

Directed flow of Λ from heavy-ion collisions and hyperon puzzle of neutron stars

Akira Ohnishi¹, A. Jinno², K. Murase¹, Y. Nara³

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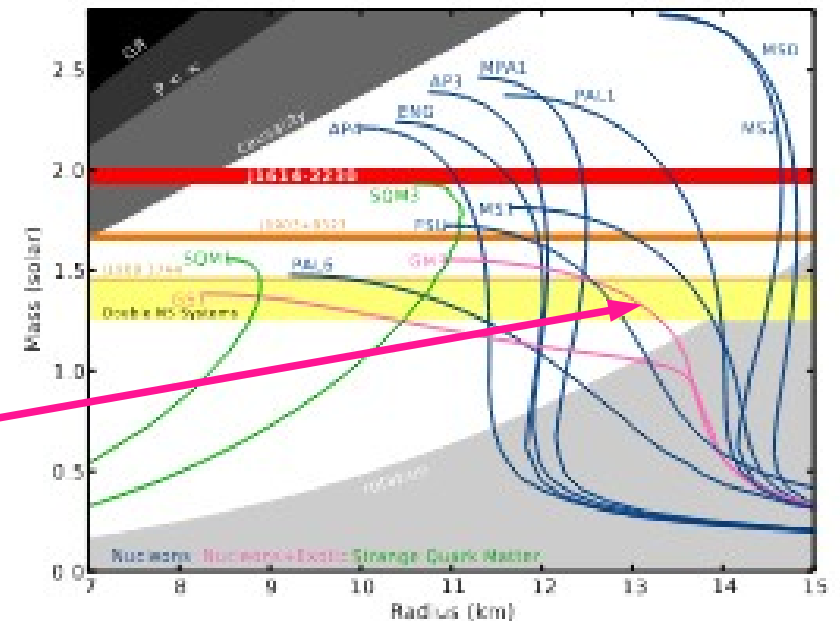


- Introduction – Hyperon puzzle
- Λ potential from chiral effective field theory (chiral EFT)
- Balance energy of proton directed flow slope
- Λ directed flow
- Summary

Y.Nara, A. Jinno, K. Murase, AO, in prep.

Hyperon Puzzle of Neutron Stars

- Observation of massive neutron stars rules out hyperonic EOS ?
 - Attractive $U_{\Lambda}(\rho)$ causes hyperon mixing in NS at $(2-4)\rho_0$, softens the EOS, and reduces $M_{\max} = (1.3-1.6) M_{\odot}$
- Proposed solutions
 - Three-body Λ NN repulsion \rightarrow repulsive $U_{\Lambda}(\rho)$ at high density
 - Transition to quark matter before Λ appears
 - General relativity \rightarrow Modified gravity

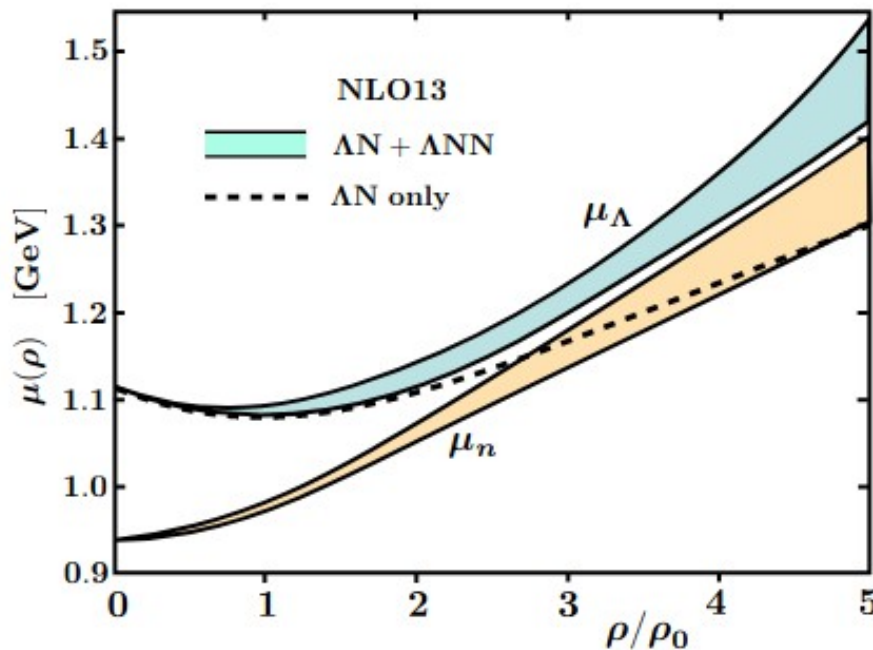


Hyperons

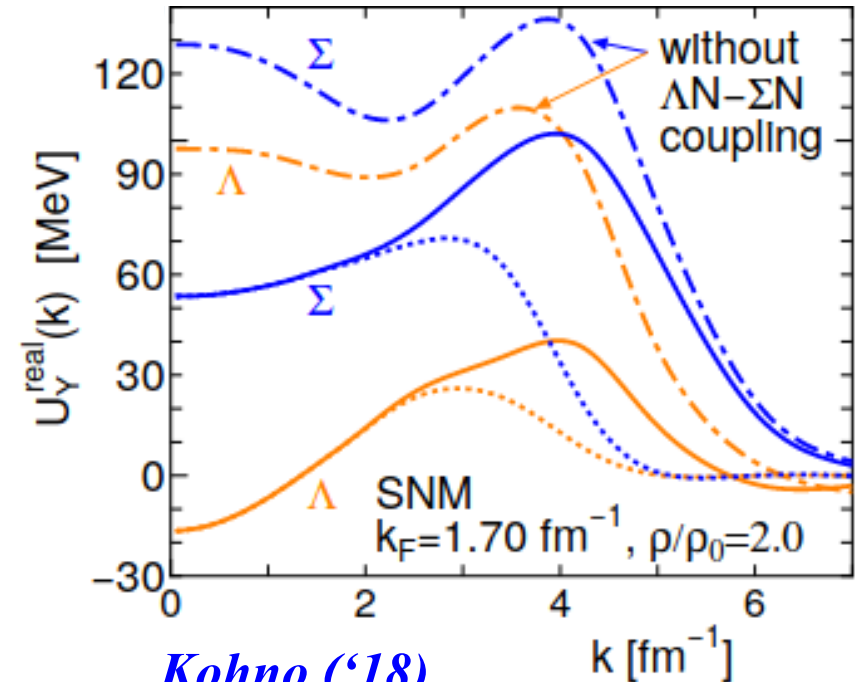
Demorest+(1010.5788)

Repulsive $U_\Lambda(\rho)$ at high density in chiral EFT

- Chiral effective field theory (chiral EFT) may cause repulsive Λ potential at high densities
Gerstung, Kaiser, Weise (2001.10563), Kohno (1802.05388)
- Yet unknown parameters are tuned to support $2 M_\odot$ neutron stars.
 - Repulsion at high densities needs to be verified !
 - Directed flow in HIC



Gerstung+('20)



Kohno ('18)

Directed flow (v_1)

- Directed flow (v_1 or $\langle p_x \rangle$) has been utilized to constrain EOS

E.g. Sahu, Cassing, Mosel, AO (nucl-th/9907002), Snellings+(nucl-ex/9908001)

- Proton v_1 slope problem *STAR (1401.3043)*

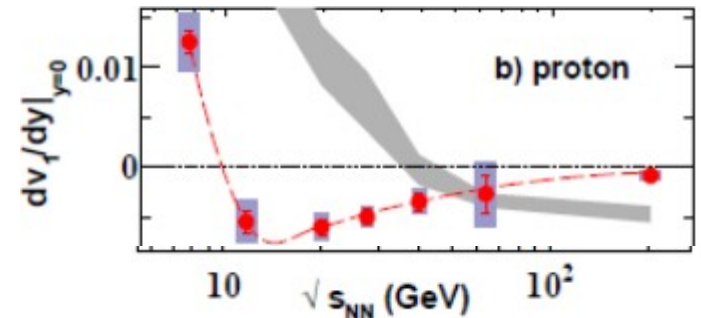
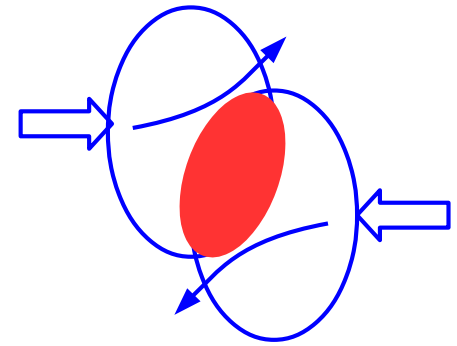
- Non-monotonic beam E. dep. of v_1 slope
- Sign change of v_1 slope at $\sqrt{s_{NN}} \sim 10$ GeV
- None of fluid and hybrid models explain the colliding energy dependence using a single EOS

Nara+(JAM, 1601.07692, 1611.08023, 1708.05617),

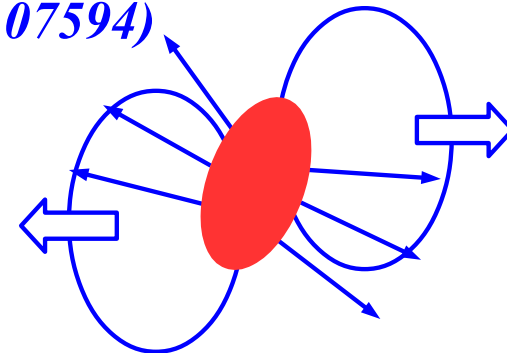
Ivanov+(3FD, 1412.1669, 1601.03902), Konchakovski+(PHSD, 1404.2765)

- An answer has been found ! *Nara, AO (PRC('22), 2109.07594)*

- Repulsion during compression \rightarrow positive
- Expansion of tilted matter \rightarrow negative
- Balance of two causes non-monotonic behavior

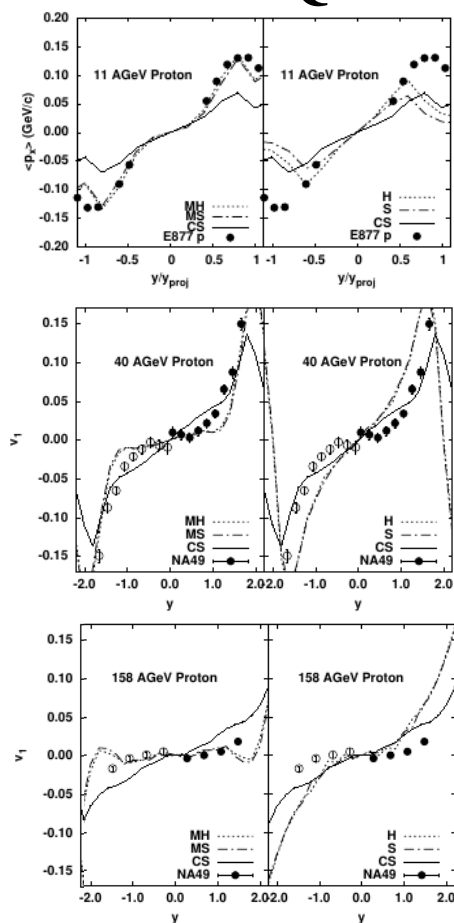


STAR, PRL112('14)162301 (1401.3043)



Past tries

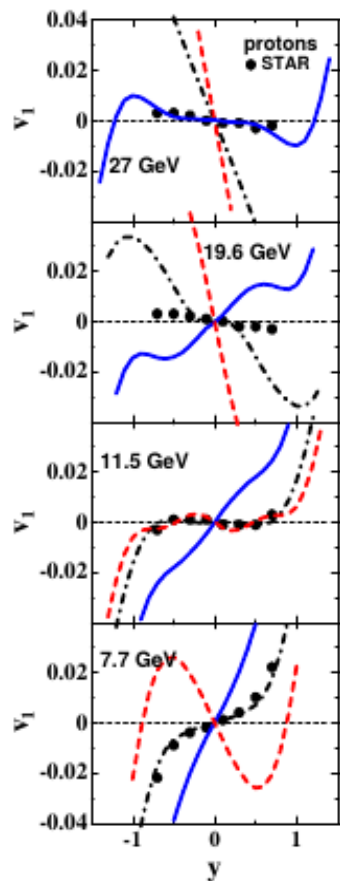
JAM-RQMD



p-dep. p-indep.

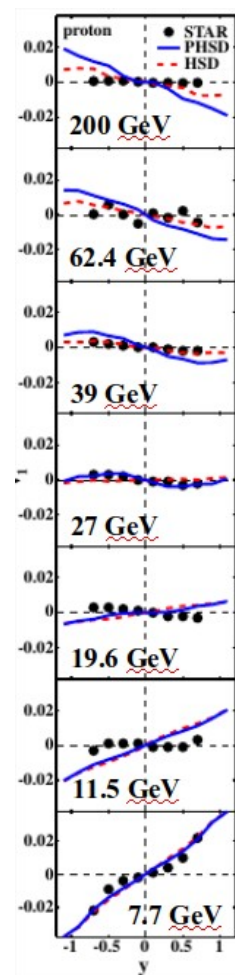
*M.Isse, AO, N.Otuka,
P.K.Sahu, Y.Nara,
PRC72('05)064908
(There was a mistake...)*

3FD



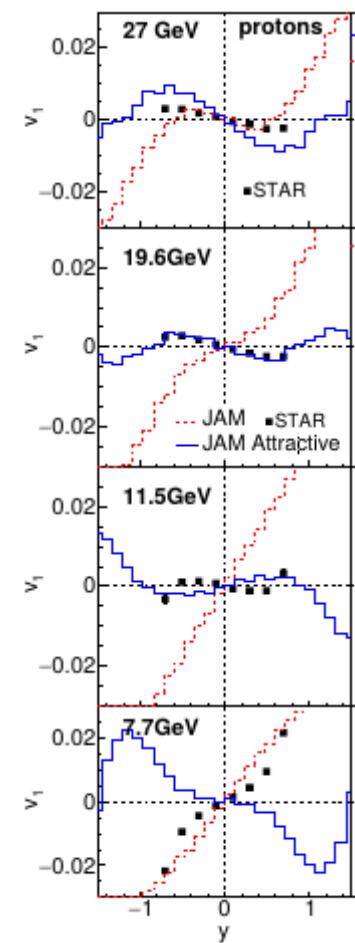
*Y.B.Ivanov,
A.A.Soldatov,
PRC91('15)
024915*

HSD/PHSD



*V.P.Konchakovski,
W.Cassing, Y.B.Ivanov,
V. D. Toneev,
PRC90('14)014903*

JAM+Att.

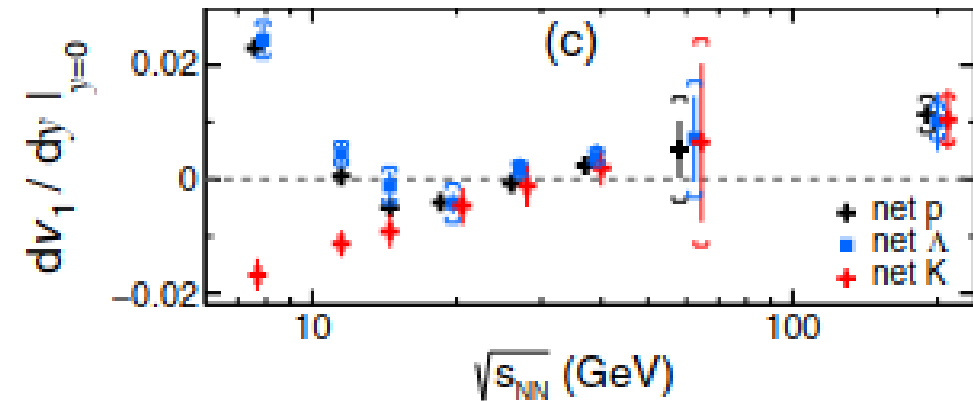


*Y.Nara, H.Niemi,
AO, H.Stoecker,
PRC94 ('16)034906*

Why Directed flow (v_1) of Λ

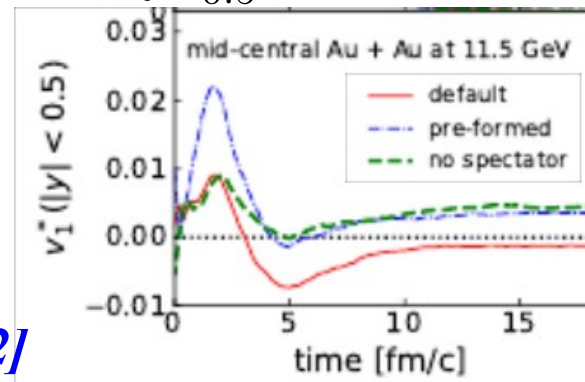
Directed flow of Λ

- In the compression+tilted expansion mechanism, directed flow of Λ should be smaller than p.
- Data show $v_1(\Lambda) \sim v_1(p)$ *STAR (1708.07132)*
 → Stronger repulsion for Λ at high densities ?

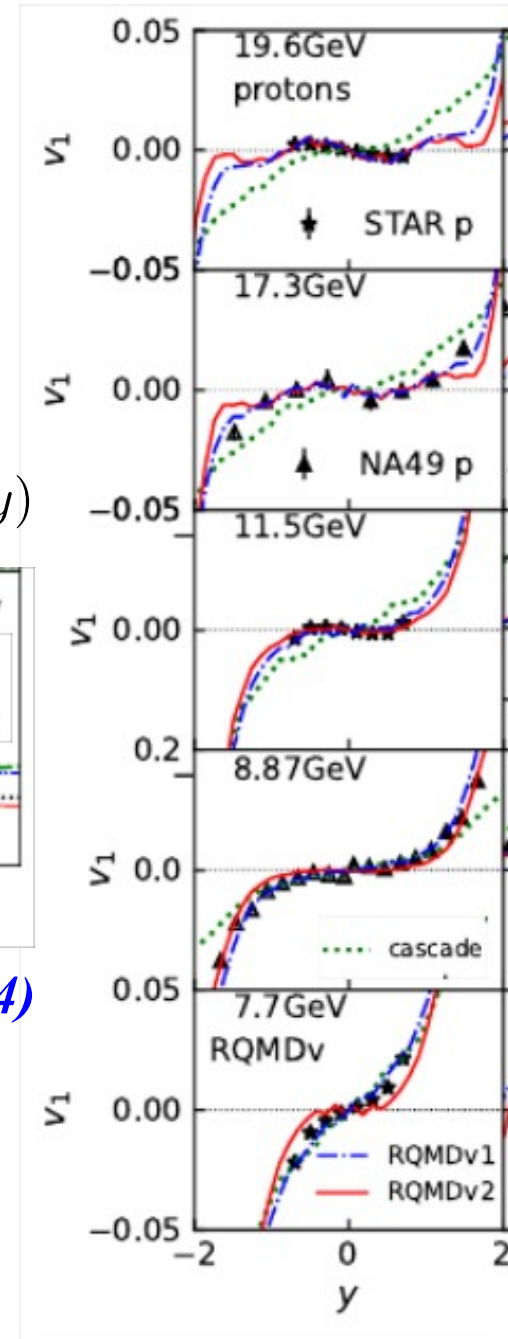


STAR, PRL120 ('18),062301 [1708.07132]

$$v_1^* = \int_{-0.5}^{0.5} dy v_1(y) \text{sign}(y)$$



Y. Nara, AO (2109.07594)



Let us examine Λ directed flow using $U_\Lambda(\rho)$ from chiral EFT !

JAM2/RQMDv+chiral EFT U_Λ

■ JAM2/RQMDv

- JAM2 = Update of JAM1 (fortran \rightarrow C++, pythia6 \rightarrow pythia8)
+ improvement of resonance exc. cross sections
+ expanding box to reduce the CPU time
- JAM2/RQMDv *Nara+(2109.07594)*
 - ◆ Lorentz vector type implementation of ρ - and p-dep. potential, which operates also on high p particles

■ U_Λ from chiral EFT

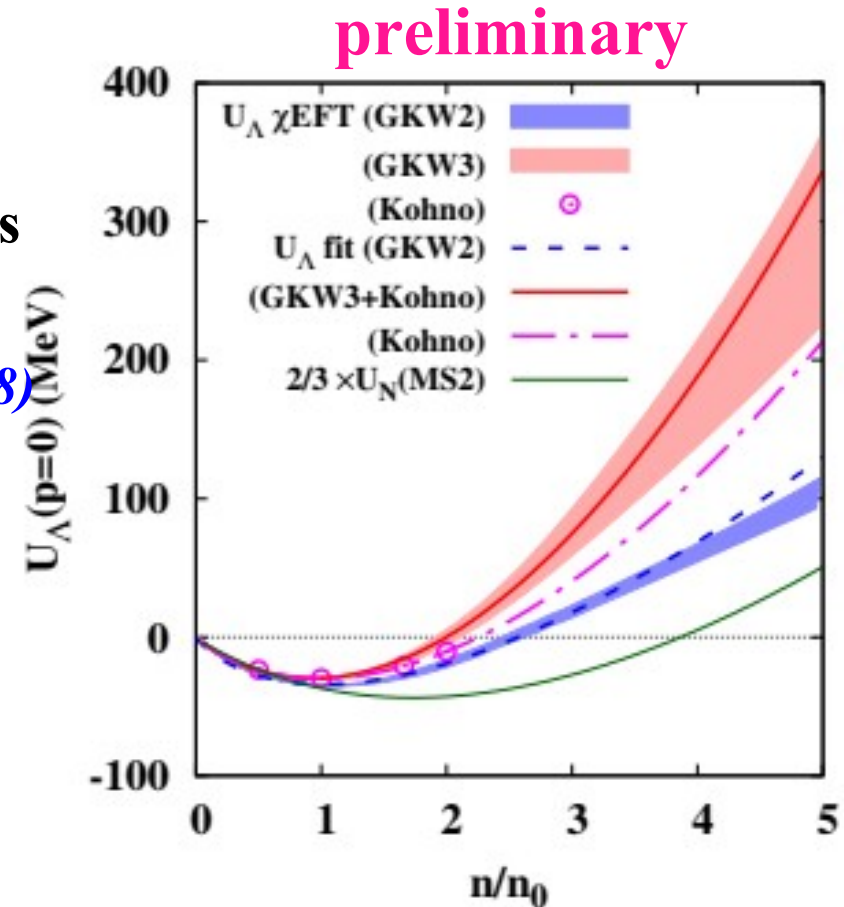
Gerstung+(2001.10563)(GKW), Kohno (1802.05388)

- ρ -dep. potential using Fermi mom. expansion *Tews+(1611.07133)*

$$U_{\text{sk}}(\rho) = a(\rho/\rho_0) + b(\rho/\rho_0)^{4/3} + c(\rho/\rho_0)^{5/3}$$

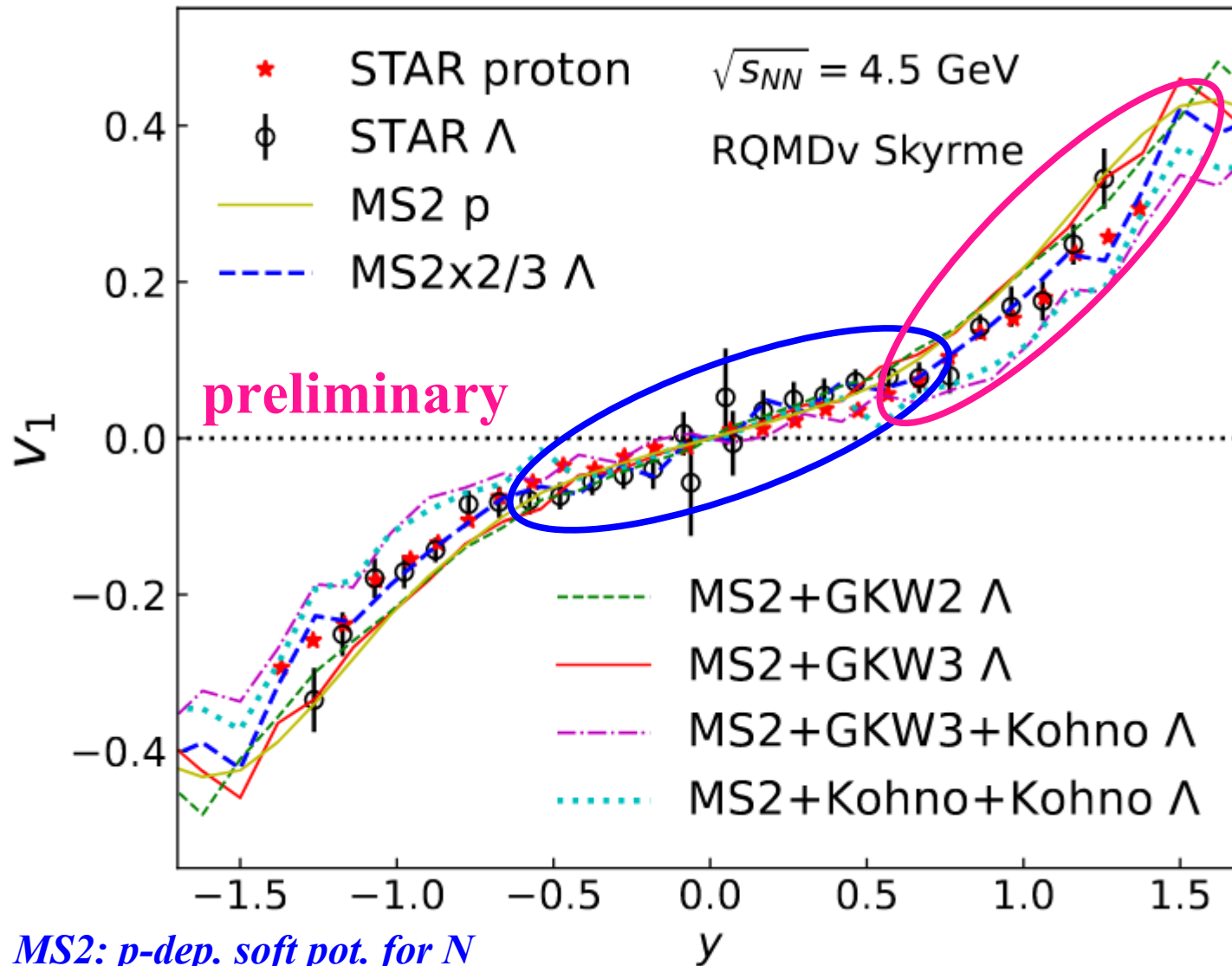
- Momentum dep. fit to *Kohno('18)*

$$U_{\text{m}}^0(\mathbf{p}) = \frac{C}{\rho_0} \int \frac{d\mathbf{p}'}{(2\pi)^3} \frac{f(\mathbf{r}, \mathbf{p}')}{1 + (\mathbf{p} - \mathbf{p}')^2/\mu^2}$$



Nara, Jinno, Murase, AO, in prep.

$$\sqrt{s_{NN}} = 4.5 \text{ GeV}$$



- Slope ($y=0$) is OK with
 - chiral EFT U_{Λ} (p-indep.)
 - $U_{\Lambda} = 2/3 U_N$
- v_1 at large $|y|$ needs stiffer U_{Λ}
 - chiral EFT (p-indep.)
- p-dep. U_{Λ} seems to underestimate v_1

MS2: p-dep. soft pot. for N

GKW2: chiral EFT with 2-body int.

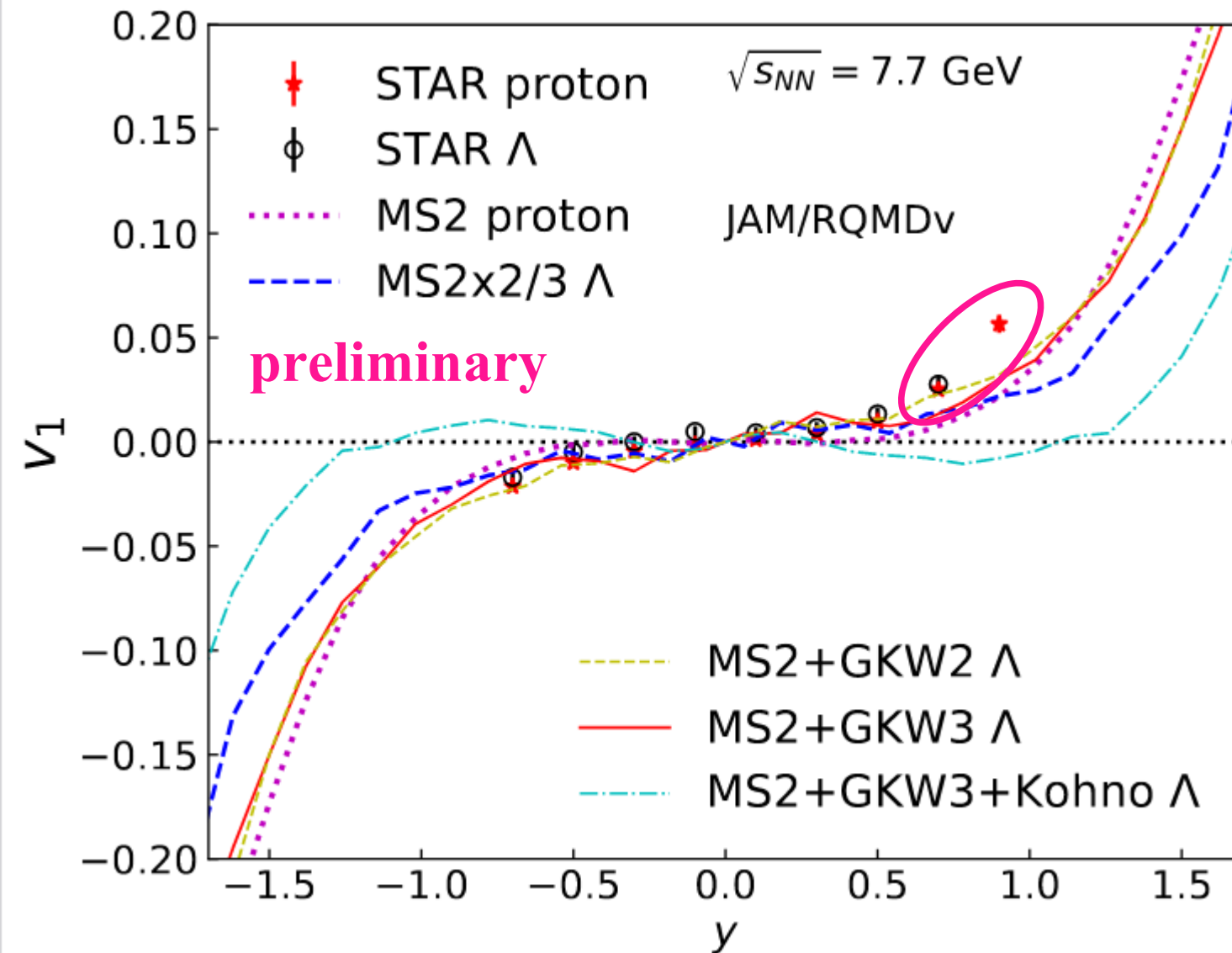
GKW3: chiral EFT with 2+3 body int.

GKW3+Kohno: GKW3 with p-dep. from Kohno

Kohno+Kohno: p- and p-dep. from Kohno

Nara, Jinno, Murase, AO, in prep.

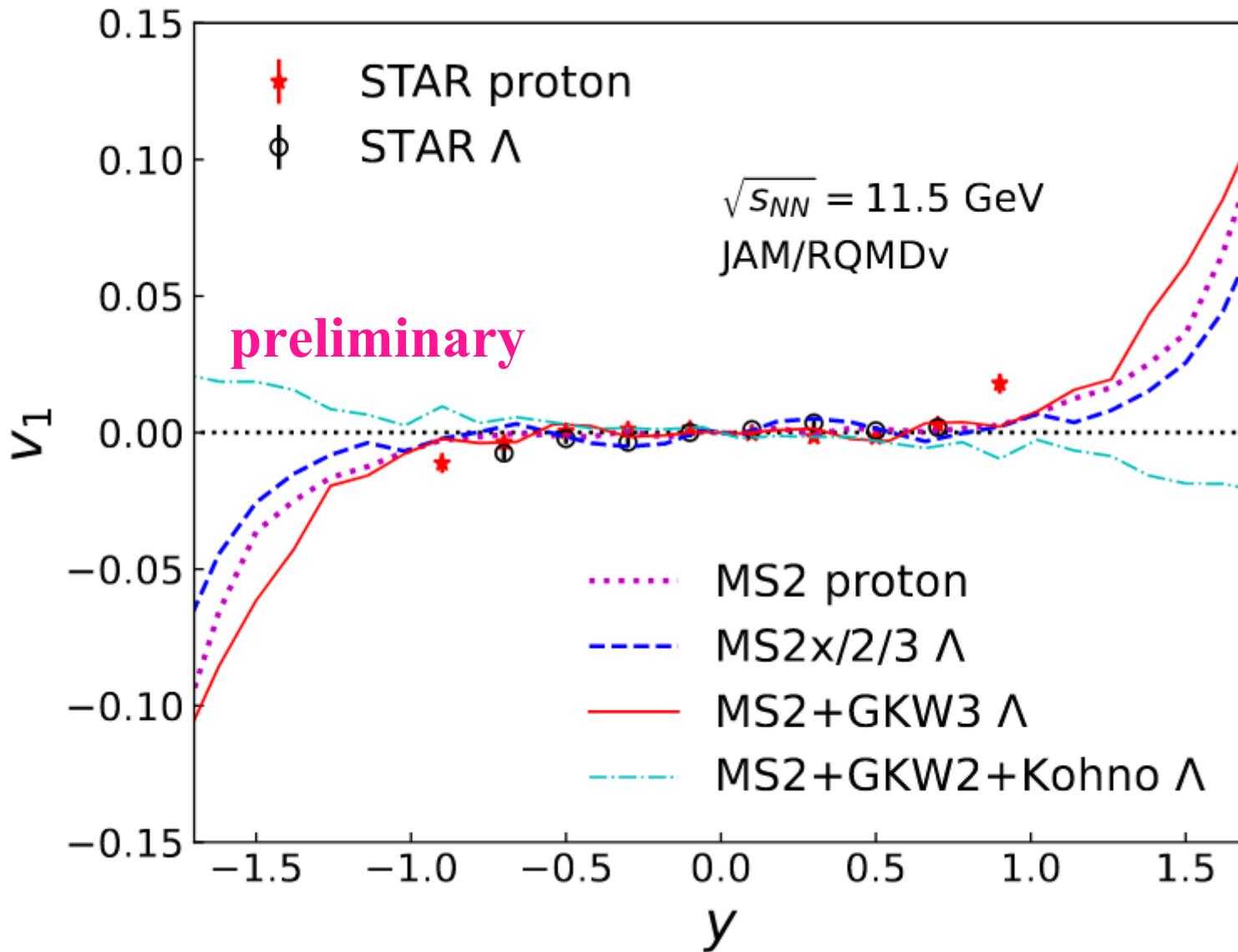
$$\sqrt{s_{NN}} = 7.7 \text{ GeV}$$



- v_1 at large $|y|$ is underestimated
- chiral EFT U_Λ s (p-indep.) are closer

Nara, Jinno, Murase, AO, in prep.

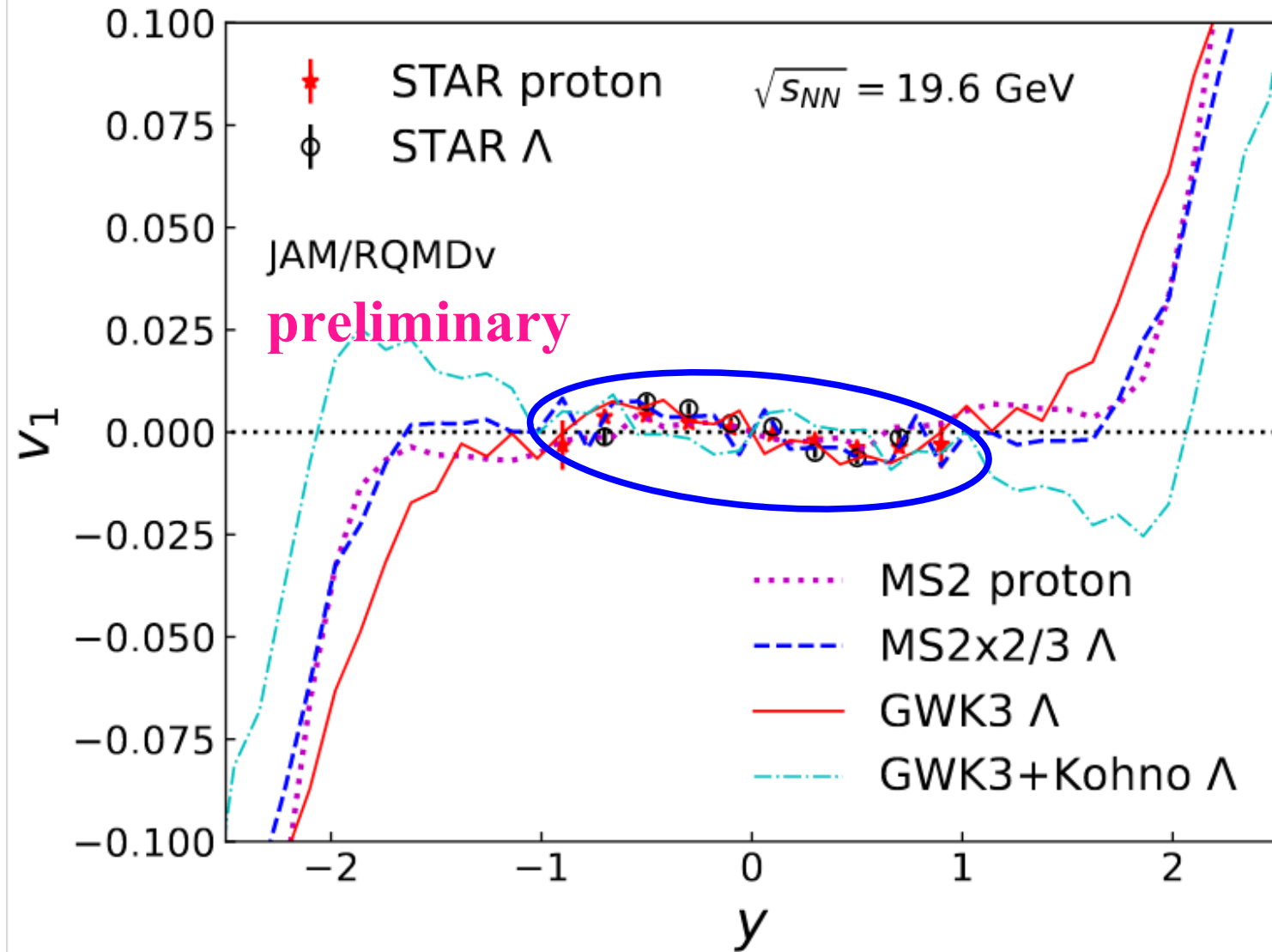
$\sqrt{s_{NN}}=11.5$ GeV



- $\sqrt{s_{NN}}=11.5$ GeV is around the "balance energy" (slope changes sign)

Nara, Jinno, Murase, AO, in prep.

$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$



- Negative dv_1/dy at $y \sim 0$

Nara, Jinno, Murase, AO, in prep.

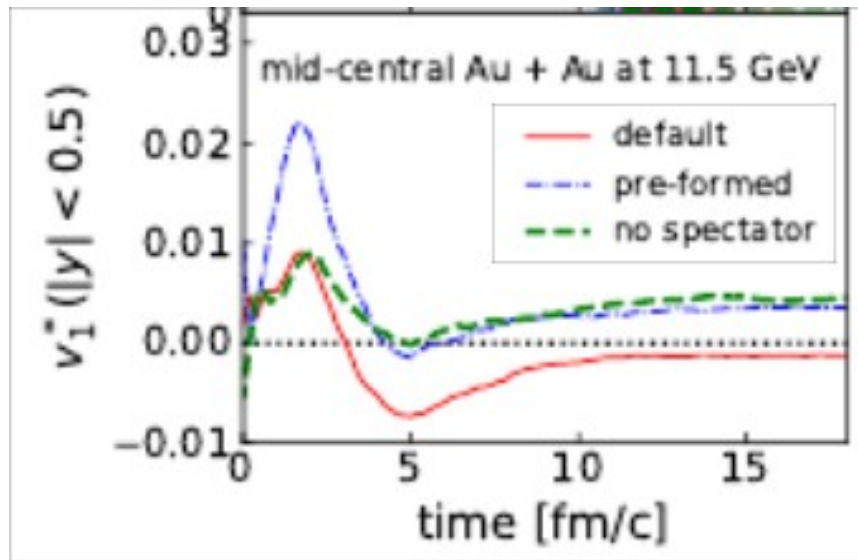
Summary

- The directed flow (v_1) of Λ from HICs at $\sqrt{s_{NN}}=(4.5-19.6)$ GeV is studied by using the Λ potential from chiral EFT.
 - U_Λ from chiral EFT contains strong repulsion from the 3-body interactions and suppresses Λ to appear in neutron stars.
 - The Λ v_1 slopes at midrapidity are roughly explained by the Λ potentials having ρ -dep. from the chiral EFT with 2+3 body int. [Similar results for $\langle px \rangle$ at $\sqrt{s_{NN}}=3.0$ GeV are obtained by D.C. Zhang+ (2107.00277)]
 - $U_\Lambda = 2/3 U_N$ and U_Λ from 2-body chiral EFT also explains the slopes. (Strong repulsion in U_Λ at high densities is not verified yet.)
 - The ρ -dep. of U_Λ seems to reduce the slopes significantly. While the ρ -dep. of U_Λ enhances the slope in the compression stage, it reduces the slope in the tilted matter expansion stage. (The simultaneous fit to ρ - and p -dep. would be necessary.)
 - The forward and backward v_1 values seem to be sensitive to the Λ potential at high densities.

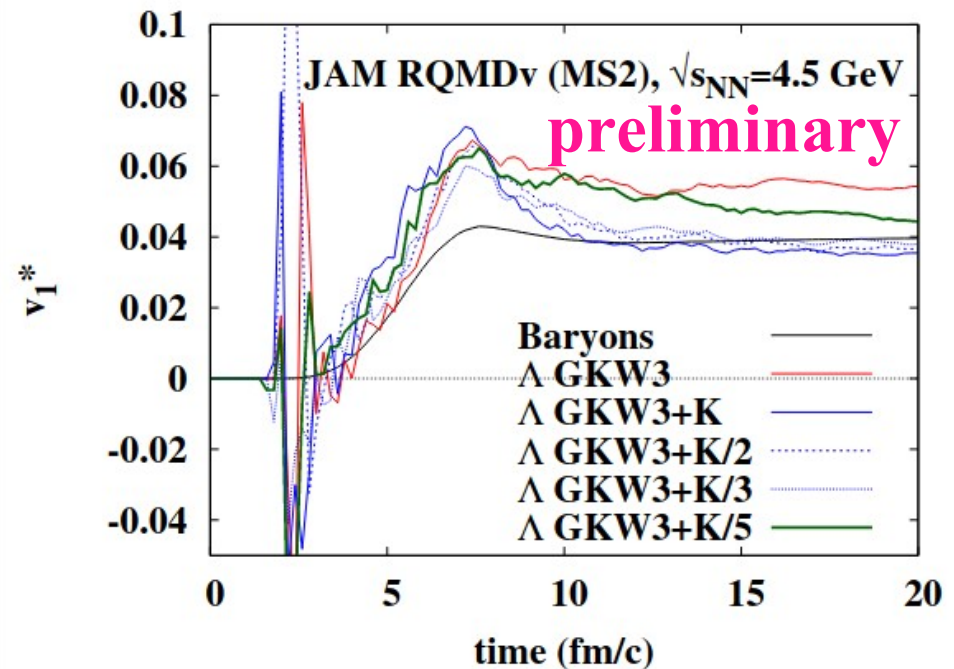
Nara, Jinno, Murase, AO, in prep.

p-dep. potential works on the way and back

- Momentum dependent potential works repulsively to high-momentum particles even at low densities. (p-indep. potential is attractive at $\rho < 2 \rho_0$)
 - It increases v_1 slope in the compression stage more strongly, but it also strongly reduces the slope in the tilted matter expansion stage.



Y. Nara, AO (2109.07594)



Nara, Jinno, Murase, AO, in prep.

Thank you for your attention !

Short (5 slide-)version follows.

Directed flow of Λ from HICs and hyperon puzzle of NSs

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Y. Nara, A. Jinnō,
K. Murase, AO, in prep.

■ Does three-body Λ NN repulsion solve the hyperon puzzle?

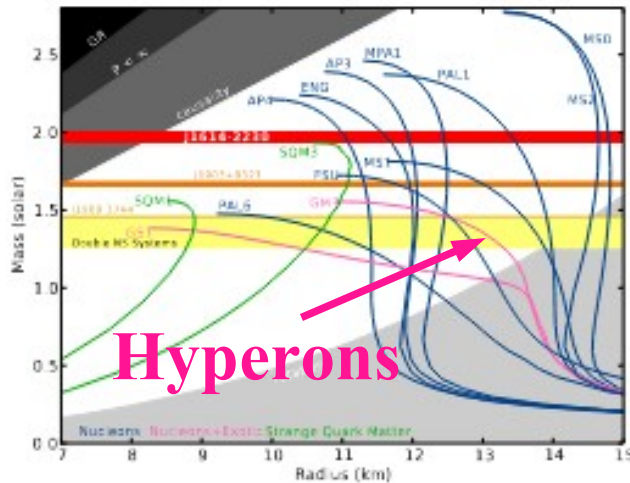
- E.g. Chiral EFT w/ 2+3-body int. can give repulsive potential.

Gerstung, Kaiser, Weise (2001.10563)(GKW), Kohno (1802.05388)

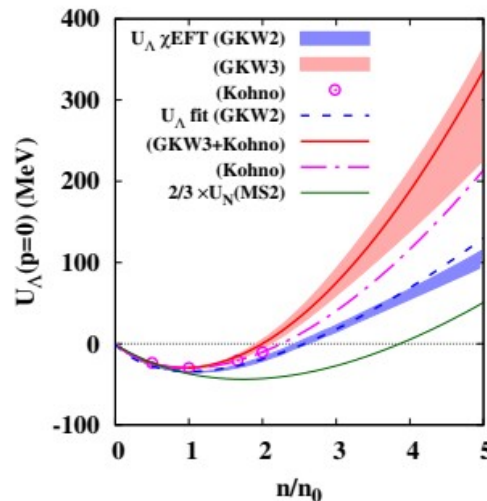
■ Examination of U_Λ via directed flow (v_1) of Λ *Data: STAR (1708.07132)*

- Studied in JAM2/RQMDv (explains v_1 of p) + U_Λ from chiral EFT

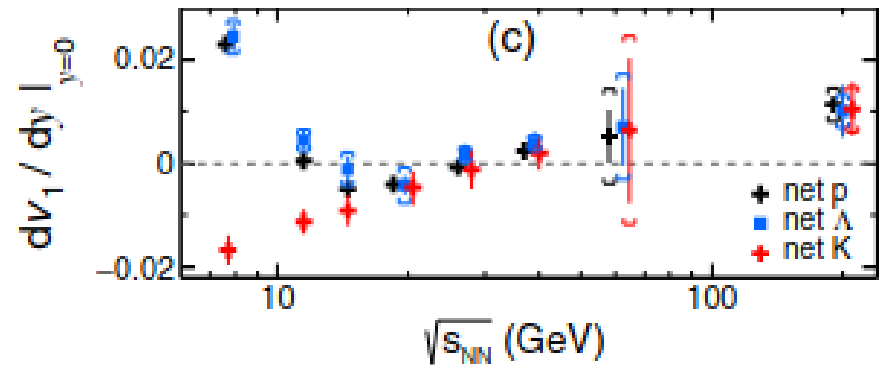
$$U_{\text{sk}}(\rho) = a(\rho/\rho_0) + b(\rho/\rho_0)^{4/3} + c(\rho/\rho_0)^{5/3} \quad U_m^0(\mathbf{p}) = \frac{C}{\rho_0} \int \frac{d\mathbf{p}'}{(2\pi)^3} \frac{f(\mathbf{r}, \mathbf{p}')}{1 + (\mathbf{p} - \mathbf{p}')^2/\mu^2}$$



Demorest+(1010.5788)



Nara+(in prep.)



STAR, PRL('18) [1708.07132]

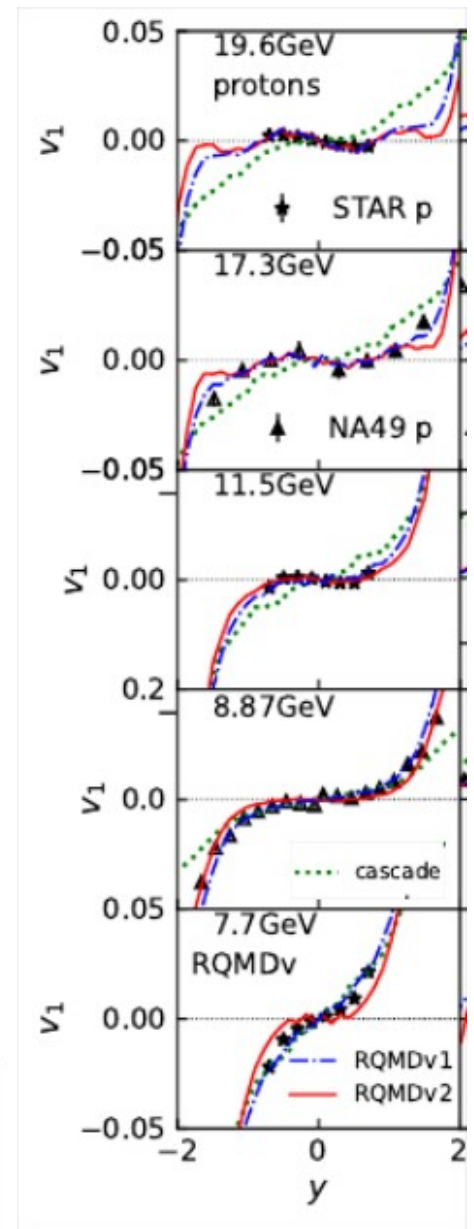
Directed flow (v_1) of protons

- Directed flow (v_1 or $\langle p_x \rangle$) has been utilized to constrain EOS

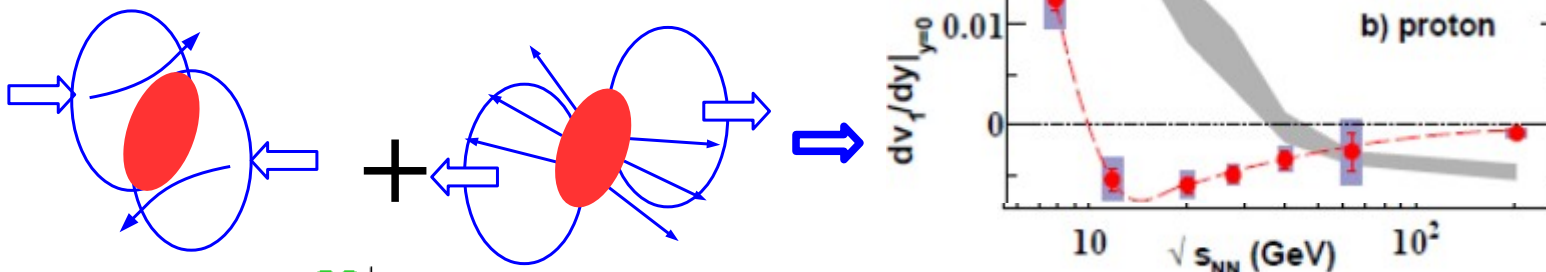
E.g. Sahu, Cassing, Mosel, AO (nucl-th/9907002), Snellings+(nucl-ex/9908001)

- Proton v_1 slope *STAR (1401.3043), Nara+(2109.07594)*

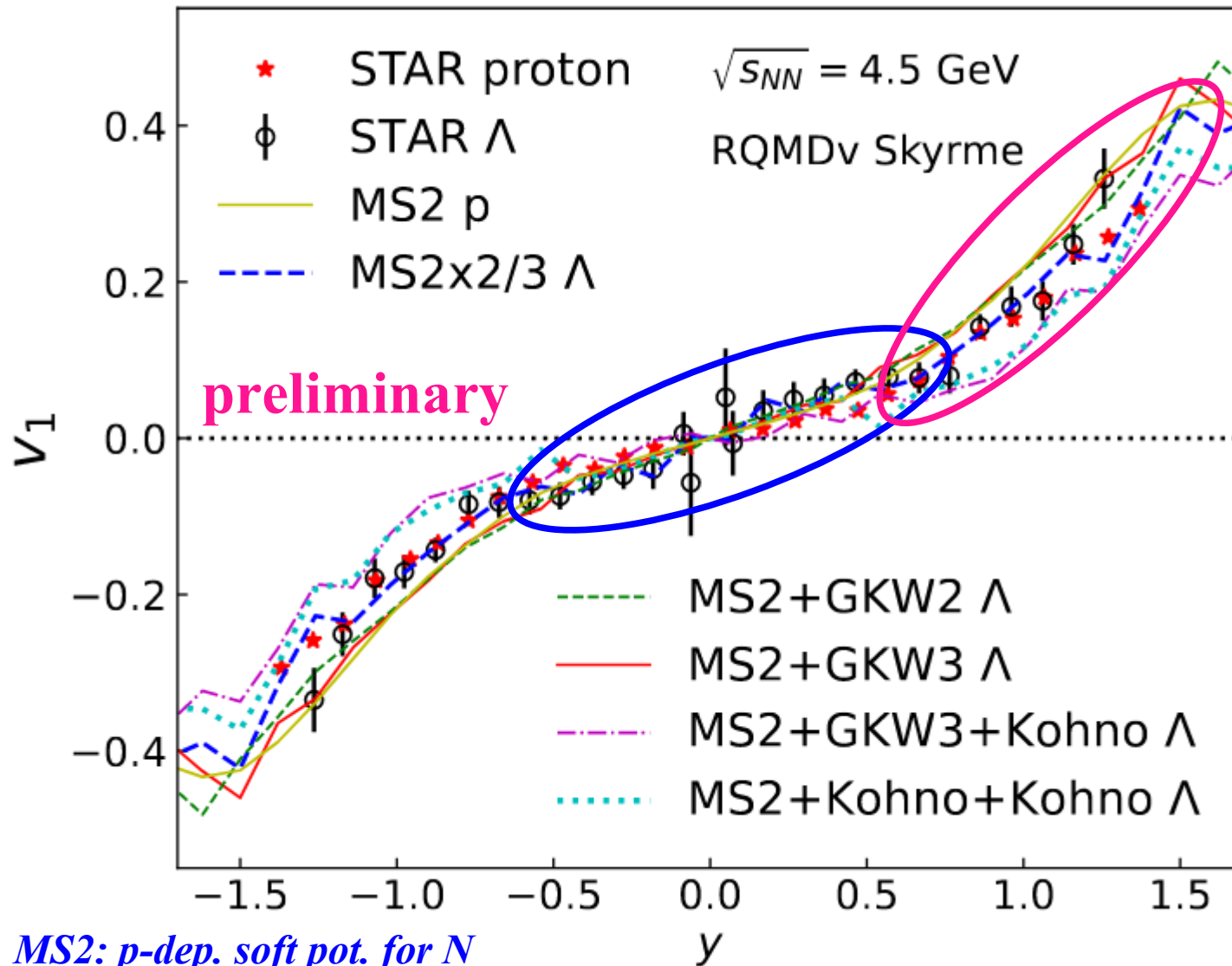
- Non-monotonic beam E. dep. of v_1 slope
- Sign change of v_1 slope at $\sqrt{s_{NN}} \sim 10$ GeV
- None of fluid and hybrid models explain the colliding energy dependence using a single EOS
- An answer *Nara, AO (PRC('22), 2109.07594)*
 - Compression (positive) and expansion (negative) contributions cause non-monotonicity.



STAR, PRL112('14) 162301 (1401.3043)



Directed flow (v_1) of Λ at $\sqrt{s_{NN}}=4.5$ GeV



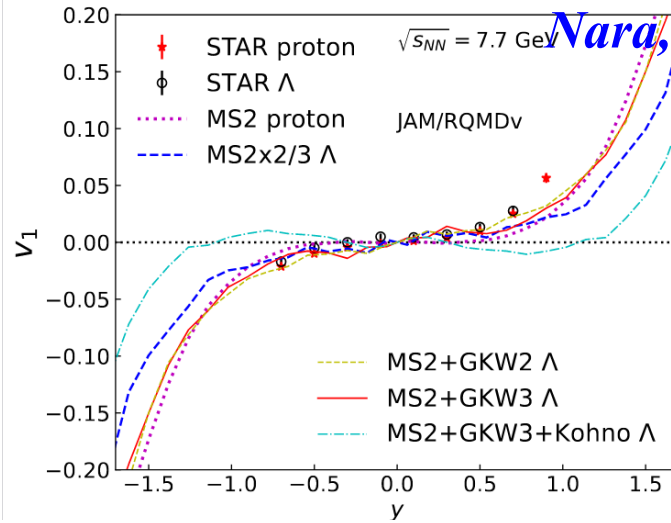
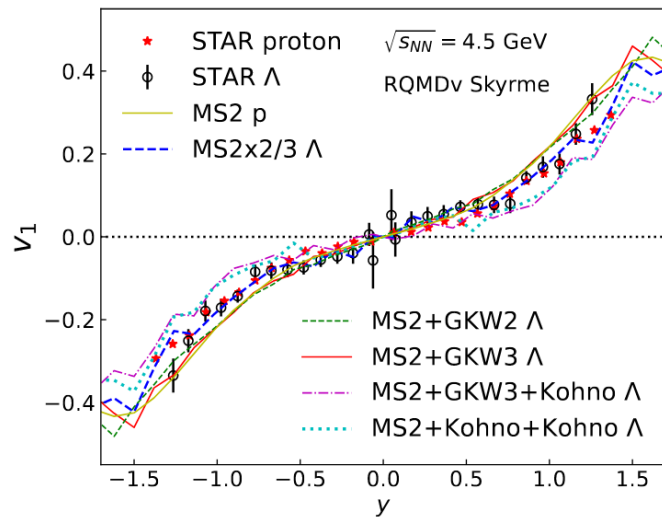
Calculation with JAM2/RQMDv

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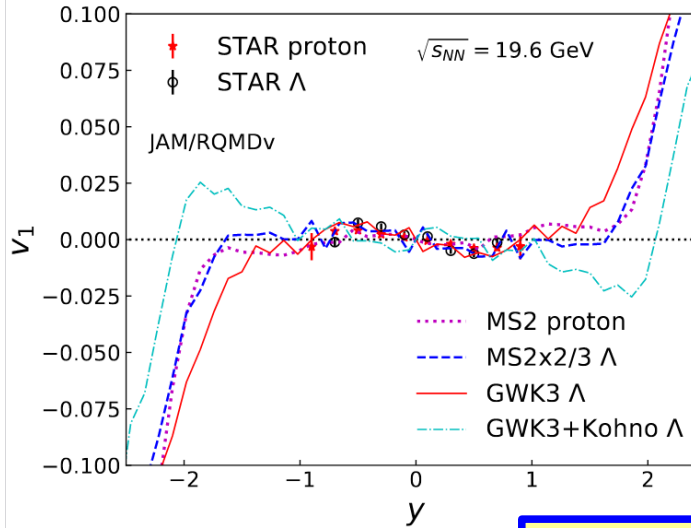
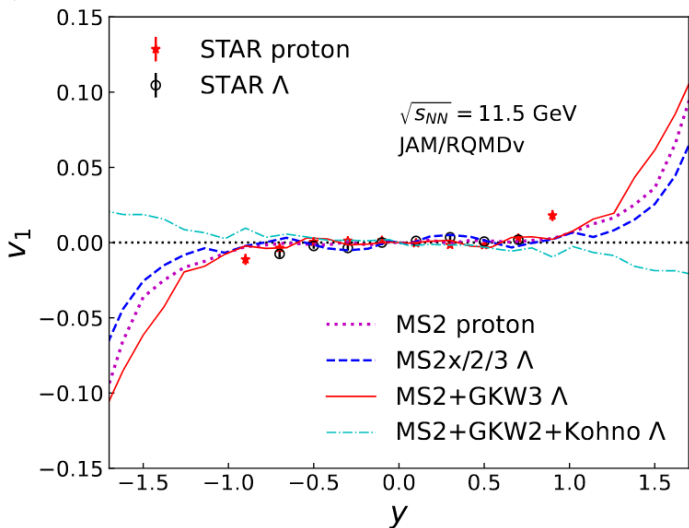
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Nara, Jinno, Murase, AO, in prep.

Directed flow of Λ at $\sqrt{s_{NN}}=(4.5-19.6)$ GeV



Nara, Jinno, Murase, AO, in prep.



- MS2: p-dep. soft pot. for N*
- GKW2: chiral EFT with 2-body int.*
- GKW3: chiral EFT with 2+3 body int.*
- GKW3+Kohno: GKW3 with p-dep. from Kohno*
- Kohno+Kohno: ρ - and p-dep. from Kohno*

U_{Λ} having the ρ -dep. in chiral EFT roughly explains the v_1 slopes.

Summary

- The directed flow (v_1) of Λ from HICs at $\sqrt{s_{NN}}=(4.5-19.6)$ GeV is studied by using the Λ potential from chiral EFT.
 - U_Λ from chiral EFT contains strong repulsion from the 3-body interactions and suppresses Λ to appear in neutron stars.
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