

Directed flow in heavy-ion collisions as a probe of the first order phase transition

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in collaboration with

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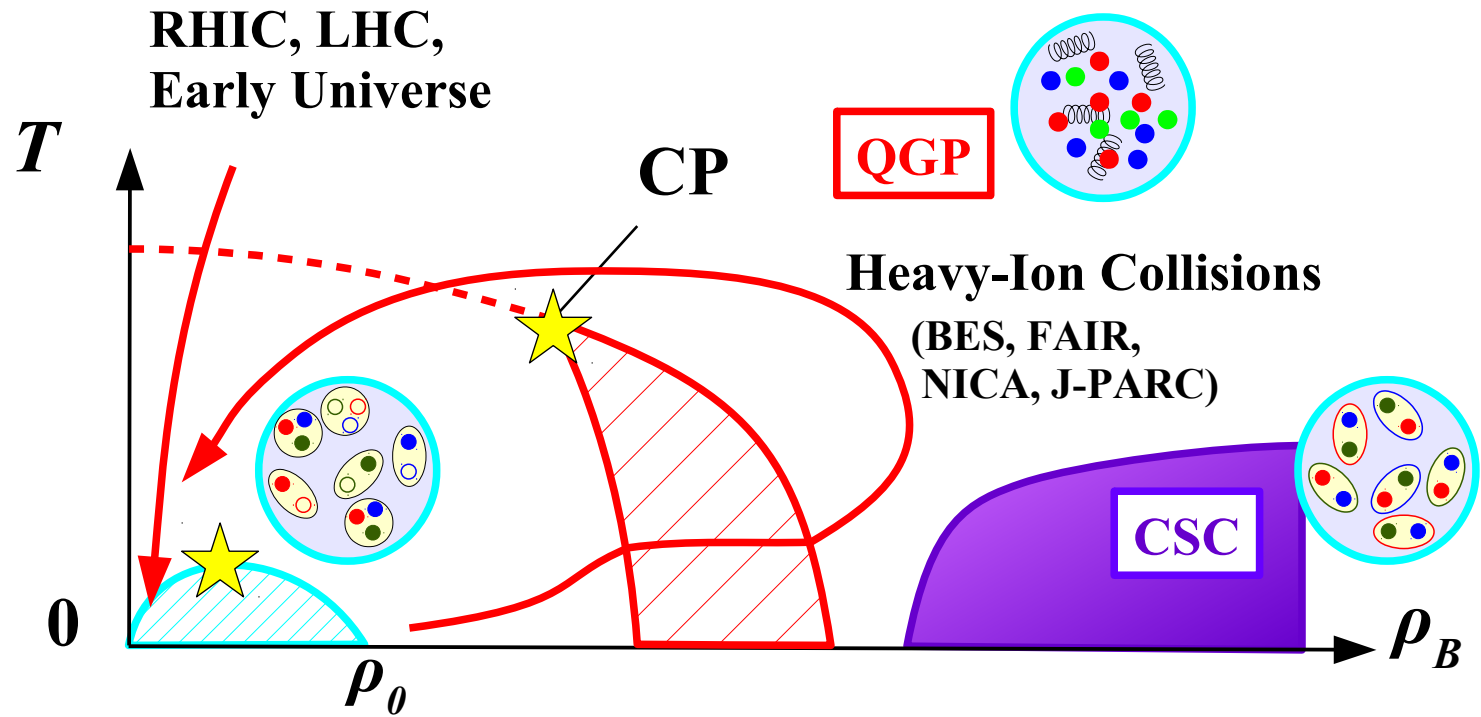
*The 34th Reimei WorkShop on
"Physics of Heavy-Ion Collisions at J-PARC",
Aug.8-9, 2016, J-PARC, Japan*

- **Introduction: Negative dv_1/dy at $\sqrt{s_{NN}} \sim 10$ GeV**
- **Hadronic transport model with Softening Effects**
- **Summary**

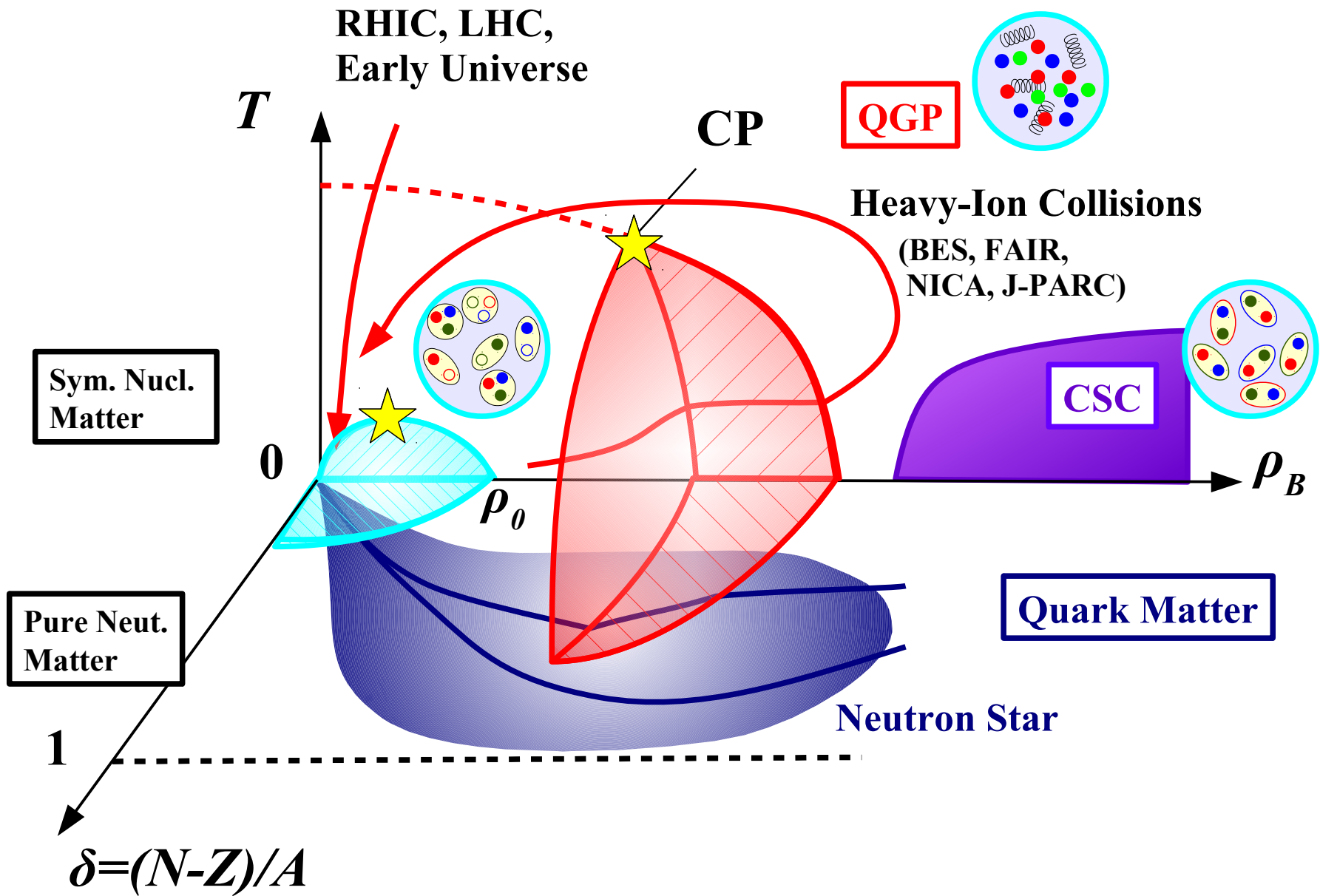
Y. Nara, A. Ohnishi, arXiv:1512.06299 [nucl-th] (QM2015 proc., to appear)

Y. Nara, H. Niemi, A. Ohnishi, H. Stoecker, arXiv:1601.07692 [hep-ph]

QCD Phase Diagram



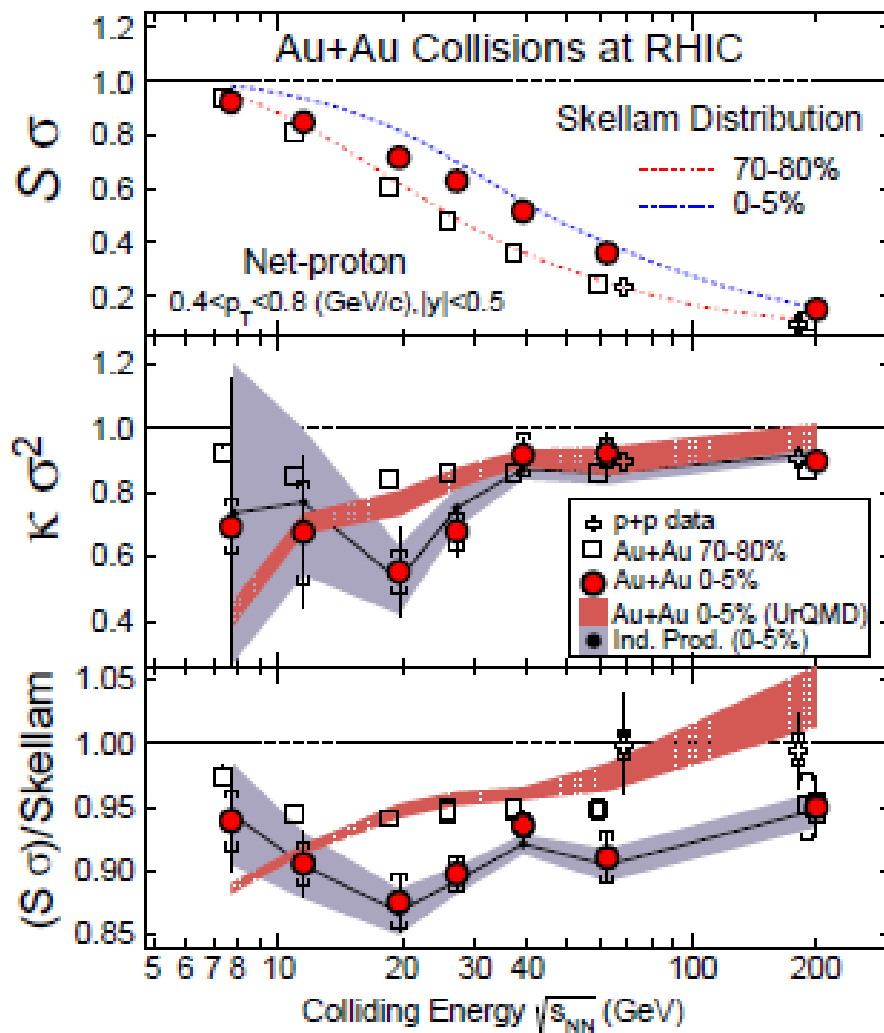
QCD Phase Diagram



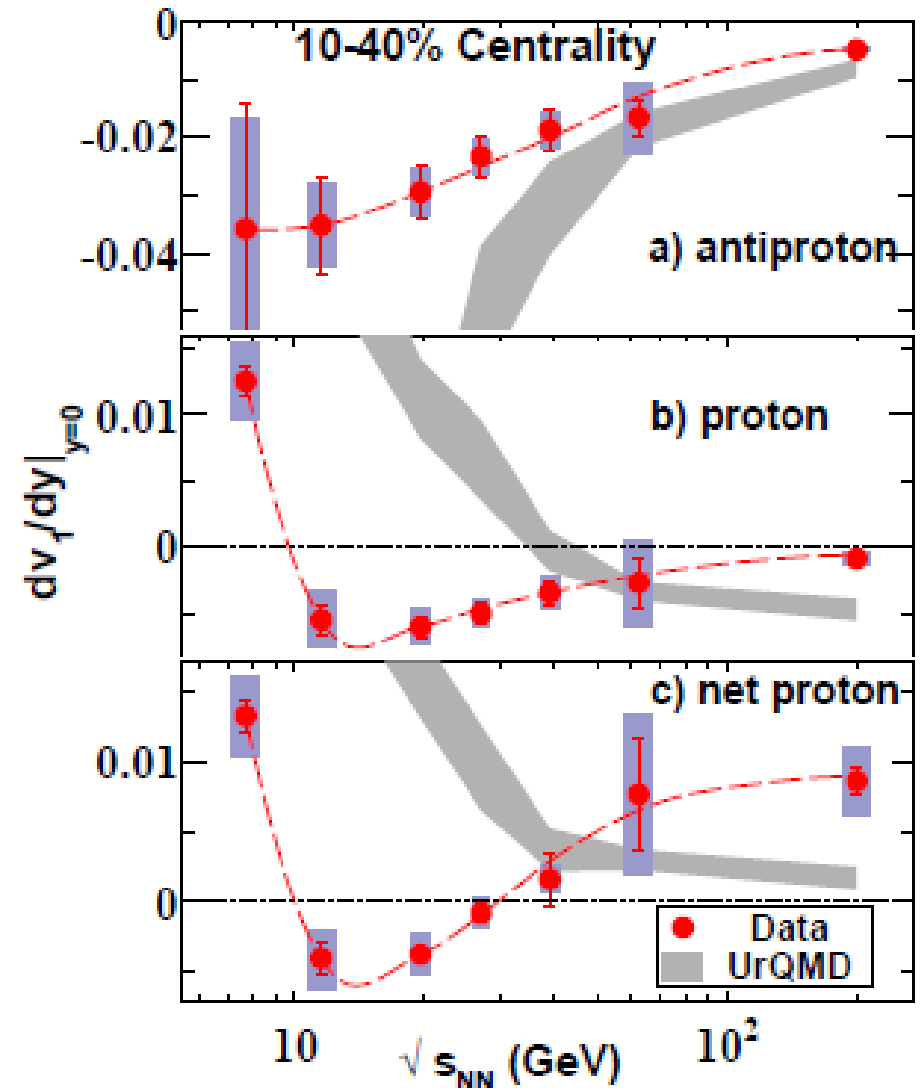
QCD phase transition

- **QCD phase transition at top RHIC & LHC energies = Crossover**
→ **One of Next Grand Challenges**
= **Discovery of 1st or 2nd order phase transition in QCD**
 - **Signals of QCD phase transition at J-PARC energies ($\sqrt{s_{NN}}=5-10$ GeV)?**
 - **(Partial) Chiral restoration** → **Modification of hadron properties**
 - **Critical Point** → **Large fluctuation of conserved charges**
 - **First-order phase transition** → **Softening of EOS**
- **Non-monotonic behavior of proton number moment ($\kappa\sigma^2$) and collective flow (dv_1/dy)**

Net-Proton Number Cumulants & Directed Flow

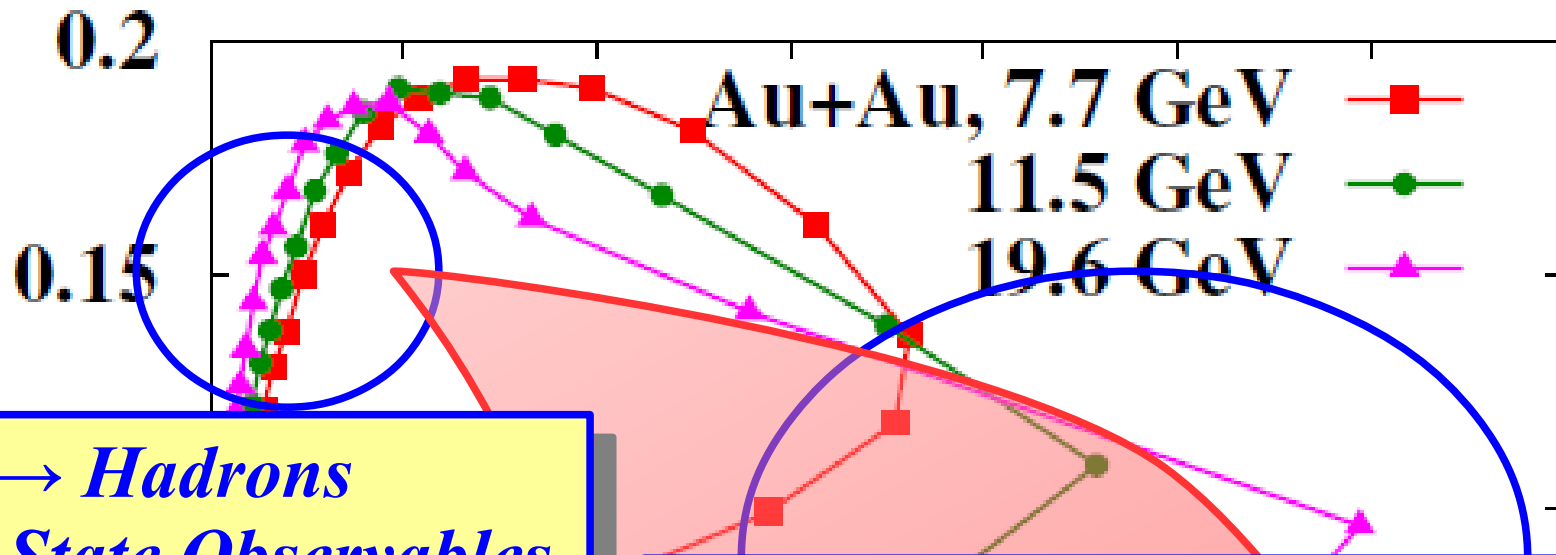


STAR Collab. PRL 112('14)032302



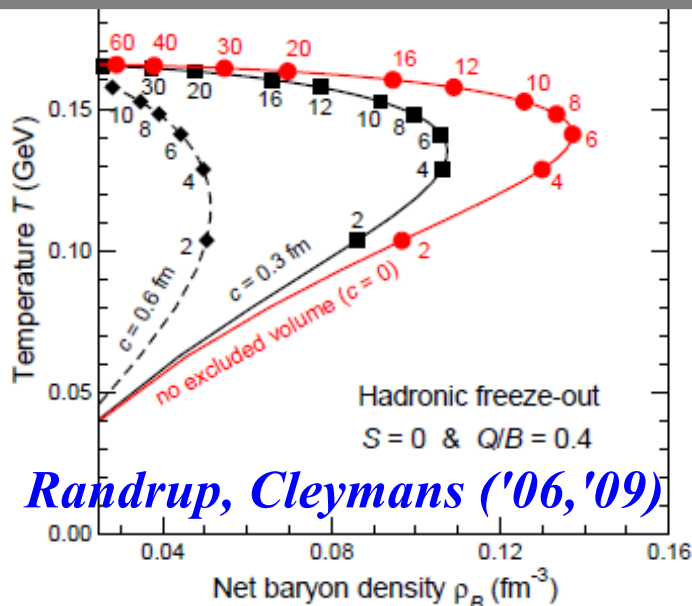
STAR Collab., PRL 112('14)162301.

Two ways to probe QCD phase transition



QGP → Hadrons
Final State Observables
Cumulants, ...

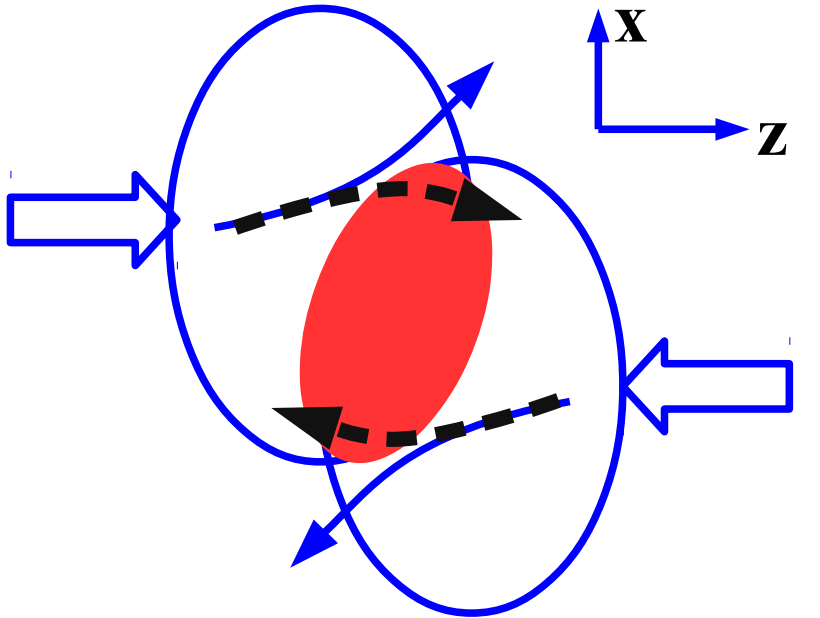
Hadrons → QGP
Early Stage Observables
Caution: (Partial) Equilibration
is necessary!



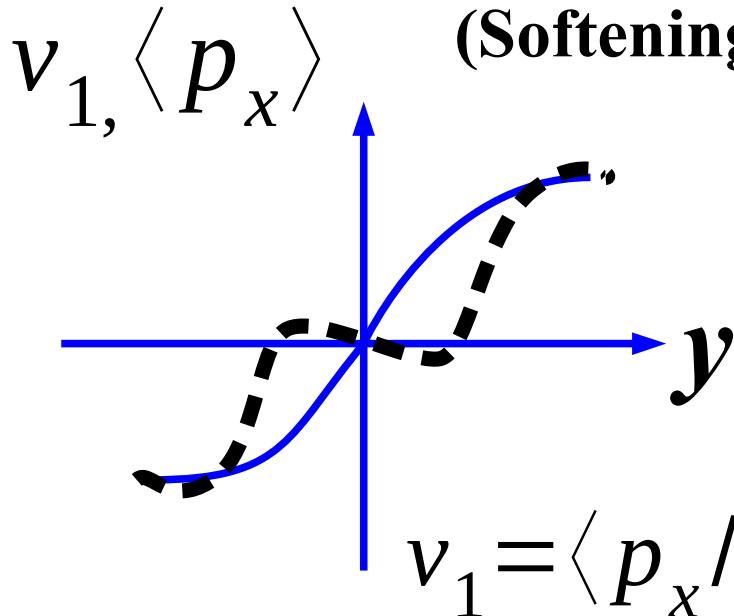
Randrup, Cleymans ('06,'09)

4 6 8 10 12 14
 ρ_B / ρ_0

What is directed flow ?



Attraction
(Softening)



$$v_1 = \langle p_x / p \rangle = \langle \cos \varphi \rangle$$

- v_1 or $\langle p_x \rangle$ as a function of y is called directed flow.
- Created in the overlapping stage of two nuclei
→ Sensitive to the EOS in the early stage.
- Becomes smaller at higher energies.

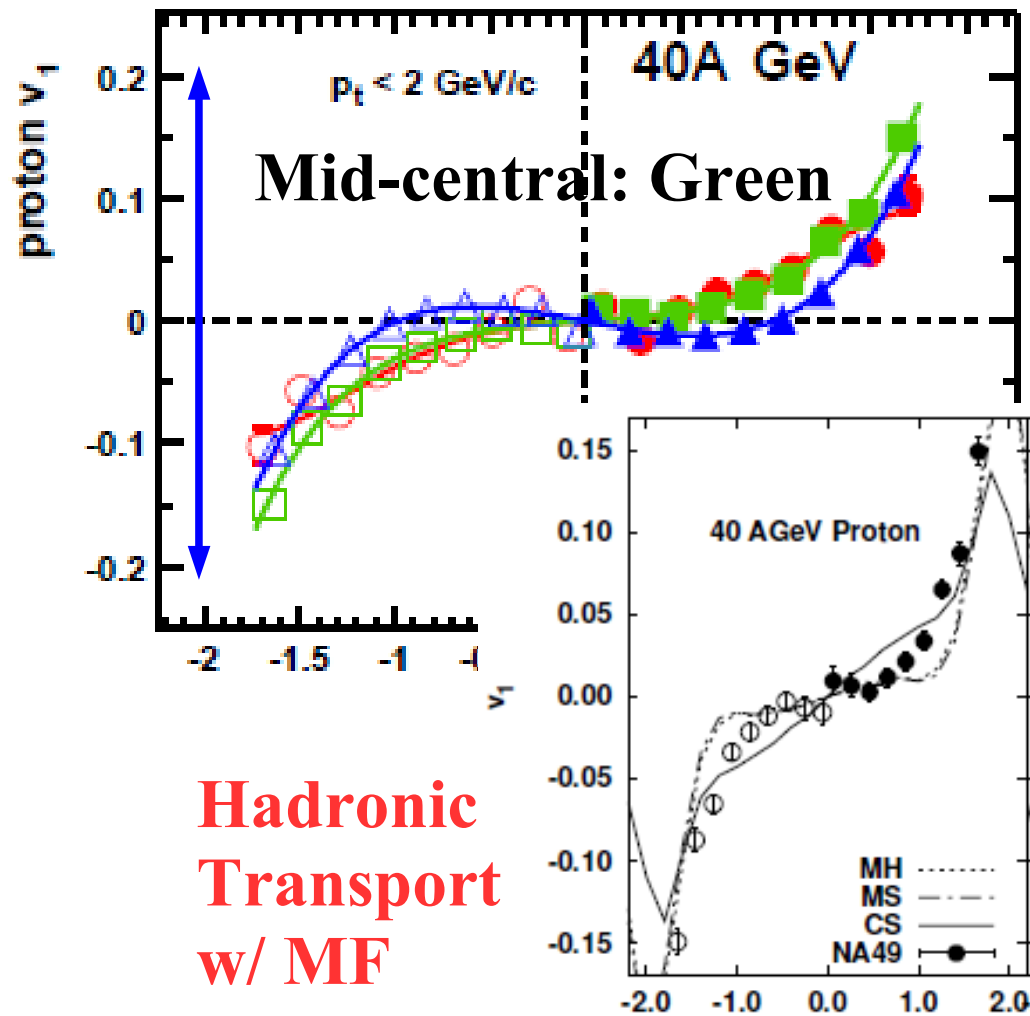
How can we explain non-monotonic dependence of dv_1/dy ?
→ *Softening or Geometry*

SPS(NA49) vs RHIC(STAR)

■ SPS (NA49), $\sqrt{s_{NN}} = 8.9$ GeV

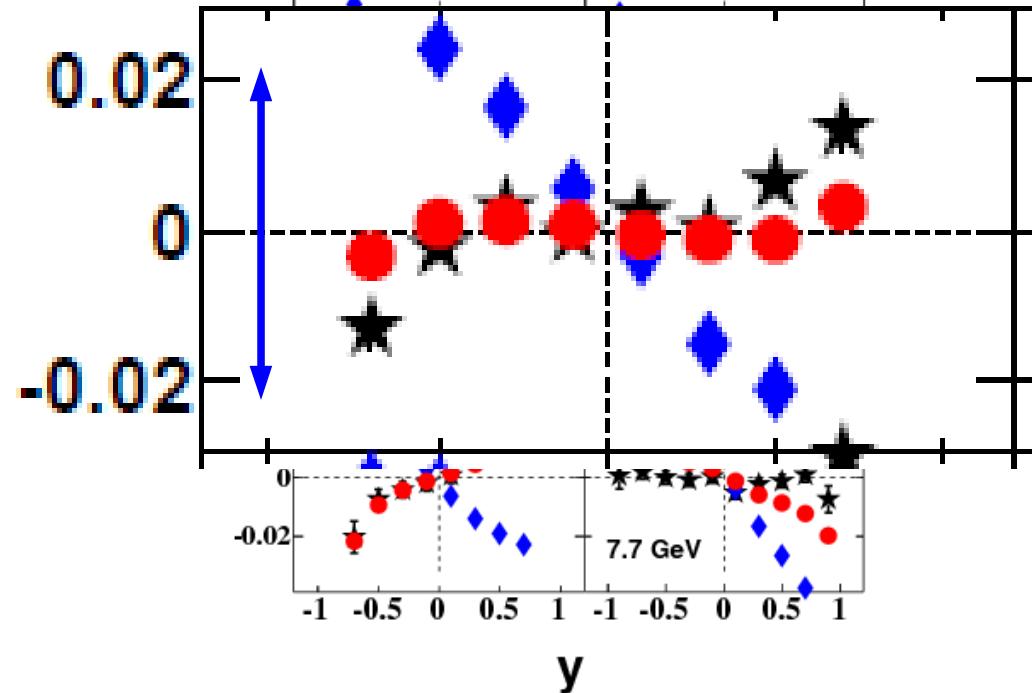
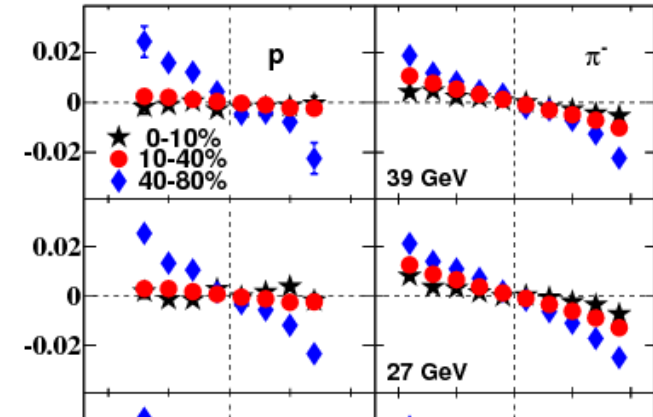
■ RHIC(STAR), 7.7-39 GeV

C. Alt et al. (NA49), PRC68 ('03) 034903



**Hadronic
Transport
w/ MF**

*M.Isse, A.O, N.Otuka, P.K.Sahu, Y.Nara,
PRC72 ('05)064908*

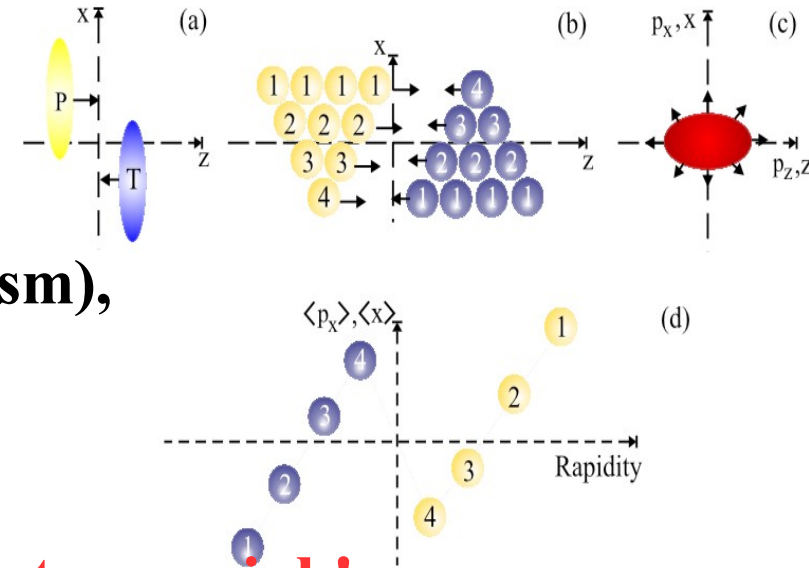


*L. Adamczyk et al. (STAR),
PRL 112(2014)162301*

A. Ohnishi @ Reimei-HI 2016, Aug.9, 2016

Does Directed Flow Collapse Signal Phase Tr. ?

- **Negative dv_1/dy at high-energy**
($\sqrt{s}_{NN} > 20$ GeV)



- **Geometric origin (bowling pin mechanism), not related to FOPT**

R. Snellings, H. Sorge, S. Voloshin, F. Wang, N. Xu, PRL84,2803('00)

- **Negative dv_1/dy at $\sqrt{s}_{NN} \sim 10$ GeV \rightarrow **Controversial !****

- **Yes, in three-fluid simulations. \rightarrow Thermalization ?**

Y. B. Ivanov and A. A. Soldatov, PRC91('15)024915; P. Batyuk et al., 1608.00965.

- **No (for semi-central collisions), in transport models incl. hybrid.**

E.g. J. Steinheimer, J. Auvinen, H. Petersen, M. Bleicher, H. Stoecker, PRC89('14)054913.

Exception: B.A.Li, C.M.Ko ('98) with *FOPT EOS*

We investigate the directed flow at J-PARC energies in hadronic transport model with / without mean field effects and with / without softening effects via attractive orbit.

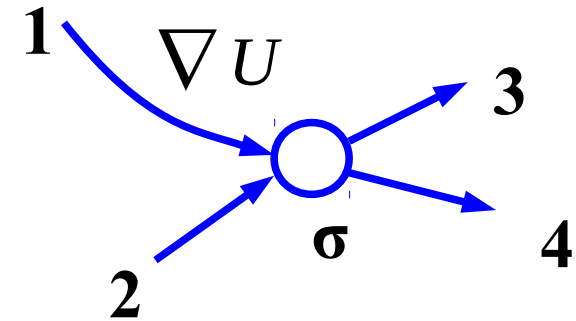
Hadronic Transport with Softening Effects

Transport Model

- **Microscopic Transport Models**
= Boltzmann equation
with (optional) potential effects

E.g. Bertsch, Das Gupta, Phys. Rept. 160(88), 190

- UrQMD 3.4 (Frankfurt), PHSD Giessen (Cassing), GiBUU 1.6 Giessen (Mosel), AMPT (Texas A&M), JAM (Y. Nara)



- **Hadron-string transport model JAM**

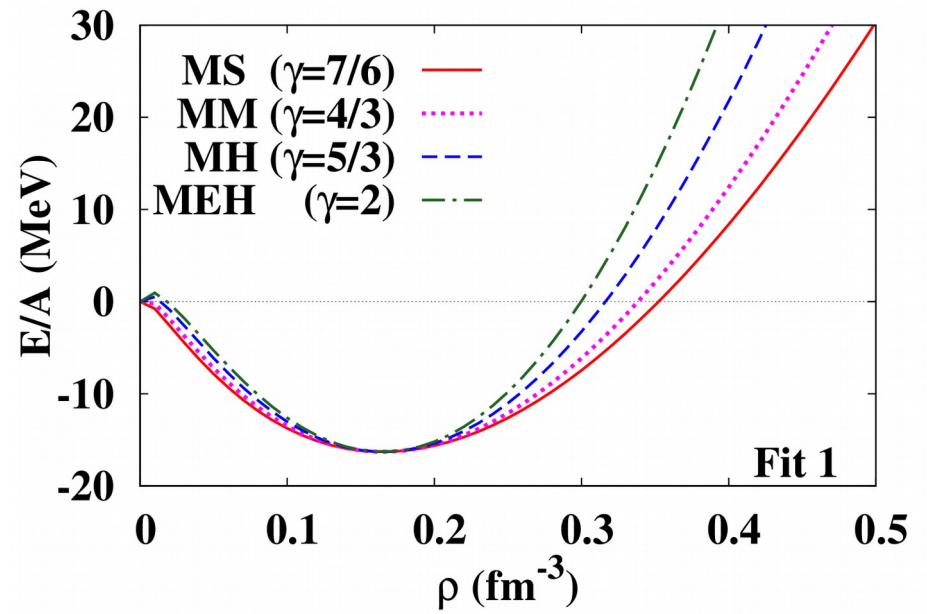
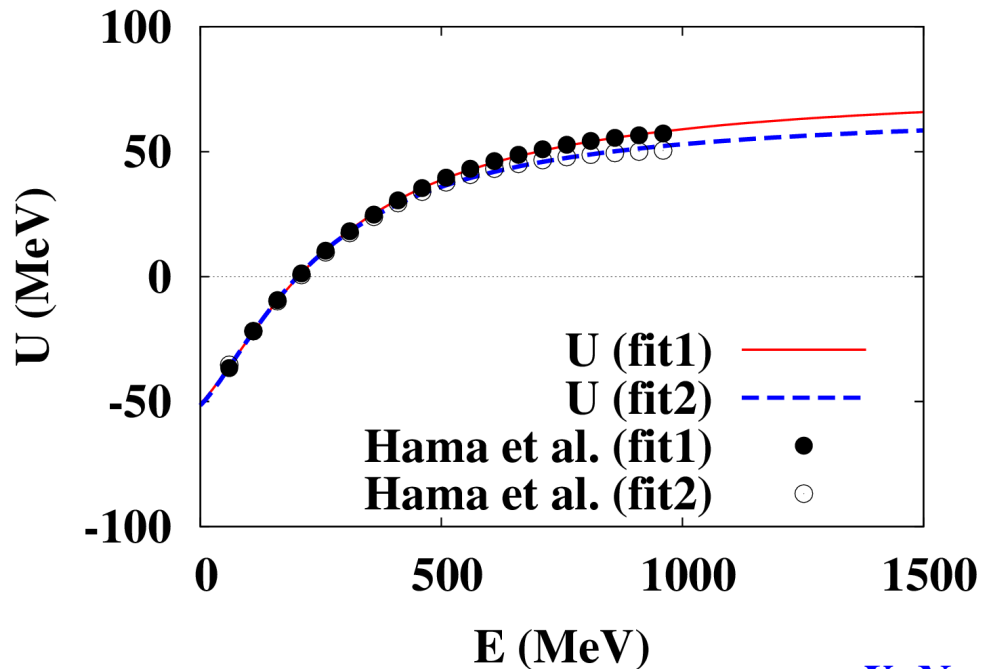
- Hadronic cascade with resonance and string excitation
Nara, Otuka, AO, Niita, Chiba, Phys. Rev. C61 (2000), 024901.
- Potential term → Mean field effects in the framework of RQMD/S
Sorge, Stocker, Greiner, Ann. of Phys. 192 (1989), 266.
Tomoyuki Maruyama et al., Prog. Theor. Phys. 96(1996), 263.
Isse, AO, Otuka, Sahu, Nara, Phys.Rev. C 72 (2005), 064908.

Mean Field Potential

■ Skyrme type density dependent + momentum dependent potential

$$V = \sum_i V_i = \int d^3r \left[\frac{\alpha}{2} \left(\frac{\rho}{\rho_0} \right)^2 + \frac{\beta}{\gamma+1} \left(\frac{\rho}{\rho_0} \right)^{\gamma+1} \right] + \sum_k \int d^3r d^3p d^3p' \frac{C_{ex}^{(k)}}{2\rho_0} \frac{f(\mathbf{r}, \mathbf{p})f(\mathbf{r}, \mathbf{p}')}{1 + (\mathbf{p} - \mathbf{p}')^2 / \mu_k^2}$$

Type	α (MeV)	β (MeV)	γ	$C_{ex}^{(1)}$ (MeV)	$C_{ex}^{(2)}$ (MeV)	μ_1 (fm ⁻¹)	μ_2 (fm ⁻¹)	K (MeV)
MH1	-12.25	87.40	5/3	-383.14	337.41	2.02	1.0	371.92
MS1	-208.89	284.04	7/6	-383.14	337.41	2.02	1.0	272.6



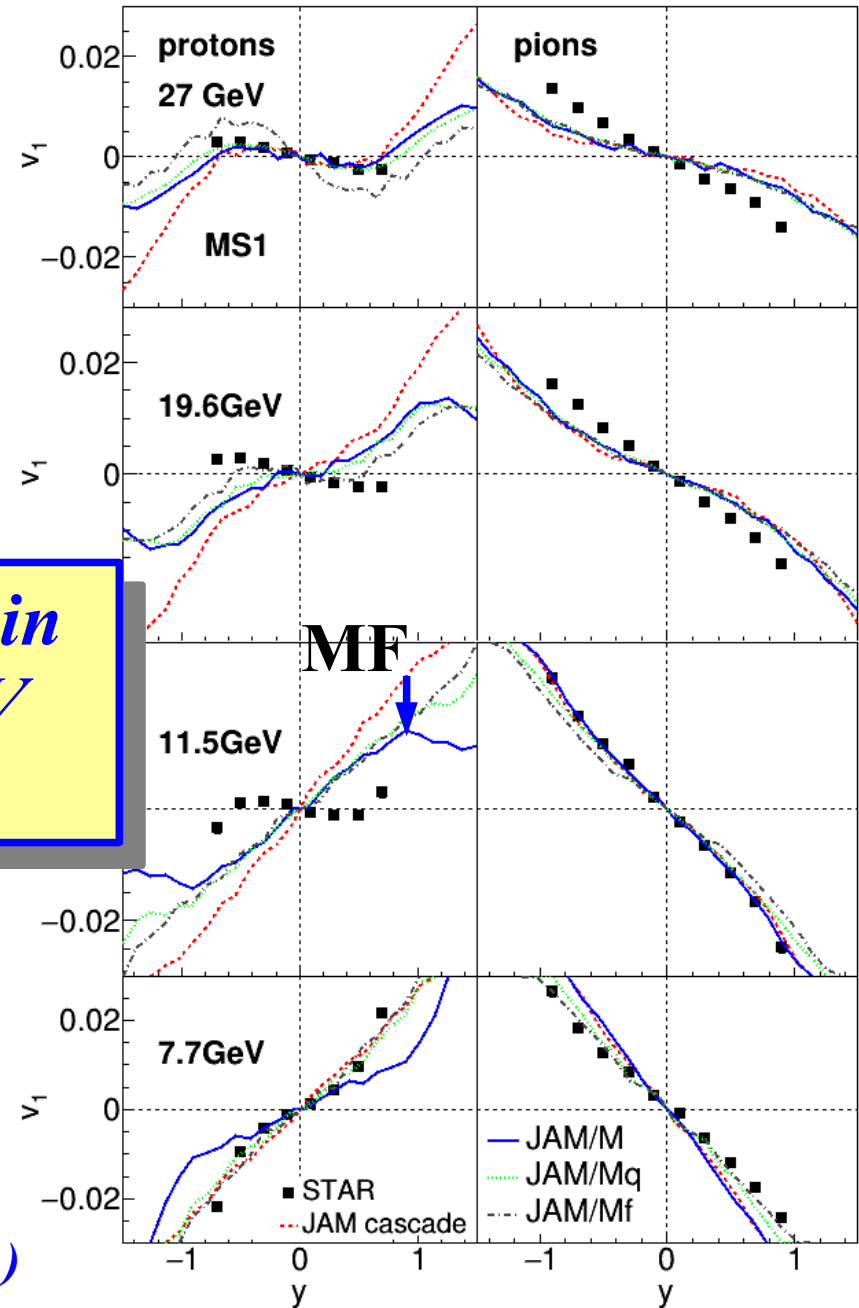
*Y. Nara, AO, arXiv:1512.06299 [nucl-th] (QM2015 proc.)
Isse, AO, Otuka, Sahu, Nara, PRC 72 (2005), 064908.*

Comparison with RHIC data on v_1

- Pot. Eff. on the v_1 is significant, but dv_1/dy becomes negative only at $\sqrt{s_{NN}} > 20$ GeV.

Hadronic approach does not explain directed flow collapse at 10-20 GeV even with potential effects.

- JAM/M: only formed baryons feel potential forces
- JAM/Mq: pre-formed hadron feel potential with factor 2/3 for diquark, and 1/3 for quark
- JAM/Mf: both formed and pre-formed hadrons feel potential forces.



Y. Nara, AO, arXiv:1512.06299 [nucl-th] (QM2015 proc.)

Softening Effects via Attractive Orbit Scattering

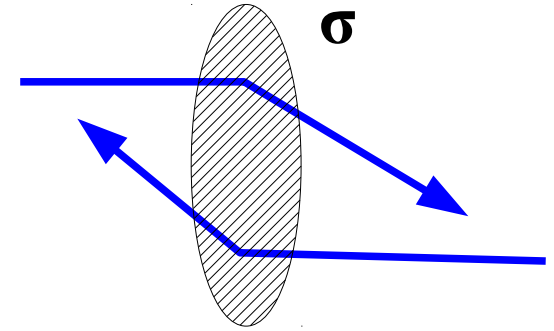
- Attractive orbit scattering simulates softening of EOS

P. Danielewicz, S. Pratt, PRC 53, 249 (1996)

H. Sorge, PRL 82, 2048 (1999).

$$P = P_f + \frac{1}{3TV} \sum_{(i,j)} (\mathbf{q}_i \cdot \mathbf{r}_i + \mathbf{q}_j \cdot \mathbf{r}_j)$$

(Virial theorem)



- With attractive orbit, particle trajectories are bended toward denser region.

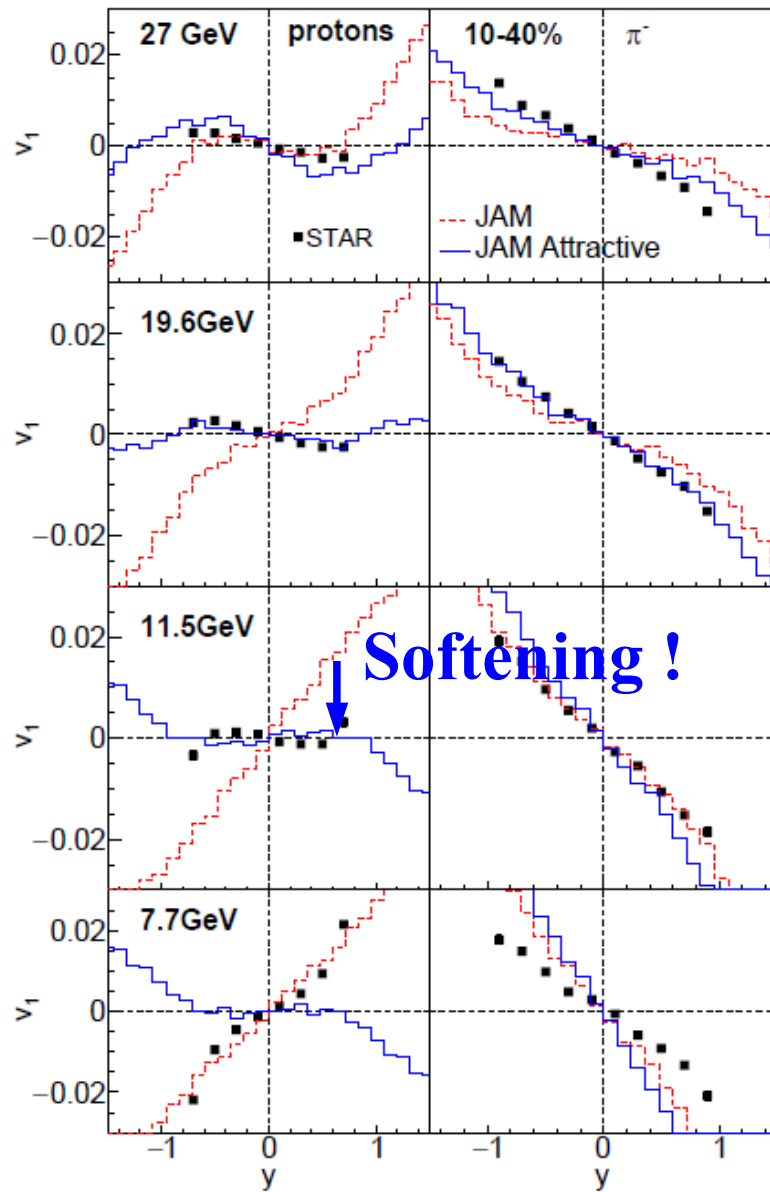
→ *Attractive orbit scattering simulates time evolution with softer EOS !*

Let us examine the EOS softening effects, which cannot be explained in hadronic mean field potential, by using attractive orbit scatterings !

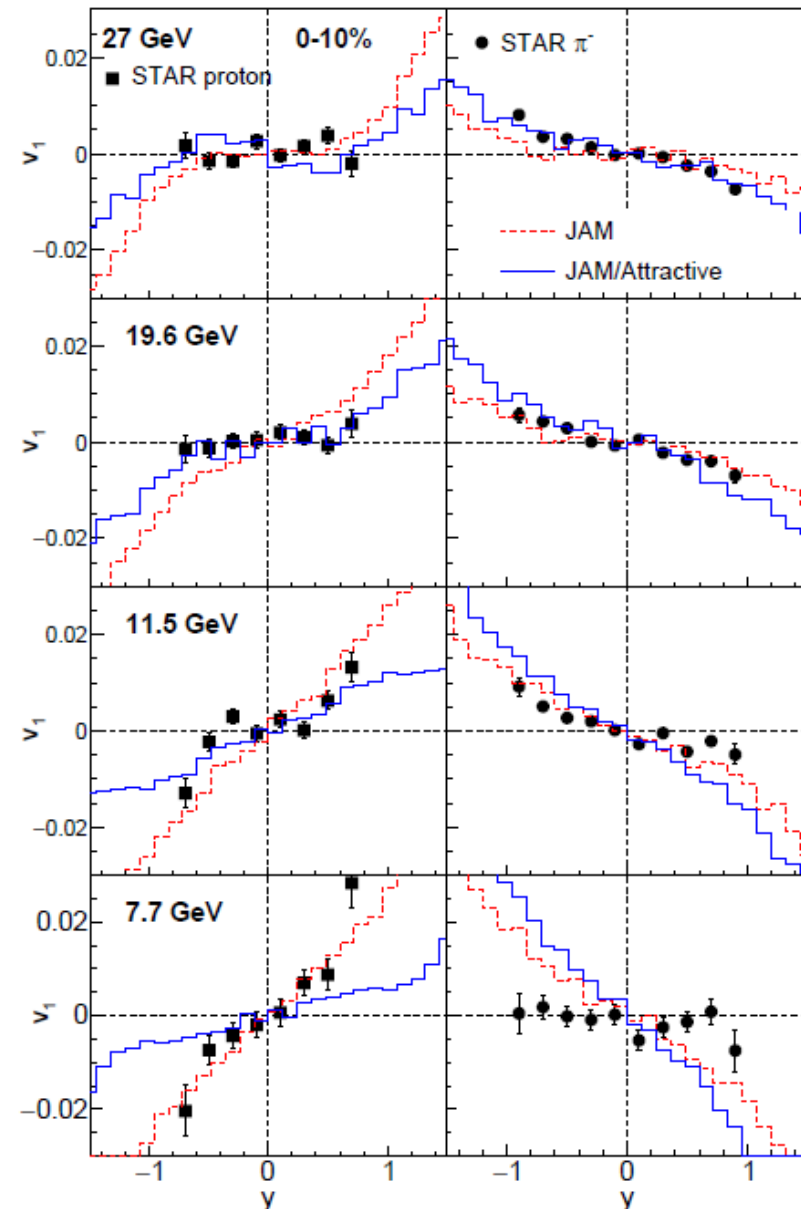
Y. Nara, H. Niemi, AO, H. Stöcker, arXiv:1601.07692 [hep-ph]

Directed Flow with Attractive Orbits

Nara, Niemi, AO, Stöcker ('16)

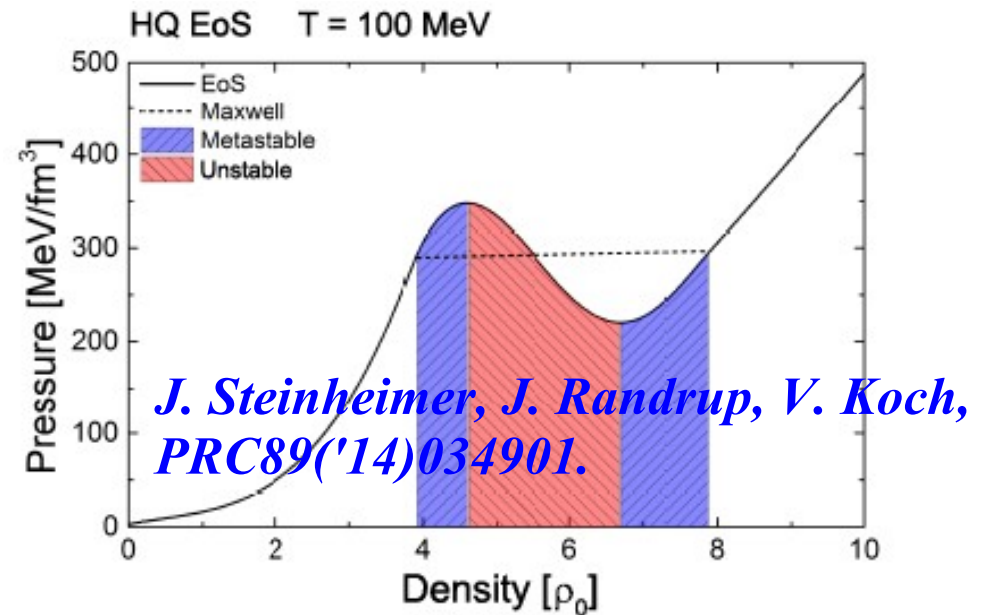
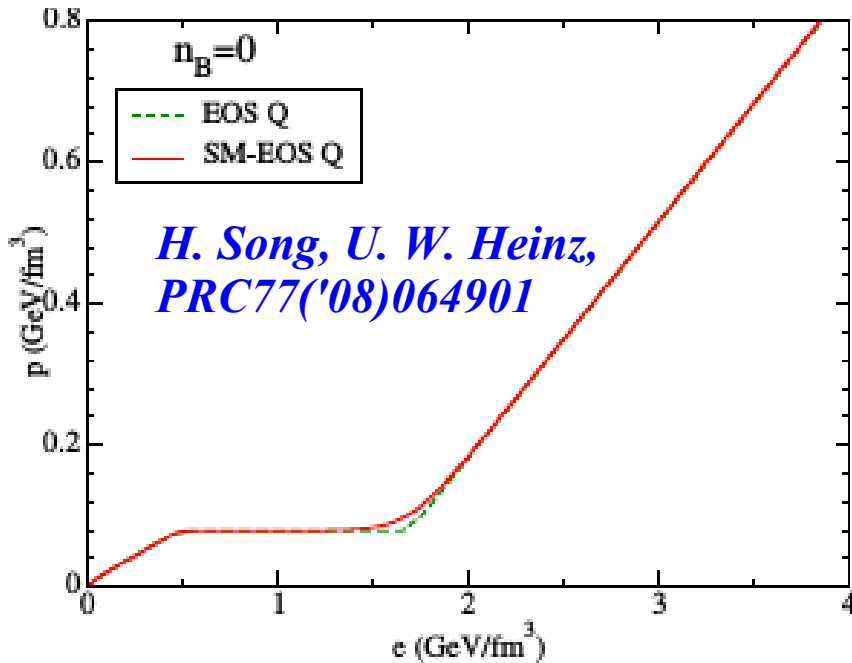
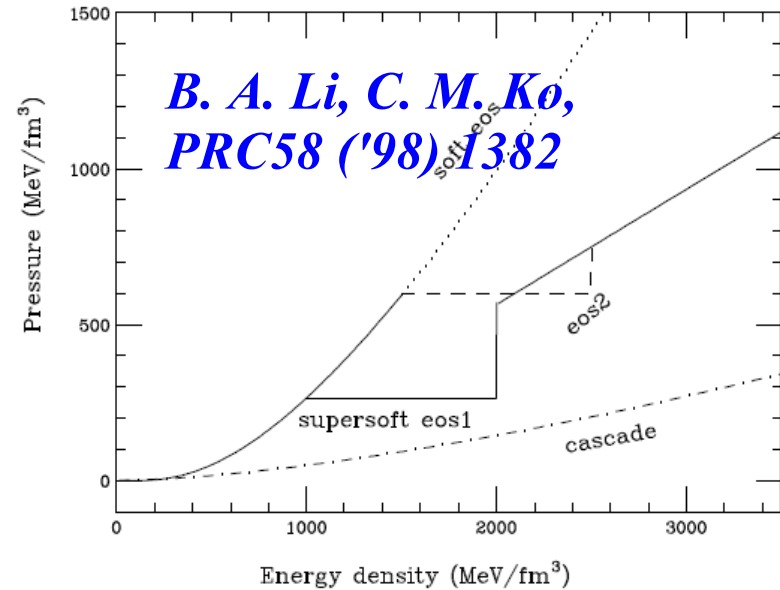
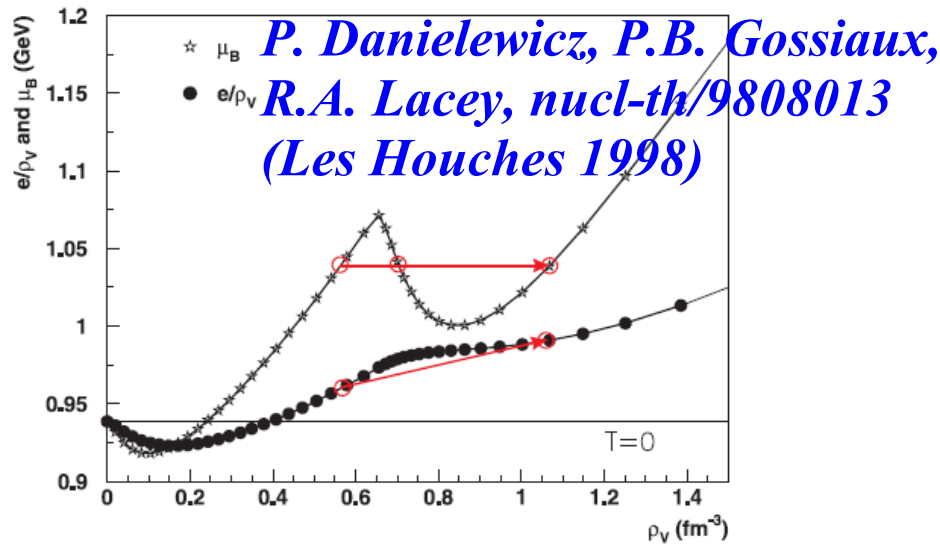


mid-central (10-40 %)



central (0-10 %)

Softening: Where and How much ?

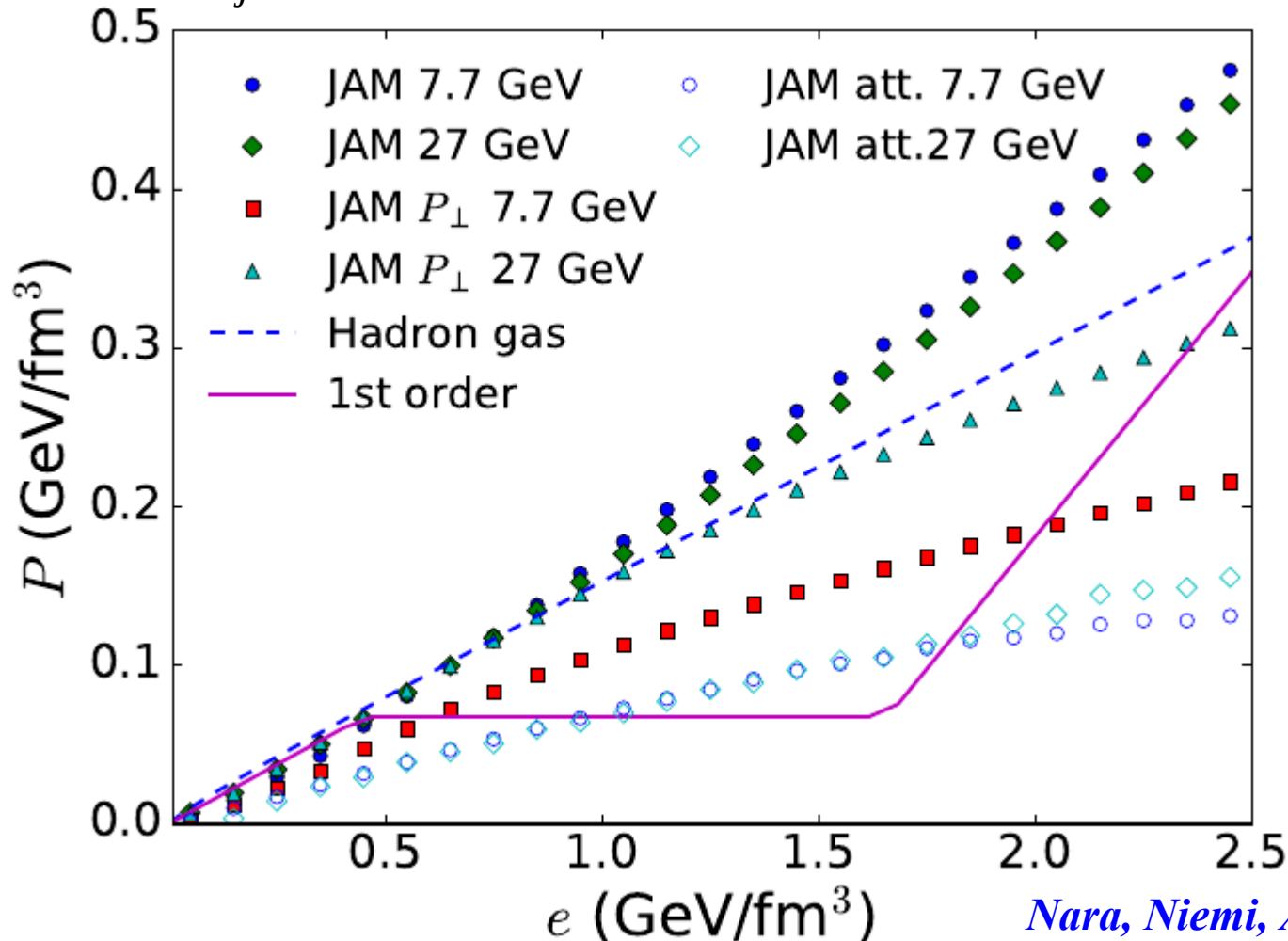


Previous analyses: $\rho_B = (3-10) \rho_0$, $P = (80-700) \text{ MeV/fm}^3$

Softening of EOS by Attractive Orbits

$$\Delta P = -\frac{\rho}{3(\delta\tau_i + \delta\tau_j)} (p_i' - p_i)^\mu (x_i - x_j)_\mu$$

H. Sorge, PRL82('99)2048.



Nara, Niemi, AO, Stöcker ('16)

Pressure in simulated EOS ~ EOS-Q (e.g. Song, Heinz ('08))

Summary

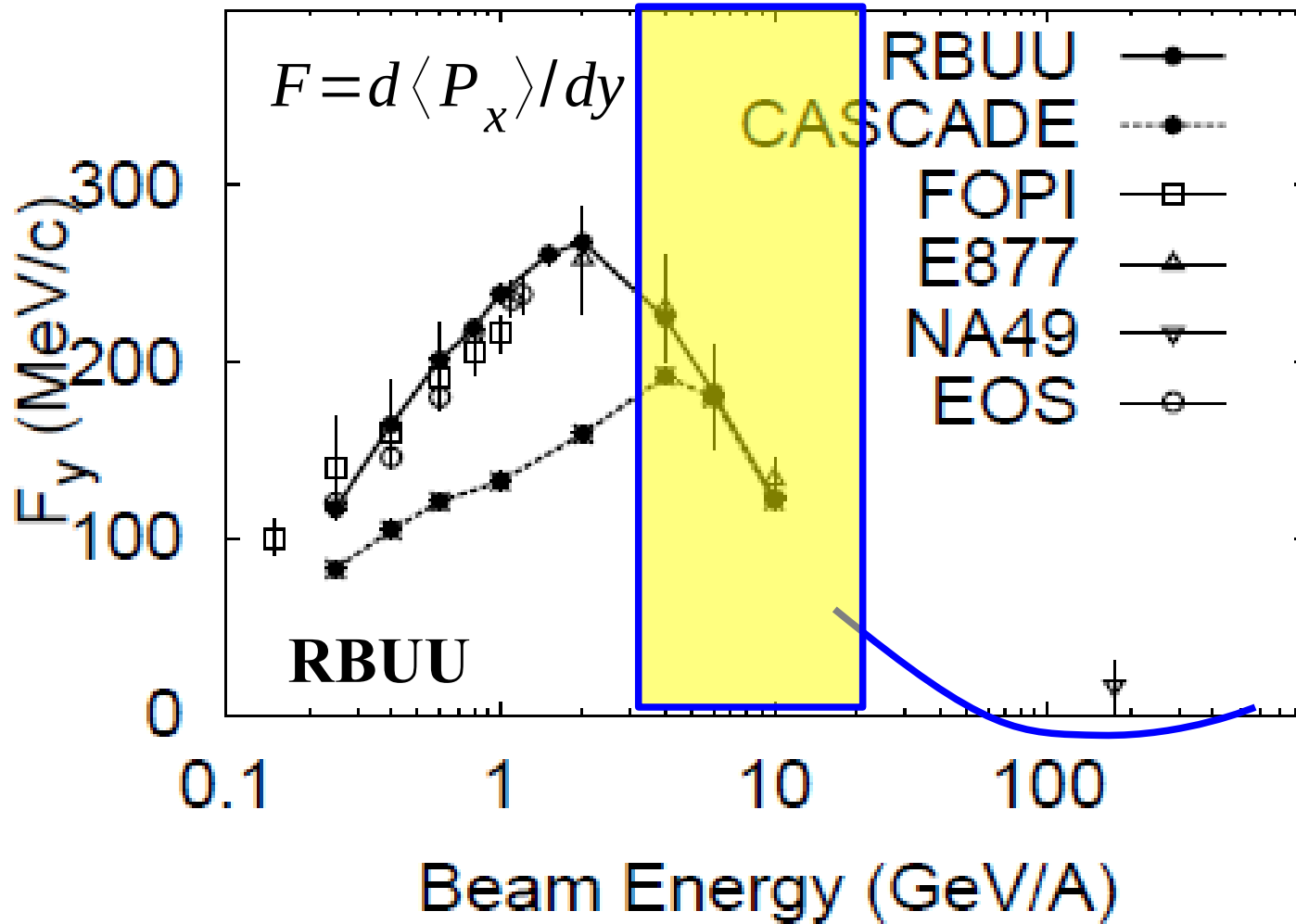
- We may see **QCD phase transition (1st or 2nd) signals at BES (or J-PARC) energies** in baryon number cumulants and v_1 slope.
- Hadronic transport models cannot explain negative v_1 slope below $\sqrt{s_{NN}} = 20$ GeV.
 - Geometric (bowling pin) mechanism becomes manifest at higher energies (JAM, JAM-MF, HSD, PHSD, UrQMD,).
- **Hadronic transport with EOS softening can describe negative v_1 slope below $\sqrt{s_{NN}} = 20$ GeV.**

Y. Nara, H. Niemi, A. Ohnishi, H. Stoecker, arXiv:1601.07692 [hep-ph]

 - **Attractive orbit scattering** simulates EOS softening (virial theorem).
 - We need more studies to confirm its nature.
First-order phase transition ? Crossover ? Forward-backward rapidities ? MF leading to softer EOS ?
- ***We need “re-hardening” at higher energies, e.g. $\sqrt{s_{NN}} = 27$ GeV.***

Thank you !

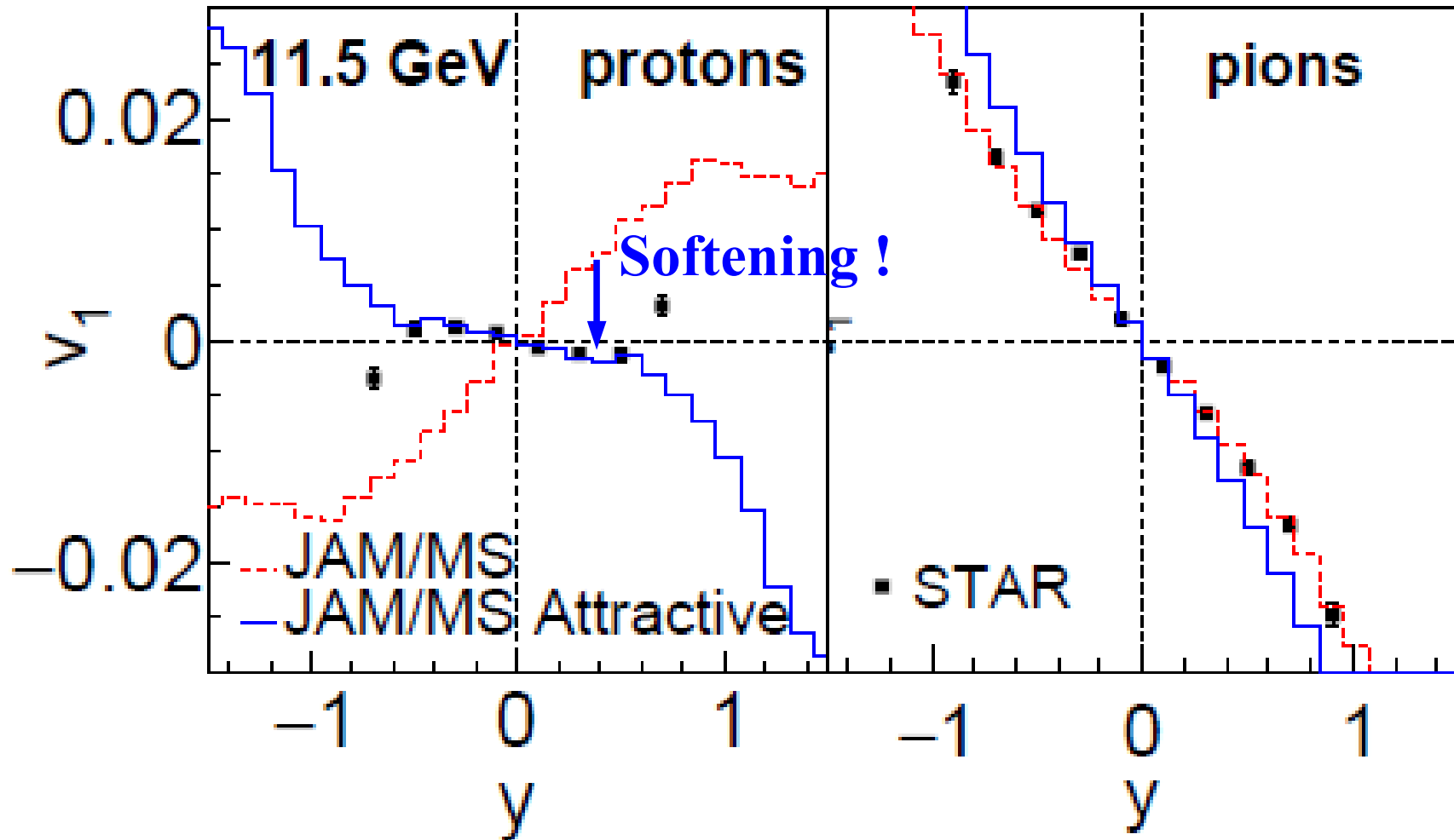
Directed Flow



P. K. Sahu, W. Cassing, U. Mosel, AO, Nucl. Phys. A 672 (2000),376

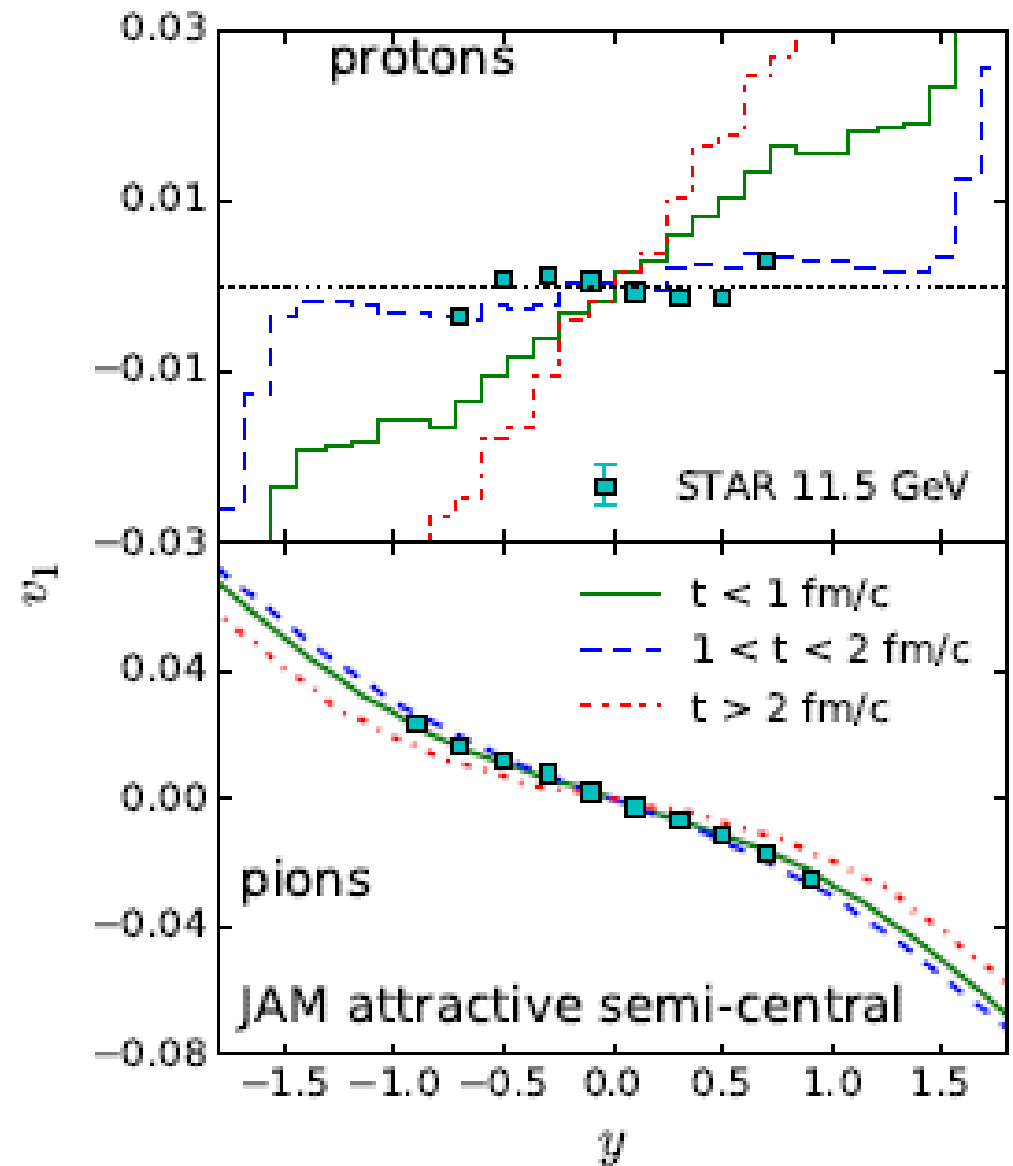
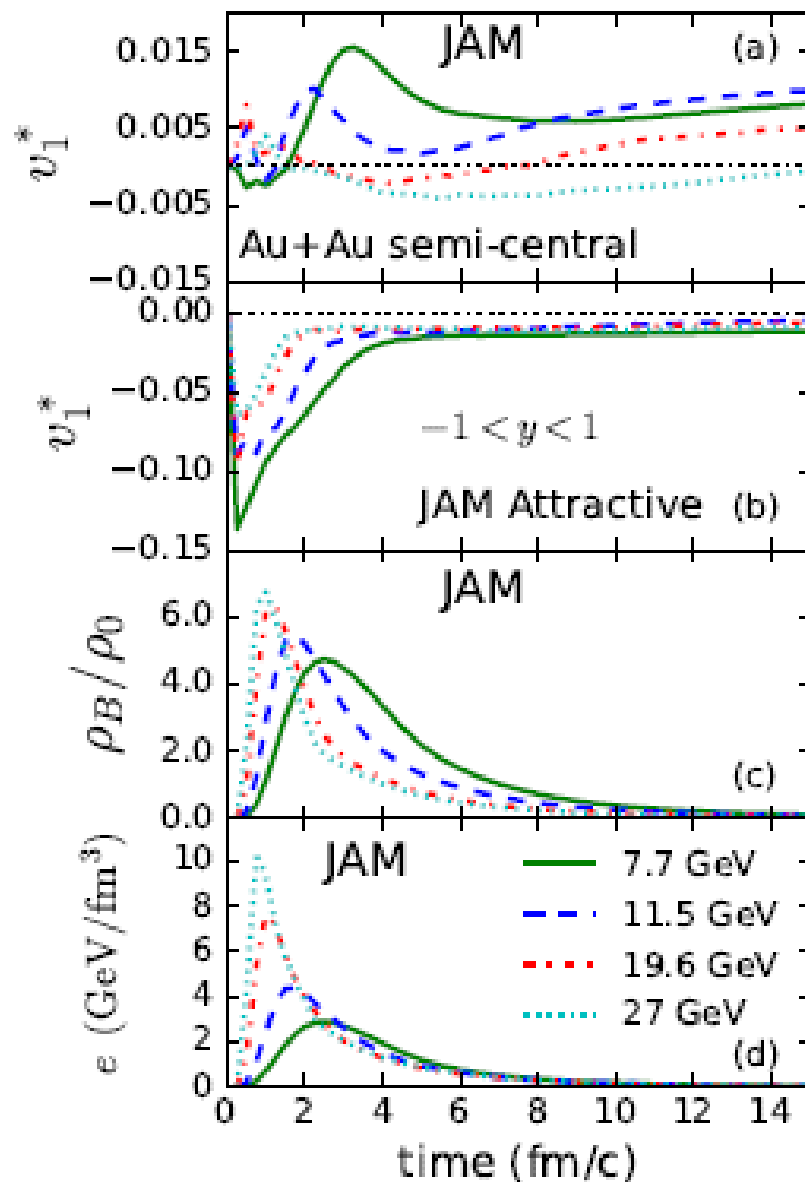
Mean Field + Attractive Orbit

Nara, Niemi, AO, Stöcker ('16)



MF+Attractive Orbit make dv_1/dy negative at $\sqrt{s_{NN}} \sim 10$ GeV

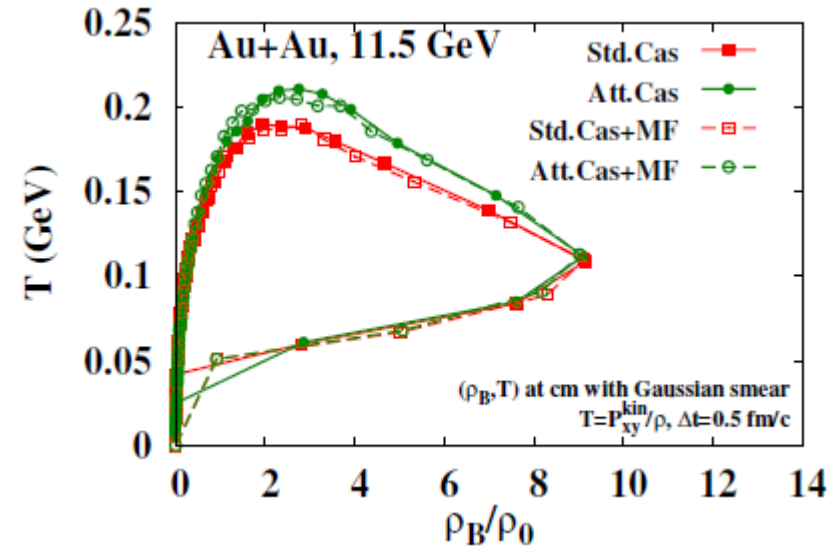
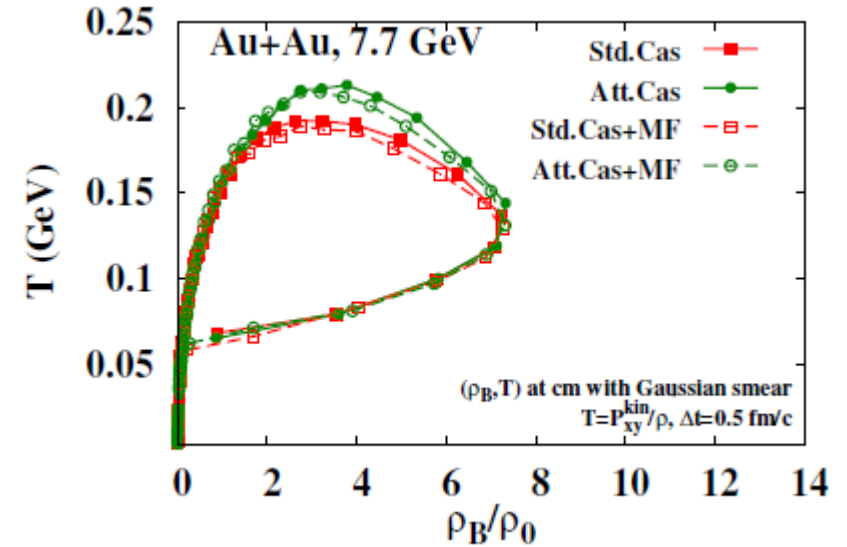
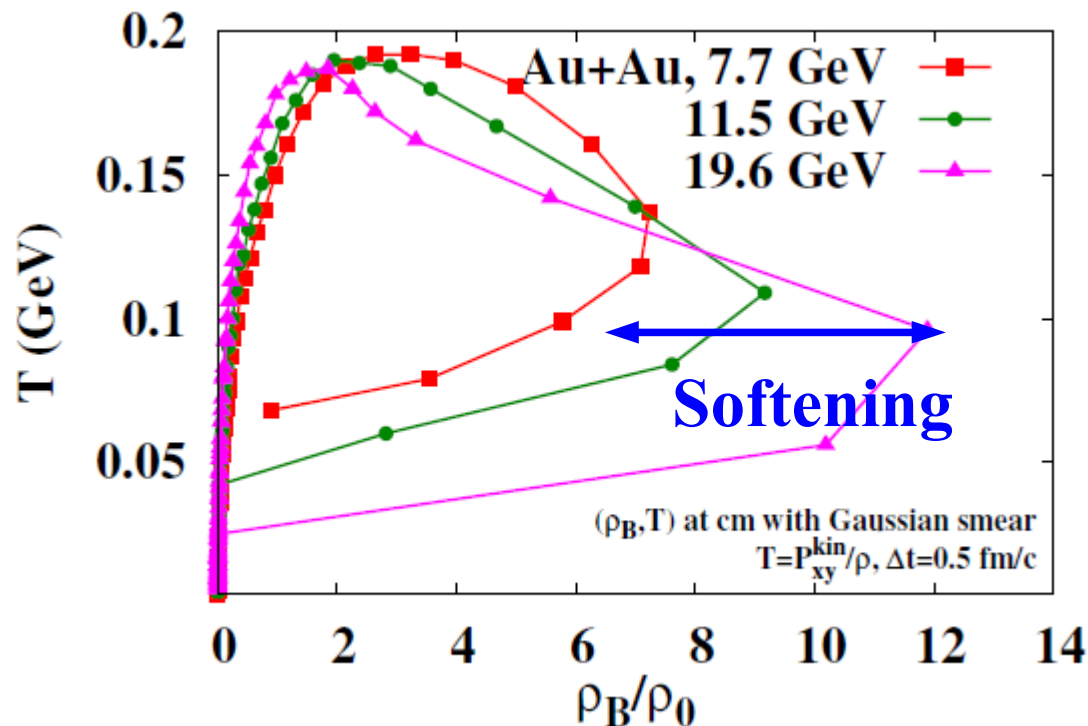
v_1 is sensitive to highest density regime



Nara, Niemi, AO, Stöcker ('16)

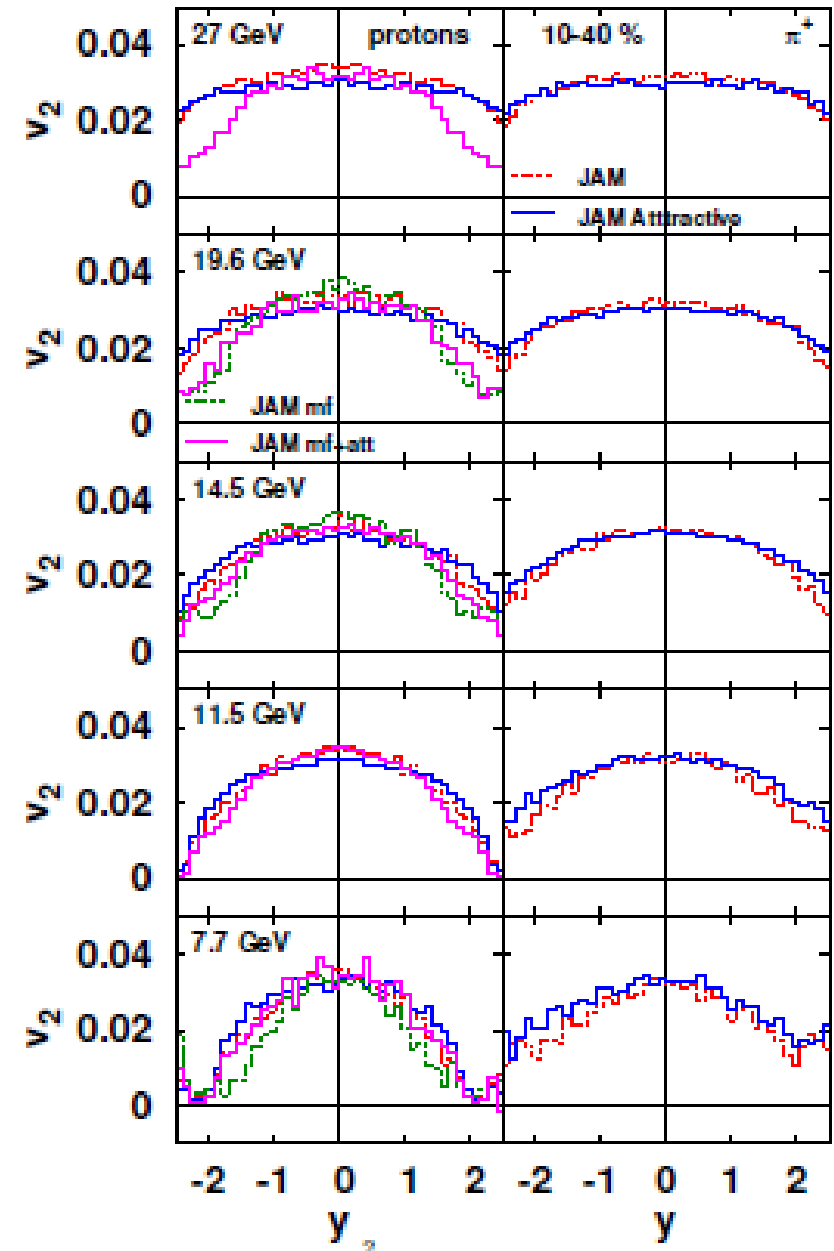
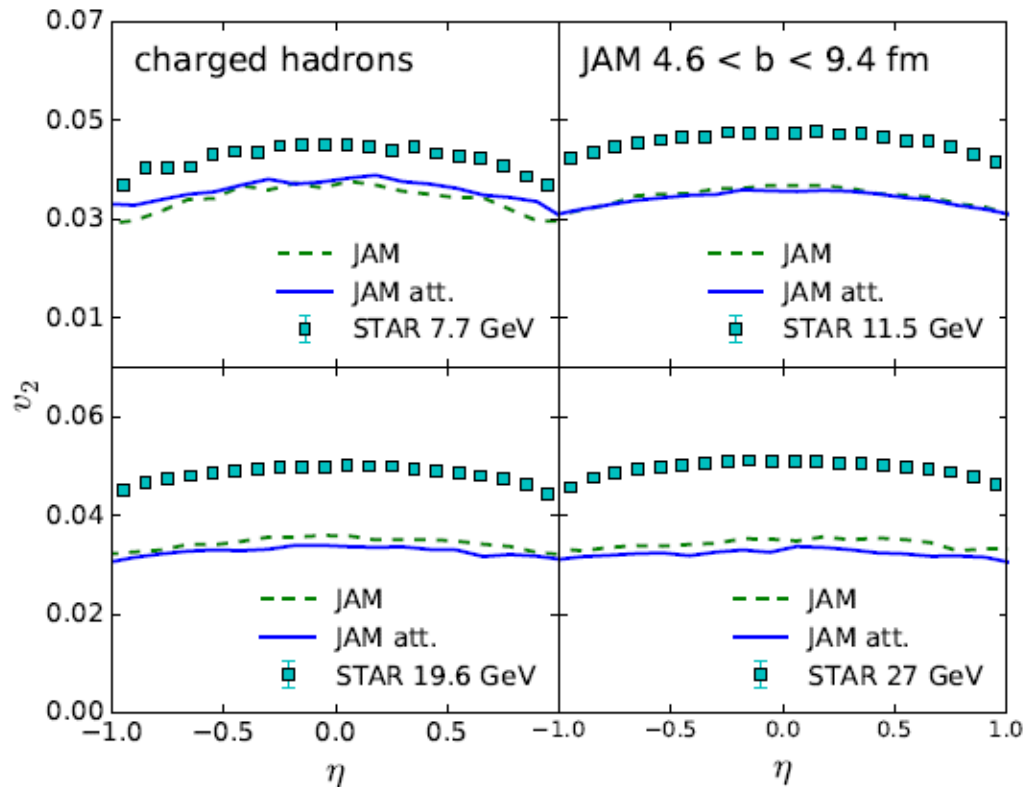
Softening of EOS: Where and How much ?

- “Softening” should take place at $\sqrt{s_{NN}}=11.5 \text{ GeV} \rightarrow \rho/\rho_B \sim (6-10)$
- Attractive orbit
 → Larger interactions
 & Higher T at later times



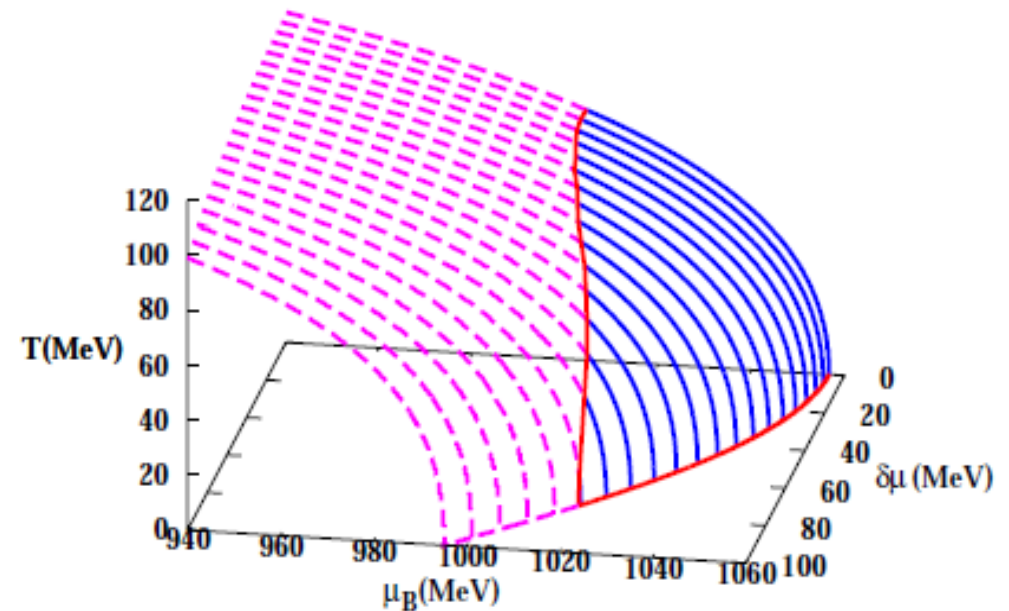
How about v_2 ?

- Do we see softening effects in other observables, e.g. v_2 ?
- Yes, attractive orbits reduces proton v_2 by ~ 0.2 %.
(but there is no qualitative change.)



Relation to Neutron Star Matter

- We may need early transition ($2-5 \rho_0$) to quark matter to solve the hyperon puzzle. Contradicting ?
 - Temperature effects ($T \sim 0 \text{ MeV} \text{ \& } 100 \text{ MeV}$)
 - Isospin chem. pot. (Weaker transition with finite $\delta\mu$)**
 - Hyperon repulsion may push up the transition density.

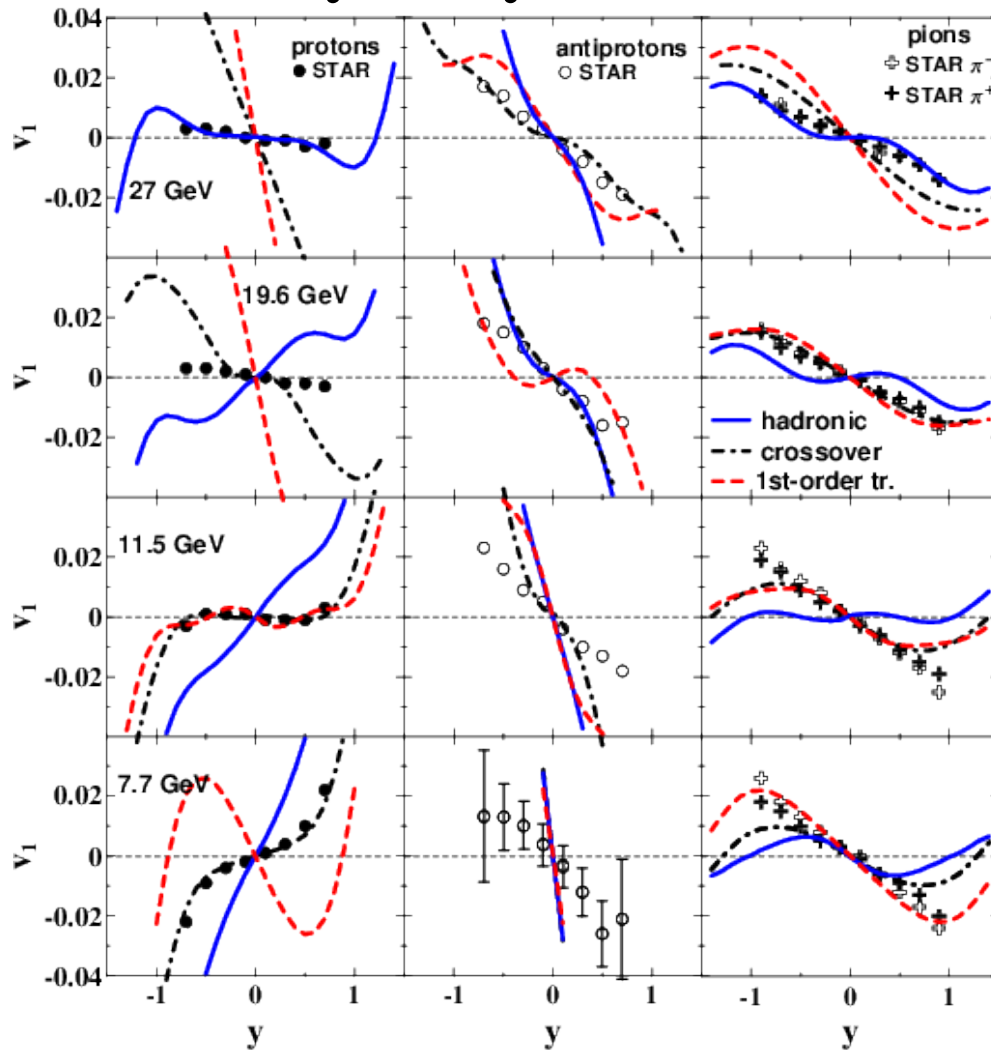


AO, Ueda, Nakano, Ruggieri, Sumiyoshi, PLB704('11),284

H. Ueda, T. Z. Nakano, AO, M. Ruggieri, K. Sumiyoshi, PRD88('13),074006

Negative dv_1/dy around $\sqrt{s_{NN}} \sim 10$ GeV

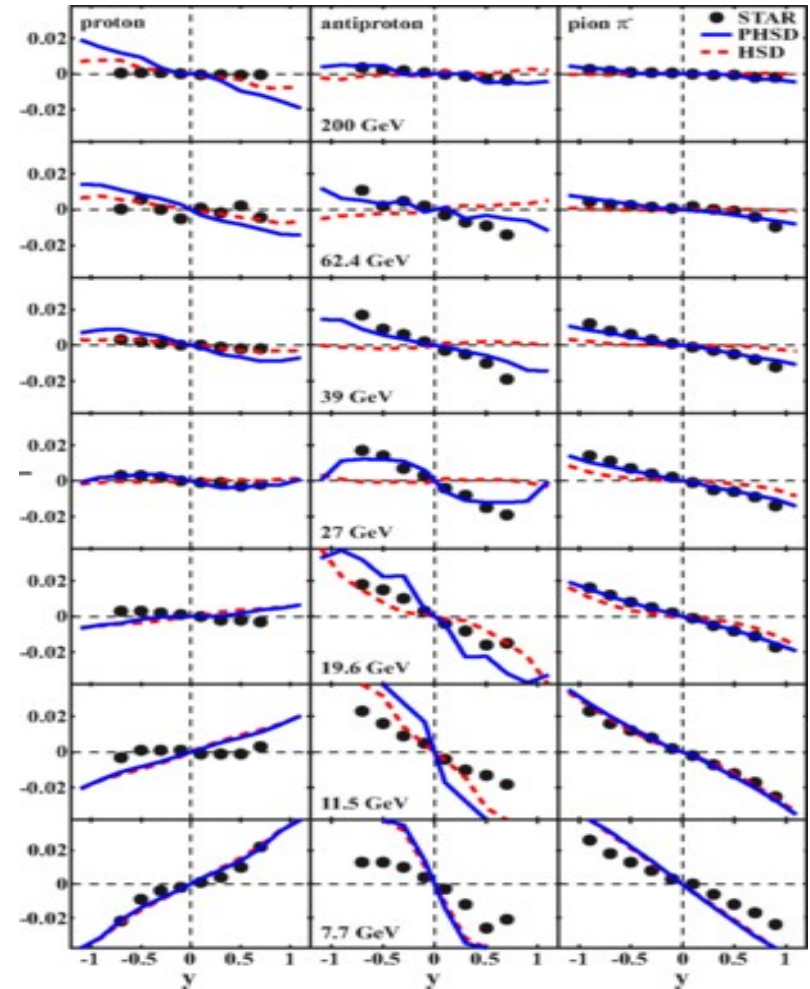
Yes in Hydrodynamics



Black: Crossover, Red: 1st

Y. B. Ivanov and A. A. Soldatov,
PRC91 (2015)024915

No at around $\sqrt{s_{NN}} \sim 10$ GeV in transport models.



V. P. Konchakovski, W. Cassing, Y. B. Ivanov,
V. D. Toneev, PRC90('14)014903