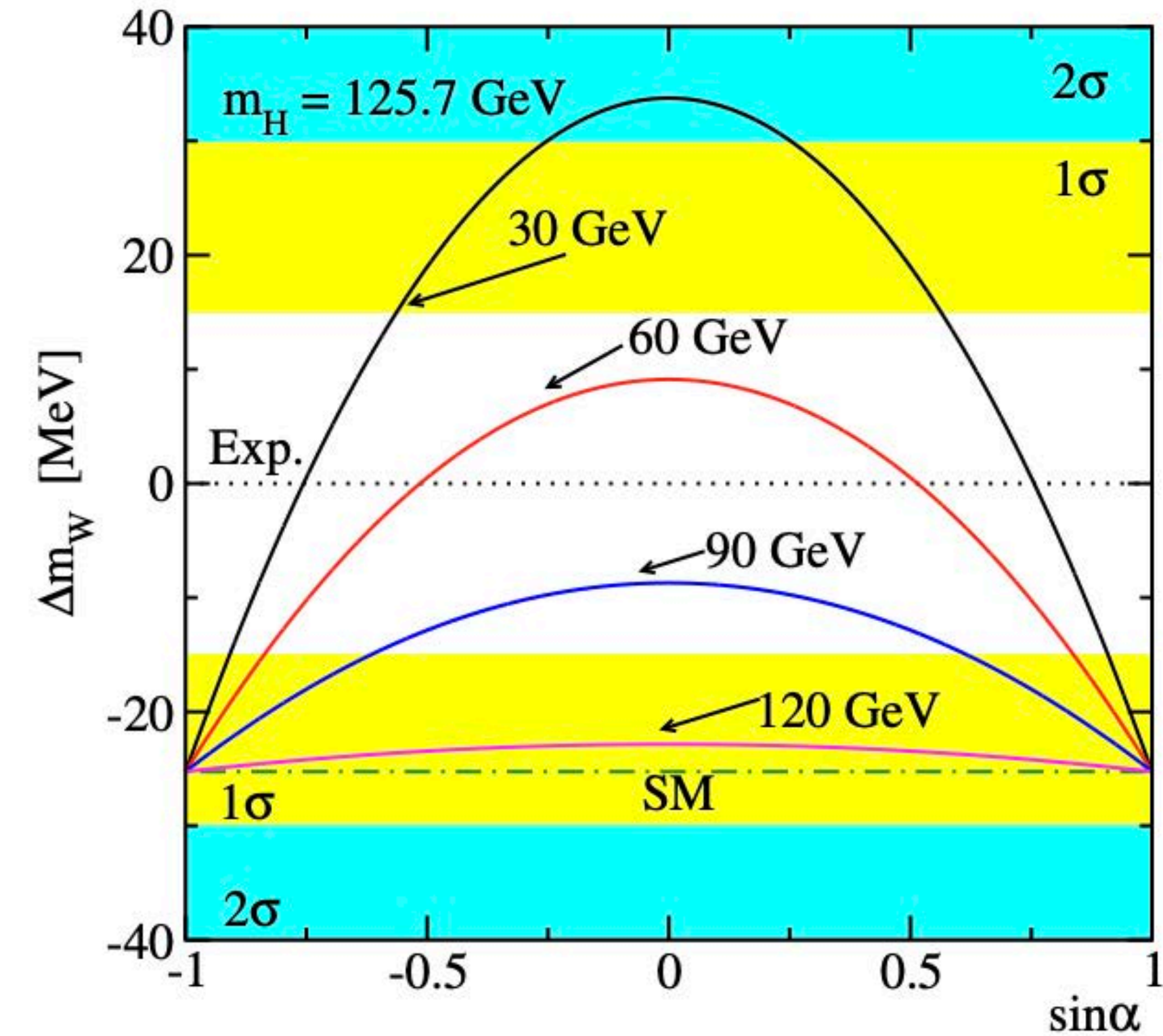
A good news?

CDF collaboration, Science 376, 170–176 (2022)

(26, 27). An example of a nonsupersymmetric SM extension is a modified Higgs sector that includes an additional scalar field with no SM gauge interactions, which predicts an M_W shift of up to ~ 100 MeV (17), depending on the mass of the additional scalar particle and its interaction with the SM Higgs boson. A light (heavy) additional scalar particle would induce a positive (negative) M_W shift. Similar but smaller shifts of 20 to 40 MeV have been calculated in an extension that contains a second Higgs-like field with the same gauge charges as the SM Higgs field (18). Implications of very weakly interacting new particles such as “dark

$$V \simeq m_{\text{mix}}^2 sh + m_1^2 \frac{s^2}{2} + m_2^2 \frac{h^2}{2}, \cos \alpha \equiv m_{\text{mix}}^2 / |m_1^2 - m_2^2|$$

[ref 17] L'opez-Val, Robens 1406.1043

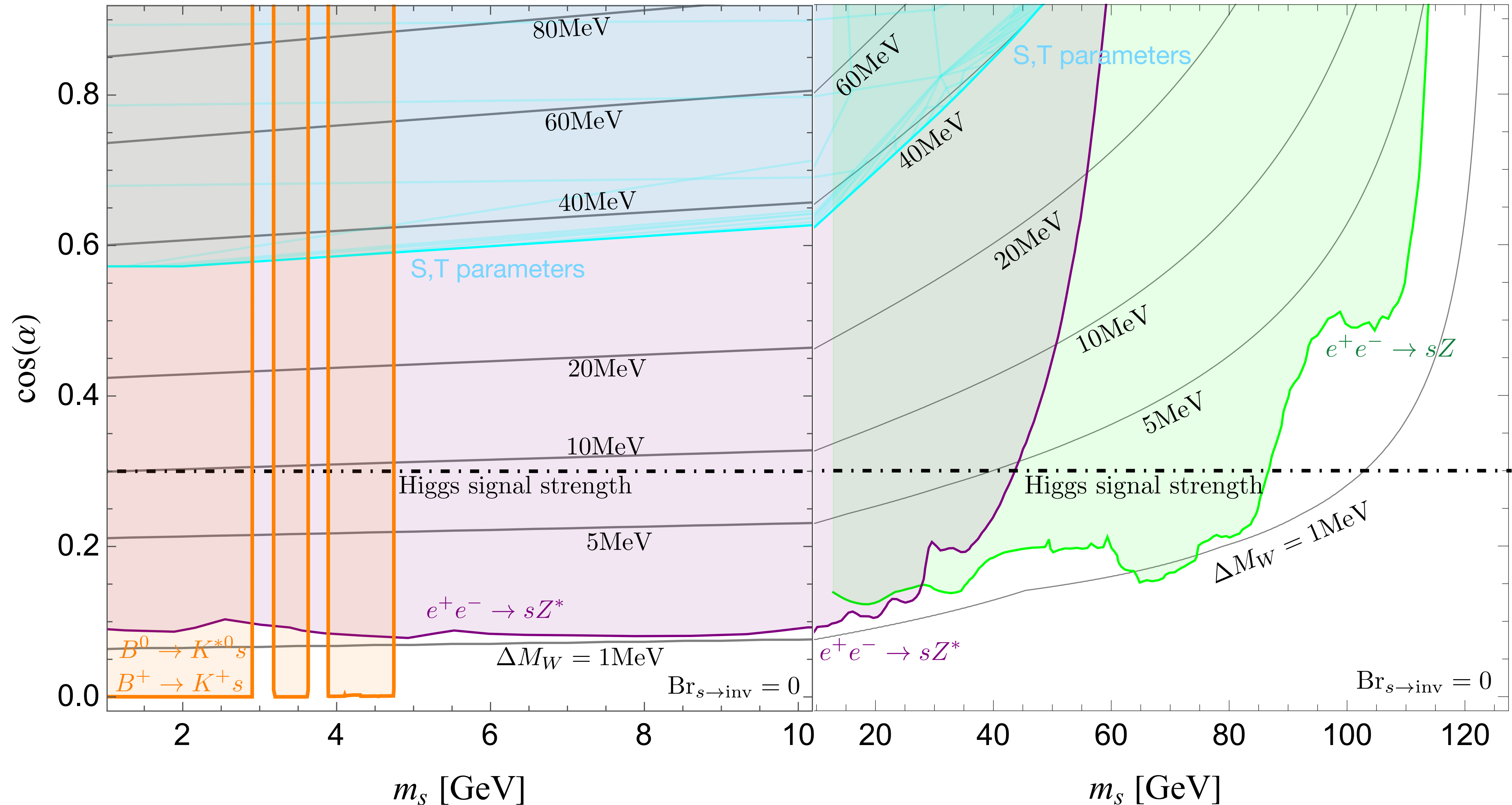


Compatibility with the LHC signal strength measurements requires $|\sin \alpha| > 0.91$.

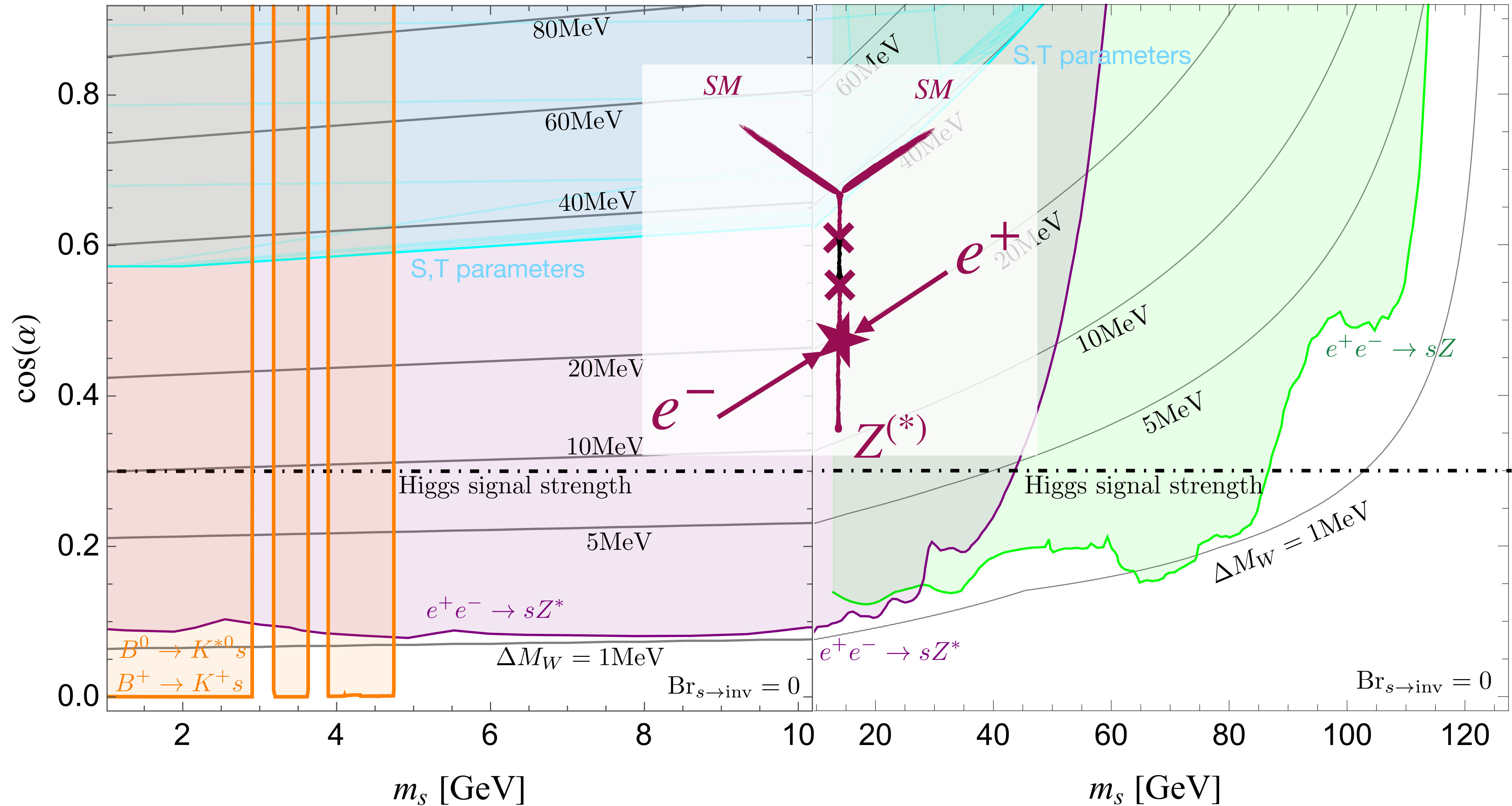
**2. Singlet scalar extension
cannot explain the CDF-II
result at all.**

Sakurai, Takahashi, WY 2204.04770

W-boson mass shift is at most **2 MeV** because of $e + \bar{e} \rightarrow sZ^{(*)} \rightarrow SMs$ at LEP and signal strength constraints.

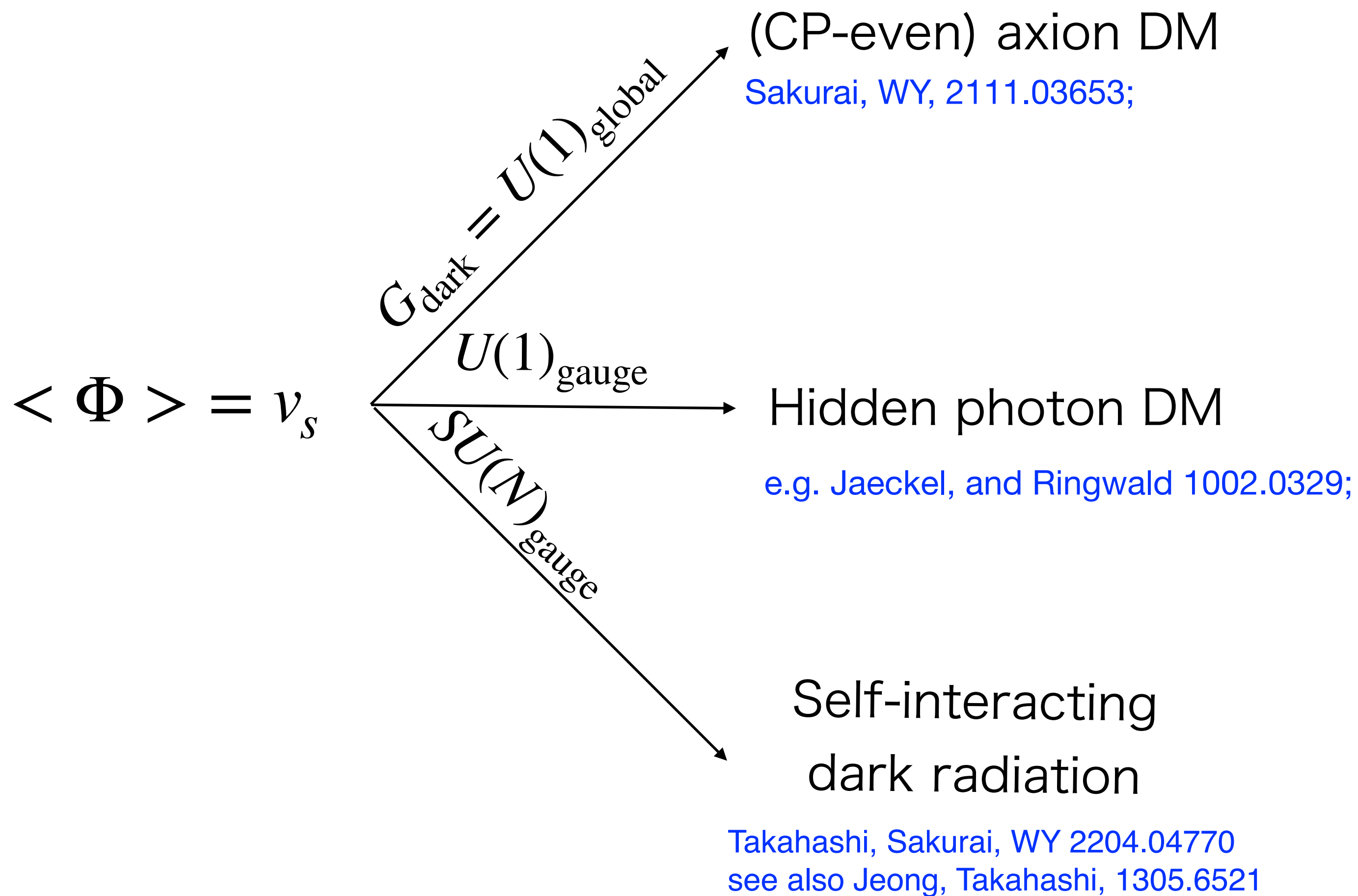


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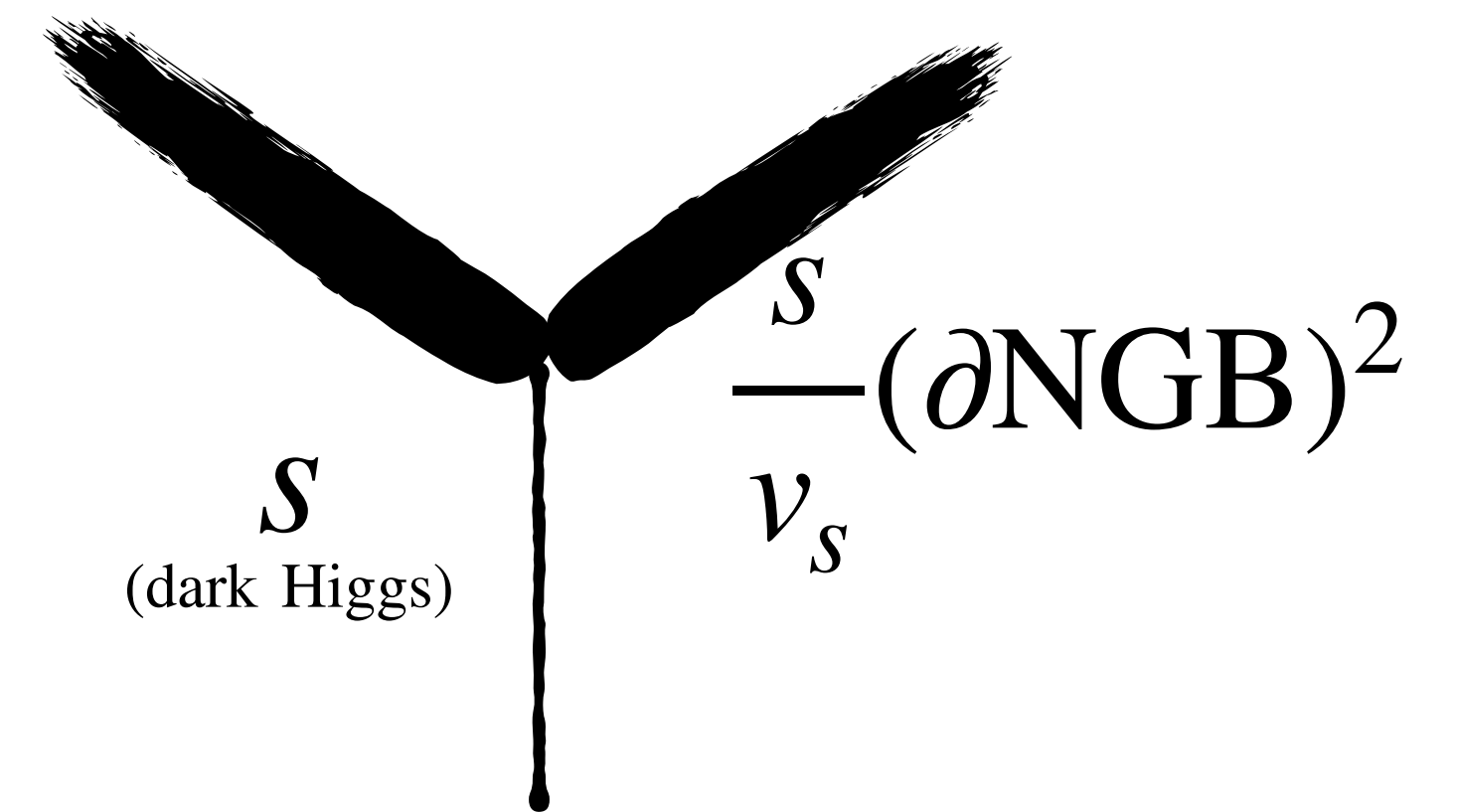


s invisible decay $\not\to$ prediction.

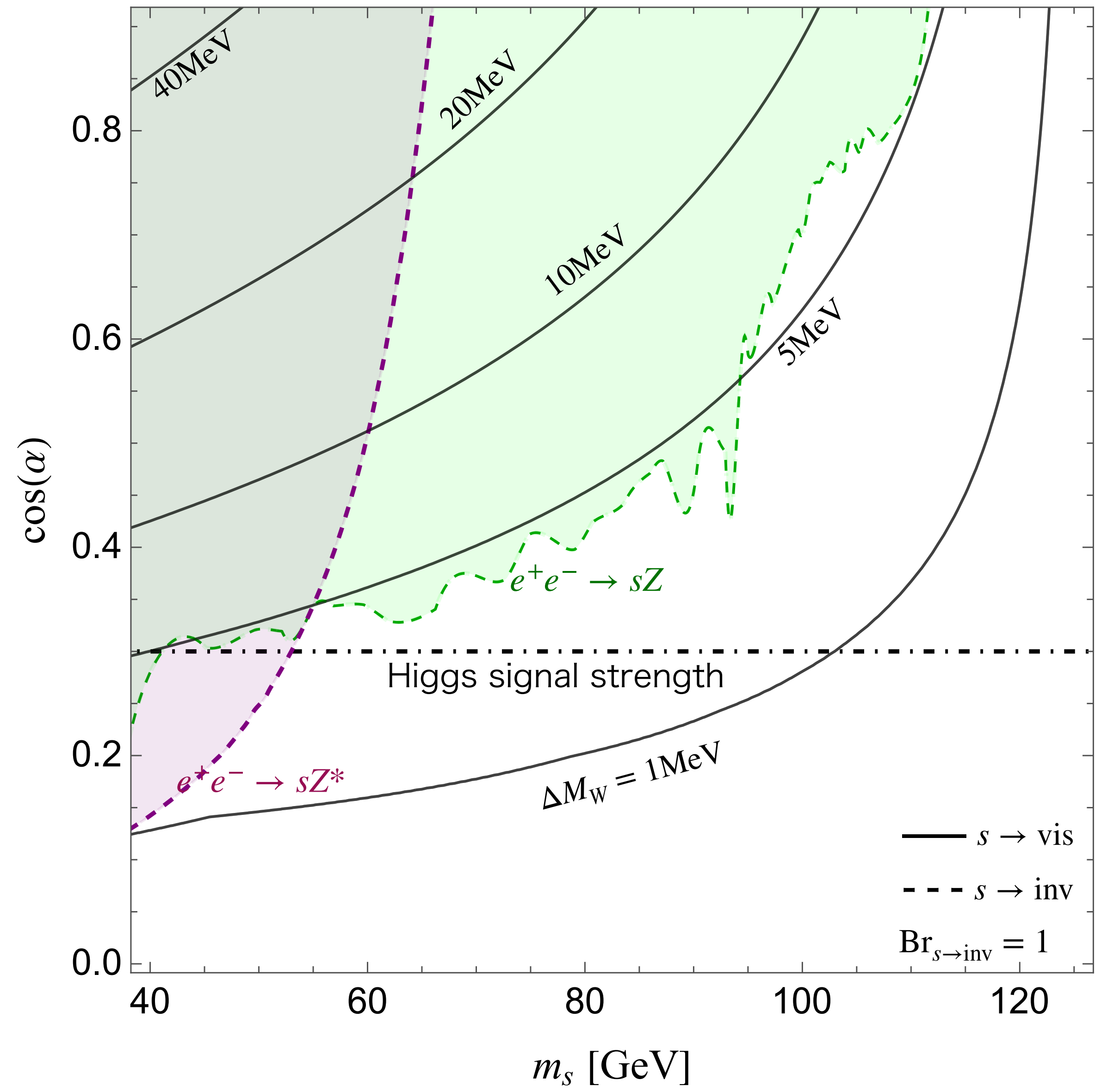
$$V = -m_{\Phi}^2 |\Phi|^2 + \lambda |\Phi|^4 + \lambda_P |\Phi|^2 |H|^2 + \lambda_H |H|^4 - \mu_H^2 |H|^2.$$



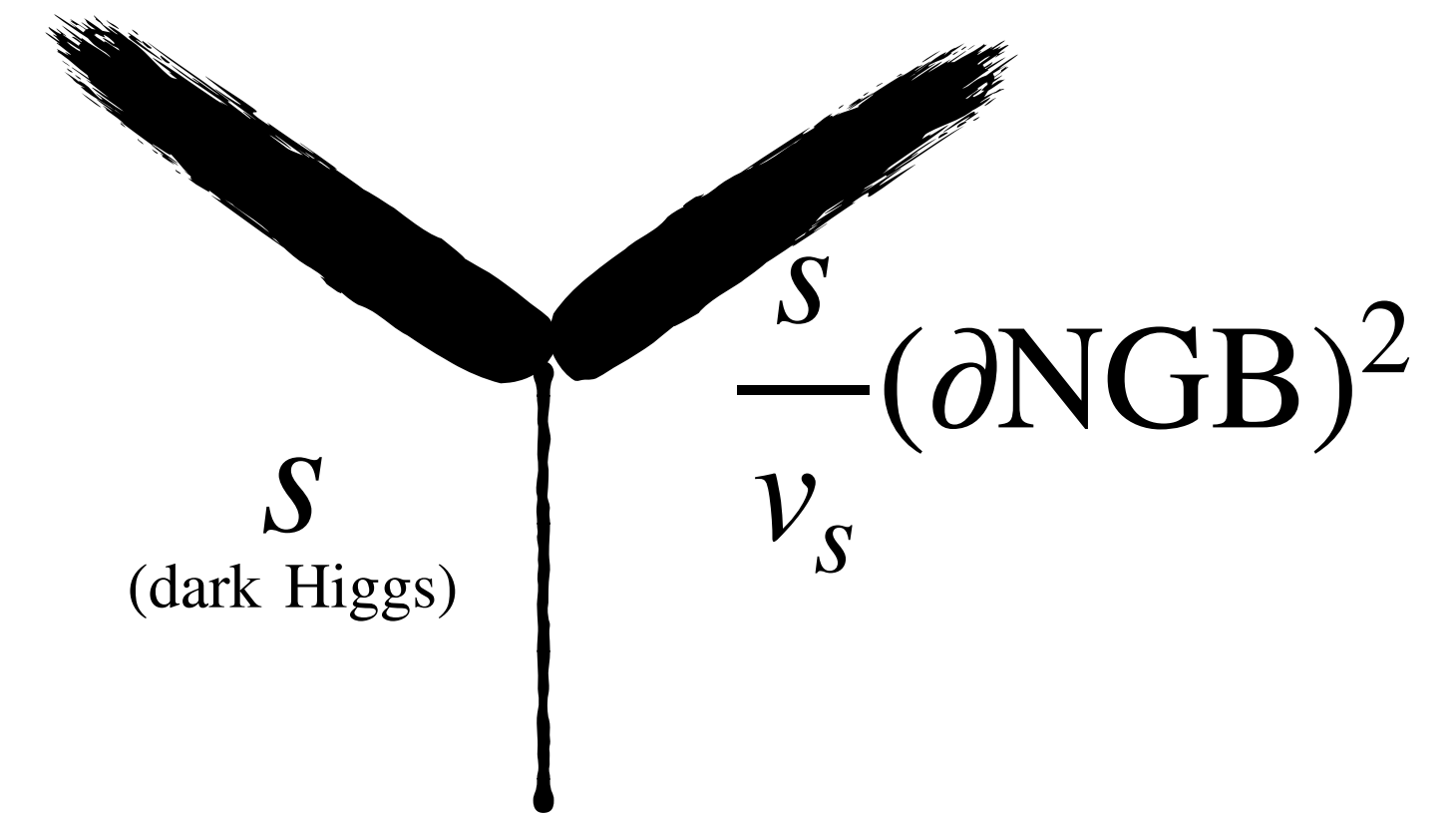
(would-be) NGB $\times 2$



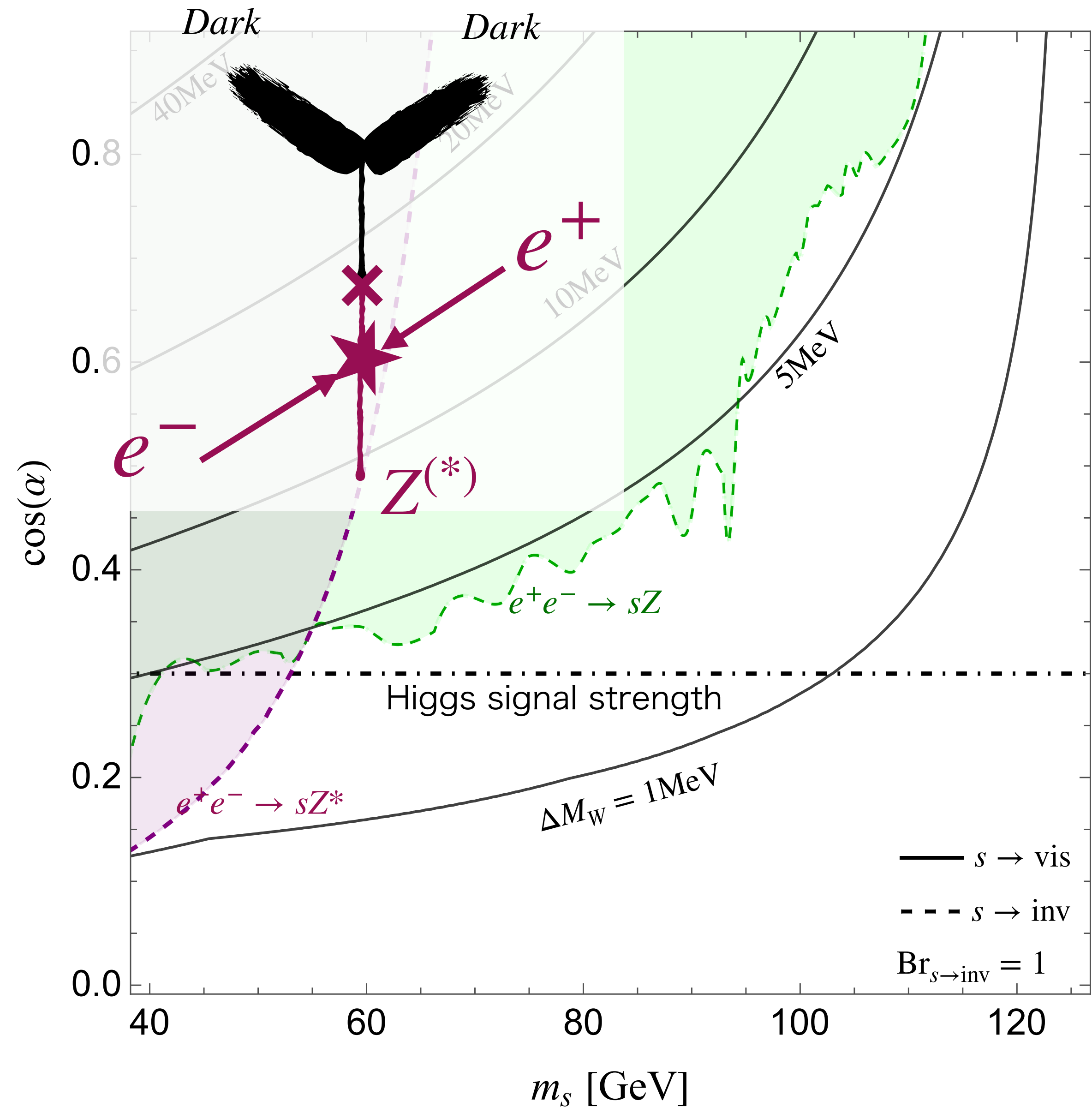
s may decay invisibly.



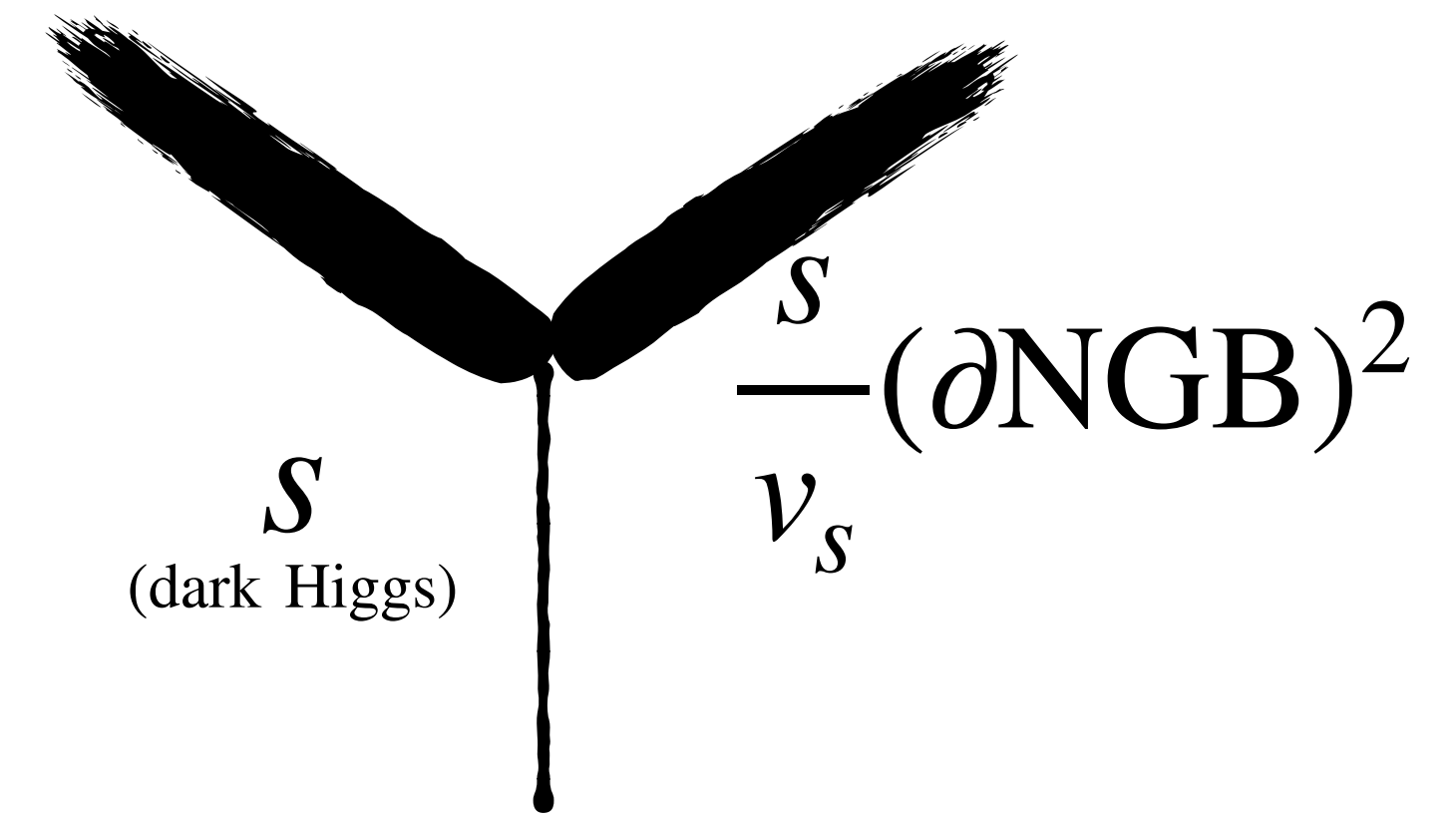
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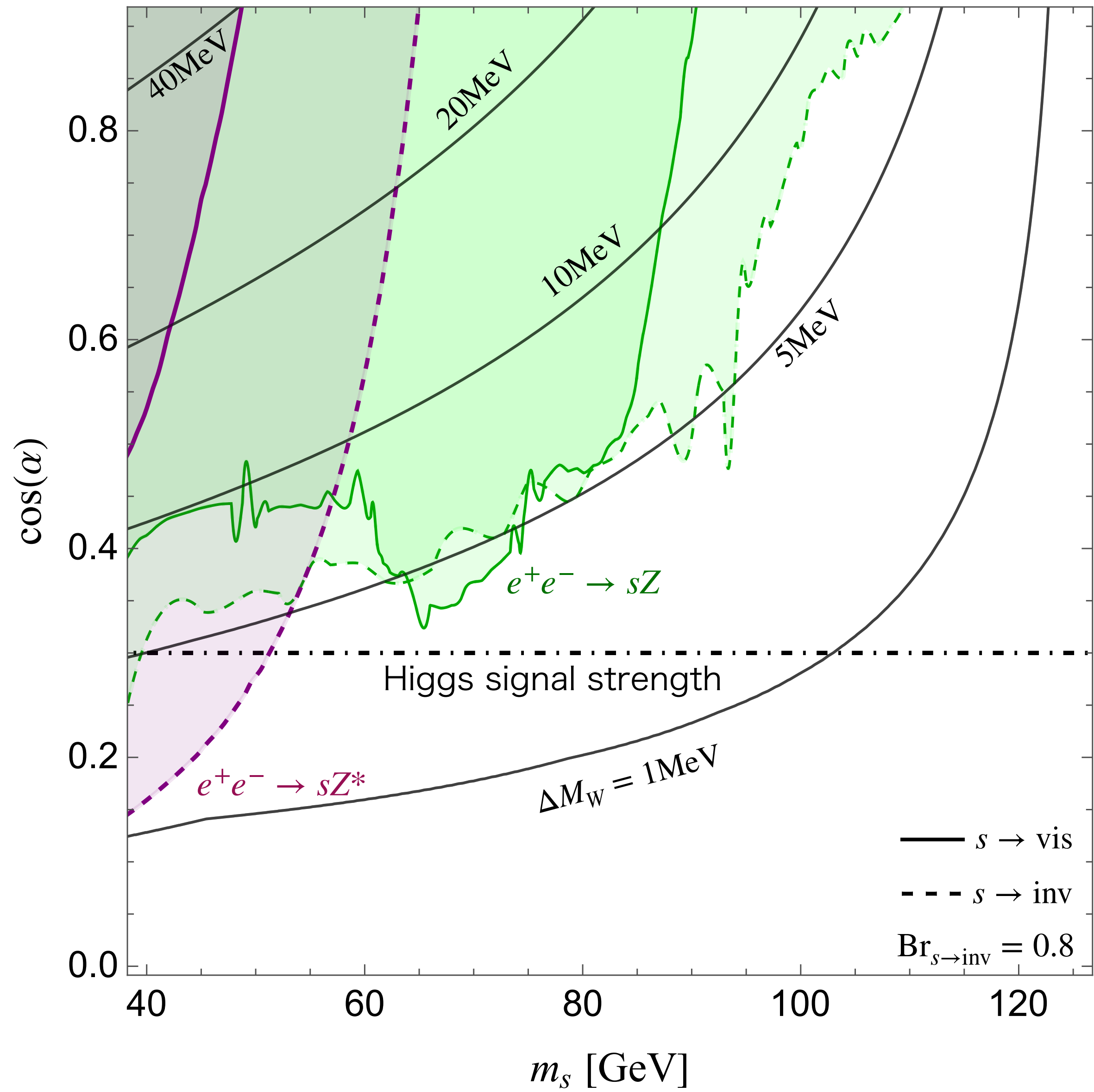
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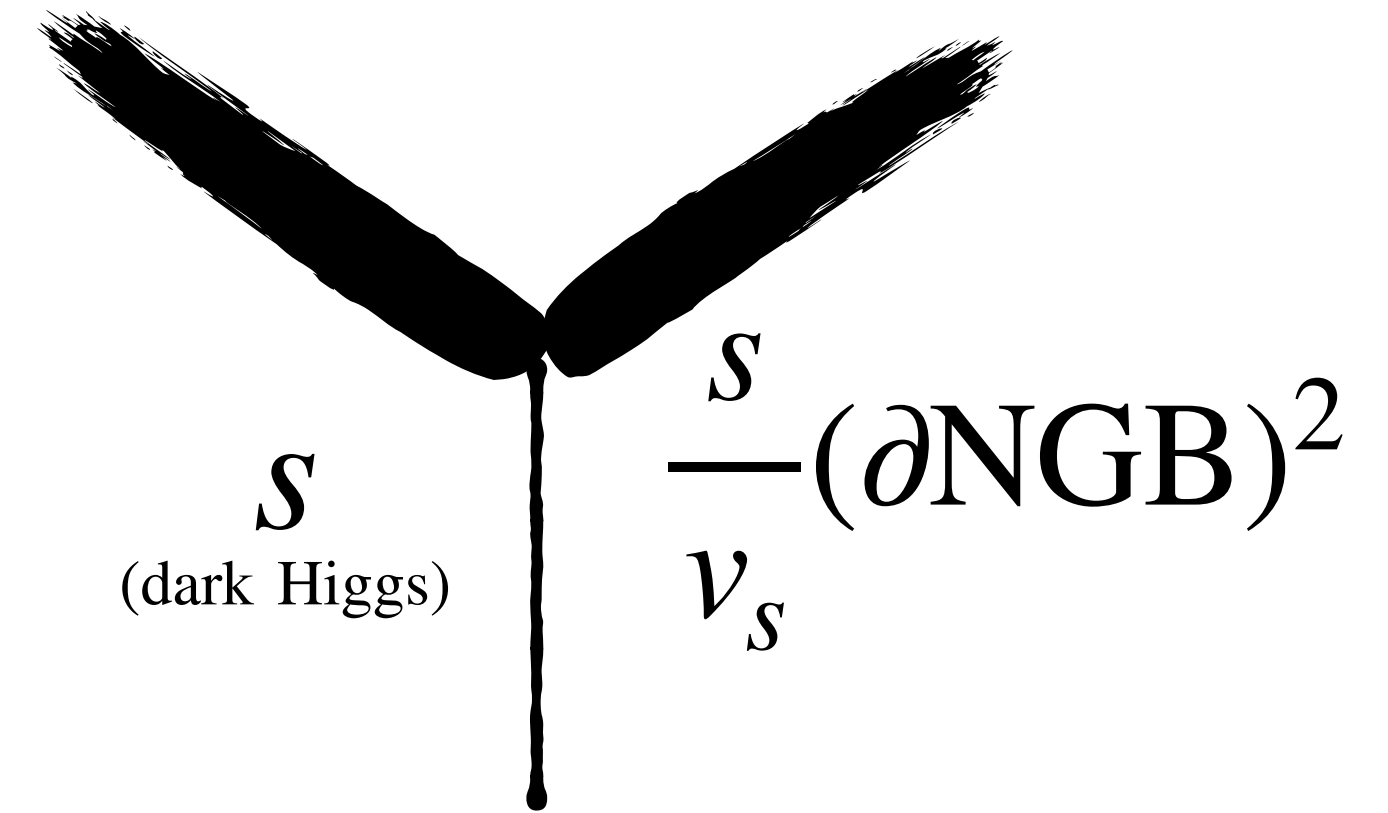
(would-be) NGB $\times 2$



s may both decay visibly and invisibly.

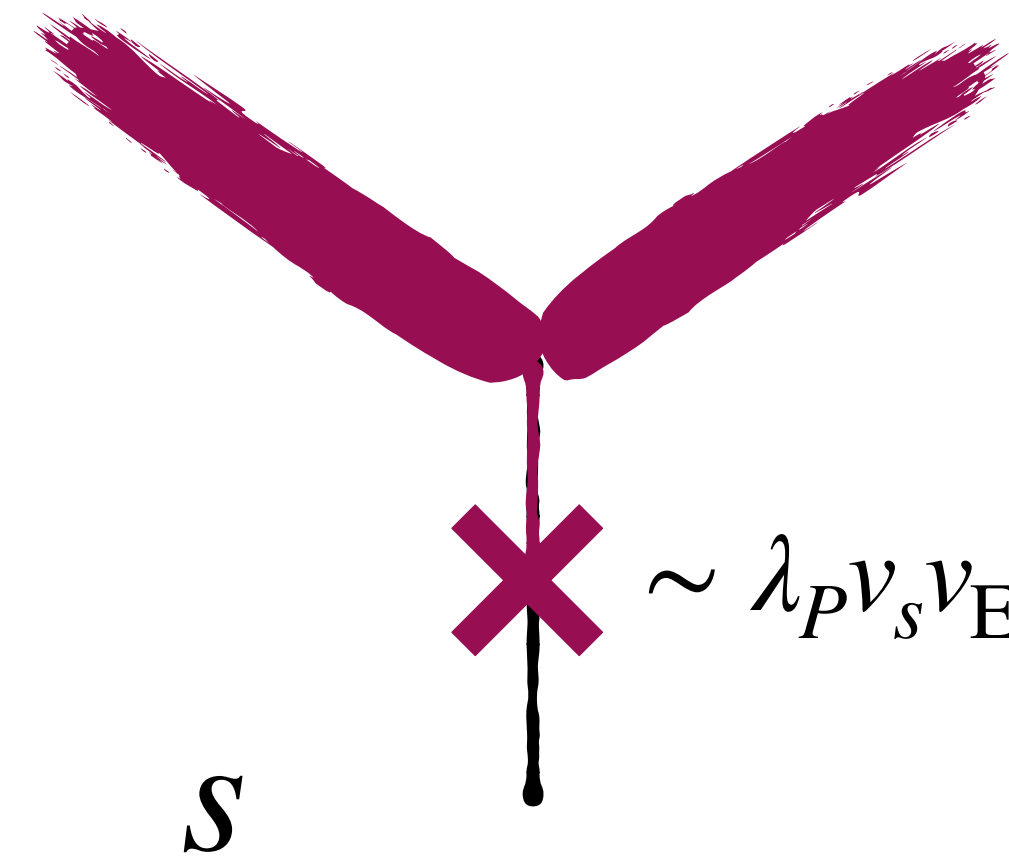


(would-be) NGB $\times 2$

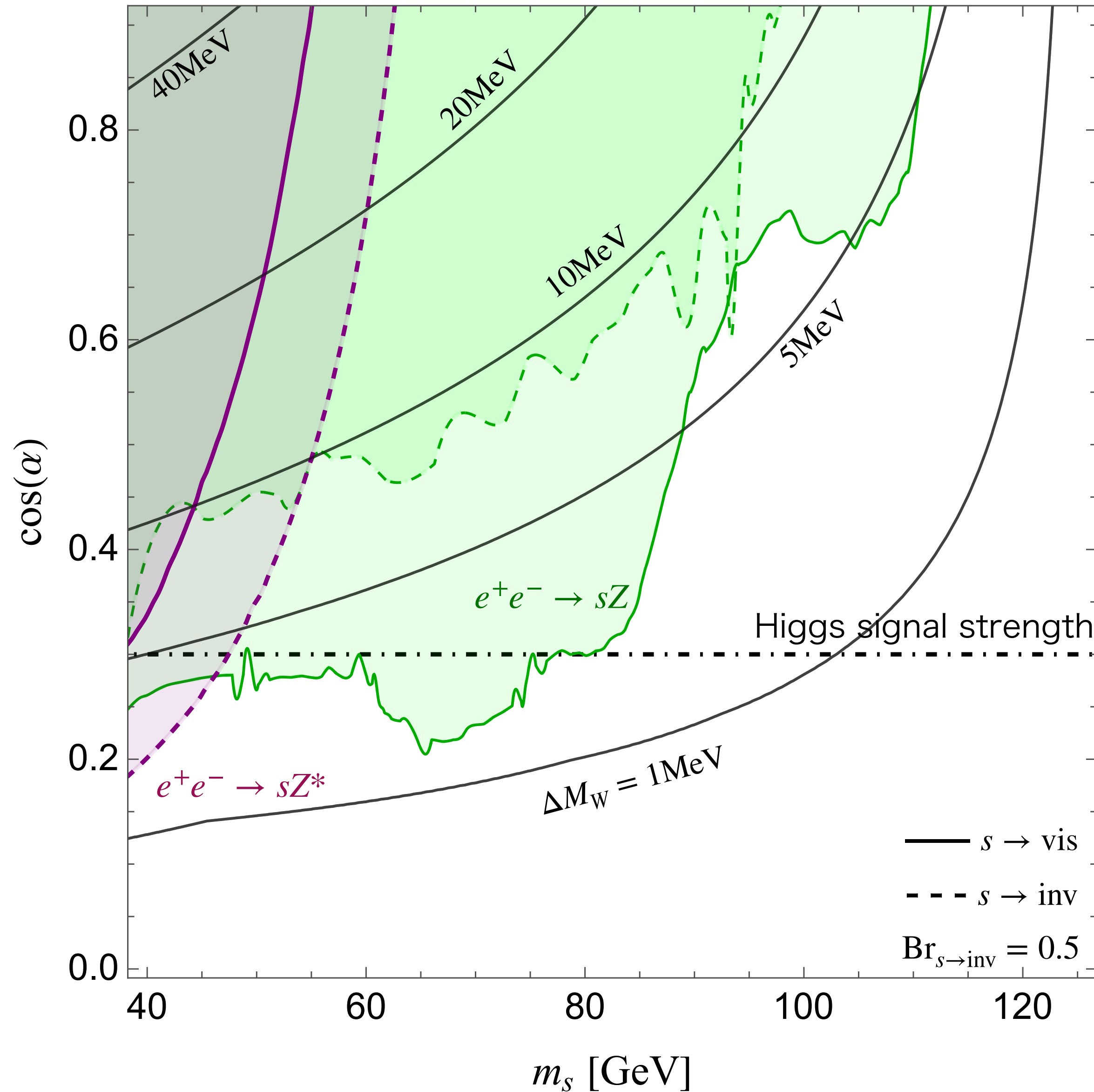


SM

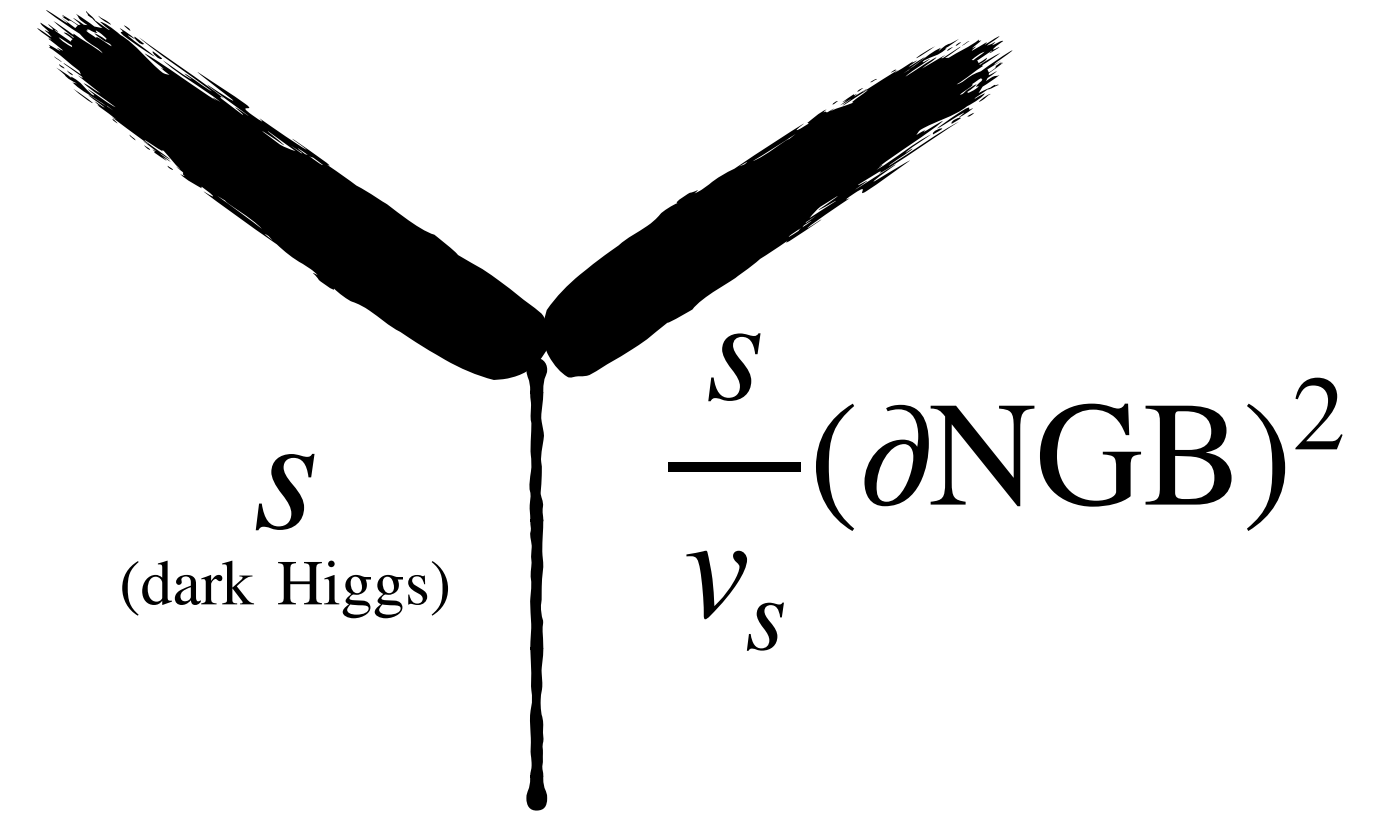
SM



s may both decay visibly and invisibly. m_W shift is at most 4-5 MeV.

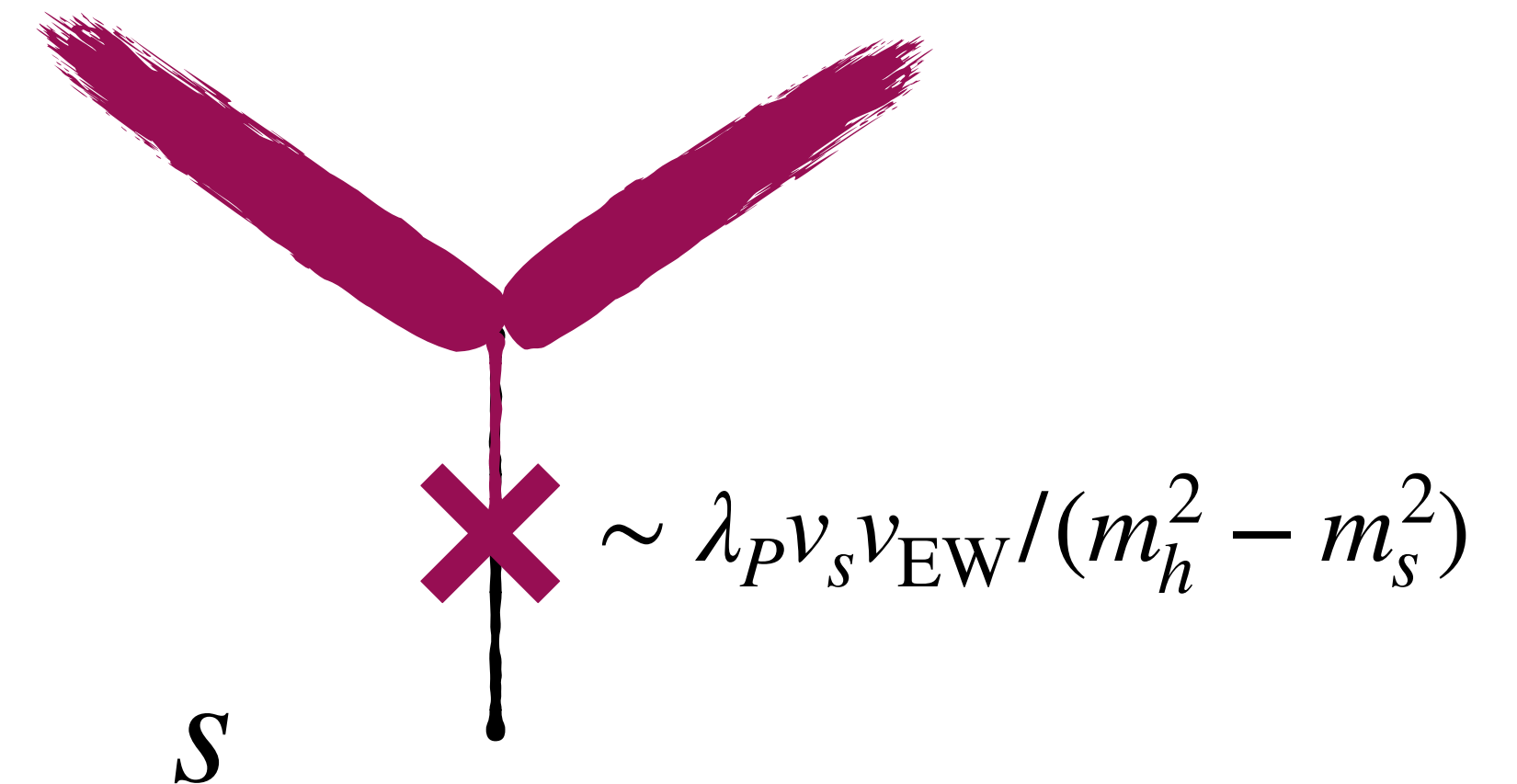


(would-be) NGB $\times 2$



SM

SM



3. Conclusions.

Sakurai, Takahashi, WY 2204.04770

- Singlet scalar extensionsでは余分な仮定なしに、フレーバー、CP的に安全でかつ、陽子崩壊は起こらない。標準模型の歴史的な成功例を犠牲にするが、暗黒物質と関連し、広く応用できる。
- Wボソン質量がどれほどズれるかははっきり指摘されてなかった。
- $\Delta M_w \lesssim 2$ MeV for s visible decay, and
 $\Delta M_w < 4 - 5$ MeV for s both visible and invisible decay
を指摘した。
- Singlet scalar extension cannot explain the CDF-II result at all.

backup

3. Conclusions.

- Singlet scalar extension cannot explain the CDF-II result at all. $\Delta M_w < 2 \text{ MeV}$ for s visible decay.

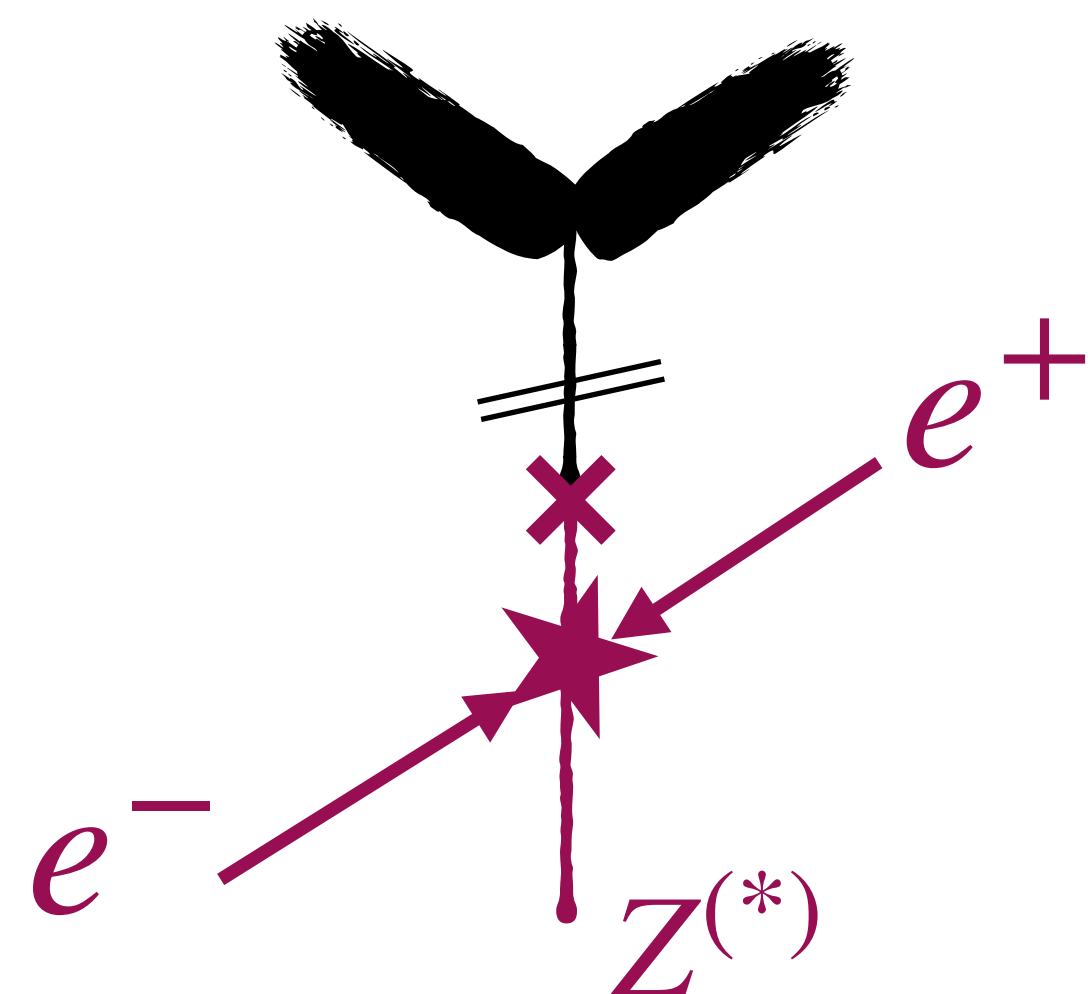
$\Delta M_w < 4 - 5 \text{ MeV}$ for s both visible and invisible decay.

- It may explain the slight preference without the CDF data.

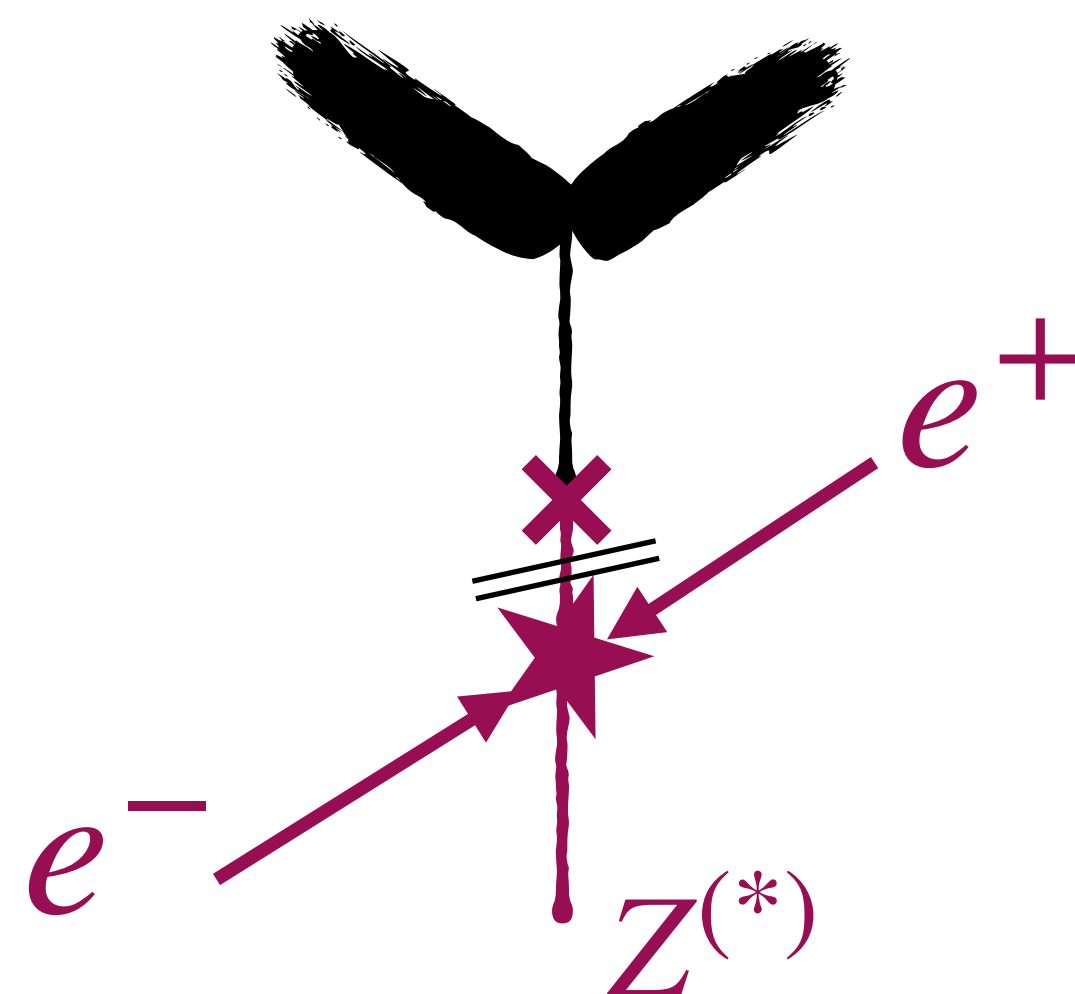
ダークヒッグスインビジブル崩壊

125GeVヒッグスインビジブル崩壊

Dark Dark



Dark Dark



ヒッグスインビジブル崩壊が
 $\Delta M_w > (1 - 2) \text{ MeV}$ ならば
ILCで必ずチェックできる。

Crazy注意!

2002-2006年 or 2009-2011年にHiggs が軽くなった説

LEP 1989-2000 1997-2000 data taking

D0 1983-2011, 2006 to 2009 data taking.

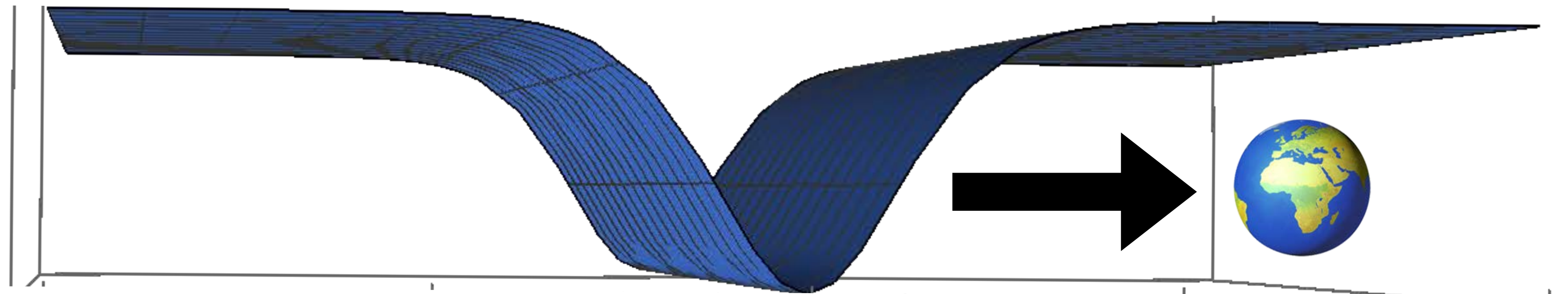
CDF II 1983-2011. 2002 and 2011 data taking

Tevatron Higgs search 2001-2011 <https://arxiv.org/pdf/1209.1586.pdf>

$$V = f(\phi) V_{SM}[|H|^2]$$

$f(\phi[x])$

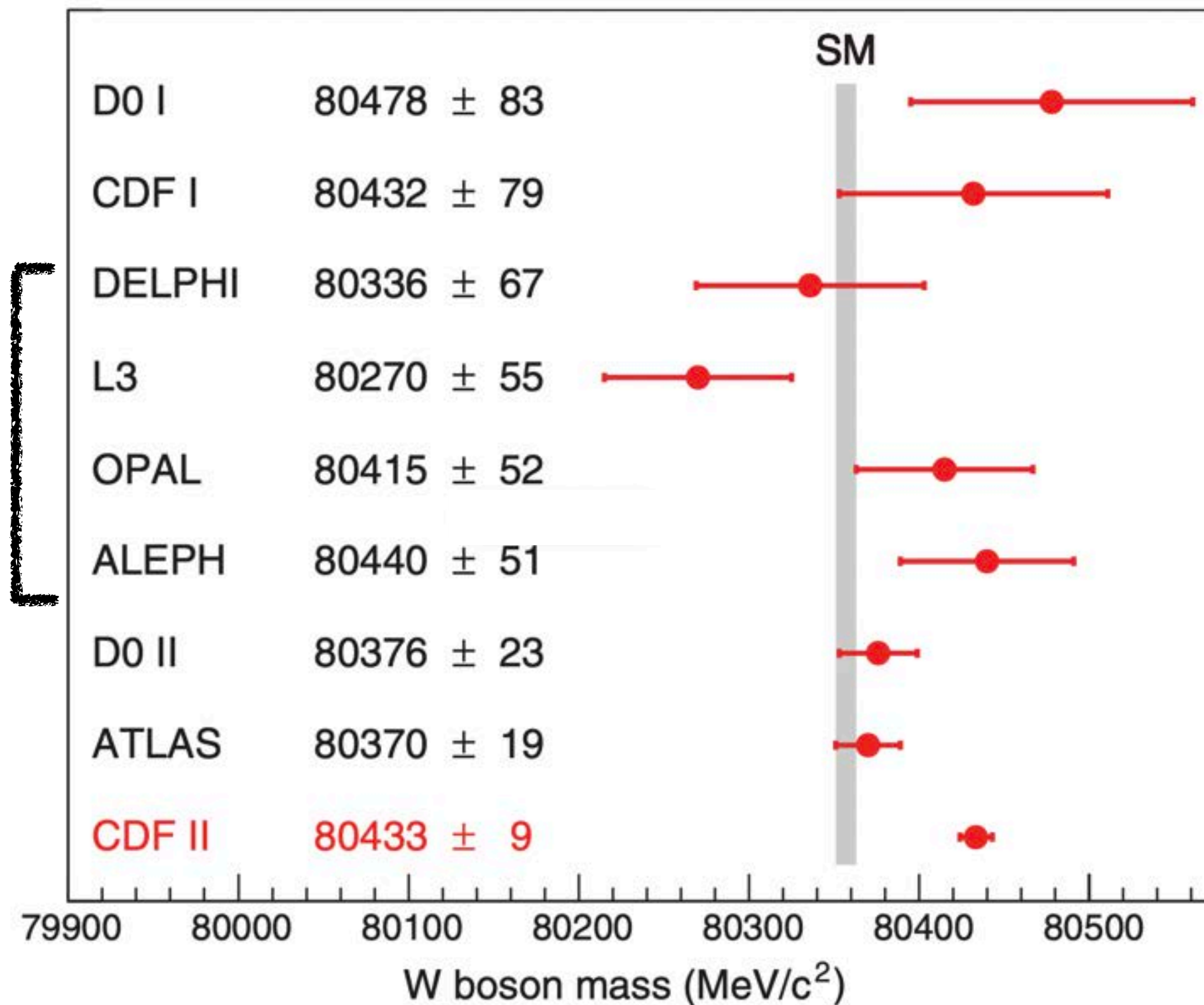
<4光年



CDF IIのWボソン質量 M_W の測定結果 (Kodai's slide)

[CDF Collaboration et al., Science 376, 170–176 (2022)]

LEP combined result
[Phys.Rept. 532 (2013) 119]
 80376 ± 33 [MeV]



Tevatronのdetector:
D0, CDF

LEPのdetector:
DELPHI, L3, OPAL,
ALEPH

SMの理論予言: 80357 ± 6 [MeV]
[PDG2021]

← EW parameterの global fit
により得られた値

[J. Erler, M. Sch

CDF IIの結果: 標準理論予言から7σ (~80 MeV) のずれ

Possible applications.

Degenerate scalar scenario

When dark higgs masses are similar, some parameter space for WIMP and EWPT will open.

WIMP DM, Abe, Cho, Mawatari, 2101.04887

EWPT, Cho, Idegawa, Senaha 2105.11830

Degenerate scalars with $\Delta m \gtrsim 0.1\text{GeV}$

can be distinguished at ILC Abe, Cho, Mawatari, 2101.04887

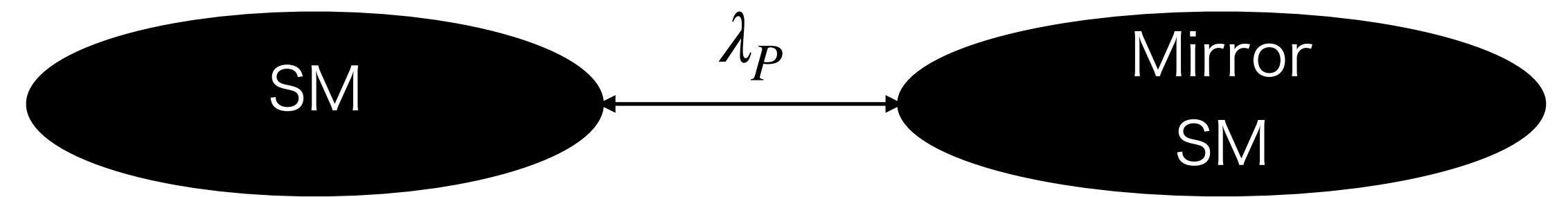
In the fuzzy Higgs region, the degenerate scalars cannot be distinguished. But it is probed by the Higgs invisible decay also at the ILC. Sakurai WY 2204.01739

Strongly coupled dark sector

Go beyond perturbative unitarity. If $\Gamma_{s \rightarrow \text{dark}}$ can be arbitrarily large, fuzzy Higgs boson is realized with arbitrary $|m_1 - m_2|$ and thus generically realized.

Exact Z_2 mirror symmetry

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{mirror SM}} - \lambda_P |H|^2 |H_{\text{mirror}}|^2$$



Relevant to fine-tuning problems or lighter QCD axion. Hook 1802.10093

$$\Gamma_h = \Gamma_{\text{mirror}h} \sim 4\text{MeV}, m_h \simeq m_{\text{mirror}h} = 125.25\text{GeV}, \alpha = \pi/4.$$

In this case, the Higgs coupling deviation of $\kappa_X = \cos(2\alpha_{\text{eff}})$ can be also probed together with invisible decay.

Sakurai WY 2204.01739

Extention to dark scalar phenomena

Mixed axions with one component decay very fast has the other component stabilized.

3. CP-even ALP from generic CPV

Kodai Sakurai, WY 2111.03653

In the following I take for simplicity $\theta_{CP} = \theta_{CKM} = 0$, which does not change our conclusions.

If we do not impose CP symmetry in the dark sector,

$$V = -m_{\Phi}^2 |\Phi|^2 + \lambda |\Phi|^4 + \lambda_P |H|^2 |\Phi|^2 + \lambda_H |H|^4 - \mu_H^2 |H|^2.$$

Accidental discrete symmetry in dark global U(1) symmetric limit:

C_{dark} symmetry: SM fields do not transform, $\Phi(t, \vec{x}) \rightarrow \Phi^*(t, \vec{x})$

CP symmetry: SM fields transform as in the SM, $\Phi(t, \vec{x}) \rightarrow \Phi^*(t, -\vec{x})$.

If we do not impose CP symmetry in the dark sector,

Explicit breaking of dark $U(1)$ controlled by κ is

$$\delta V = \kappa \left(\sum_{j=1}^4 c_j m_{\Phi}^{4-j} \Phi^j + \sum_{j=1}^2 (\tilde{c}_j^H m_{\Phi}^{2-j} \Phi^j |H|^2 + \tilde{c}_j^{\Phi} m_{\Phi}^{2-j} \Phi^j |\Phi|^2) \right) + \text{h.c.}$$

$\arg c, \tilde{c} \neq 0$

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$$\arg c, \tilde{c} \neq 0$$

But $C_{\text{dark}} \cdot CP$ remains: $SM \rightarrow CP SM$,

$$\Phi(t, \vec{x}) \rightarrow \Phi(t, -\vec{x}) \quad (\text{a parity for dark Higgs}).$$

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$$\arg c, \tilde{c} \neq 0$$

$$C_{\text{dark}} \cdot CP: \text{SM} \rightarrow \text{CP SM}$$

$$\Phi(t, \vec{x}) \rightarrow \Phi(t, -\vec{x}), \text{ thus } a[t, \vec{x}] (\equiv -i \arg \Phi) \rightarrow a[t, -\vec{x}]$$

If we do not impose CP symmetry in the dark sector,

Explicit breaking of dark $U(1)$ controlled by κ is

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$\arg c, \tilde{c} \neq 0$

$$CP_{\text{EFT}} \equiv C_{\text{dark}} \cdot CP: \quad \text{SM} \rightarrow \text{CP SM}, \quad a[t, \vec{x}] \rightarrow a[t, -\vec{x}]$$

A simple UV completion of axion without imposing CP symmetry has accidental CP_{EFT} with **ALP being CP-even.**

Couplings of the CP-even ALP

$$V = -m_\Phi^2 |\Phi|^2 + \lambda |\Phi|^4 + \lambda_P |H|^2 |\Phi|^2 + \lambda_H |H|^4 - \mu_H^2 |H|^2.$$

$$\mathcal{L}_{\text{eff}} \sim \frac{\mathcal{O}_{SM}}{m_\Phi^{d_{\mathcal{O}_{SM}}}} (\partial a)^2$$

- Induced from U(1) symmetric part, and thus $C_{\text{dark}} \times CP$ symmetric.
- Non-renormalizable (dim 6 or 8).
i.e. very weak at low energy.

$$\delta V = \kappa \left(\sum_{j=1}^4 c_j m_\Phi^{4-j} \Phi^j + \sum_{j=1}^2 (\tilde{c}_j^H m_\Phi^{2-j} \Phi^j |H|^2 + \tilde{c}_j^\Phi m_\Phi^{2-j} \Phi^j |\Phi|^2) \right) + \text{h.c.}$$

$$\mathcal{L}_{\text{eff}} \sim \frac{m_a^2 m_h}{m_\Phi} a h, \theta_{ah} \sim \frac{m_a^2}{m_h m_\Phi}$$

- Induced from U(1) breaking part.
- At $\kappa \rightarrow 0$, (i.e. $m_a^2 \rightarrow 0$), it vanishes, i.e.
amplitude $\propto m_a^2$
- Renormalizable, dominant at low energy.

3. Phenomenology of CP-even ALP

- Probing CP-even ALP in Higgs factory
- CP-even ALP DM

CP-even ALP can be naturally produced via Higgs boson decay

$$V = -m_{\Phi}^2 |\Phi|^2 + \lambda |\Phi|^4 + \lambda_P |\Phi|^2 |H|^2 + \lambda_H |H|^4 - \mu_H^2 |H|^2.$$

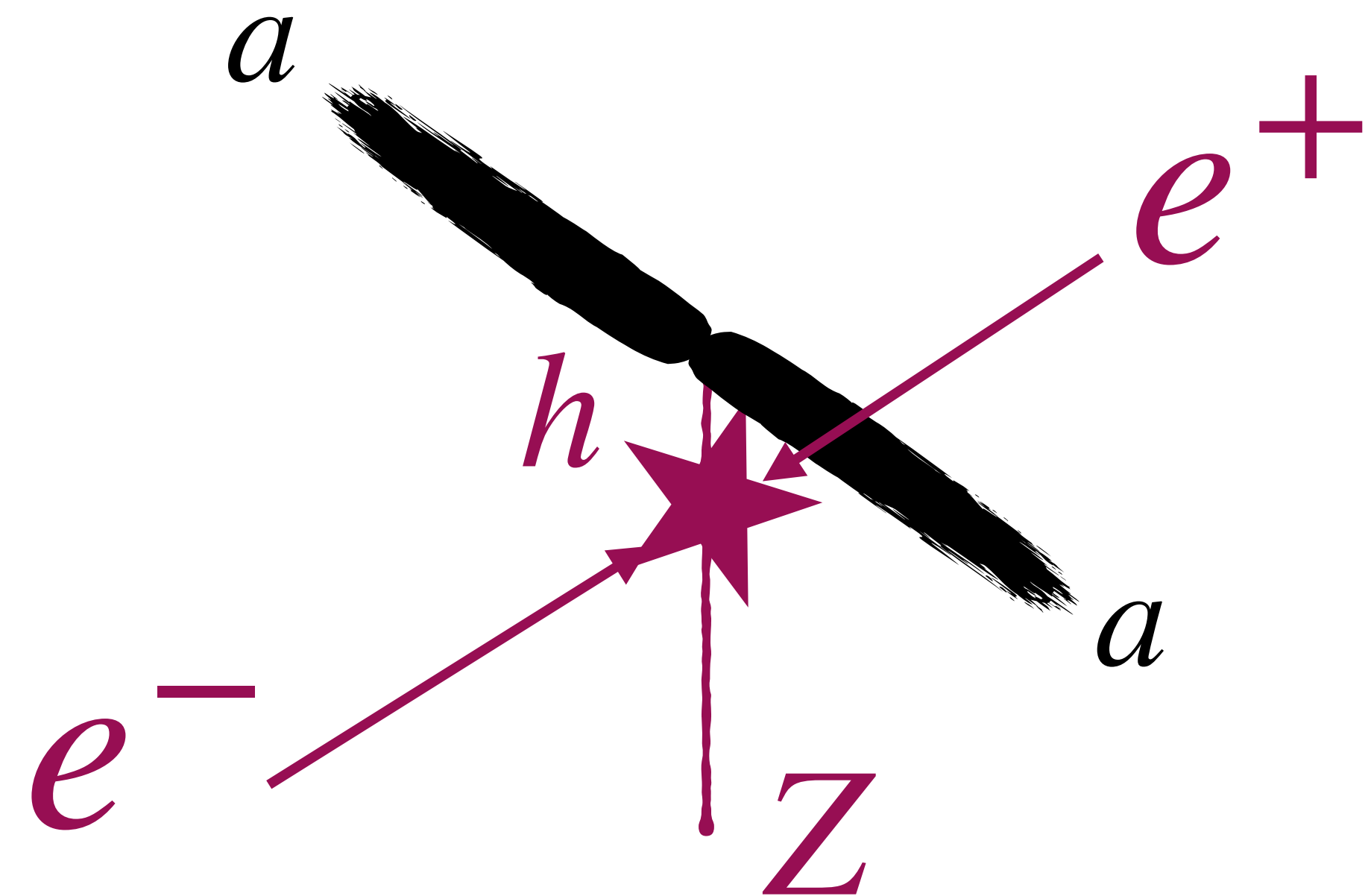
Higgs portal coupling

Φ : U(1) Higgs field
 H : SM Higgs doublet

$$\mathcal{L}_{\text{eff}} \sim \frac{\sqrt{2}v}{\Lambda_H^2} h(\partial a)^2 \quad \frac{1}{\Lambda_H^2} \equiv -\frac{\lambda_P}{m_s^2 - m_h^2}.$$

$$\Gamma_{h \rightarrow aa} \simeq \frac{1}{16\pi} \frac{v^2 m_h^3}{\Lambda_H^4}.$$

$$\text{Br}_{h \rightarrow aa} = 2\% \left(\frac{2\text{TeV}}{\Lambda_H} \right)^4$$



Couplings of the CP-even ALP

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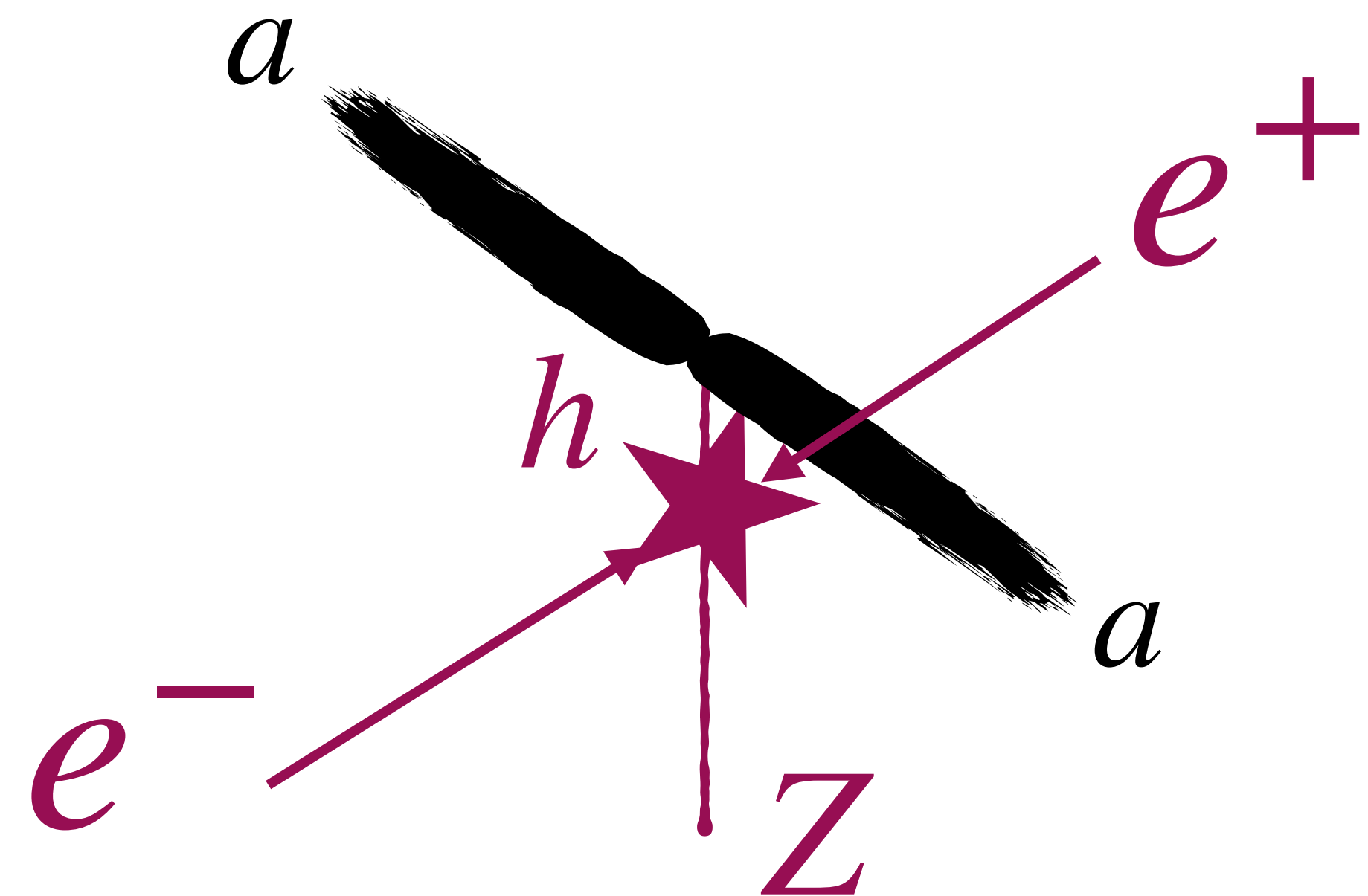
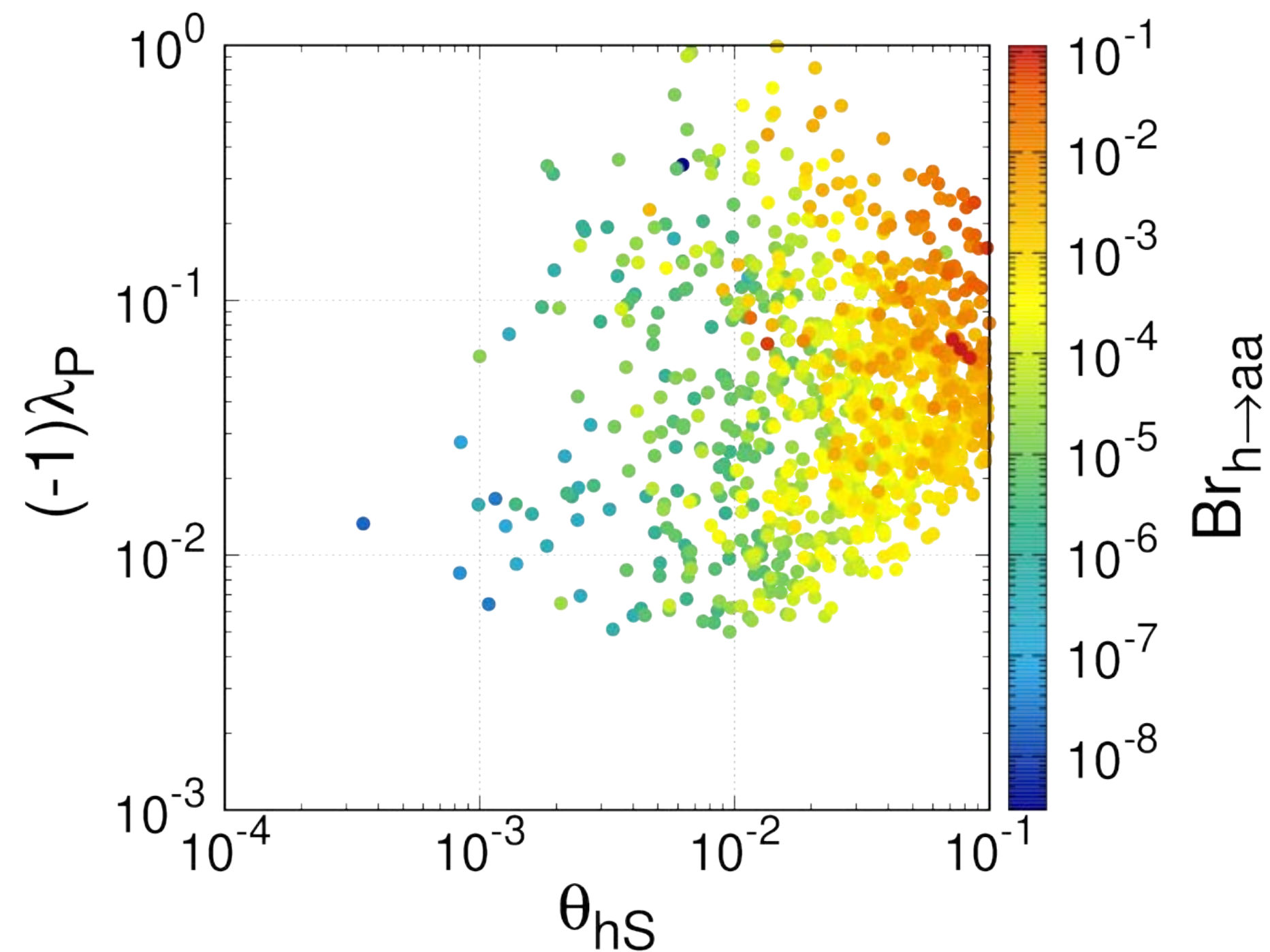
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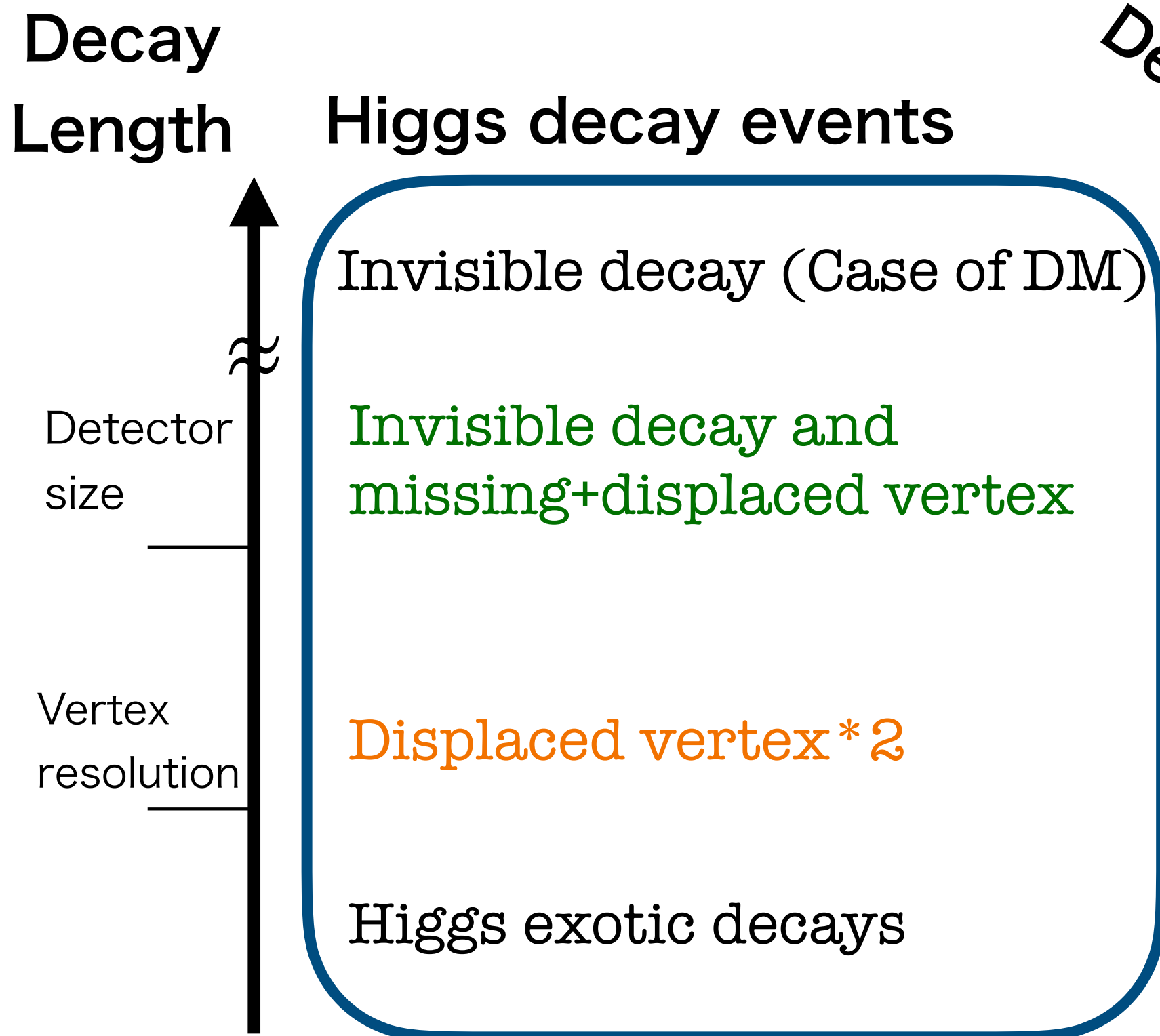
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 H : SM Higgs doublet

Production from Higgs decay



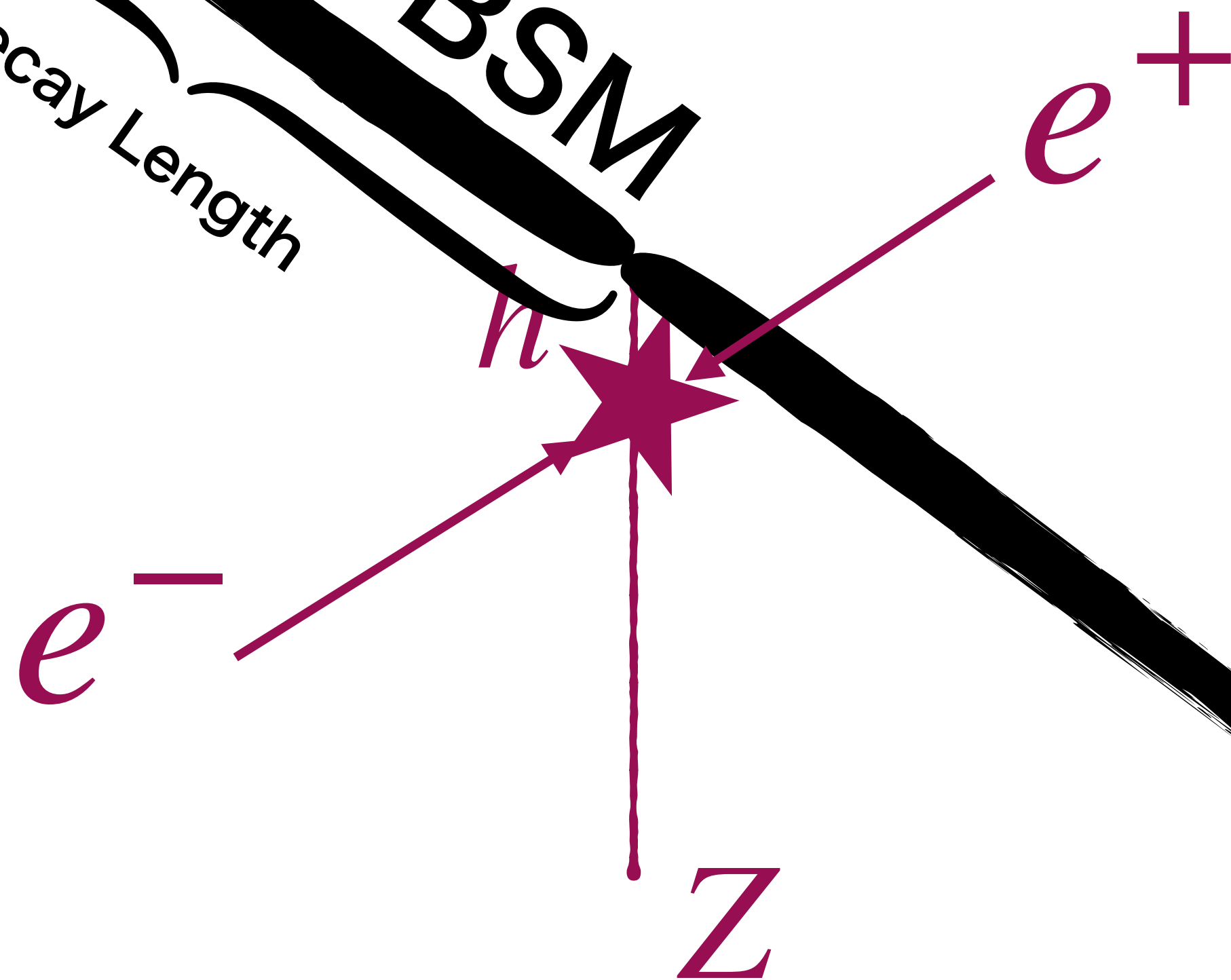
After the production, it is long-lived if light.

$$\delta V = \kappa \left[\sum_{j=1}^4 c_j m_{\Phi}^{4-j} \Phi^j + \sum_{j=1}^2 (\tilde{c}_j^H m_{\Phi}^{2-j} \Phi^j |H|^2 + \dots) \right]$$



Light BSM

Decay Length



$$\mathcal{L}_{\text{eff}} \sim \frac{m_a^2 m_h}{m_{\Phi}} ah, \theta_{ah} \sim \frac{m_a^2}{m_h m_{\Phi}}$$

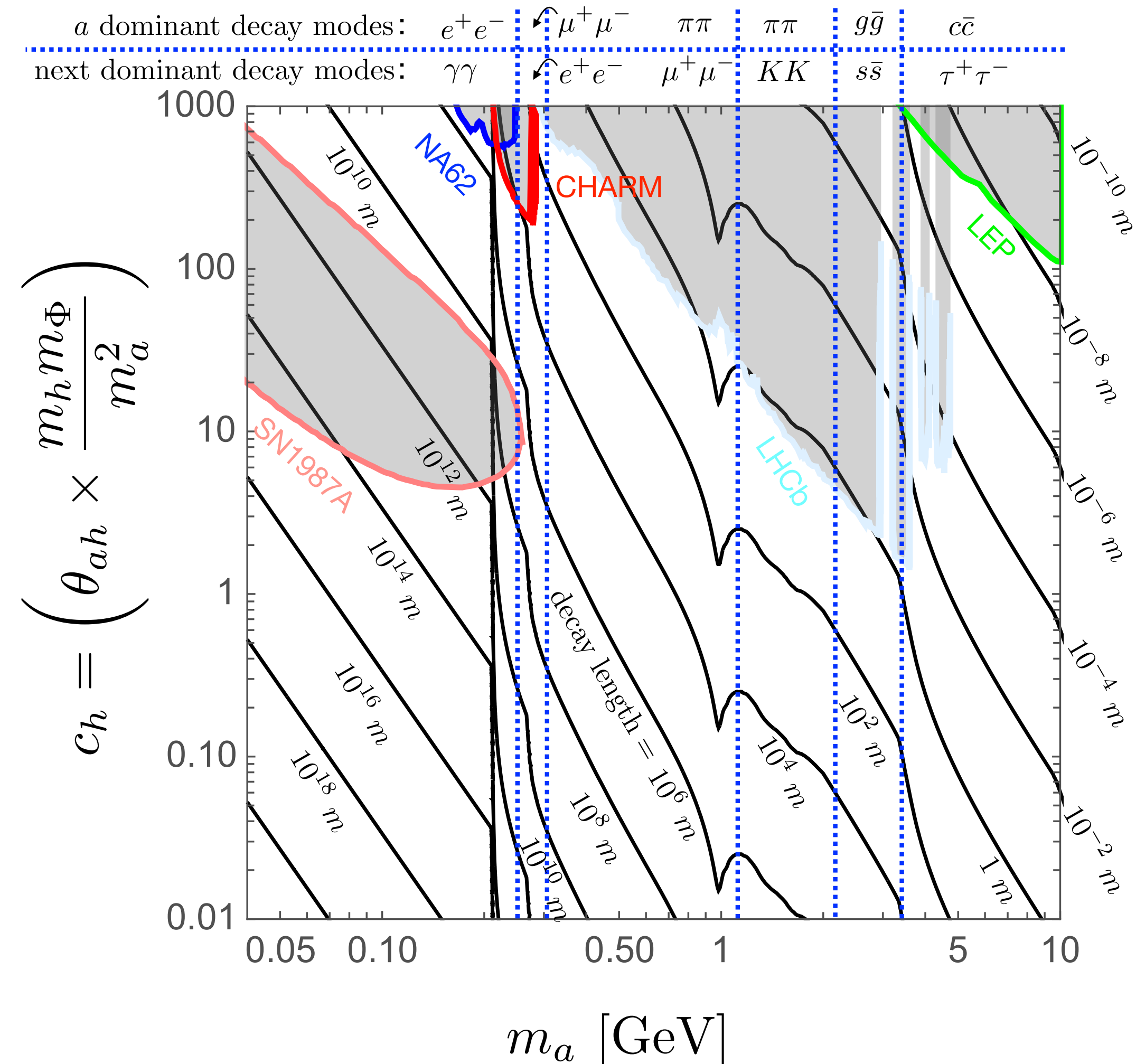
Visible particles

Probing CP-even ALP at e.g. ILC 250GeV

Kodai Sakurai, WY 2111.03653

Decay length and product of a from Higgs decay and signature at ILC

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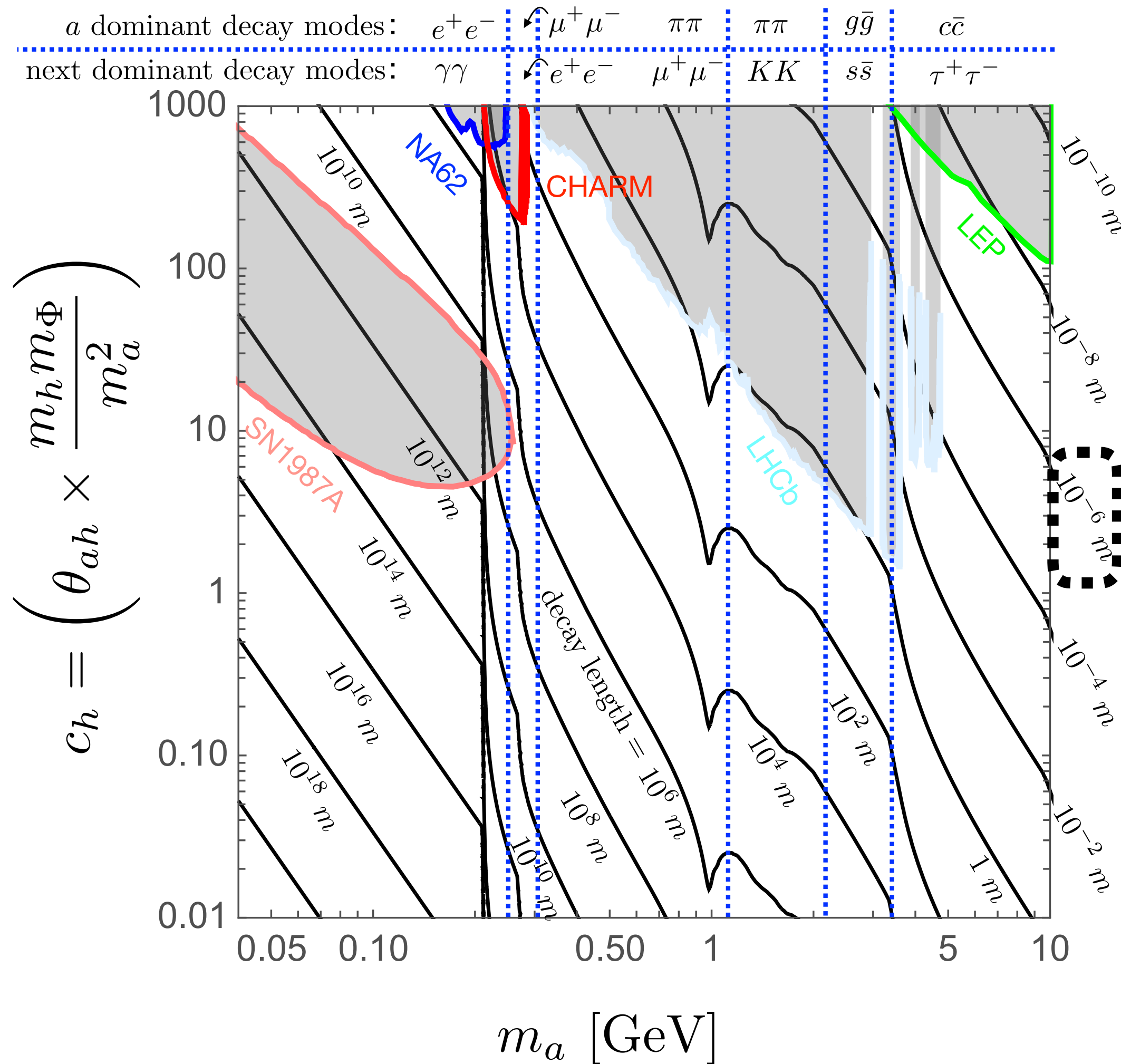
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2111.02437
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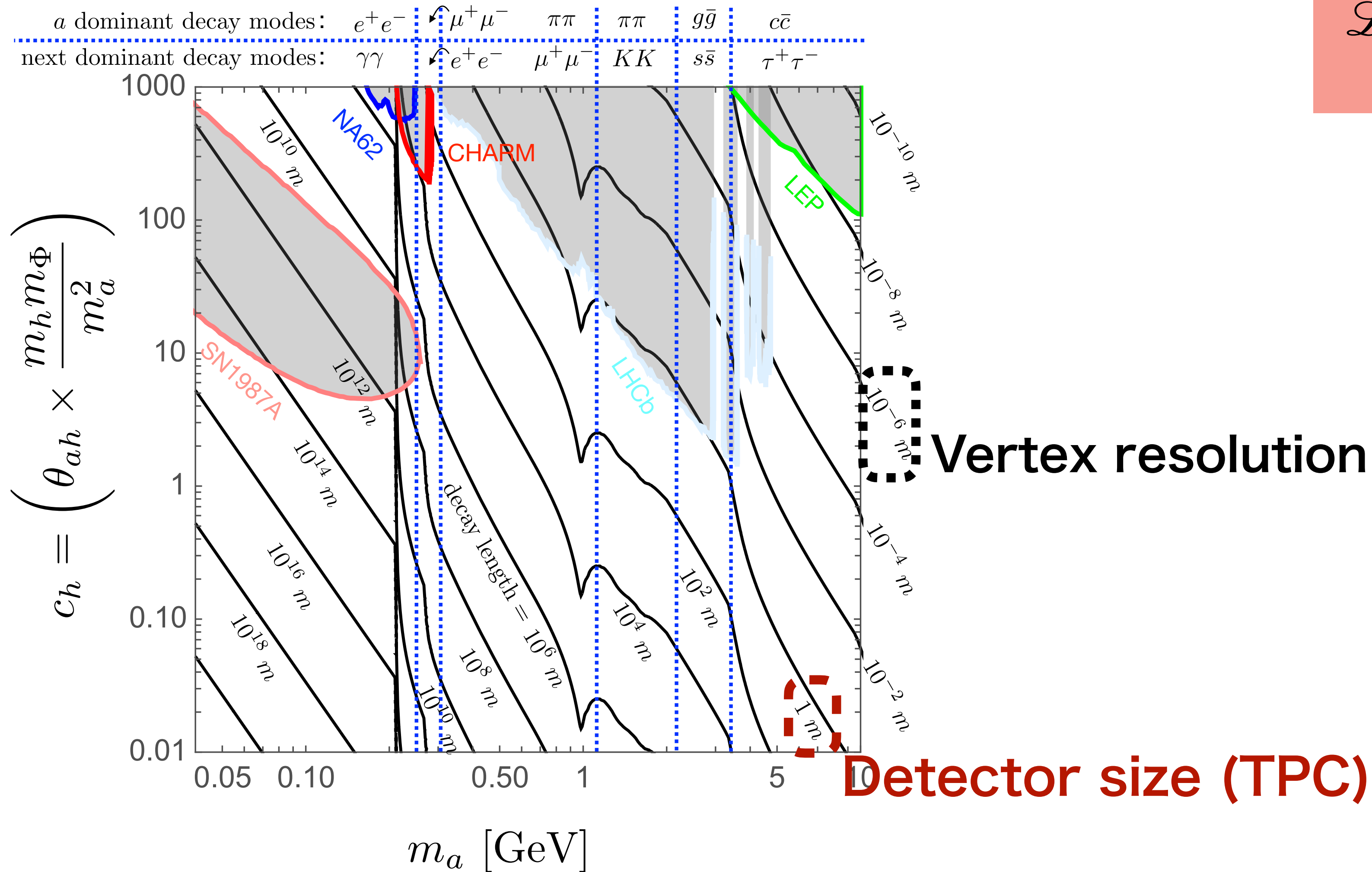
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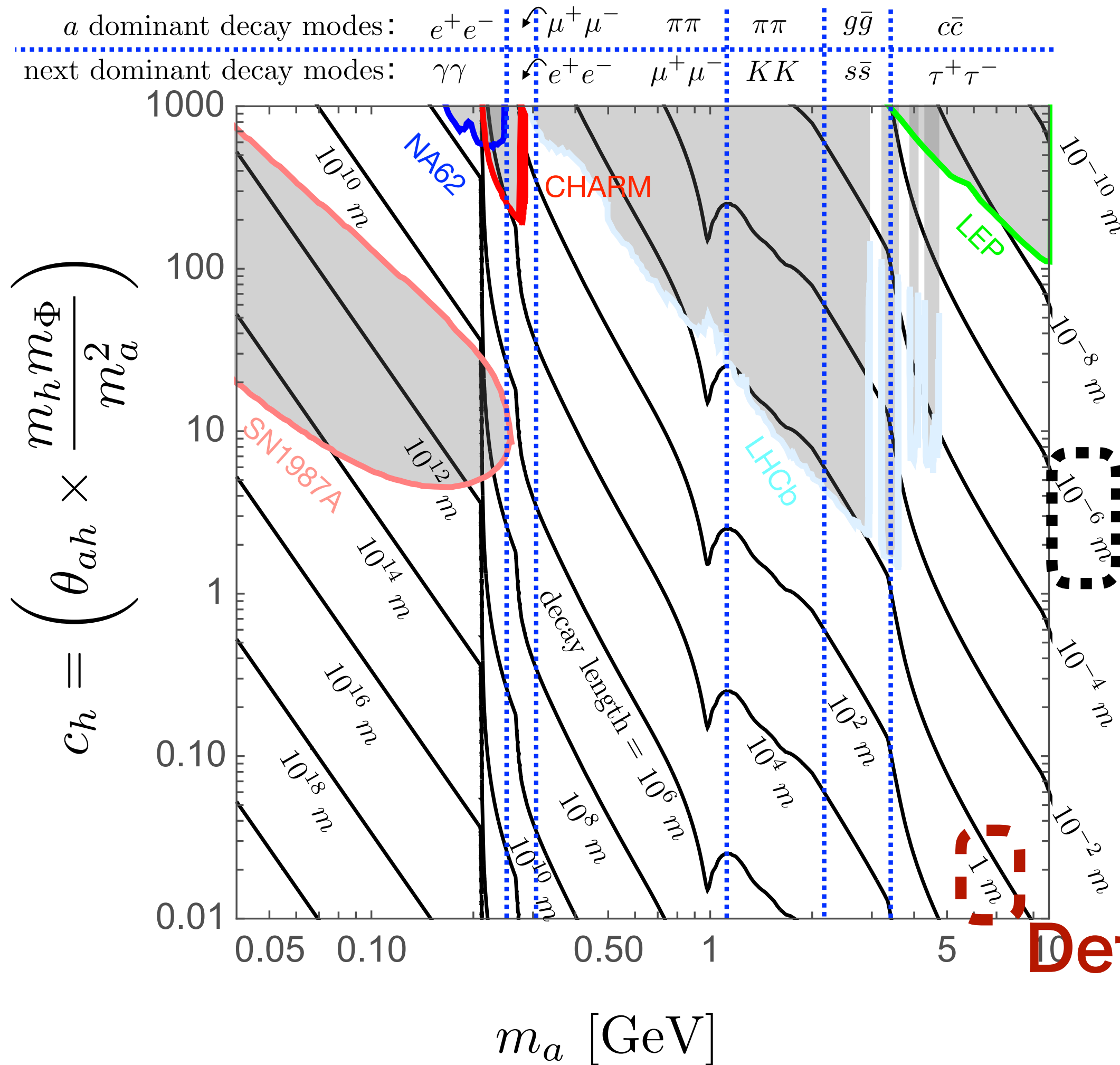
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Higgs exotic decay

Vertex resolution

Detector size (TPC)

$$Br_{h \rightarrow aa(\rightarrow c\bar{c}c\bar{c})} \gtrsim 10^{-3} \quad (2\sigma \text{CL}, 2 \text{ ab}^{-1})$$

Liu et al, 1612.09284

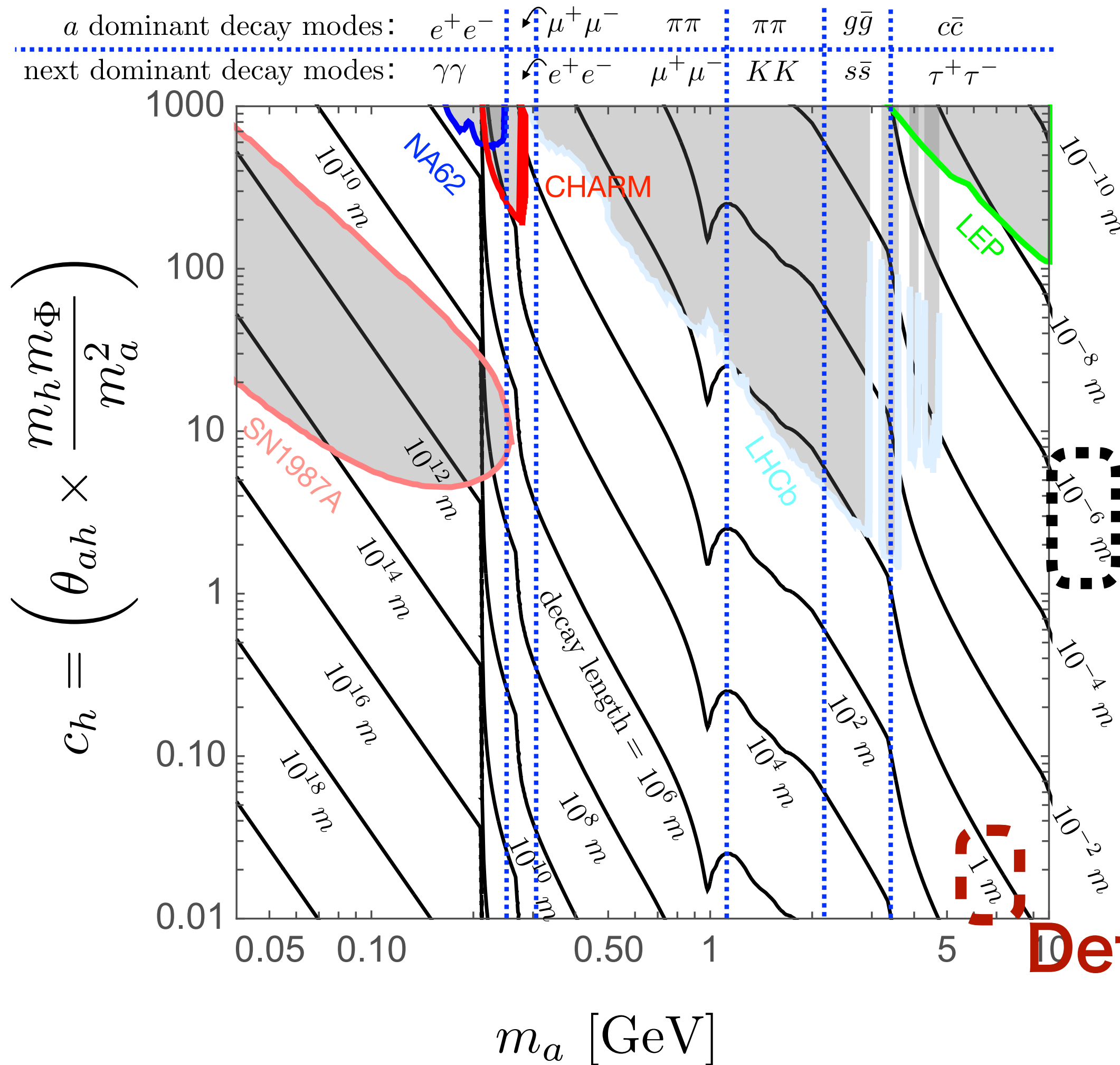
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Vertex resolution

Displaced vertex $\times 2$

$$Br_{h \rightarrow aa} > 10^{-6}, \quad (2\sigma\text{CL}, 3 \text{ ab}^{-1})$$

Detector size (TPC)

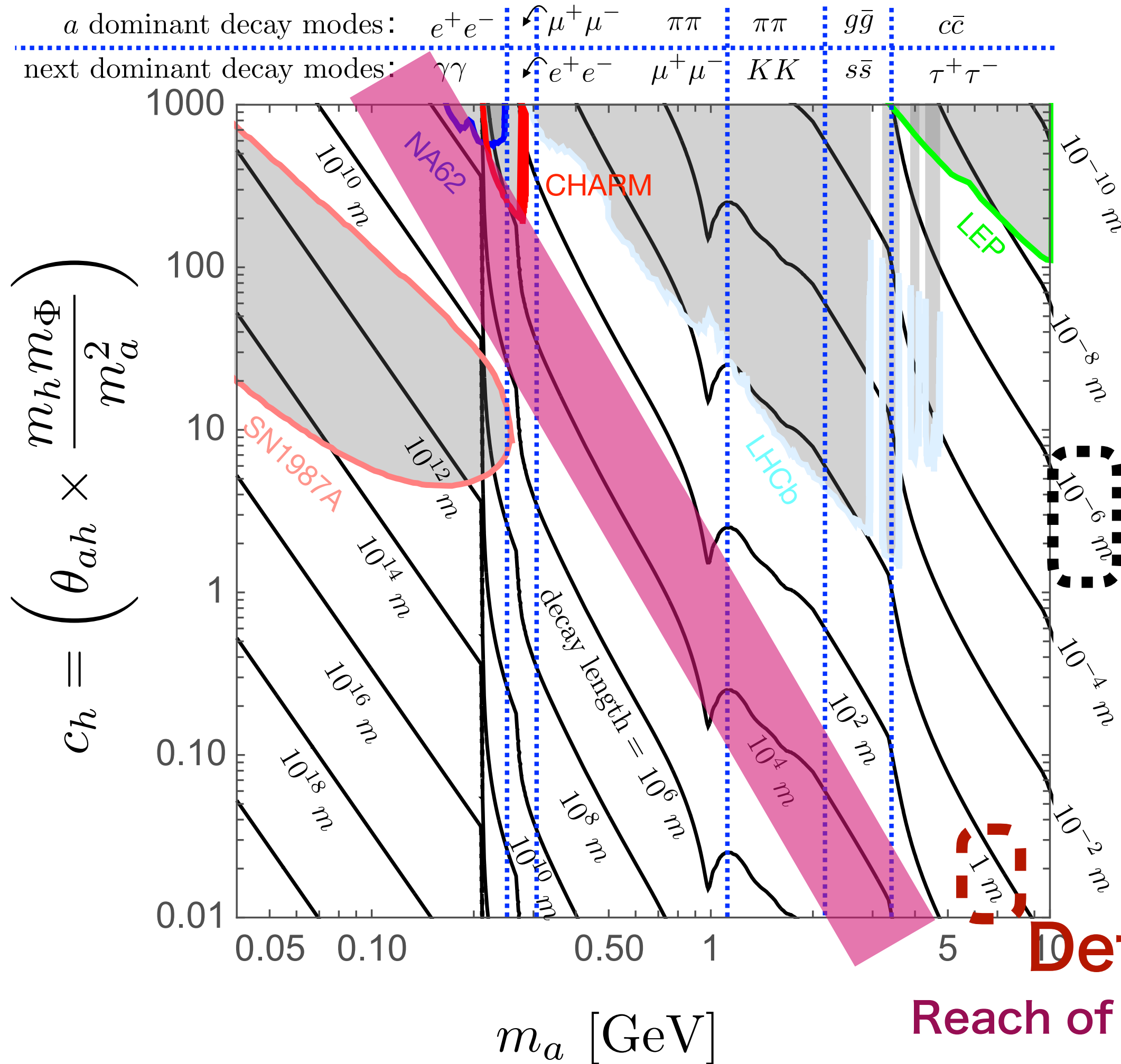
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Reach of rare displaced vertex ($Br_{h \rightarrow aa} \sim 1\%, 3 \text{ ab}^{-1}$)

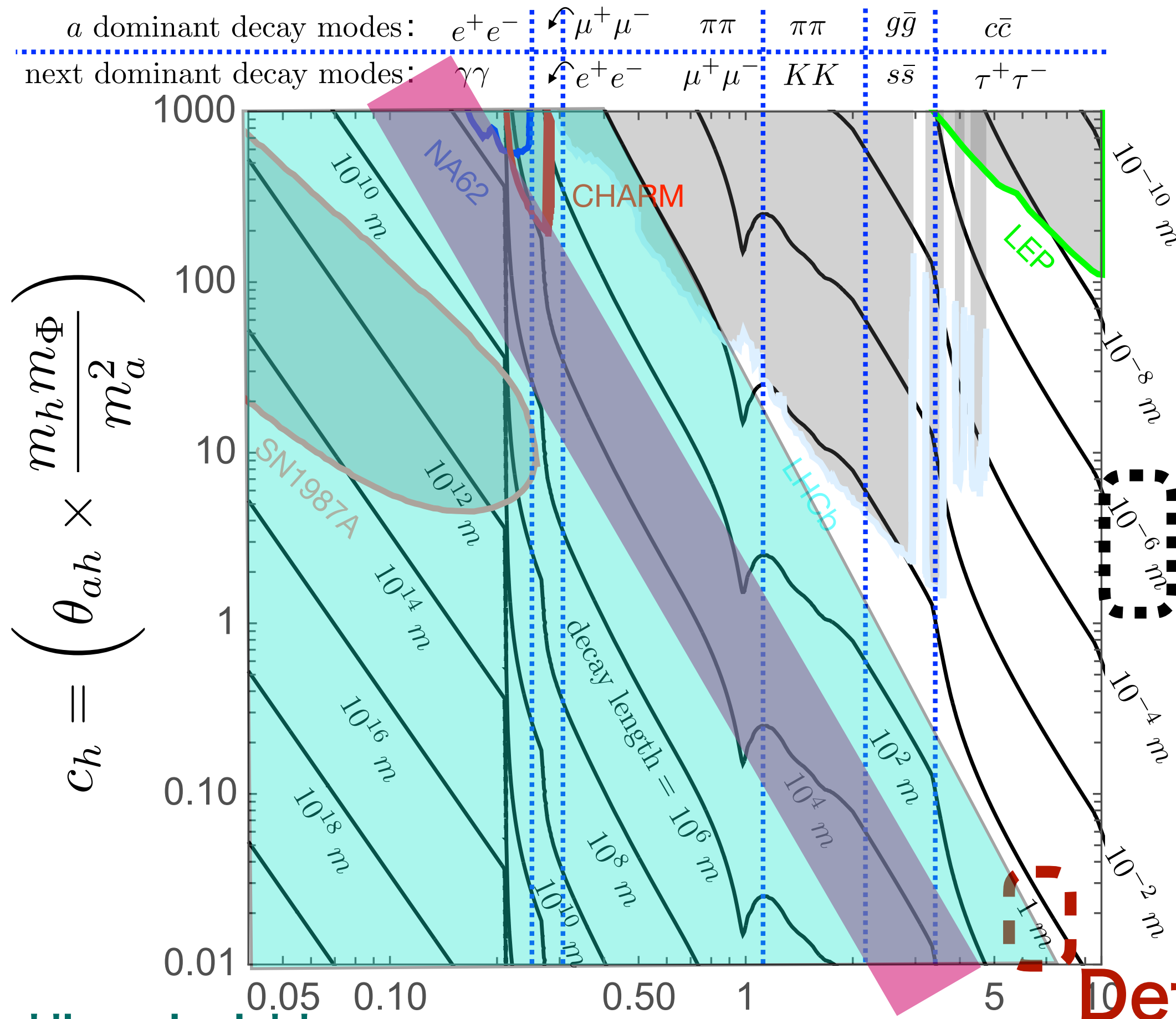
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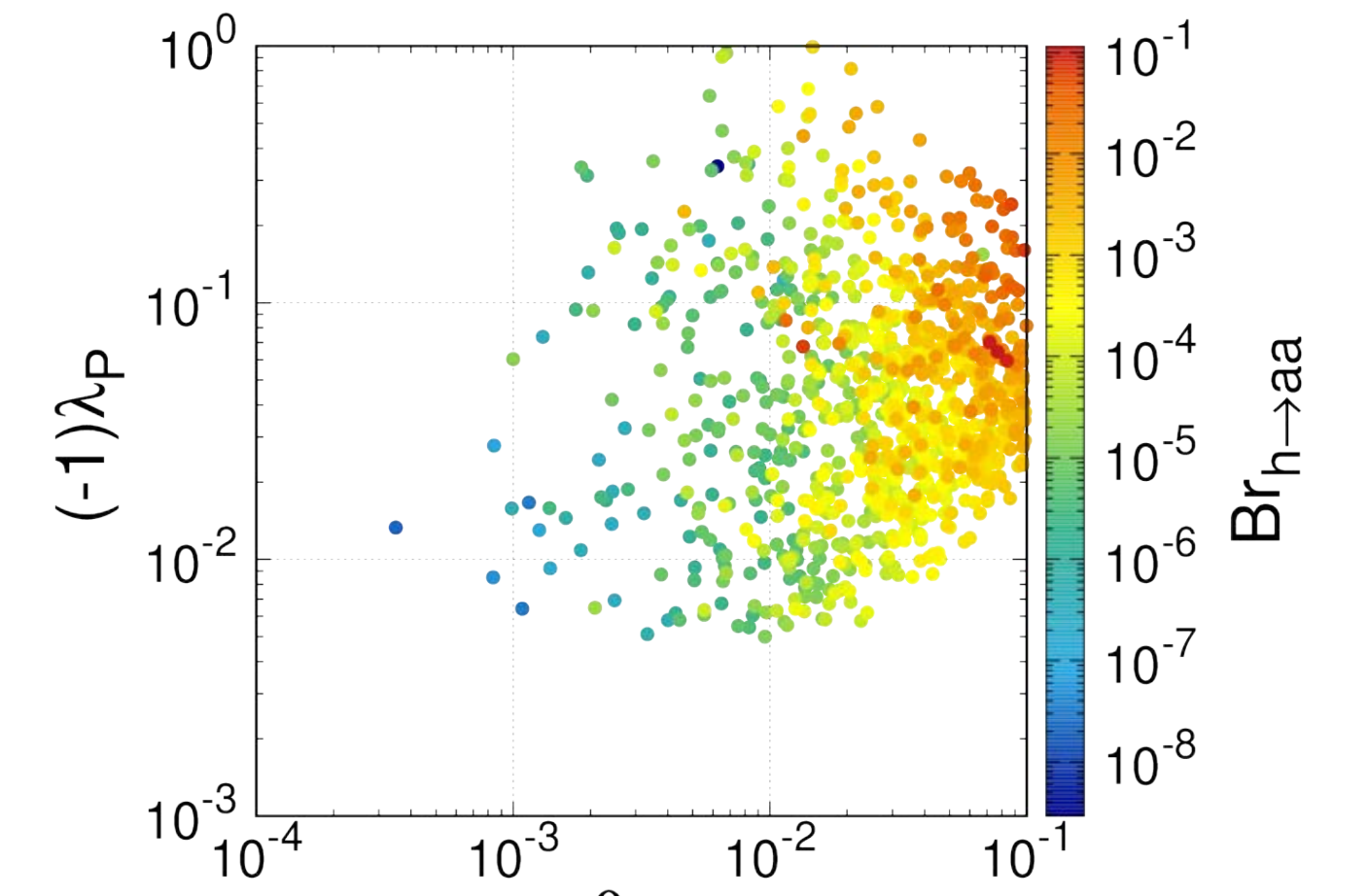
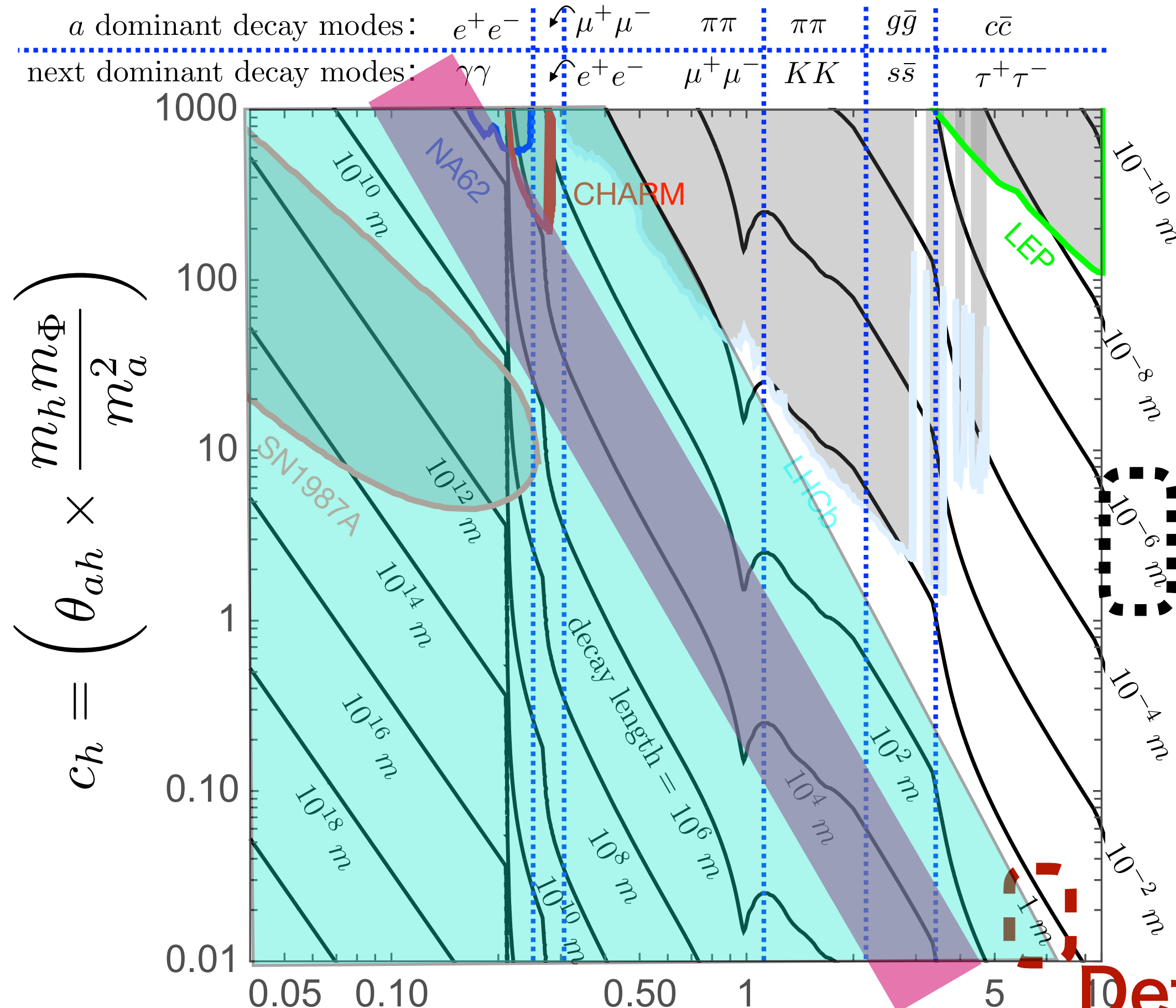
Higgs invisible

$$\text{decay } Br_{h \rightarrow aa} \gtrsim 0.1\% \quad m_a [\text{GeV}] \quad \text{ILC whitepaper}$$

Probing CP-even ALP at e.g. ILC 250GeV

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What roles does CP-even ALP play in the early Universe?

- Light mediator to DM with $\mathcal{L} \supset \Phi \bar{\Psi}_{\text{DM}}^c \Psi_{\text{DM}}$.

ALP couples SM fermion weakly but strongly with DM, which is the desired property of a light mediator.

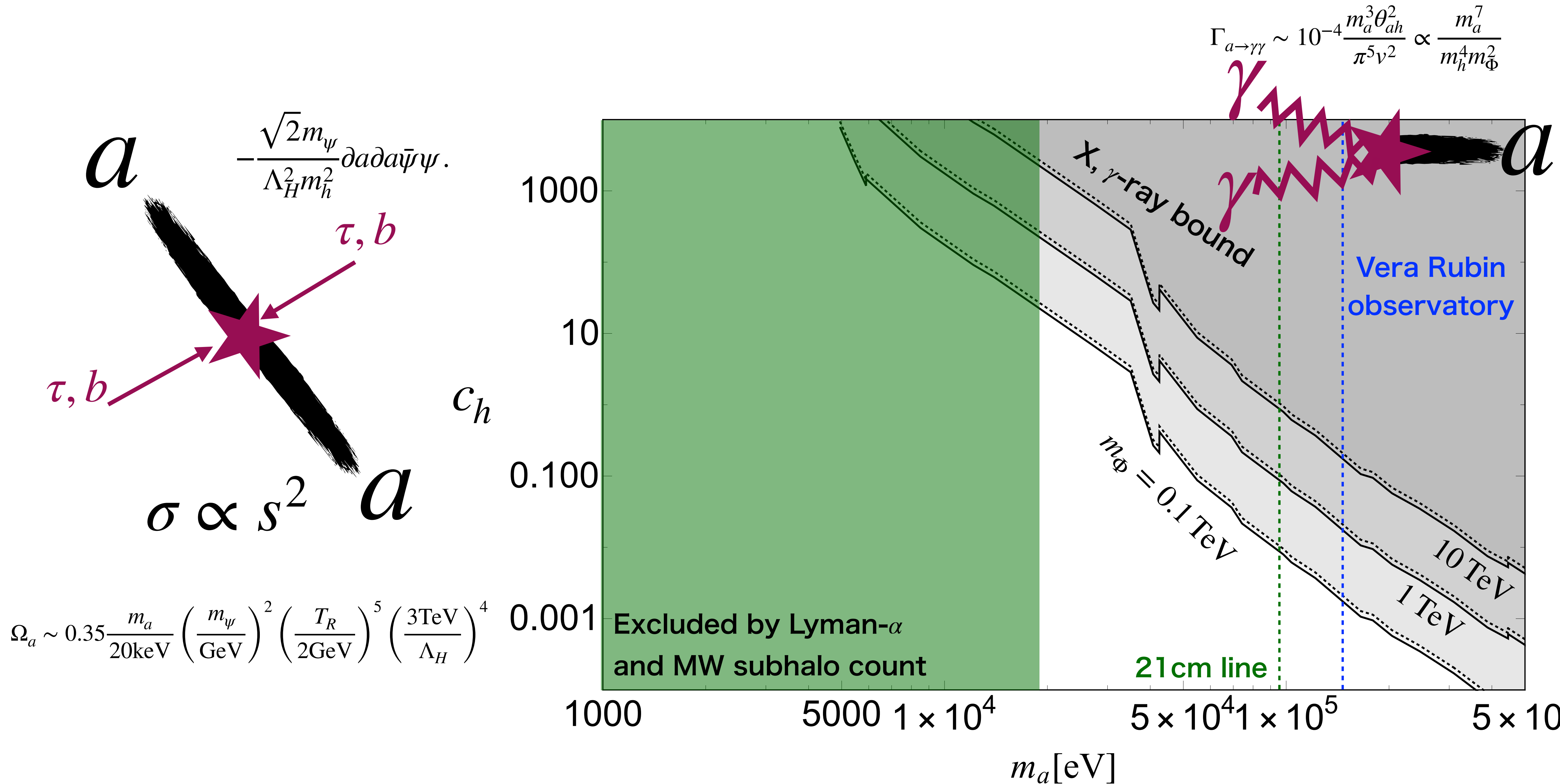
Please study it with WIMP, which should be an interesting topic!

- **CP-even ALP DM.** This talk.

CP-even ALP is a good DM candidate if it is lighter than MeV.

$$\Gamma_{a \rightarrow \gamma\gamma} \sim 10^{-4} \frac{m_a^3 \theta_{ah}^2}{\pi^5 v^2} \propto \frac{m_a^7}{m_h^4 m_\Phi^2}$$

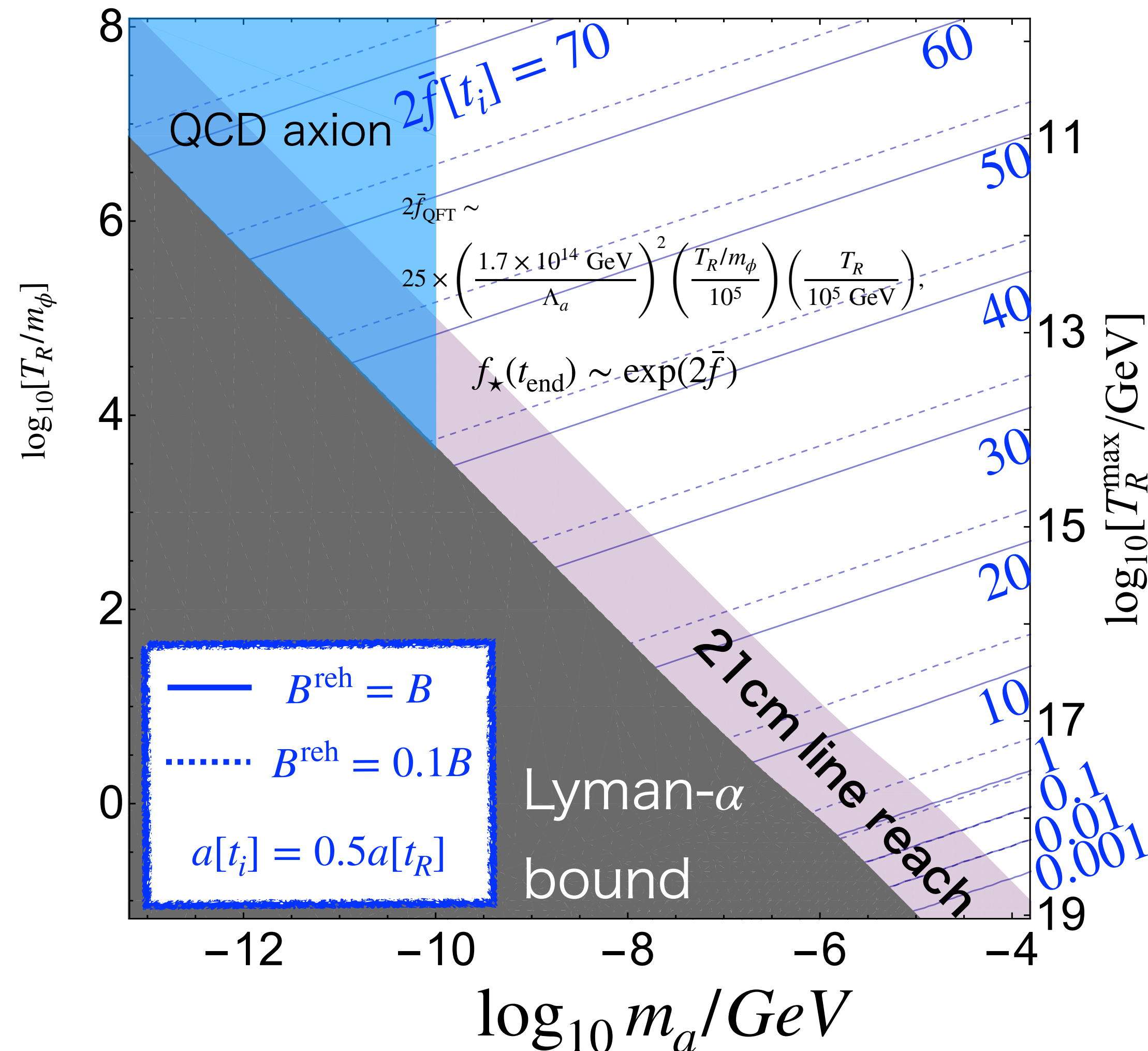
Thermally produced CP-even ALP DM



Non-thermal production scenario: lighter mass range.

Light bosonic DM can be produced during reheating if $T_R > m_{\text{inflaton}}$ as laser.

Moroi, WY, 2011.09475, 2011.12285



$$\mathcal{L}_{\text{int}} = \frac{\phi}{\Lambda_a} \partial_\mu a \partial^\mu a + \frac{\phi}{\Lambda_G} G_{\mu\nu}^{(a)} G^{(a)\mu\nu}$$

For CP-even ALP, we need $T_R \ll m_\Phi$ for the produced ALP not to be thermalized. Probed by inflaton search, 21 cm line.

There naturally exists the portal coupling between Φ and H .

$$V = -m_{\Phi}^2 |\Phi|^2 + \lambda |\Phi|^4 + \lambda_P |\Phi|^2 |H|^2 + \lambda_H |H|^4 - \mu_H^2 |H|^2.$$

Higgs portal coupling

Φ : U(1) Higgs field
 H : SM Higgs doublet

$$\Gamma_{h \rightarrow aa} \simeq \frac{1}{16\pi} \frac{v^2 m_h^3}{\Lambda_H^4}.$$

$$\text{Br}_{h \rightarrow aa} = 30\% \left(\frac{1 \text{TeV}}{\Lambda_H} \right)^4$$

$$\frac{1}{\Lambda_H^2} \equiv \frac{\lambda_P}{m_{\Phi}^2}.$$

