
Three Baryon Interaction in the Quark Cluster Model

– 3B Interaction from Determinant Interaction of Quarks as an example –

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YITP, Kyoto U.

QH seminar, Nov. 11, 2016

AO, K. Kashiwa, K. Morita, arXiv:1610.06306

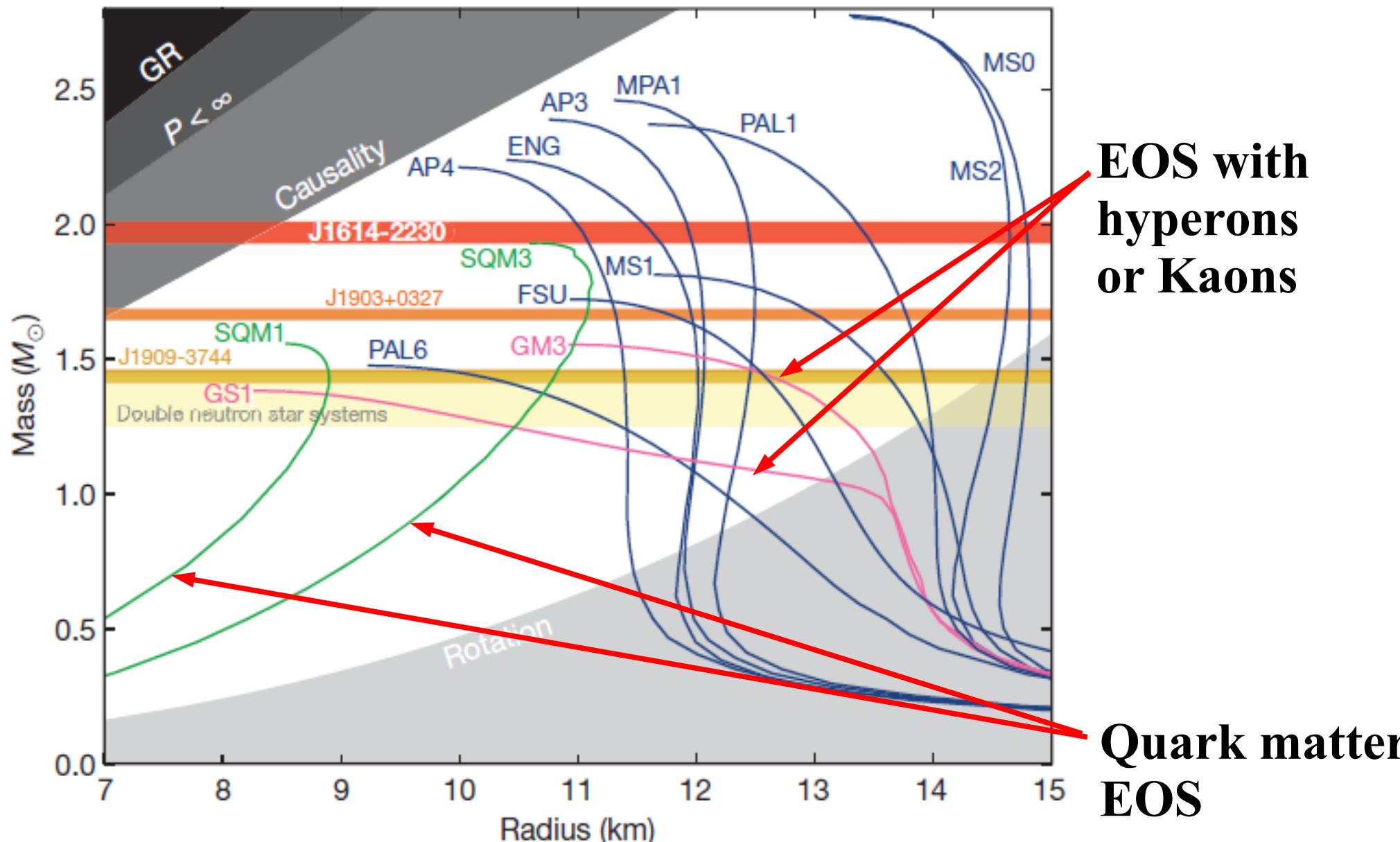


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from quark cluster model**
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- **Summary**

Hyperon Puzzle (or Hyperon Crisis)

Demorest et al., Nature 467 (2010) 1081 (Oct.28, 2010).



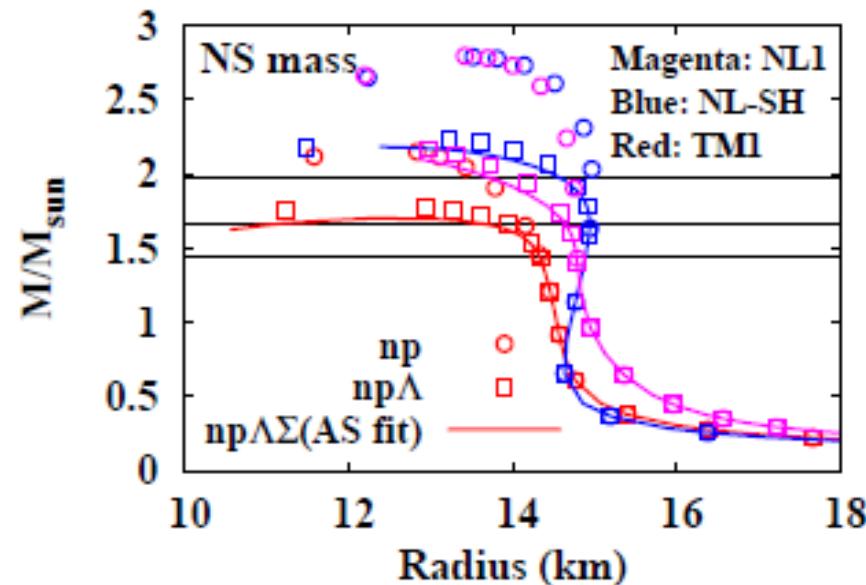
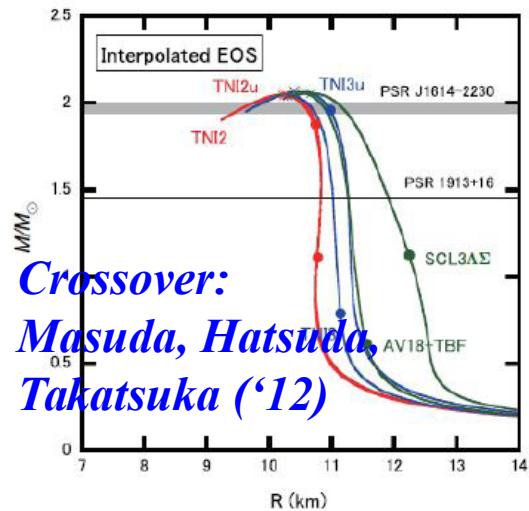
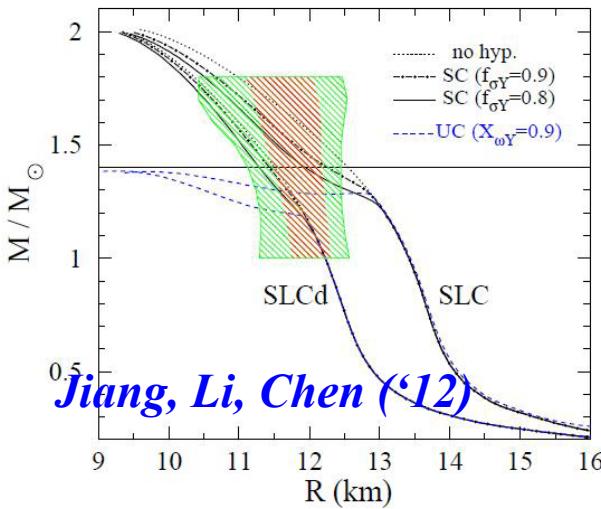
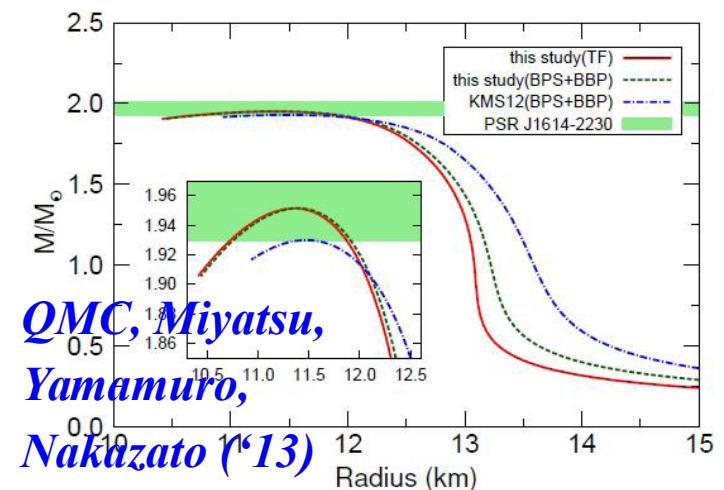
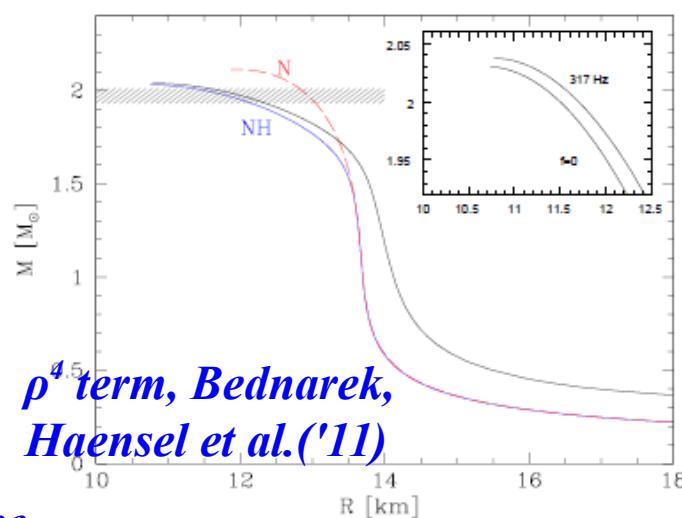
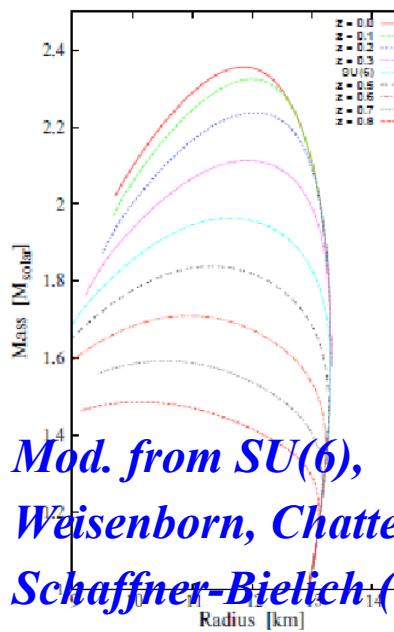
Proposed Solutions

- Hyperonic EOS cannot support massive NS ($M \sim 2 M_{\odot}$).
Demorest et al. (2010), Antoniadis et al. (2013)

- Proposed Solutions

- **Hyperonic Three-Body Force (or density dep. coupling)**
Bednarek et al. ('12), Jiang et al. ('12); Long et al. ('12); Yamamoto et al. ('14); Lonardoni et al. ('15); Tsubakihara et al. ('13), T. Miyatsu et al. ('13), ...
- **Crossover Transition to Quark Matter**
Bonanno et al. ('12); Masuda et al. ('13); Bejger et al. ('16), ...
- **Modified Gravity**
Astashenok et al. ('14)
- **Three-nucleon interaction is known to be necessary.**
How can we determine YNN (+YYN, YYY) potential ?

Massive Neutron Stars with Hyperons

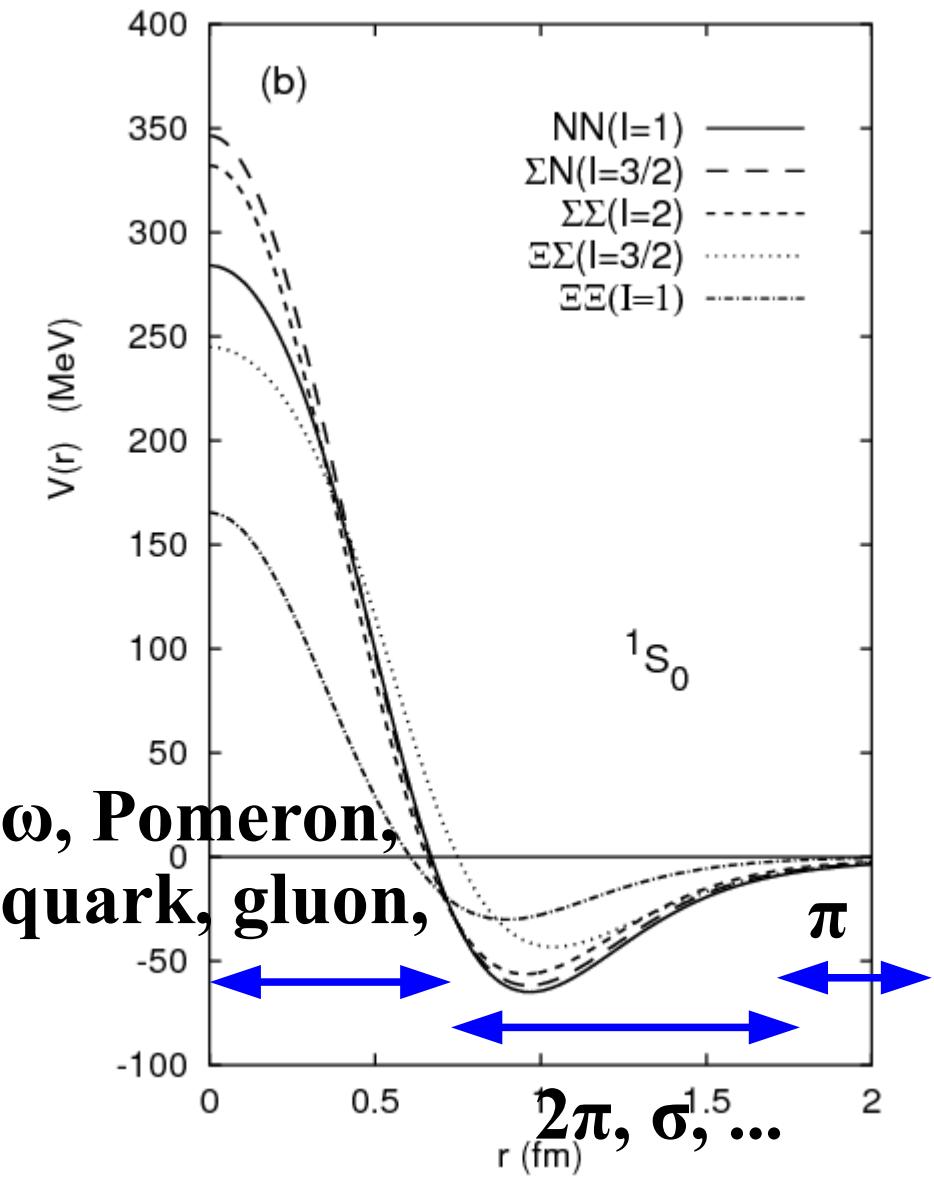


Tsubakihara, Harada, AO, arXiv:1402.0979

Baryon-Baryon Force

- Long-range ($r > 2$ fm): π exch.
 - Intermediate ($r \sim 1$ fm):
multi π exch., boson exch., ...
 - Short range ($r < 0.6$ fm):
vector boson exch.,
Pomeron exch.,
quark exclusion + one gluon exch.,
...
- V.G. Neudachin, Yu.F. Smirnov, R. Tamagaki,
PTP 58 ('77) 1072; M. Oka, K. Yazaki,
PTP ('81) 572.*

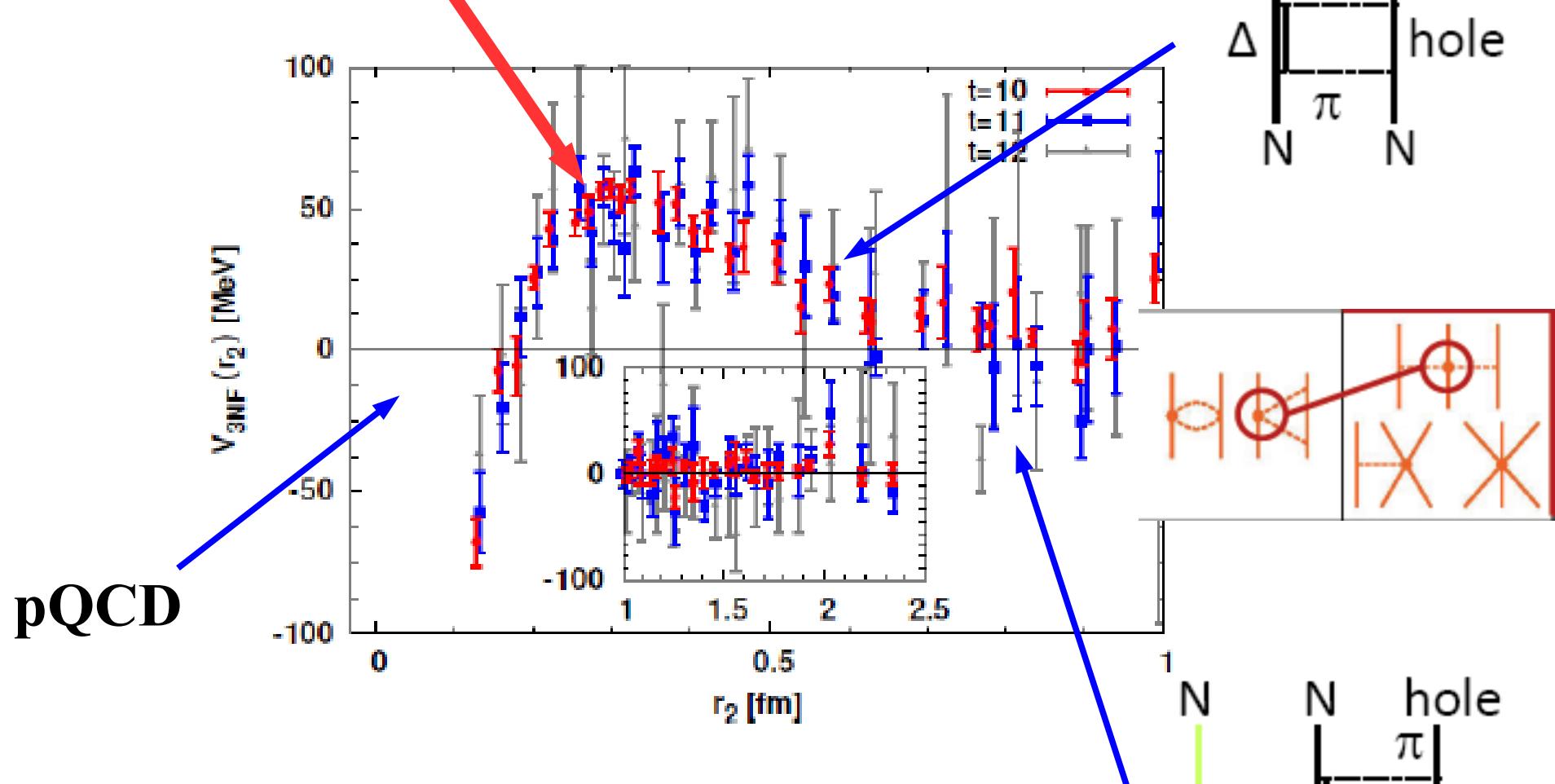
*Quark model description
of 3B repulsion should be
a promising approach !*



Fujiwara, Suzuki, Nakamoto ('07)

Three-Baryon force

What makes 3B repulsion at $r \sim 0.5$ fm ?



Taken from NPCSM 2016 talks,
Doi (Wed), Kohno (Thu), Tews (Thu)

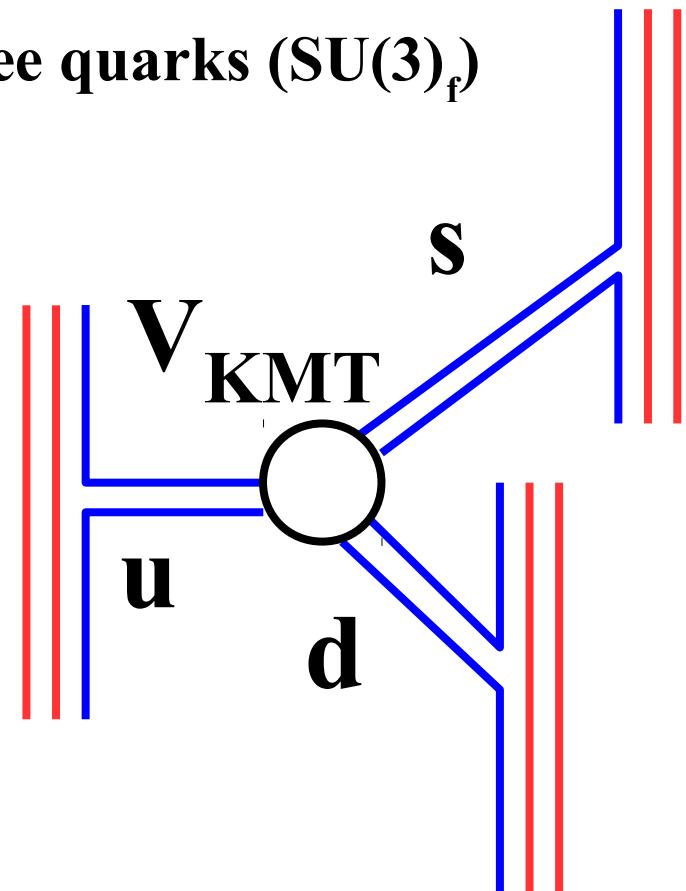
Kobayashi-Maskawa-'t Hooft (KMT) interaction

■ KMT interaction

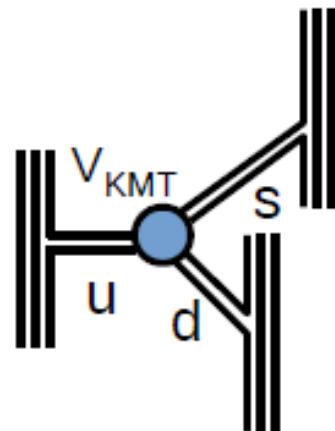
Kobayashi, Maskawa ('70), 't Hooft ('76)

$$\mathcal{L} = g_D \ (\det \Phi + \text{h.c.}) , \quad \Phi_{ij} = \bar{q}_j (1 - \gamma_5) q_i$$

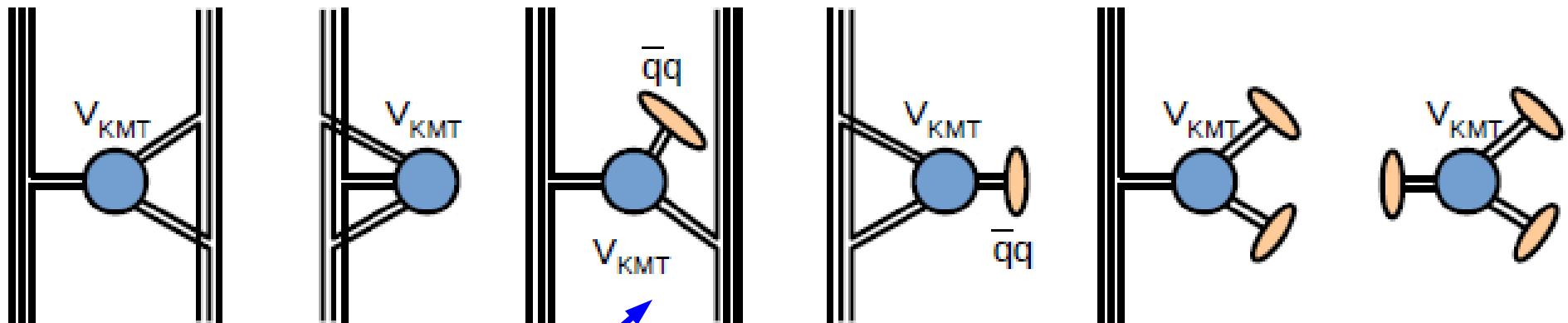
- Determinant interaction in flavor for three quarks ($SU(3)_f$)
- Responsible for $U(1)_A$ anomaly
 $\eta - \eta'$ mass diff.
→ $g_D = -9.29$ *Hatsuda, Kunihiro ('94)*
– 12.36 *Rehberg, Klevanski, Hufner ('96)*
- KMT interaction should generate
2B and 3B interaction
when hyperons are involved.
- Repulsive in $\Lambda\Lambda$ system
→ Pushes up H particle energy.
Takeuchi, Oka ('91)



Kobayashi-Maskawa-'t Hooft (KMT) interaction



3B force
(AO, Kashiwa, Morita)



Repulsion in $\Lambda\Lambda$ int.
Takeuchi, Oka ('91)

quark mass, vac. E.
(Hatsuda, Kunihiro)

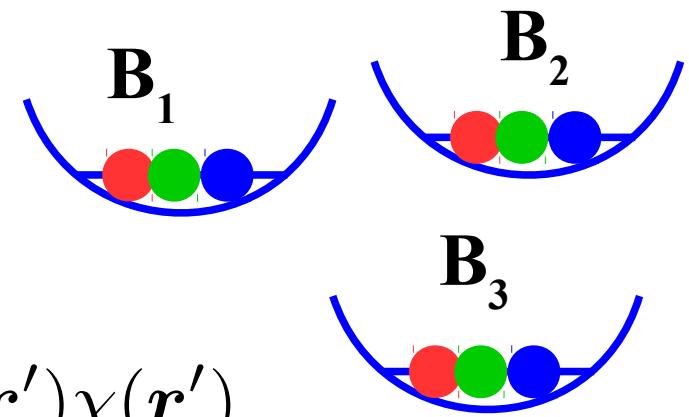
Does anomaly support massive NS ?

Quark Cluster model

- **Totally anti-symmetrized wave function of baryons**

$$|\Psi\rangle = \mathcal{A}|\chi_{12}B_1B_2\rangle$$

$$|\Psi\rangle = \mathcal{A}|\chi_{123}B_1B_2B_3\rangle$$



- **Resonating Group Method**

$$\int d\mathbf{r}' H(\mathbf{r}, \mathbf{r}') \chi(\mathbf{r}') = E \int d\mathbf{r}' N(\mathbf{r}, \mathbf{r}') \chi(\mathbf{r}')$$

$$\rightarrow -\frac{\hbar^2}{2\mu} \nabla^2 \chi^{(N)} + (V\chi^{(N)}) = E\chi^{(N)} \quad (\chi^{(N)} = \mathcal{N}^{1/2}\chi)$$

$$H(\mathbf{r}, \mathbf{r}') = \langle \mathbf{r} B_1 B_2 \dots | H | \mathcal{A}(\mathbf{r}' B_1 B_2 \dots) \rangle$$

$$N(\mathbf{r}, \mathbf{r}') = \langle \mathbf{r} B_1 B_2 \dots | \mathcal{A}(\mathbf{r}' B_1 B_2 \dots) \rangle$$

- **When (wave length of χ) \gg (baryon size),**

$$V(\mathbf{r}) \simeq \Delta K + \langle V\mathcal{A} \rangle / \langle \mathcal{A} \rangle$$

Norm Kernel

■ Antisymmetrizer makes the calculation complicated !

$$\begin{aligned} \mathcal{A} = & [1 - 9(P_{36} + P_{39} + P_{69}) + 27(P_{369} + P_{396}) \\ & + 54(P_{25}P_{39} + P_{35}P_{69} + P_{38}P_{69})] \mathcal{A}_B \\ & - 216P_{25}P_{38}P_{69}, \end{aligned}$$

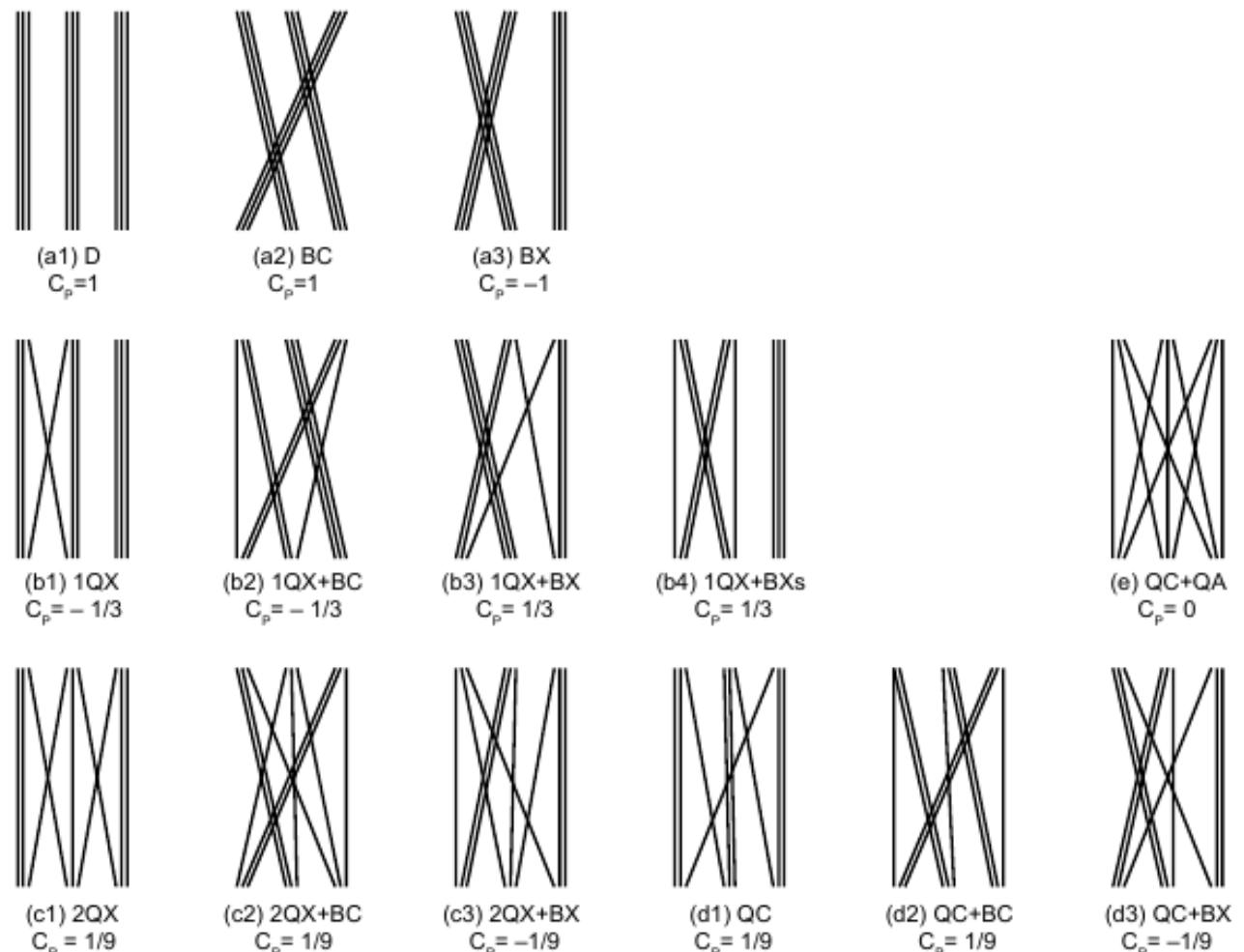
$$\mathcal{A}_B = \sum_{\mathcal{P}} (-1)^{\pi(\mathcal{P})} \mathcal{P}$$

Toki, Suzuki, Hecht ('82)

■ Recent work

Nakamoto, Suzuki ('16)

→ Norm kernel
of 3 octet B



Norm Kernel

■ **Single baryon w.f.** $|\psi_A\rangle = \mathcal{A}/\sqrt{3!} \times \epsilon_{abc}/\sqrt{3!} \times [|\text{Flavor}\rangle \otimes |\text{Spin}\rangle \otimes |\text{Spatial w.f.}\rangle]^{abc}$.

■ **Two barvon w.f.**

$$|\psi_A(n_\uparrow, n_\downarrow)\rangle = \frac{1}{\sqrt{6!}} |\mathcal{A}[\psi(n_\uparrow)\psi(n_\downarrow)]\rangle$$

Norm

$$\begin{aligned} \mathcal{N}_A &= \langle \psi_A(n_\uparrow, n_\downarrow) | \psi_A(n_\uparrow, n_\downarrow) \rangle \\ &= \langle \psi(n_\uparrow)\psi(n_\downarrow) | \mathcal{A}[\psi(n_\uparrow)\psi(n_\downarrow)] \rangle \\ &= \frac{1}{(3!)^2} \sum_{i,j,k,l} c_i^*(n_\uparrow)c_j^*(n_\downarrow)c_k(n_\uparrow)c_l(n_\downarrow) \epsilon_{abc} \epsilon_{def} \epsilon_{a'b'c'} \epsilon_{d'e'f'} \\ &\quad \times \langle \phi_i^{abc}(n_\uparrow)\phi_j^{def}(n_\downarrow) | \mathcal{A}[\phi_k^{a'b'c'}(n_\uparrow)\phi_l^{d'e'f'}(n_\downarrow)] \rangle \\ &= \sum_{i,j,k,l} c_i^*(n_\uparrow)c_j^*(n_\downarrow)c_k(n_\uparrow)c_l(n_\downarrow) \sum_P C_P(\phi_i\phi_j, \phi_k\phi_l) F_P(\phi_i\phi_j, \phi_k\phi_l) \end{aligned}$$

B	$ \text{Flavor}\rangle$	$ \text{Spin}\rangle$
n_\uparrow	$ ddu\rangle/\sqrt{2}$	$ \uparrow\downarrow\uparrow + \downarrow\uparrow\uparrow - 2\uparrow\uparrow\downarrow\rangle/\sqrt{6}$
p_\uparrow	$ uud\rangle/\sqrt{2}$	$ \uparrow\downarrow\uparrow + \downarrow\uparrow\uparrow - 2\uparrow\uparrow\downarrow\rangle/\sqrt{6}$
Λ_\uparrow	$ uds\rangle$	$ \uparrow\downarrow\uparrow - \downarrow\uparrow\uparrow\rangle/\sqrt{2}$
Σ_\uparrow^-	$ dds\rangle/\sqrt{2}$	$ \uparrow\downarrow\uparrow + \downarrow\uparrow\uparrow - 2\uparrow\uparrow\downarrow\rangle/\sqrt{6}$
Σ_\uparrow^0	$ uds\rangle$	$ \uparrow\downarrow\uparrow + \downarrow\uparrow\uparrow - 2\uparrow\uparrow\downarrow\rangle/\sqrt{6}$
Σ_\uparrow^+	$ uus\rangle/\sqrt{2}$	$ \uparrow\downarrow\uparrow + \downarrow\uparrow\uparrow - 2\uparrow\uparrow\downarrow\rangle/\sqrt{6}$
Ξ_\uparrow^-	$ ssd\rangle/\sqrt{2}$	$ \uparrow\downarrow\uparrow + \downarrow\uparrow\uparrow - 2\uparrow\uparrow\downarrow\rangle/\sqrt{6}$
Ξ_\uparrow^0	$ ssu\rangle/\sqrt{2}$	$ \uparrow\downarrow\uparrow + \downarrow\uparrow\uparrow - 2\uparrow\uparrow\downarrow\rangle/\sqrt{6}$

Anti-symmetrization

$$\mathcal{A}[1^a 1^b 1^c 2^d 2^e 2^f] = 1^a 1^b 1^c 2^d 2^e 2^f - 1^a 1^b 2^d 1^c 2^e 2^f + 1^a 2^e 2^d 1^c 1^b 2^f + \dots$$

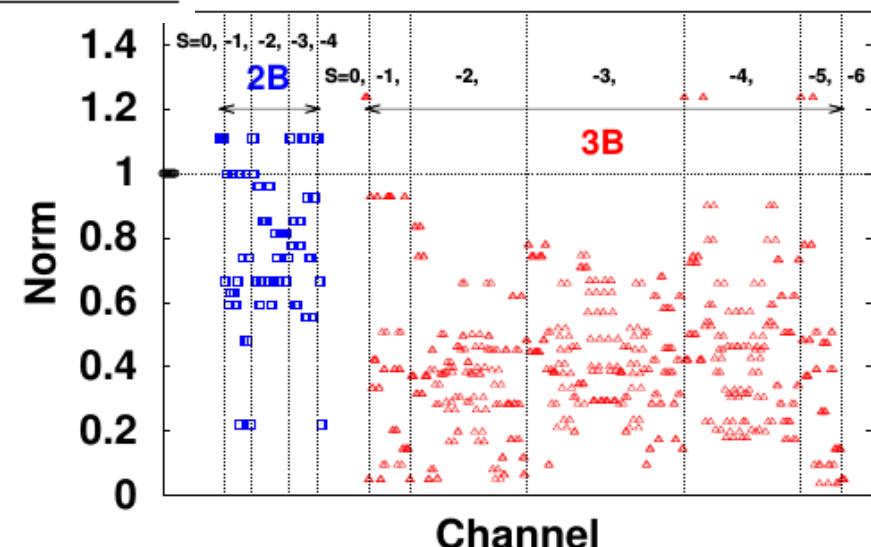
$$C_P = -\frac{1}{(3!)^2} \epsilon_{abd} \epsilon_{cef} \epsilon_{abc} \epsilon_{def} = -\frac{1}{36} 2\delta_{dc} 2\delta_{cd} = -\frac{1}{3}$$

$$F_P(\phi_i\phi_j, \phi_k\phi_l) = \langle \phi_i(n_\uparrow)\phi_j(n_\downarrow) | P[\phi_k(n_\uparrow)\phi_l(n_\downarrow)] \rangle_{\text{fss}} = 0 \text{ or } 1 \quad \leftarrow \text{Numerical}$$

Norm Kernel

Baryon(s)	$\mathcal{N}_{\mathcal{A}}$	$\mathcal{T}_{\mathcal{A}}$	\mathcal{T}	$\mathcal{T}_{nB}(n = 2, 3)$
$(NN)_{(S,T)=(0,1),(1,0)}$	10/9	0	0	0
$N_{\uparrow}\Lambda_{\uparrow}, N_{\downarrow}\Lambda_{\downarrow}$	1	20/3	20/3	20/3
$N_{\uparrow}\Lambda_{\downarrow}, N_{\downarrow}\Lambda_{\uparrow}$	1	10/3	10/3	10/3
$(\Lambda\Lambda)_{S=0}$	1	18/3	18/3	18/3
$(NNN)_{(S,T)=(1/2,1/2)}$	100/81	0	0	0
$n_{\uparrow}n_{\downarrow}\Lambda, p_{\uparrow}p_{\downarrow}\Lambda$	25/27	350/27	14	12/3
$n_{\uparrow}p_{\uparrow}\Lambda_{\uparrow}, n_{\downarrow}p_{\downarrow}\Lambda_{\downarrow}$	25/27	750/27	30	50/3
$n_{\uparrow}p_{\uparrow}\Lambda_{\downarrow}, n_{\downarrow}p_{\downarrow}\Lambda_{\uparrow}$	25/27	250/27	10	10/3
$n_{\uparrow}p_{\downarrow}\Lambda, n_{\downarrow}p_{\uparrow}\Lambda$	25/27	425/27	17	21/3
$N\Lambda_{\uparrow}\Lambda_{\downarrow}$	45/54	1035/54	23	21/3

Not very small



AO, Kashiwa, Morita ('16)

KMT matrix element

■ Reduction of KMT interaction to 3 quark pot.

$$V_{\text{KMT}} \simeq -2g_{\text{D}} \int d^3x \varepsilon_{ijk} u^\dagger(\mathbf{x}) q_i(\mathbf{x}) d^\dagger(\mathbf{x}) q_j(\mathbf{x}) s^\dagger(\mathbf{x}) q_k(\mathbf{x})$$
$$= -2g_{\text{D}} \varepsilon_{ijk} \sum_{\{\alpha, \beta, \gamma\}} \hat{T}_\alpha^{u,i} \hat{T}_\beta^{d,j} \hat{T}_\gamma^{s,k} \delta(x_\alpha - x_\beta) \delta(x_\beta - x_\gamma)$$

■ Flavor exchanging operator

$$\hat{\mathcal{T}}^{\text{KMT}} = \sum_{\{\alpha, \beta, \gamma\}} \varepsilon_{ijk} \hat{T}_\alpha^{u,i} \hat{T}_\beta^{d,j} \hat{T}_\gamma^{s,k}$$

$$\mathcal{T}_A \equiv \langle \psi_A | \hat{\mathcal{T}}^{\text{KMT}} | \psi_A \rangle \quad \quad \mathcal{T} = \mathcal{T}_A / \mathcal{N}_A$$

■ Subtract the two-body part

$$\mathcal{T}_{3B}(n_\uparrow n_\downarrow \Lambda_\uparrow) = \mathcal{T}(n_\uparrow n_\downarrow \Lambda_\uparrow) - \mathcal{T}(n_\uparrow \Lambda_\uparrow) - \mathcal{T}(n_\downarrow \Lambda_\uparrow)$$

KMT matrix element

$$\begin{aligned}
 \langle \phi | V_{\text{KMT}} | \phi' \rangle &= \sum_{\{\alpha, \beta, \gamma\}} \langle q_\alpha q_\beta q_\gamma | V_{\text{KMT}} | q'_\alpha q'_\beta q'_\gamma \rangle \prod_{i \neq \{\alpha, \beta, \gamma\}} \langle q_i | q'_i \rangle \quad \text{irrelevant} \\
 &\quad \text{quarks} \\
 \text{product w.f.} &= -2g_D \langle \sigma | \sigma' \rangle \sum_{\{\alpha, \beta, \gamma\}} F_{\alpha\beta\gamma}^{\text{KMT}}(f, f') R_{\alpha\beta\gamma}^{\text{KMT}}(\varphi, \varphi') ,
 \end{aligned}$$

$$\langle \sigma | \sigma' \rangle = \prod_\alpha \langle \sigma_\alpha | \sigma'_\alpha \rangle ,$$

**flavor matching
(numerical)**

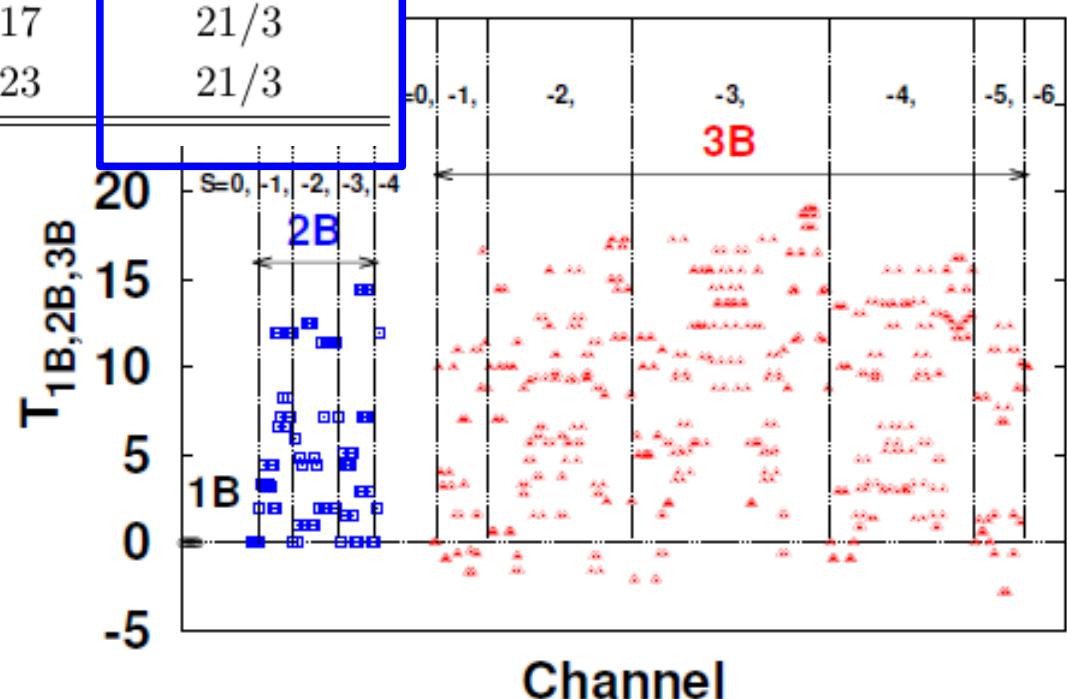
$$\begin{aligned}
 F_{\alpha\beta\gamma}^{\text{KMT}}(f, f') &= \langle f | \varepsilon_{ijk} \hat{T}_\alpha^{u,i} \hat{T}_\beta^{d,j} \hat{T}_\gamma^{s,k} | f' \rangle \\
 &= \delta_{u,f_u} \delta_{d,f_d} \delta_{s,f_s} \sum_{ijk} \varepsilon_{ijk} \boxed{\delta_{i,f'_\alpha} \delta_{j,f'_\beta} \delta_{k,f'_\gamma}} \prod_{\mu \neq \{\alpha, \beta, \gamma\}} \delta_{f_\mu, f'_\mu} ,
 \end{aligned}$$

$$\begin{aligned}
 R_{\alpha\beta\gamma}^{\text{KMT}}(\varphi, \varphi') &= \langle \varphi | \delta(x_\alpha - x_\beta) \delta(x_\beta - x_\gamma) | \varphi' \rangle \\
 &= \int d^3x \varphi_\alpha^*(x) \varphi_\beta^*(x) \varphi_\gamma^*(x) \varphi'_\alpha(x) \varphi'_\beta(x) \varphi'_\gamma(x) \prod_{\mu \neq \alpha, \beta, \gamma} \langle \varphi_\mu | \varphi'_\mu \rangle .
 \end{aligned}$$

KMT matrix element

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**Big for np Λ
(S=3/2)**



KMT matrix elements
strongly depend
on the channel

3B potential from KMT interaction

■ 3q int. → 3B potential

$$V_{3B}^{KMT} = -2g_D T_{3B} \int d^3x \varphi_{R_a}^*(x) \varphi_{R_b}^*(x) \varphi_{R_c}^*(x) \varphi_{R_d}(x) \varphi_{R_e}(x) \varphi_{R_f}(x)$$

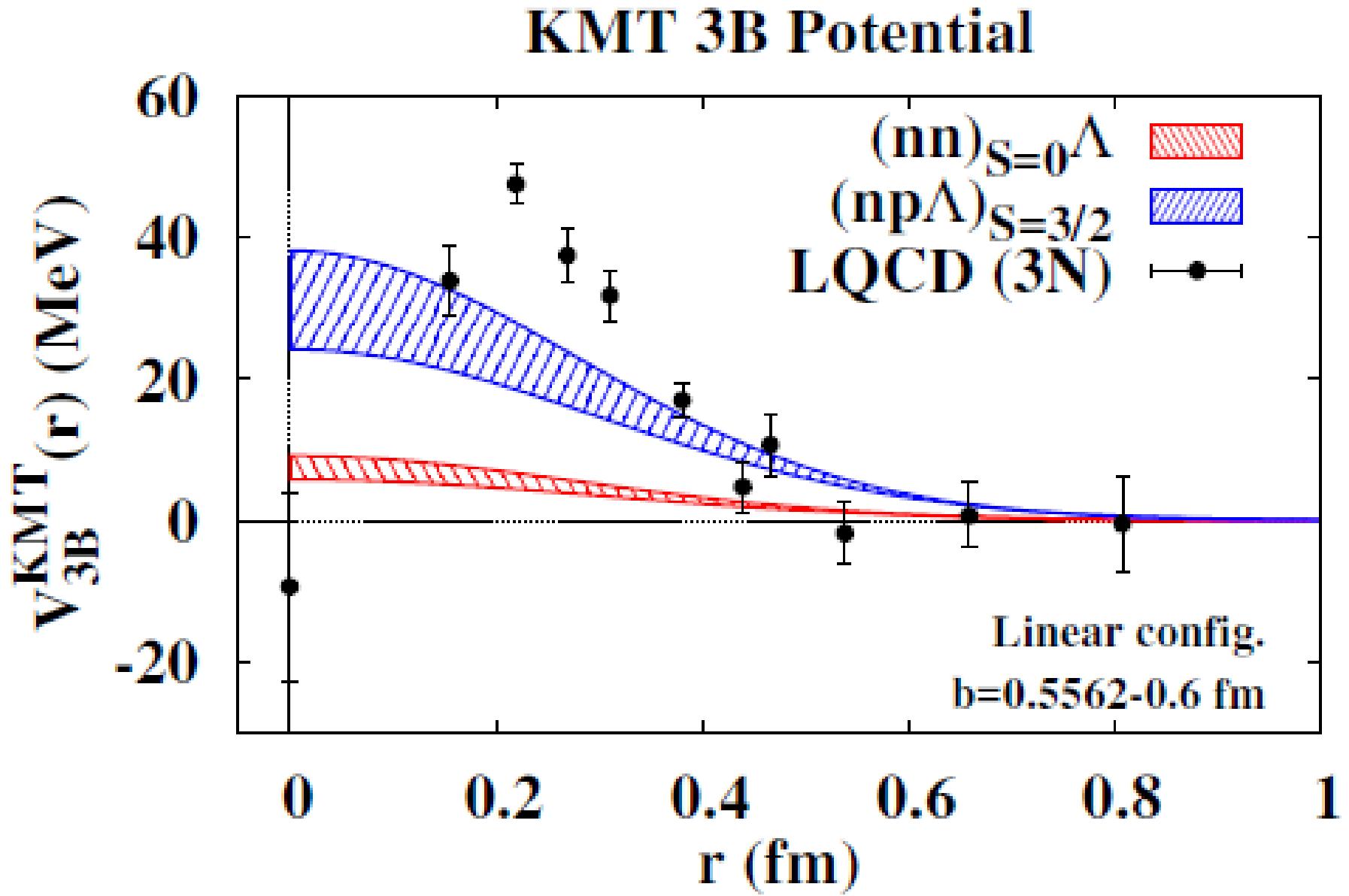
$$V_{3B}^{KMT}(R_1, R_2, R_3) \simeq V_0 T_{3B} \exp \left[-\frac{2\nu}{3} (R_{12}^2 + R_{23}^2 + R_{31}^2) \right]$$

$$V_0 \equiv \frac{-2g_D}{(\sqrt{3}\pi b^2)^3} = \frac{-2g_D \Lambda^5}{(\sqrt{3}\pi b^2 \Lambda^2)^3} \quad \Lambda = \begin{cases} 1.45 \text{ MeV} & (b = 0.6 \text{ fm}) , \\ 2.29 \text{ MeV} & (b = 0.5562 \text{ fm}) . \end{cases}$$

Parameters are taken from

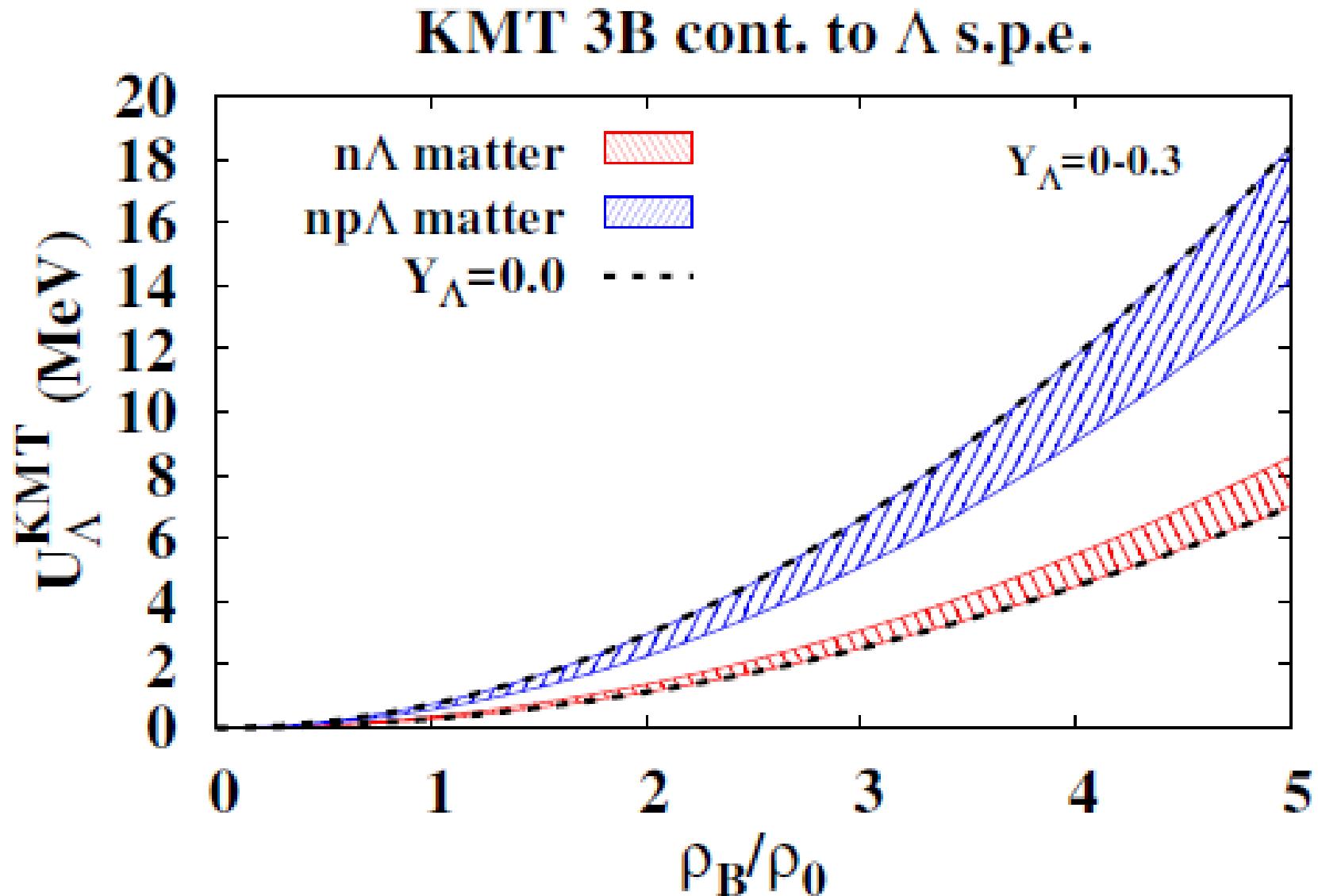
*Hatsuda, Kunihiro ('94), Rehberg, Klevanski, Hufner ('96),
Fujiwara, Suzuki, Nakamoto ('07), Oka, Yazaki ('81)*

3B potential from KMT interaction



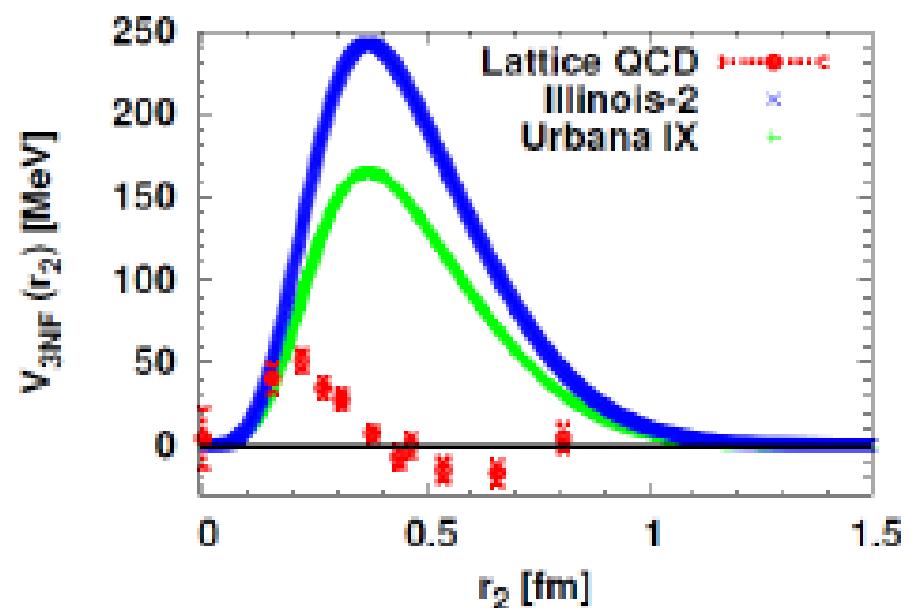
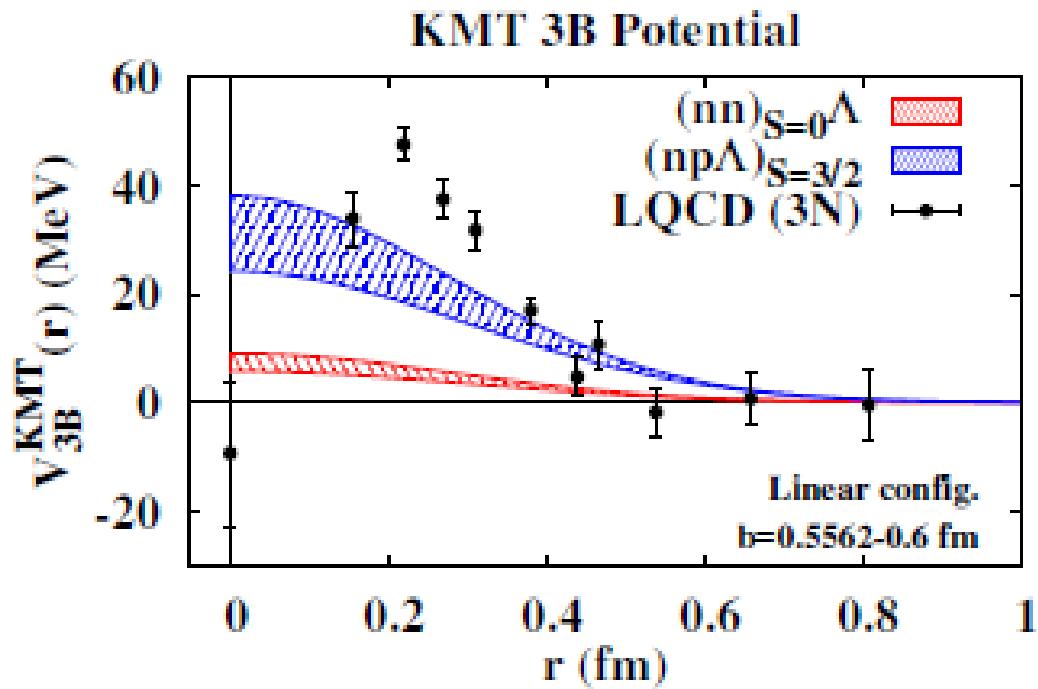
Lattice data: Doi et al. (HAL QCD) ('07)

KMT-3B Contribution to Λ potential



Density is assumed to be uniform. No correlation effects.

3B potential from KMT: Repulsive enough ?



Summary

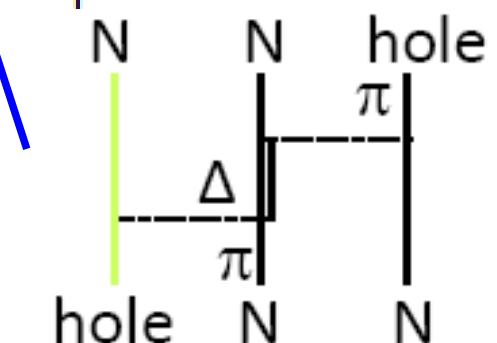
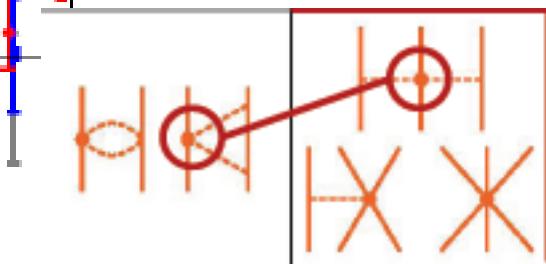
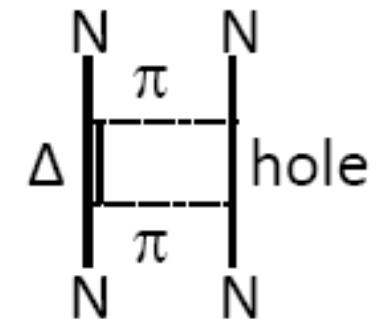
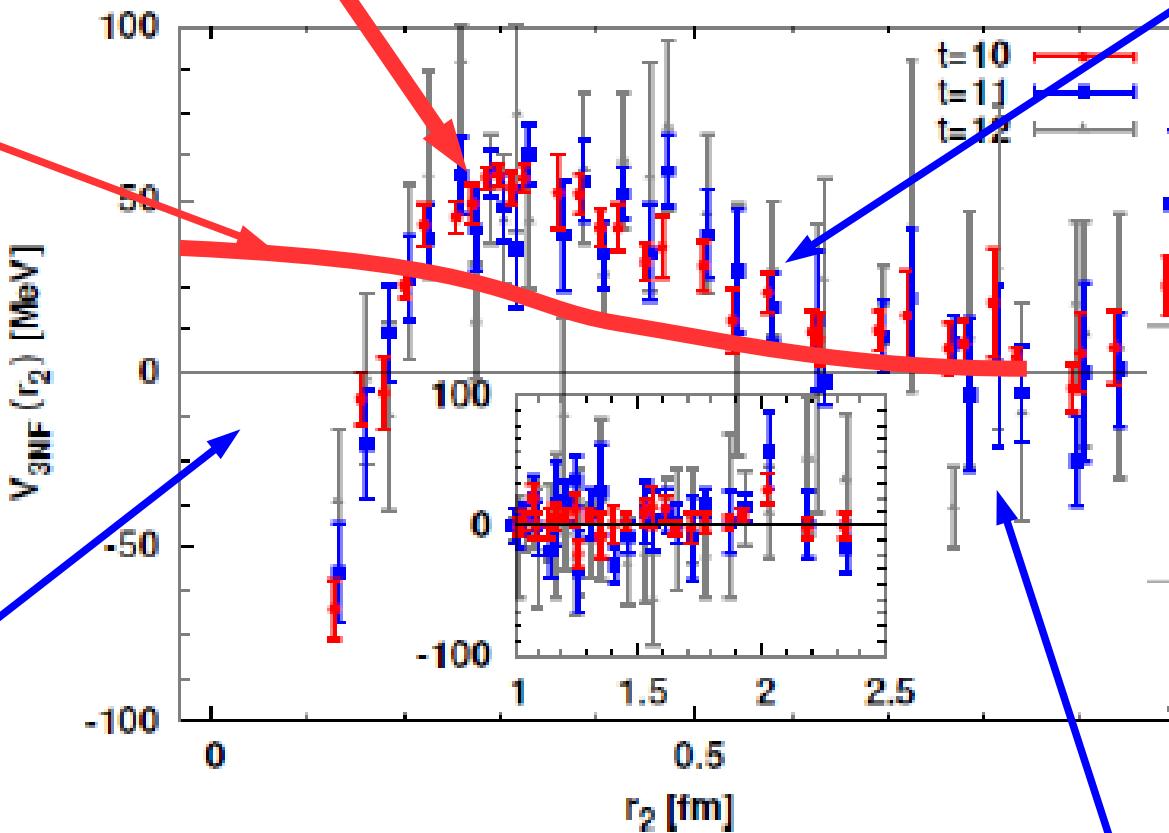
- Quark model three-baryon (3B) potential may be a promising method to evaluate the 3B potential at short distances.
- Kobayashi-Maskawa-'t Hooft (KMT) interaction generates 3q potential among u,d,s quarks, and generates 3B potential only when hyperons are involved.
- Expectation value of the KMT interaction is evaluated in the cases where 3B are located at the same spatial point.
Matrix elements strongly depend on the baryon trio.
- 3B potential from KMT interaction is obtained.
 - It is comparable in strength to the lattice 3N potential.
 - More repulsive in np Λ than in nn Λ
(Negative contribution to symmetry energy.)
- 3B pot. from KMT is not strong enough to solve the hyperon puzzle, but contributes to hyperon suppression.

Three-Baryon force

What makes 3B repulsion at $r \sim 0.5$ fm ?

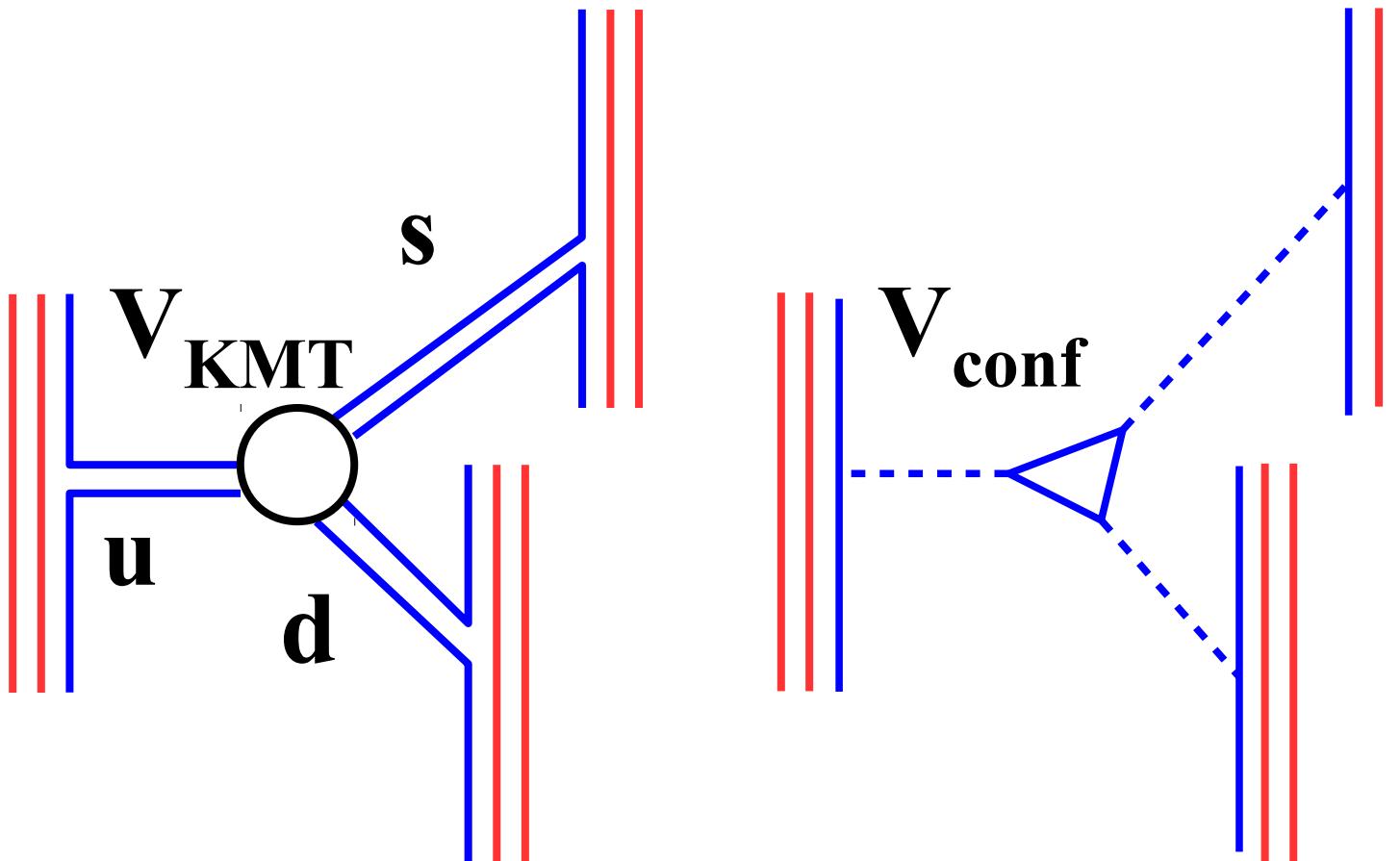
KMT
(Λ NN)

pQCD



Taken from NPCSM 2016 talks,
Doi (Wed), Kohno (Thu), Tews (Thu)

Confinement Potential \rightarrow 3B Potential ?



$$V_{\text{conf}} = \sum_{\{\alpha, \beta, \gamma\}} \varepsilon_{abc} \varepsilon_{a'b'c'} f(x_\alpha, x_\beta, x_\gamma)$$

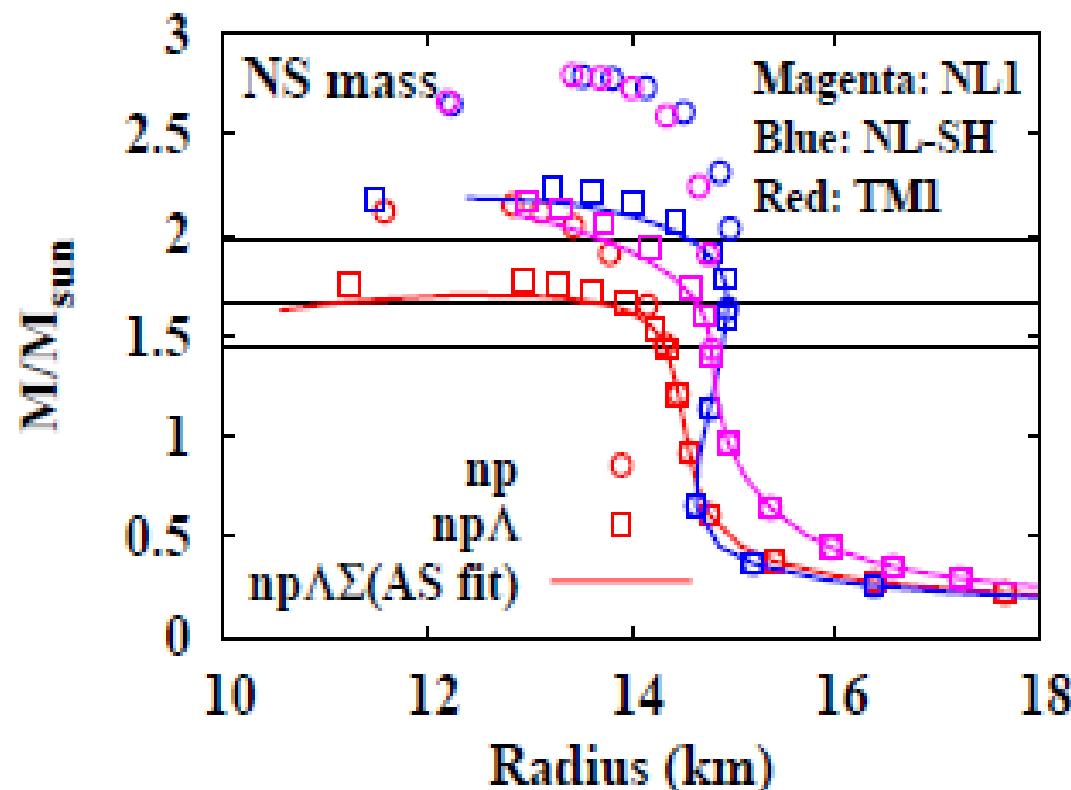
Takahashi, Suganuma, Nemoto, Matsufuru ('02)

Thank you for your attention !

Massive Neutron Stars with Hyperons

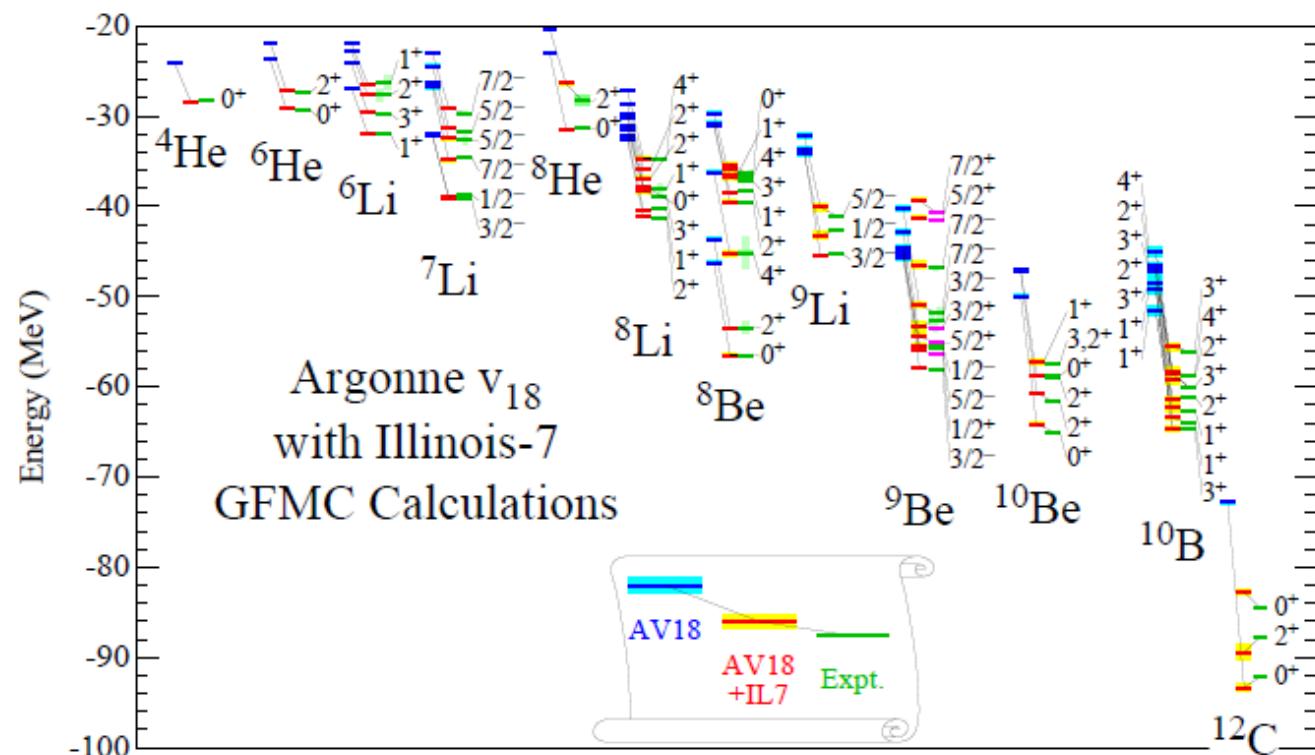
Tsubakihara, Harada, AO, arXiv:1402.0979

- Ruled-out EOS with hyperons = GM3
Glendenning & Moszkowski (1991)
- We did NOTHING special and find $2 M_{\odot}$ NS can be supported.
 - “Typical” RMF for nucl. matter
NL1, NL-SH, TM1
Reinhardt et al. ('86); Sharma, Nagarajan, Ring ('93); Sugahara, Toki ('94).
 - $s\bar{s}$ mesons are introduced
 - Hypernuclear data
 $\Lambda, \Lambda\Lambda$ hypernuclei
 Σ atomic shifts
SU(3) relation to isoscalar -vector couplings



What is necessary to solve the massive NS puzzle ?

- There are many “model” solutions.
- Ab initio calculation including three-baryon force (3BF)
 - Bare 2NF+Phen. 3NF(UIX, IL2-7) + many-body theory (verified in light nuclei).
 - Chiral EFT (2NF+3NF) + many-body theory
 - Dirac-Bruckner-HF (no 3NF)



J. Carlson et al. ('14)

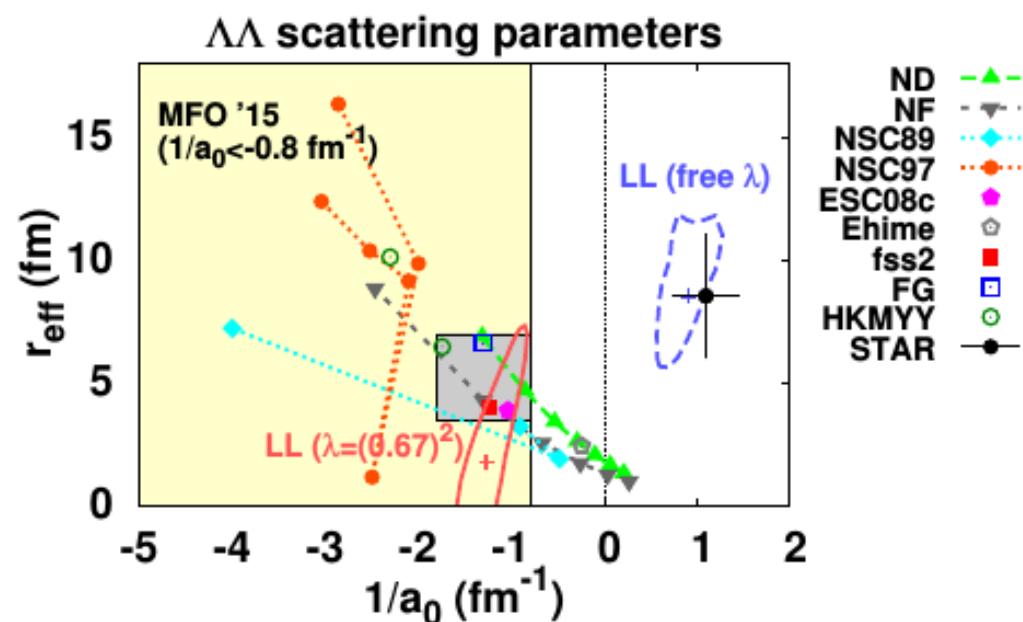
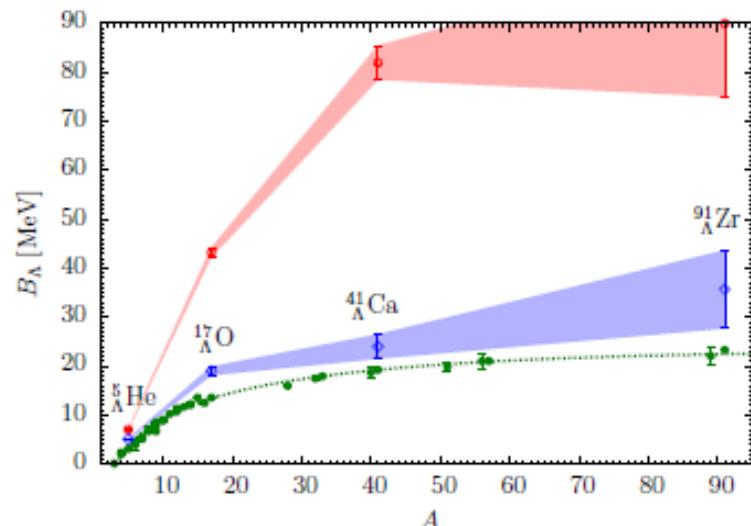
3BF including Hyperons

- 3BF incl. YNN, YYN and YYY should exist and contribute to EOS.

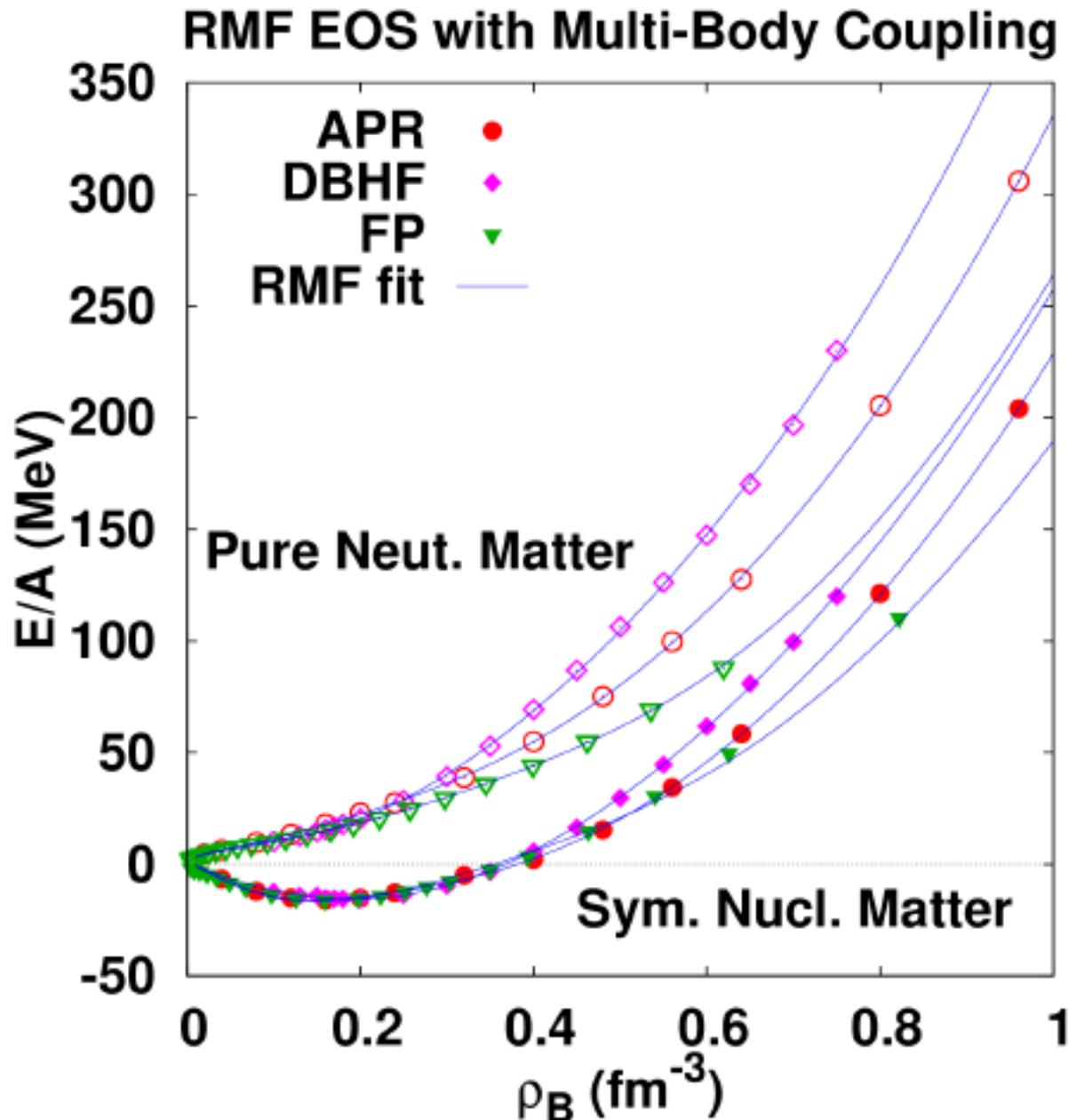
Nishizaki, Takatsuka, Yamamoto ('02)

- Chiral EFT, Multi-Pomeron exch., Quark Pauli, Lattice 3BF, SJ, .. Kohno('10); Heidenbauer+'13); Yamamoto+'14; Nakamoto, Suzuki; Doi+(HALQCD,'12); Tamagaki('08); ...
- Quant. MC study Lonardoni *et al.* ('14)
- Quark Meson Coupling Miyatsu *et al.*; Thomas (HHIQCD)
- $\Lambda\Lambda N$ K. Morita, T. Furumoto, AO, PRC91('15)024916

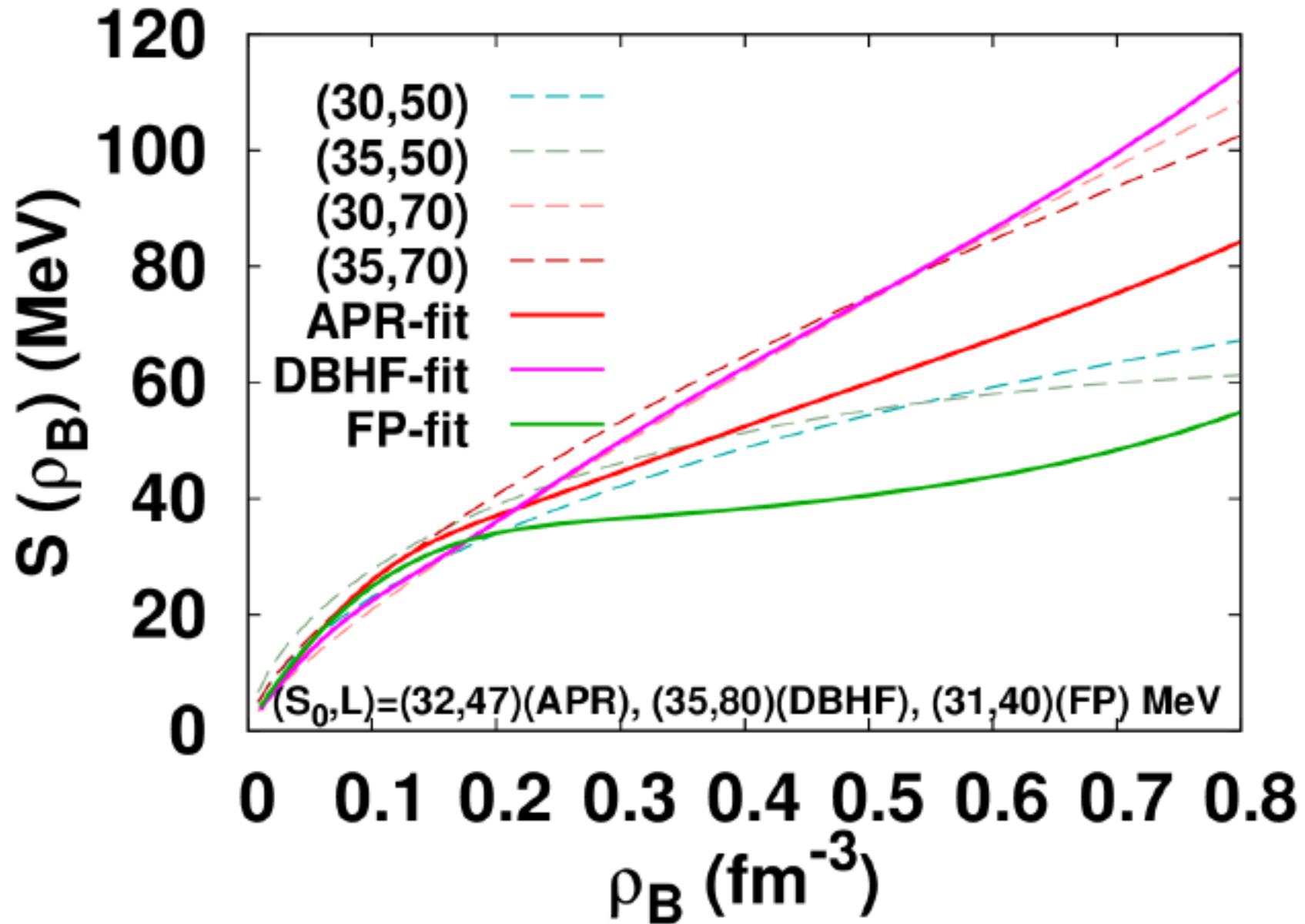
Caveat: Missing data



Fitting “Ab initio” EOS via RMF

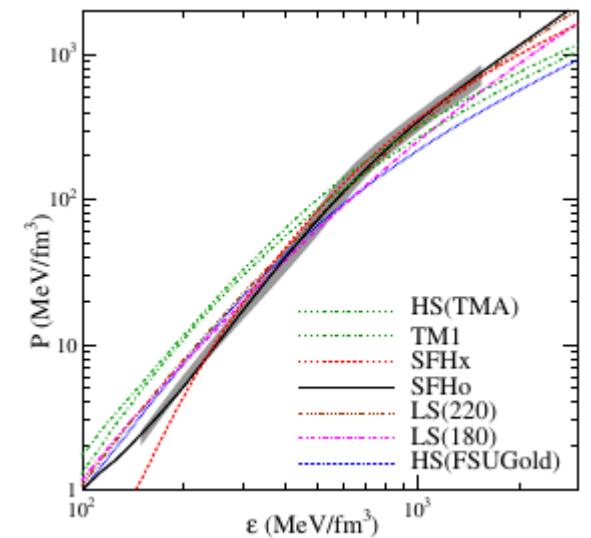
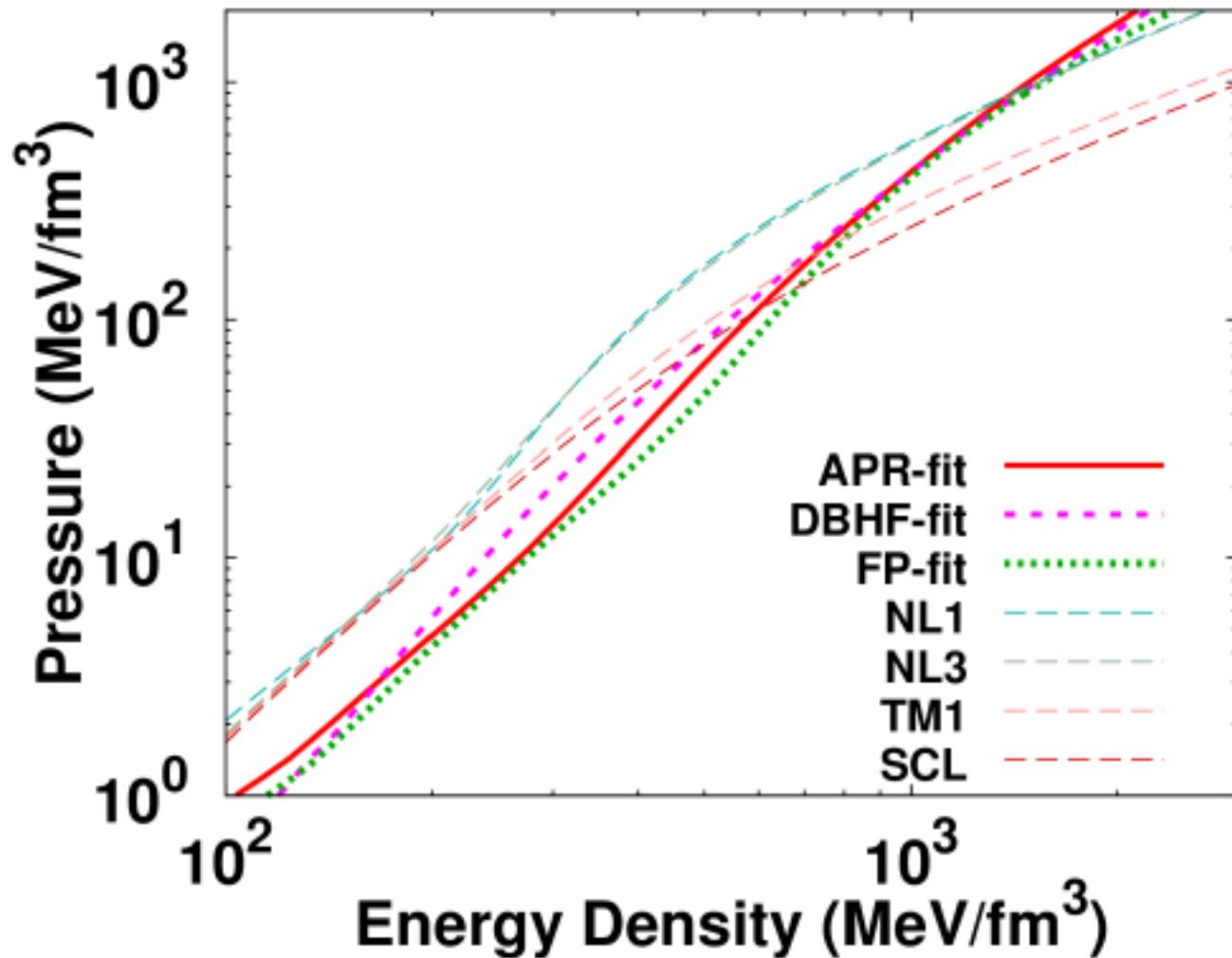


Symmetry Energy



Neutron Star Matter EOS

Neutron Star Matter EOS

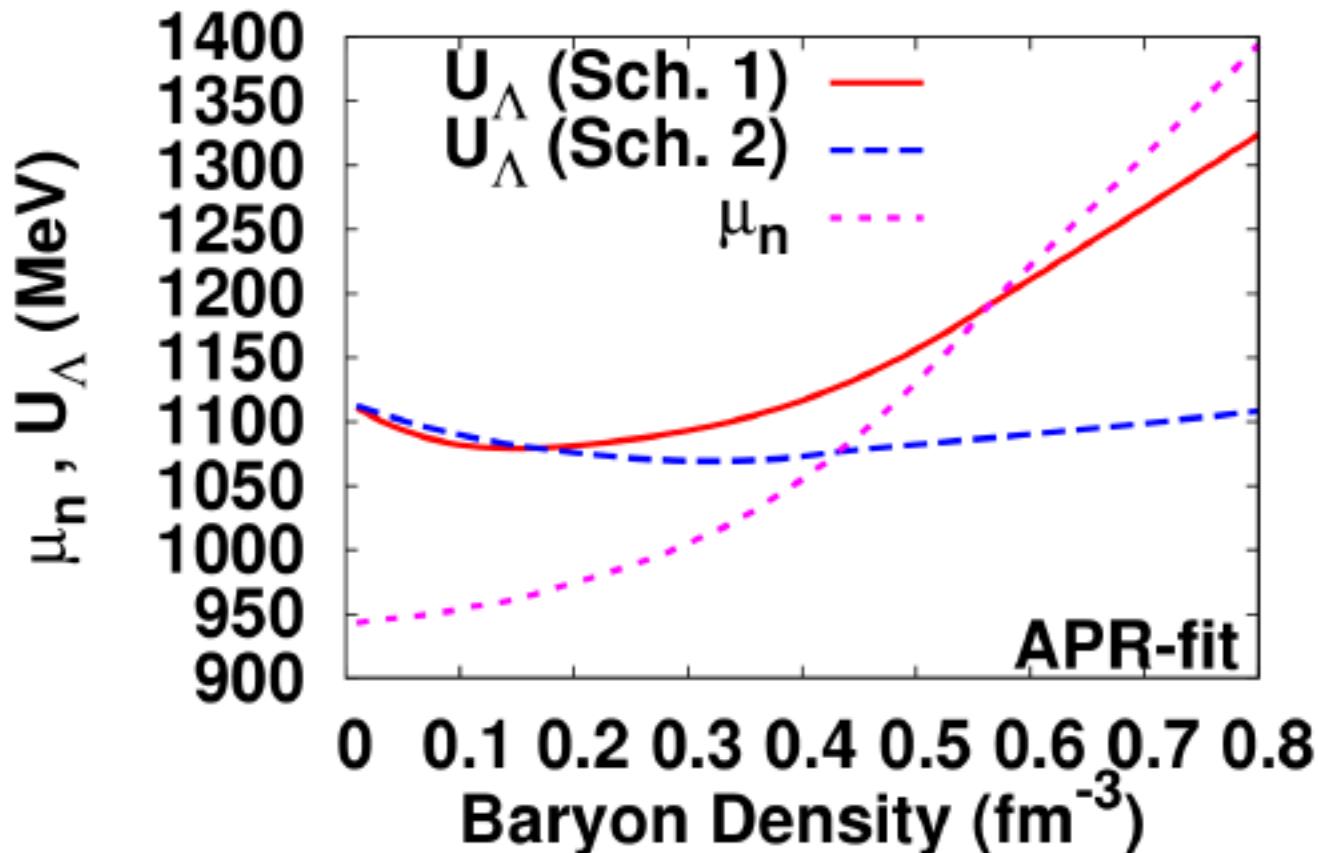


*A. W. Steiner, M. Hempel,
T. Fischer,
ApJ 774 (2013) 17
(TMA+NSE w/ excl. vol.)*

NS matter in “ab initio”-fit + Λ

■ Λ potential in nuclear matter at $\rho_0 \sim -30$ MeV

- Scheme 1: $U_\Lambda(\rho) = \alpha U_N(\rho)$
- Scheme 2: $U_\Lambda(\rho) = 2/3 U_{N=2}(\rho) + \beta U_{N>2}(\rho)$



M-R curve of Neutron Stars

