

*KEK & Osaka University*

*Spectrum of Eleven-dimensional Supergravity  
on a PP-wave Background*

*Tetsuji Kimura*

hep-th/0307193

*in collaboration with Kentaroh Yoshida (KEK)*

**Abstract**

We calculate the spectrum of the linearized supergravity around the maximally supersymmetric pp-wave background in eleven dimensions. The resulting spectrum agrees with that of zero-mode Hamiltonian of a supermembrane theory on the pp-wave background. We also discuss the connection with the Kaluza-Klein zero modes of  $AdS_4 \times S^7$  background.

## Introduction

### *M-theory, or Supermembrane Theory in Eleven Dimensions*

- strong coupling limit of type IIA string theory:  $R_{10} \propto g_{\text{IIA}}^{2/3}$
- eleven-dimensional supergravity as low energy theory
- three-form gauge field couples M2-brane (electrically) and M5-brane (magnetically)
- BFSS matrix theory (0 + 1-dim.  $U(N)$  supersymmetric Yang-Mills)
- supermembrane theory

supermembrane on flat background

de Wit, Hoppe, Nicolai, Nucl. Phys. B305 (1988) 545



but **unstable (continuous spectrum)**

de Wit, Lüscher, Nicolai, Nucl. Phys. B320 (1989) 135

supermembrane on pp-wave background

discretized energy spectrum

Nakayama, Sugiyama, Yoshida, Phys. Rev. D68 (2003) 026001

We would like to investigate fluctuation fields in eleven-dimensional **supergravity** and compare with the zero-modes of supermembrane **on the pp-wave background**.

## Maximally Supersymmetric Spaces in Eleven-dimensional Supergravity



### Kowalski-Glikman Solution (PP-wave Background)

$$ds^2 = -2dx^+dx^- + G_{++}(dx^+)^2 + \sum_{I=1}^9(dx^I)^2 \quad x^\pm = \frac{1}{\sqrt{2}}(x^0 \pm x^{10})$$

$$G_{++} = -\left[\left(\frac{\mu}{3}\right)^2 \sum_{\tilde{I}=1}^3(x^{\tilde{I}})^2 + \left(\frac{\mu}{6}\right)^2 \sum_{I'=4}^9(x^{I'})^2\right] \quad \mu = F_{+123} \neq 0$$

Affine connection  $\Gamma_{MN}^P$ , curvature tensor  $R^R_{PMN}$  and spin connection  $\omega_M^{AB}$ :

$$\Gamma_{++}^{\tilde{I}} = \Gamma_{\tilde{I}+}^- = \omega_+^{\tilde{I}-} = \left(\frac{\mu}{3}\right)^2 x^{\tilde{I}} \quad \Gamma_{++}^{I'} = \Gamma_{I'+}^- = \omega_+^{I'-} = \left(\frac{\mu}{6}\right)^2 x^{I'}$$

$$R_{+\tilde{I}+}^{\tilde{J}} = \delta_{\tilde{I}\tilde{J}} \left(\frac{\mu}{3}\right)^2 \quad R_{+I'+}^{J'} = \delta_{I'J'} \left(\frac{\mu}{6}\right)^2$$

$$\mathcal{R}_{++} = \mathcal{R}^{--} = \frac{1}{2}\mu^2 \quad \mathcal{R} = 0$$

## Setup

### Eleven-dimensional Supergravity

$e_M^A$  : vielbein

$E_A^M$  : inverse vielbein

$\Psi_M$  : gravitino (Majorana spinor)

$$\bar{\Psi} = i\Psi^\dagger \hat{\Gamma}^0 = \Psi^T C$$

$C_{MNP}$  : three-form gauge field

Lagrangian of eleven-dimensional supergravity (without torsion):

$$\begin{aligned} \mathcal{L} = & e\mathcal{R} - \frac{1}{2}e\bar{\Psi}_M \hat{\Gamma}^{MNP} D_N \Psi_P - \frac{1}{48}e F_{MNPQ} F^{MNPQ} \\ & + \frac{1}{192}e\bar{\Psi}_M \tilde{\Gamma}^{MNPQRS} \Psi_N F_{PQRS} \\ & + \frac{1}{(144)^2} \epsilon^{MNPQRSUVWXY} F_{MNPQ} F_{RSUV} C_{WXY} \end{aligned}$$

various conventions:

$$D_N \Psi_P = \partial_N \Psi_P - \frac{1}{4} \omega_N^{AB} \hat{\Gamma}_{AB} \Psi_P$$

$$\tilde{\Gamma}^{NPQR}_M = \hat{\Gamma}^{NPQR}_M - 8\delta_M^{[N} \hat{\Gamma}^{PQR]}$$
$$\tilde{\Gamma}^{MNPQRS} = \hat{\Gamma}^{MNPQRS} + 12g^{M[P} \hat{\Gamma}^{QR} g^{S]N}$$

$$\epsilon^{012\dots 10} = 1$$

$SO(10, 1)$  Levi-Civita symbol

## Classical Field Equations

$$\begin{aligned} 0 &= \frac{1}{2}g_{MN}\mathcal{R} - \mathcal{R}_{MN} - \frac{1}{96}g_{MN}F_{PQRS}F^{PQRS} + \frac{1}{12}F_{MPQR}F_N{}^{PQR} \\ 0 &= \hat{\Gamma}^{MNP}D_N\Psi_P - \frac{1}{96}\tilde{\Gamma}^{MNPQRS}\Psi_N F_{PQRS} \\ 0 &= \nabla^Q\{e F_{QMNP}\} + \frac{18}{(144)^2}g_{MZ}g_{NK}g_{PL}\varepsilon^{ZKLQRSUVWXY}F_{QRSU}F_{VWXY} \end{aligned}$$

We neglected the terms derived from torsion, etc.

They contribute to higher-order fluctuations.

## Hamiltonian

We will encounter Klein-Gordon type equations of motion and have to evaluate its energy spectrum. Klein-Gordon type equation of motion for a field  $\phi(x)$ :

$$(\square - \alpha \mu i \partial_-) \phi(x^+, x^-, x^I) = 0$$

$\alpha$  : arbitrary constant       $x^+$  : evolution parameter

d'Alembertian  $\square$  on the pp-wave background:

$$\square = -\frac{1}{\sqrt{-g}} \partial_M (\sqrt{-g} g^{MN} \partial_N) = 2\partial_+ \partial_- + G_{++} (\partial_-)^2 - (\partial_K)^2$$

Fourier transformed expression of  $\phi(x)$ :

$$\phi(x^+, x^-, x^I) = \int \frac{d p_- d^9 p_I}{\sqrt{(2\pi)^{10}}} e^{i(p_- x^- + p_I x^I)} \tilde{\phi}(x^+, p_-, p_I)$$

Then the operator  $(\square - \alpha \mu i \partial_-)$  is represented as:

$$0 = 2p_- i\partial_+ - \left\{ \sum_{\tilde{I}=1}^3 \left(\frac{\mu}{3}\right)^2 (\partial_{p_{\tilde{I}}})^2 + \sum_{I'=4}^9 \left(\frac{\mu}{6}\right)^2 (\partial_{p_{I'}})^2 \right\} (p_-)^2 + (p_I)^2 + \alpha \mu p_-$$

Introduce the standard technique of harmonic oscillators:

$$\begin{aligned}
 a^{\tilde{I}} &\equiv \frac{1}{\sqrt{2\tilde{m}}} \{p_{\tilde{I}} + \tilde{m}\partial_{p_{\tilde{I}}}\} & \bar{a}^{\tilde{I}} &\equiv \frac{1}{\sqrt{2\tilde{m}}} \{p_{\tilde{I}} - \tilde{m}\partial_{p_{\tilde{I}}}\} & \tilde{m} &\equiv -\frac{1}{3}\mu p_- \\
 a^{I'} &\equiv \frac{1}{\sqrt{2m'}} \{p_{I'} + m'\partial_{p_{I'}}\} & \bar{a}^{I'} &\equiv \frac{1}{\sqrt{2m'}} \{p_{I'} - m'\partial_{p_{I'}}\} & m' &\equiv -\frac{1}{6}\mu p_-
 \end{aligned}$$

whose commutation relations are represented by

$$[a^{\tilde{I}}, \bar{a}^{\tilde{J}}] = \delta^{\tilde{I}\tilde{J}} \quad [a^{I'}, \bar{a}^{J'}] = \delta^{I'J'} \quad [a^{\tilde{I}}, \bar{a}^{J'}] = [a^{I'}, \bar{a}^{\tilde{J}}] = 0$$

Thus we express the Hamiltonian  $H = i\partial_+$ :

$$H = \frac{1}{3}\mu \sum_{\tilde{I}} \bar{a}^{\tilde{I}} a^{\tilde{I}} + \frac{1}{6}\mu \sum_{I'} \bar{a}^{I'} a^{I'} + \frac{1}{2}\mu (2 - \alpha)$$

Last term = **zero-point energy**  $E_0$  of the system (eigenvalue of  $H$ ):

$$E_0 = \frac{1}{2}\mu \mathcal{E}_0(\phi) \quad \mathcal{E}_0(\phi) = 2 - \alpha$$

## Bosonic/Fermionic Spectrum

Fluctuation fields are expanded around classical fields as follows:

$$g_{MN} \rightarrow g_{MN} + h_{MN}$$

$g_{MN}$  : pp-wave background

$$\Psi_M \rightarrow 0 + \psi_M$$

$$C_{MNP} \rightarrow C_{MNP} + \mathcal{C}_{MNP}$$

$$4\partial_{[+}C_{123]} = F_{+123} = \mu$$

Constraints:

$$\left. \begin{aligned} h_{-M} &= h^{+N} = 0 \\ \psi_- &= 0 \\ \mathcal{C}_{-NP} &= 0 \end{aligned} \right\}$$

light-cone gauge-fixing



## Bosonic Fields

(-- ) component of (A.1):

$$0 = h_{II} \implies \text{traceless condition}$$

Substitute this into (-I) component of (A.1):

$$h_{I+} = \frac{1}{\partial_-} \partial_J h_{IJ} \implies \text{non-dynamical}$$

(-IJ) component of (A.3):

$$\mathcal{C}_{+IJ} = \frac{1}{\partial_-} \partial_K \mathcal{C}_{IJK} \implies \text{non-dynamical}$$

Trace of (A.1) under the above conditions:

$$h_{++} = \frac{1}{(\partial_-)^2} \partial_I \partial_J h_{IJ} - \frac{1}{3\partial_-} \mu \mathcal{C}_{123} \implies \text{non-dynamical}$$

Let us substitute the previous conditions into the other components of (A.1) and (A.3).

Non-trivial equations:

$$(\tilde{I}\tilde{J}) \text{ of (A.1) : } \quad 0 = \square h_{\tilde{I}\tilde{J}} + \frac{2}{3}\mu \delta_{\tilde{I}\tilde{J}} \partial_- \mathcal{C} \quad (1a)$$

$$(\tilde{I}J') \text{ of (A.1) : } \quad 0 = \square h_{\tilde{I}J'} + \mu \partial_- \mathcal{C}_{\tilde{I}J'} \quad (1b)$$

$$(I'J') \text{ of (A.1) : } \quad 0 = \square h_{I'J'} - \frac{1}{3}\mu \delta_{I'J'} \partial_- \mathcal{C} \quad (1c)$$

$$(\tilde{I}\tilde{J}\tilde{K}) \text{ of (A.3) : } \quad 0 = \square \mathcal{C} - 2\mu \partial_- h_{\tilde{I}\tilde{I}} \quad (1d)$$

$$(\tilde{I}\tilde{J}K') \text{ of (A.3) : } \quad 0 = \square \mathcal{C}_{\tilde{I}J'} - \mu \partial_- h_{\tilde{I}J'} \quad (1e)$$

$$(\tilde{I}J'K') \text{ of (A.3) : } \quad 0 = \square \mathcal{C}_{\tilde{I}J'K'} \quad (1f)$$

$$(I'J'K') \text{ of (A.3) : } \quad 0 = \square \mathcal{C}_{I'J'K'} - \frac{1}{6}\mu \varepsilon^{I'J'K'W'X'Y'} \partial_- \mathcal{C}_{W'X'Y'} \quad (1g)$$

$$\varepsilon^{I'J'K'W'X'Y'} : \text{ SO(6) Levi-Civita symbol} \quad (\varepsilon_{456789} = \varepsilon^{456789} = 1)$$

$$\varepsilon_{\tilde{I}\tilde{J}\tilde{K}} : \text{ SO(3) Levi-Civita symbol} \quad (\varepsilon_{123} = \varepsilon^{123} = 1)$$

$$\mathcal{C}_{\tilde{I}J'} \equiv \frac{1}{2} \varepsilon_{\tilde{I}\tilde{K}\tilde{L}} \mathcal{C}_{\tilde{K}\tilde{L}J'} \quad \mathcal{C} \equiv 2\mathcal{C}_{123}$$

(1f) leads to the zero-mode energy  $\mathcal{E}_0(\mathcal{C}_{\tilde{I}J'K'})$  and degrees of freedom  $\mathcal{D}(\mathcal{C}_{\tilde{I}J'K'})$ :

$$\mathcal{E}_0(\mathcal{C}_{\tilde{I}J'K'}) = 2 \quad \mathcal{D}(\mathcal{C}_{\tilde{I}J'K'}) = 45$$

Diagonalize  $h_{\tilde{I}J'}$  and  $\mathcal{C}_{\tilde{I}J'}$ :

$$H_{\tilde{I}J'} = h_{\tilde{I}J'} + i\mathcal{C}_{\tilde{I}J'} \quad \bar{H}_{\tilde{I}J'} = h_{\tilde{I}J'} - i\mathcal{C}_{\tilde{I}J'}$$

Thus modified (1b) and (1e) are

$$\begin{aligned} 0 &= (\square - \mu i\partial_-)H_{\tilde{I}J'} & 0 &= (\square + \mu i\partial_-)\bar{H}_{\tilde{I}J'} \\ \implies \mathcal{E}_0(H_{\tilde{I}J'}) &= 1 & \mathcal{E}_0(\bar{H}_{\tilde{I}J'}) &= 3 & \mathcal{D}(H_{\tilde{I}J'}) &= \mathcal{D}(\bar{H}_{\tilde{I}J'}) = 18 \end{aligned}$$

Apply similar consideration to (1a), (1c) and (1d):

$$\begin{aligned} h_{\tilde{I}\tilde{J}}^\perp &\equiv h_{\tilde{I}\tilde{J}} - \frac{1}{3}\delta_{\tilde{I}\tilde{J}}h_{\tilde{K}\tilde{K}} & h_{I'J'}^\perp &\equiv h_{I'J'} - \frac{1}{6}\delta_{I'J'}h_{K'K'} \\ h &\equiv h_{\tilde{I}\tilde{I}} + i\mathcal{C} & \bar{h} &\equiv h_{\tilde{I}\tilde{I}} - i\mathcal{C} \end{aligned}$$

Then we find

$$\begin{aligned} \mathcal{E}_0(h_{\tilde{I}\tilde{J}}^\perp) &= \mathcal{E}_0(h_{I'J'}^\perp) = 2 & \mathcal{D}(h_{\tilde{I}\tilde{J}}^\perp) &= 5 & \mathcal{D}(h_{I'J'}^\perp) &= 20 \\ \mathcal{E}_0(h) &= 0 & \mathcal{E}_0(\bar{h}) &= 4 & \mathcal{D}(h) &= \mathcal{D}(\bar{h}) = 1 \end{aligned}$$

Decomposing  $\mathcal{C}_{I'J'K'}$  into self-dual part  $\mathcal{C}_{I'J'K'}^\oplus$  and anti-self-dual part  $\mathcal{C}_{I'J'K'}^\ominus$ :

$$\mathcal{C}_{I'J'K'}^\oplus \equiv \frac{i}{3!} \varepsilon^{I'J'K'W'X'Y'} \mathcal{C}_{W'X'Y'}^\oplus \quad \mathcal{C}_{I'J'K'}^\ominus \equiv -\frac{i}{3!} \varepsilon^{I'J'K'W'X'Y'} \mathcal{C}_{W'X'Y'}^\ominus$$

They satisfy the following equations:

$$\begin{aligned} (\square + \mu i \partial_-) \mathcal{C}_{I'J'K'}^\oplus &= 0 & (\square - \mu i \partial_-) \mathcal{C}_{I'J'K'}^\ominus &= 0 \\ \Rightarrow \mathcal{E}_0(\mathcal{C}_{I'J'K'}^\oplus) &= 3 & \mathcal{E}_0(\mathcal{C}_{I'J'K'}^\ominus) &= 1 & \mathcal{D}(\mathcal{C}_{I'J'K'}^\oplus) &= \mathcal{D}(\mathcal{C}_{I'J'K'}^\ominus) &= 10 \end{aligned}$$

Now we have fully solved the field equations for bosonic fluctuations and have derived the spectrum of graviton  $h_{MN}$  and three-form gauge field  $\mathcal{C}_{MNP}$ . The resulting spectrum is splitting with a certain energy difference in contrast to the flat case. The resulting spectrum is **completely identical** with that of zero-mode Hamiltonian in the supermembrane theory on the pp-wave background.

We summarize this result with fermion spectrum.

## Fermionic Fields

$M = -$  component of (A.2):

$$\hat{\Gamma}^P \psi_P = 0 \quad \Longrightarrow \quad \text{Lorentz-type gauge-fixing condition}$$

$M = +$  component of (A.2):

$$\partial^P \psi_P = 0 \quad \rightarrow \quad \psi_+ = \frac{1}{\partial_-} \partial_I \psi_I \quad \Longrightarrow \quad \text{non-dynamical}$$

$M = \tilde{I}$  component of (A.2):

$$\psi_{\tilde{I}}^{\ominus} = \frac{1}{2\partial_-} \hat{\Gamma}^+ \hat{\Gamma}^K \partial_K \psi_{\tilde{I}}^{\oplus} \quad \Longrightarrow \quad \text{non-dynamical}$$

$$0 = \square \psi_{\tilde{I}}^{\oplus} - \frac{1}{2} \mu \hat{\Gamma}^{123} (\delta_{\tilde{I}\tilde{J}} - \hat{\Gamma}_{\tilde{I}} \hat{\Gamma}_{\tilde{J}}) \partial_- \psi_{\tilde{J}}^{\oplus} \quad (2)$$

where  $\psi_{\tilde{I}} \equiv \psi_{\tilde{I}}^{\oplus} + \psi_{\tilde{I}}^{\ominus}$ :  $\psi_{\tilde{I}}^{\oplus} \equiv -\frac{1}{2} \hat{\Gamma}^- \hat{\Gamma}^+ \psi_{\tilde{I}}$  and  $\psi_{\tilde{I}}^{\ominus} \equiv -\frac{1}{2} \hat{\Gamma}^+ \hat{\Gamma}^- \psi_{\tilde{I}}$

In order to solve this equation, we shall introduce the following fields:

$$\begin{aligned} \psi_{\tilde{I}}^{\oplus\perp} &\equiv \left( \delta_{\tilde{I}\tilde{J}} - \frac{1}{3} \hat{\Gamma}_{\tilde{I}} \hat{\Gamma}_{\tilde{J}} \right) \psi_{\tilde{J}}^{\oplus} && \hat{\Gamma}\text{-transverse mode} \\ \psi_{\tilde{I}}^{\oplus\parallel} &\equiv \hat{\Gamma}^{\tilde{I}} \psi_{\tilde{I}}^{\oplus} && \hat{\Gamma}\text{-parallel mode} \end{aligned}$$

Acting  $\widehat{\Gamma}^{\widetilde{I}}$  or  $(\delta_{\widetilde{K}\widetilde{I}} - \frac{1}{3}\widehat{\Gamma}_{\widetilde{K}}\widehat{\Gamma}_{\widetilde{I}})$  on (2):

$$0 = \square \psi_1^{\oplus\parallel} - \mu \widehat{\Gamma}^{123} \partial_- \psi_1^{\oplus\parallel} \quad 0 = \square \psi_{\widetilde{K}}^{\oplus\perp} - \frac{1}{2} \mu \widehat{\Gamma}^{123} \partial_- \psi_{\widetilde{K}}^{\oplus\perp} \quad (3)$$

Decompose  $\psi_{\widetilde{I}}^{\oplus\perp}$  and  $\psi_1^{\oplus\parallel}$  according to the chirality in terms of  $i\widehat{\Gamma}^{123}$ :

$$\begin{aligned} \psi_{\widetilde{I}\text{R}}^{\oplus\perp} &\equiv \frac{1 + i\widehat{\Gamma}^{123}}{2} \psi_{\widetilde{I}}^{\oplus\perp} & \psi_{\widetilde{I}\text{L}}^{\oplus\perp} &\equiv \frac{1 - i\widehat{\Gamma}^{123}}{2} \psi_{\widetilde{I}}^{\oplus\perp} \\ \psi_{1\text{R}}^{\oplus\parallel} &\equiv \frac{1 + i\widehat{\Gamma}^{123}}{2} \psi_1^{\oplus\parallel} & \psi_{1\text{L}}^{\oplus\parallel} &\equiv \frac{1 - i\widehat{\Gamma}^{123}}{2} \psi_1^{\oplus\parallel} \end{aligned}$$

Multiplying chiral projection operator  $\frac{1}{2}(1 \pm i\widehat{\Gamma}^{123})$  to (3):

$$\begin{aligned} 0 &= (\square + \mu i\partial_-) \psi_{1\text{R}}^{\oplus\parallel} & 0 &= (\square - \mu i\partial_-) \psi_{1\text{L}}^{\oplus\parallel} \\ 0 &= (\square + \frac{1}{2}\mu i\partial_-) \psi_{\widetilde{I}\text{R}}^{\oplus\perp} & 0 &= (\square - \frac{1}{2}\mu i\partial_-) \psi_{\widetilde{I}\text{L}}^{\oplus\perp} \end{aligned}$$

Zero-mode energies and degrees of freedom of  $\psi_{\widetilde{I}\text{R}}^{\oplus\perp}$  and  $\psi_{\widetilde{I}\text{L}}^{\oplus\perp}$ :

$$\mathcal{E}_0(\psi_{\widetilde{I}\text{R}}^{\oplus\perp}) = \frac{5}{2} \quad \mathcal{E}_0(\psi_{\widetilde{I}\text{L}}^{\oplus\perp}) = \frac{3}{2} \quad \mathcal{D}(\psi_{\widetilde{I}\text{R}}^{\oplus\perp}) = \mathcal{D}(\psi_{\widetilde{I}\text{L}}^{\oplus\perp}) = 8 \times (3 - 1) = 16$$

We will discuss  $\mathcal{E}_0$  and  $\mathcal{D}$  of  $\widehat{\Gamma}$ -parallel modes later on.

$M = I'$  component of (A.2):

$$0 = \left\{ \hat{\Gamma}^+ \left( \partial_+ + \frac{1}{2} \mathbf{G}_{++} \partial_- \right) + \hat{\Gamma}^- \partial_- + \hat{\Gamma}^K \partial_K \right\} \psi_{I'} + \frac{1}{4} \mu \hat{\Gamma}^{+123} \left( \delta_{I'J'} - \hat{\Gamma}_{I'} \hat{\Gamma}_{J'} \right) \psi_{J'}$$

Decompose  $\psi_{I'}$  into the  $\hat{\Gamma}$ -parallel mode and  $\hat{\Gamma}$ -transverse mode:

$$\begin{aligned} 0 &= \left( \square - \frac{5}{2} \mu i \partial_- \right) \psi_{2R}^{\oplus\parallel} & 0 &= \left( \square + \frac{5}{2} \mu i \partial_- \right) \psi_{2L}^{\oplus\parallel} \\ 0 &= \left( \square - \frac{1}{2} \mu i \partial_- \right) \psi_{I'R}^{\oplus\perp} & 0 &= \left( \square + \frac{1}{2} \mu i \partial_- \right) \psi_{I'L}^{\oplus\perp} \end{aligned}$$

where the  $\hat{\Gamma}$ -transverse mode and  $\hat{\Gamma}$ -parallel mode are defined as

$$\begin{aligned} \psi_{I'}^{\oplus} &= -\frac{1}{2} \hat{\Gamma}^- \hat{\Gamma}^+ \psi_{I'} & \psi_{I'R}^{\oplus\perp} &= \frac{1 + i \hat{\Gamma}^{123}}{2} \psi_{I'}^{\oplus\perp} & \psi_{I'L}^{\oplus\perp} &= \frac{1 - i \hat{\Gamma}^{123}}{2} \psi_{I'}^{\oplus\perp} \\ \psi_{2R}^{\oplus\parallel} &= \frac{1 + i \hat{\Gamma}^{123}}{2} \psi_2^{\oplus\parallel} & \psi_{2L}^{\oplus\parallel} &= \frac{1 - i \hat{\Gamma}^{123}}{2} \psi_2^{\oplus\parallel} \end{aligned}$$

We find that the zero-mode energies and degrees of freedom:

$$\mathcal{E}_0(\psi_{I'R}^{\oplus\perp}) = \frac{3}{2} \quad \mathcal{E}_0(\psi_{I'L}^{\oplus\perp}) = \frac{5}{2} \quad \mathcal{D}(\psi_{I'R}^{\oplus\perp}) = \mathcal{D}(\psi_{I'L}^{\oplus\perp}) = 8 \times (6 - 1) = 40$$

Linear combination of  $\widehat{\Gamma}$ -parallel modes:

$$\psi_{\text{R}}^{\oplus\parallel} \equiv \frac{2}{5}\psi_{1\text{R}}^{\oplus\parallel} - \psi_{2\text{R}}^{\oplus\parallel} \quad \psi_{\text{L}}^{\oplus\parallel} \equiv \frac{2}{5}\psi_{1\text{L}}^{\oplus\parallel} - \psi_{2\text{L}}^{\oplus\parallel}$$

We can easily see that the re-defined fermions satisfy:

$$\begin{aligned} 0 &= \left( \square - \frac{3}{2}\mu i\partial_- \right) \psi_{\text{R}}^{\oplus\parallel} & 0 &= \left( \square + \frac{3}{2}\mu i\partial_- \right) \psi_{\text{L}}^{\oplus\parallel} \\ \Rightarrow \mathcal{E}_0(\psi_{\text{R}}^{\oplus\parallel}) &= \frac{1}{2} & \mathcal{E}_0(\psi_{\text{L}}^{\oplus\parallel}) &= \frac{7}{2} & \mathcal{D}(\psi_{\text{R}}^{\oplus\parallel}) &= \mathcal{D}(\psi_{\text{L}}^{\oplus\parallel}) = 8 \end{aligned}$$

We have solved and derived the spectrum of gravitino in the case of pp-wave background. As a result, we have found that the spectrum is splitting with a certain energy difference in the same manner with the spectrum of bosons.



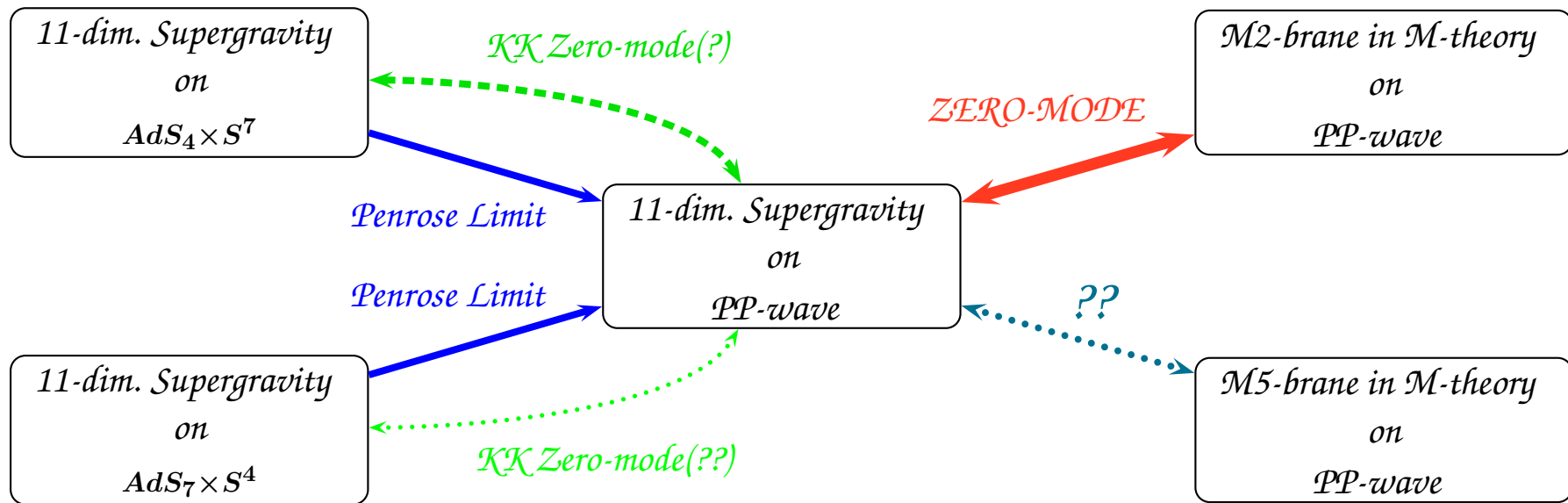
## Results

energy $\mathcal{E}_0$	bosons ( $\mathcal{D}$ )			fermions ( $\mathcal{D}$ )		degrees
4	$\bar{h}(1)$					1
7/2				$\psi_{\text{L}}^{\oplus\parallel}(8)$		8
3	$\bar{H}_{\tilde{I}\tilde{J}'}(18)$	$\mathcal{C}_{\tilde{I}'\tilde{J}'\tilde{K}'}^{\oplus}(10)$				28
5/2				$\psi_{\tilde{I}\tilde{R}}^{\oplus\perp}(16)$	$\psi_{\tilde{I}'\tilde{L}}^{\oplus\perp}(40)$	56
2	$\mathcal{C}_{\tilde{I}\tilde{J}'\tilde{K}'}(45)$	$h_{\tilde{I}\tilde{J}}^{\perp}(5)$	$h_{\tilde{I}'\tilde{J}'}^{\perp}(20)$			70
3/2				$\psi_{\tilde{I}\tilde{L}}^{\oplus\perp}(16)$	$\psi_{\tilde{I}'\tilde{R}}^{\oplus\perp}(40)$	56
1	$H_{\tilde{I}\tilde{J}'}(18)$	$\mathcal{C}_{\tilde{I}'\tilde{J}'\tilde{K}'}^{\ominus}(10)$				28
1/2				$\psi_{\text{R}}^{\oplus\parallel}(8)$		8
0	$h(1)$					1

*Energy spectrum and degrees of freedom of physical modes in 11-dimensional supergravity on the pp-wave background.*

*This result corresponds to that of the zero-mode Hamiltonian in the **supermembrane theory on the pp-wave background**.*

## Conclusion and Discussions



## *Work in Progress*

- Propagators and energy-momentum tensors of  $h_{MN}$ ,  $\mathcal{C}_{MNP}$  and  $\psi_M$
- Dimensional reduction to type IIA supergravity (only 24 supercharges)
- Comparison with KK zero-mode and algebra of  $AdS_7 \times S^4$
- Relationship to Zero-mode spectrum and superalgebra of M5-brane

## Appendix A Penrose Limit of Anti-de Sitter Spaces

- $AdS_4 \times S^7 \rightarrow$  Kowalski-Glikman

$$ds^2 = R^2 \left\{ -(\cosh \rho)^2 dt^2 + d\rho^2 + (\sinh \rho)^2 d\Omega_2^2 \right\} + (2R)^2 \left\{ (\cos \theta)^2 d\varphi^2 + d\theta^2 + (\sin \theta)^2 d\Omega_5^2 \right\}$$

$$x^+ = \frac{\mu}{6}(t + 2\varphi) \quad x^- = \frac{3R^2}{\mu}(t - 2\varphi) \quad x = R\rho \quad y = 2R\theta, \quad R \rightarrow \infty : \text{ Penrose limit}$$

$$\Rightarrow ds^2 = -2dx^+ dx^- - \left(\frac{\mu}{3}\right)^2 \left\{ x^2 + \frac{1}{4}y^2 \right\} (dx^+)^2 + \{dx^2 + x^2 d\Omega_2^2\} + \{dy^2 + y^2 d\Omega_5^2\}$$

- $AdS_7 \times S^4 \rightarrow$  Kowalski-Glikman

$$ds^2 = (2R)^2 \left\{ -(\cosh \rho)^2 dt^2 + d\rho^2 + (\sinh \rho)^2 d\Omega_5^2 \right\} + R^2 \left\{ (\cos \theta)^2 d\varphi^2 + d\theta^2 + (\sin \theta)^2 d\Omega_2^2 \right\}$$

$$x^+ = \frac{\mu}{2} \left( t + \frac{1}{2}\varphi \right) \quad x^- = \frac{4R^2}{\mu} \left( t - \frac{1}{2}\varphi \right) \quad x = 2R\rho \quad y = R\theta, \quad R \rightarrow \infty : \text{ Penrose limit}$$

$$\Rightarrow ds^2 = -2dx^+ dx^- - \mu^2 \left\{ x^2 + 4y^2 \right\} (dx^+)^2 + \{dx^2 + x^2 d\Omega_5^2\} + \{dy^2 + y^2 d\Omega_2^2\}$$

$$x^+ \rightarrow \frac{1}{2}x^+ \quad x^- \rightarrow 2x^- \quad \mu \rightarrow \frac{1}{3}\mu \quad x \leftrightarrow y$$

$$\Rightarrow ds^2 = -2dx^+ dx^- - \left(\frac{\mu}{3}\right)^2 \left\{ x^2 + \frac{1}{4}y^2 \right\} (dx^+)^2 + \{dx^2 + x^2 d\Omega_2^2\} + \{dy^2 + y^2 d\Omega_5^2\}$$

## Appendix B

### Linearized Field Equations for Fluctuations on the PP-wave Background

$$\begin{aligned}
0 = & -\frac{1}{2}g_{MN}\left\{h^{PQ}\mathcal{R}_{PQ}-\nabla^P\nabla^Q h_{PQ}+\nabla^P\nabla_P h_Q{}^Q\right\} \\
& -\frac{1}{2}\left\{\nabla^P\nabla_M h_{NP}+\nabla^P\nabla_N h_{MP}-\nabla_M\nabla_N h_P{}^P-\nabla^P\nabla_P h_{MN}\right\} \\
& -\frac{1}{24}g_{MN}\left\{2F^{PQRS}\partial_P\mathcal{C}_{QRS}-F_{PQRS}F_U{}^{QRS}h^{PU}\right\} \\
& +\frac{1}{3}F_M{}^{PQR}\partial_{[N}\mathcal{C}_{PQR]}+\frac{1}{3}F_N{}^{PQR}\partial_{[M}\mathcal{C}_{PQR]}-\frac{1}{4}F_{MPQR}F_{NU}{}^{QR}h^{PU}
\end{aligned} \tag{A.1}$$

$$\begin{aligned}
0 = & \hat{\Gamma}^{MNP}D_N\psi_P-\frac{1}{4}\mu\hat{\Gamma}^{MN+123}\psi_N \\
& -\frac{1}{4}\mu\left\{g^{M+}(\hat{\Gamma}^{12}g^{3N}+\hat{\Gamma}^{23}g^{1N}+\hat{\Gamma}^{31}g^{2N})-g^{M1}(\hat{\Gamma}^{23}g^{+N}+\hat{\Gamma}^{3+}g^{2N}+\hat{\Gamma}^{+2}g^{3N})\right. \\
& \quad \left.+g^{M2}(\hat{\Gamma}^{3+}g^{1N}+\hat{\Gamma}^{+1}g^{3N}+\hat{\Gamma}^{13}g^{+N})-g^{M3}(\hat{\Gamma}^{+1}g^{2N}+\hat{\Gamma}^{12}g^{+N}+\hat{\Gamma}^{2+}g^{1N})\right\}\psi_N
\end{aligned} \tag{A.2}$$

$$\begin{aligned}
0 = & -\left(\frac{1}{2}g^{QR}h_U{}^U-h^{QR}\right)\left\{\Gamma_{RQ}^S F_{SMNP}+\Gamma_{RM}^S F_{QSNP}+\Gamma_{RN}^S F_{QMSP}+\Gamma_{RP}^S F_{QMNS}\right\} \\
& +4g^{QR}\left\{\partial_R\partial_{[Q}\mathcal{C}_{MNP]}-\Gamma_{RQ}^S\partial_{[S}\mathcal{C}_{MNP]}-\Gamma_{RM}^S\partial_{[Q}\mathcal{C}_{SNP]}-\Gamma_{RN}^S\partial_{[Q}\mathcal{C}_{MSP]}-\Gamma_{RP}^S\partial_{[Q}\mathcal{C}_{MNS]}\right\} \\
& -\frac{1}{2}g^{QR}\left\{F_{SMNP}(\nabla_R h_Q{}^S+\nabla_Q h_R{}^S-\nabla^S h_{RQ})+F_{QSNP}(\nabla_R h_M{}^S+\nabla_M h_R{}^S-\nabla^S h_{RM})\right. \\
& \quad \left.+F_{QMSP}(\nabla_R h_N{}^S+\nabla_N h_R{}^S-\nabla^S h_{RN})+F_{QMNS}(\nabla_R h_P{}^S+\nabla_P h_R{}^S-\nabla^S h_{RP})\right\} \\
& +\frac{1}{144}g_{MZ}g_{NK}g_{PL}\varepsilon^{ZKLQRSUVWXY}F_{QRSU}\partial_V\mathcal{C}_{WXY}
\end{aligned} \tag{A.3}$$