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Realization of AdS Vacua in Attractor Mechanism on Generalized Geometry

arXiv:0810.0937 [hep-th]

Tetsuji KIMURA Yukawa Institute for Theoretical Physics, Kyoto University We are looking for the origin of 4D physics

Physical information

- particle contents and spectra
- (broken) symmetries and interactions
- potential, vacuum and cosmological constant

10D string theories could provide information via compactifications

A typical success:

 $-E_8 \times E_8$ heterotic string compactified on Calabi-Yau three-fold

- number of generations = $|\chi(CY_3)|/2$
- E_6 gauge symmetry
- zero cosmological constant

P. Candelas, G.T. Horowitz, A. Strominger, E. Witten "Vacuum configurations for superstrings," Nucl. Phys. B 258 (1985) 46

But, this vacuum is too simple

(No non-trivial background fluxes)

Relax assumptions: introduce non-trivial background objects

NS-fluxes, fermion condensations, RR-fluxes, D-branes, etc.

Torsion, warp factor, and/or cosmological constant are generated

How should we evaluate matter contents and vacua?

Study again compactification scenarios

Decompose 10D SUSY variations:

$$\epsilon^{1} = \varepsilon_{1} \otimes (a \eta_{+}^{1}) + \varepsilon_{1}^{c} \otimes (\overline{a} \eta_{-}^{1}) \qquad \epsilon^{2} = \varepsilon_{2} \otimes (\overline{b} \eta_{-}^{2}) + \varepsilon_{2}^{c} \otimes (b \eta_{+}^{2})$$

$$\delta \Psi_{M}^{\mathcal{A}} = 0 \qquad \begin{pmatrix} \delta \psi_{\mathcal{A}\mu} = 0 & \dashrightarrow & \text{superpotential } \mathcal{W} \\ \delta \psi_{m}^{\mathcal{A}} = 0 & \dashrightarrow & \text{K\"ahler potential } K \end{pmatrix}$$

Decompose 10D SUSY variations:

$$\epsilon^1 = \varepsilon_1 \otimes (a \eta^1_+) + \varepsilon_1^c \otimes (\overline{a} \eta^1_-) \qquad \epsilon^2 = \varepsilon_2 \otimes (\overline{b} \eta^2_-) + \varepsilon_2^c \otimes (b \eta^2_+)$$

$$\delta \Psi_M^{\mathcal{A}} = 0 \quad \left\langle \begin{array}{cc} \delta \psi_{\mathcal{A}\mu} = 0 & \dashrightarrow & \text{superpotential } \mathcal{W} \\ \delta \psi_m^{\mathcal{A}} = 0 & \dashrightarrow & \text{K\"ahler potential } \mathbf{K} \end{array} \right.$$

 $\delta \psi_m^{\mathcal{A}} = 0$ gives the Killing spinor equation on the compactified space \mathcal{M} :

$$\delta \psi_m^{\mathcal{A}} = \left(\partial_m + \frac{1}{4} \omega_{mab} \gamma^{ab} \right) \eta_+^{\mathcal{A}} + \left(\text{3-form fluxes} \cdot \eta \right)^{\mathcal{A}} + \left(\text{other fluxes} \cdot \eta \right)^{\mathcal{A}} = 0$$

Information of 6D SU(3) Killing spinors η^1_+ , η^2_+ :

Calabi-Yau three-fold

$$\downarrow$$

 $SU(3)$ -structure manifold with torsion
 \downarrow
generalized geometry

► Calabi-Yau three-fold --→ Fluxes are strongly restricted

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\begin{cases} \text{type IIA: No fluxes} \\ \text{type IIB: } F_3 - \tau H \\ \text{heterotic: No fluxes} \end{cases}
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 \blacktriangleright SU(3)-structure manifold --+ Some components of fluxes can be interpreted as torsion

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type IIA
type IIB
heterotic<sup>1</sup>
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1: Piljin Yi, TK "Comments on heterotic flux compactifications," JHEP 0607 (2006) 030, hep-th/0605247

2: TK "Index theorems on torsional geometries," JHEP 0708 (2007) 048, arXiv:0704.2111

Generalized geometry --+ Any types of fluxes can be included

All the $\mathcal{N} = 1$ SUSY solutions can be classified

Consider the compactified space \mathcal{M}_6

• Ordinary complex structure $J^m{}_n$ lives in $T\mathcal{M}$:

$$J^{2} = -\mathbf{1}_{6}, \qquad J^{m}{}_{n} = -2i \eta^{\dagger}_{+} \gamma^{m}{}_{n} \eta_{+}$$
$$\eta_{+}: SU(3) \text{ invariant spinor}$$

• Generalized complex structures $\mathcal{J}^{\Lambda}{}_{\Sigma}$ in $T\mathcal{M}\oplus T^{*}\mathcal{M}$

with basis $\{ dx^m \land, \iota_{\partial_n} \}$ and (6, 6)-signature

$$\begin{aligned} \mathcal{J}^2 &= -\mathbf{1}_{12}, \qquad \mathcal{J}^{\Lambda}_{\pm \Sigma} &= \left\langle \operatorname{Re} \Phi_{\pm}, \Gamma^{\Lambda}_{\Sigma} \operatorname{Re} \Phi_{\pm} \right\rangle \\ \Phi_{\pm}: SU(3,3) \text{ invariant spinors} \end{aligned}$$

 Φ_{\pm} can be described by means of η^1_{\pm} and η^2_{\pm} in SUSY parameters

10D type IIA supergravity as a low energy theory of IIA string

compactifications on a certain compact space in the presence of fluxes 4D $\mathcal{N} = 2$ supergravity SUSY truncation 4D $\mathcal{N} = 1$ supergravity

In this talk

we focus on the search of $\mathcal{N}=1$ SUSY vacua

$$S = \int \left(\frac{1}{2}R * \mathbf{1} - \frac{1}{2}F^{a} \wedge *F^{a} - K_{\mathcal{M}\overline{\mathcal{N}}}\nabla\phi^{\mathcal{M}} \wedge *\nabla\overline{\phi}^{\overline{\mathcal{N}}} - V * \mathbf{1}\right)$$
$$V = e^{K} \left(K^{\mathcal{M}\overline{\mathcal{N}}}D_{\mathcal{M}}\mathcal{W}\overline{D_{\mathcal{N}}\mathcal{W}} - 3|\mathcal{W}|^{2}\right) + \frac{1}{2}|D^{a}|^{2}$$

K :	Kähler	potential	

$${\cal W}$$
: superpotential $igstarrow {ar \epsilon}_{--} = \delta \psi_\mu =
abla_\mu \, arepsilon - \, {
m e}^{rac{ar \kappa}{2}} \, {\cal W} \, \gamma_\mu \, arepsilon^c$

$$D^a$$
: D-term $\leftarrow - \delta \chi^a = \mathrm{Im} F^a_{\mu\nu} \gamma^{\mu\nu} \varepsilon + \mathrm{i} D^a \varepsilon$

Search of vacua $\partial_{\mathcal{P}} V \big|_* = 0$

 $V_* > 0$: de Sitter space $V_* = 0$: Minkowski space $V_* < 0$: Anti-de Sitter space

4D $\mathcal{N} = 1$ Minkowski vacua in type IIA

Classification of SUSY solutions on the SU(3) generalized geometries $(\eta_+^1 = \eta_+^2)$:

IIA	a = 0 or $b = 0$ (type A)	$a=b{ m e}^{{ m i}eta}$ (type BC)	
1	$\mathcal{W}_1 =$		
T	$F_0^{(1)} = \mp F_2^{(1)} = F_4^{(1)} = \mp F_6^{(1)}$	$F_{2n}^{(1)} = 0$	
		generic eta	$\beta = 0$
8	$\mathcal{W}_2 = F_2^{(8)} = F_4^{(8)} = 0$	$\mathrm{Re}\mathcal{W}_2 = \mathrm{e}^{\phi} F_2^{(8)}$	$\mathrm{Re}\mathcal{W}_2 = \mathrm{e}^{\phi} F_2^{(8)} + \mathrm{e}^{\phi} F_4^{(8)}$
		$\mathrm{Im}\mathcal{W}_2=0$	$\mathrm{Im}\mathcal{W}_2=0$
6	$\mathcal{W}_3 = \mp *_6 H_3^{(6)}$	$W_3 = H_3^{(6)} = 0$	
3	$\overline{\mathcal{W}}_5 = 2\mathcal{W}_4 = \mp 2\mathrm{i}H_3^{(\overline{3})} = \overline{\partial}\phi$	$F_2^{(\overline{3})} = 2\mathrm{i}\overline{\mathcal{W}}_5 = -2\mathrm{i}\overline{\partial}A = \frac{2\mathrm{i}}{3}\overline{\partial}\phi$	
	$\overline{\partial}A = \overline{\partial}a = 0$	$\mathcal{W}_4 = 0$	

M. Graña, R. Minasian, M. Petrini, A. Tomasiello hep-th/0407249 M. Graña hep-th/0509003

type A	NS-flux only (common to IIA, IIB, heterotic) $W_1 = W_2 = 0, W_3 \neq 0$: complex
type //	$\mathcal{W}_1 = \mathcal{W}_2 = 0$, $\mathcal{W}_3 eq 0$: complex
type BC	RR-flux only
	RR-flux only $W_1 = \text{Im}W_2 = W_3 = W_4 = 0$, $\text{Re}W_2 \neq 0$, $W_5 \neq 0$: symplectic

For $\mathcal{N} = 1$ AdS₄ vacua: hep-th/0403049, hep-th/0407263, hep-th/0412250, hep-th/0502154, hep-th/0609124, etc.

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8	$\mathcal{W}_2 = F_2^{(8)} = F_4^{(8)} = 0$	$\mathrm{Re}\mathcal{W}_2 = \mathrm{e}^{\phi} F_2^{(8)}$	$\text{Re}\mathcal{W}_2 = e^{\phi}F_2^{(8)} + e^{\phi}F_4^{(8)}$
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M. Graña, R. Minasian, M. Petrini, A. Tomasiello hep-th/0407249 M. Graña hep-th/0509003

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For $\mathcal{N} = 1$ AdS₄ vacua: hep-th/0403049, hep-th/0407263, hep-th/0412250, hep-th/0502154, hep-th/0609124, etc.

 $SU(3) \times SU(3)$ generalized geometries $(\eta^1_+ \neq \eta^2_+ \text{ at some points})$

would complete the classification. (But, it's quite hard to find all solutions.)

Moduli stabilization

We obtain SUSY AdS or Minkowski vacuum on an attractor point

Mathematical feature

We find a powerful rule to evaluate the attractor points:

Discriminants of the superpotential govern the cosmological constant

Stringy effects

We see that α' corrections are included in reduced configurations

cf.) heterotic string compactifications on SU(3)-structure manifolds

Contents

 $\mathbf{O} \mathcal{N} = 1$ scalar potential from generalized geometry

Search of SUSY vacua

Summary and discussions

$\mathcal{N}=1$ scalar potential from generalized geometry

Consider a compact space \mathcal{M}_6

• Ordinary complex structure $J^m{}_n$ in $T\mathcal{M}$ is given by SU(3) Weyl spinor η_+ :

$$J^2 = -\mathbf{1}_6, \qquad J^m{}_n = -2\mathrm{i}\,\eta^{\dagger}_+\,\gamma^m{}_n\,\eta_+$$

• Generalized complex structures $\mathcal{J}^{\Lambda}_{\Sigma}$ in $T\mathcal{M} \oplus T^*\mathcal{M}$: basis $\{dx^m \land, \iota_{\partial_n}\}$, (6, 6)-signature

$$\mathcal{J}_{\pm}^2 = -\mathbf{1}_{12}, \qquad \mathcal{J}_{\pm \Sigma}^{\Lambda} = \left\langle \operatorname{Re} \Phi_{\pm}, \Gamma^{\Lambda}{}_{\Sigma} \operatorname{Re} \Phi_{\pm} \right\rangle$$

 $\Phi_{\pm}: \quad SU(3,3) \text{ Weyl spinors}$

 Γ^{Λ} : Cliff(6, 6) gamma matrix

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$$\Phi_{\pm}: \quad SU(3,3) \text{ Weyl spinors } isomorphic \text{ to even/odd-forms on } T^{*}\mathcal{M}$$

$$\Gamma^{\Lambda}: \quad \operatorname{Cliff}(6,6) \text{ gamma matrix } (\operatorname{repr.} = (\operatorname{d} x^{m} \wedge, \iota_{\partial_{n}}))$$

Mukai pairing:

even forms:
$$\langle \Psi_+, \Phi_+ \rangle = \Psi_6 \wedge \Phi_0 - \Psi_4 \wedge \Phi_2 + \Psi_2 \wedge \Phi_4 - \Psi_0 \wedge \Phi_6$$

odd forms: $\langle \Psi_-, \Phi_- \rangle = \Psi_5 \wedge \Phi_1 - \Psi_3 \wedge \Phi_3 + \Psi_1 \wedge \Phi_5$

There are two SU(3) spinors on ${\mathfrak M}:$ $\eta^1_+\text{, }\eta^2_+$ with

$$\eta_{+}^{2} = c_{\parallel}(y)\eta_{+}^{1} + c_{\perp}(y)(v + iv')^{m} \gamma_{m} \eta_{-}^{1}, \qquad (v - iv')^{m} = \eta_{+}^{1\dagger} \gamma^{m} \eta_{-}^{2}$$

There are two SU(3) spinors on \mathcal{M} : η^1_+ , η^2_+ with

$$\eta_{+}^{2} = c_{\parallel}(y)\eta_{+}^{1} + c_{\perp}(y)(v + iv')^{m} \gamma_{m} \eta_{-}^{1}, \qquad (v - iv')^{m} = \eta_{+}^{1\dagger} \gamma^{m} \eta_{-}^{2}$$

▶ On the SU(3) generalized geometry $(\eta_+^1 = \eta_+^2)$:

$$\Phi_{+} = e^{-B - iJ}, \qquad \Phi_{-} = e^{-B}\Omega$$
$$J_{mn} = -2i \eta_{+}^{\dagger} \gamma_{mn} \eta_{+}, \qquad \Omega_{mnp} = -2i \eta_{-}^{\dagger} \gamma_{mnp} \eta_{+}$$

▶ On the $SU(3) \times SU(3)$ generalized geometry $(\eta_+^1 \neq \eta_+^2)$ at some points y):

$$\Phi_{+} = e^{-B} (\overline{c}_{\parallel} e^{-ij} - i\overline{c}_{\perp} w) \wedge e^{-iv \wedge v'}, \qquad \Phi_{-} = e^{-B} (c_{\parallel} e^{-ij} + ic_{\perp} w) \wedge (v + iv')$$
$$J^{\mathcal{A}} = j \pm v \wedge v', \qquad \Omega^{\mathcal{A}} = w \wedge (v \pm iv')$$

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Such a space \mathcal{M} has two moduli spaces: special Kähler geometries of local type Kähler potentials, prepotentials, projective coordinates

--> similar to 4D $\mathcal{N} = 2$ supergravity by Calabi-Yau compactifications

M. Graña, J. Louis, D. Waldram hep-th/0505264

Moduli spaces of $\mathcal M$ are the special Kähler geometries of local type Kähler potentials, prepotentials, projective coordinates

$$K_{+} = -\log i \int_{\mathcal{M}} \langle \Phi_{+}, \overline{\Phi}_{+} \rangle = -\log i \left(\overline{X}^{A} \mathcal{F}_{A} - X^{A} \overline{\mathcal{F}}_{A} \right)$$
$$K_{-} = -\log i \int_{\mathcal{M}} \langle \Phi_{-}, \overline{\Phi}_{-} \rangle = -\log i \left(\overline{Z}^{I} \mathcal{G}_{I} - Z^{I} \overline{\mathcal{G}}_{I} \right)$$

Expand the even/odd-forms Φ_{\pm} by the basis forms:

$$\Phi_{+} = X^{A}\omega_{A} - \mathcal{F}_{A}\widetilde{\omega}^{A}, \qquad \omega_{A} = (1, \omega_{a}), \qquad \widetilde{\omega}^{A} = (\widetilde{\omega}^{a}, \operatorname{vol}(\mathcal{M})) \qquad : \quad 0, 2, 4, 6 \text{-forms}$$

$$\Phi_{-} = Z^{I}\alpha_{I} - \mathcal{G}_{I}\beta^{I}, \qquad \alpha_{I} = (\alpha_{0}, \alpha_{i}), \qquad \beta^{I} = (\beta^{i}, \beta^{0}) \qquad : \quad 1, 3, 5 \text{-forms}$$

$$\int_{\mathcal{M}} \langle \omega_A, \omega_B \rangle = 0, \quad \int_{\mathcal{M}} \langle \omega_A, \widetilde{\omega}^B \rangle = \delta_A^B, \quad \int_{\mathcal{M}} \langle \alpha_I, \alpha_J \rangle = 0, \quad \int_{\mathcal{M}} \langle \alpha_I, \beta^J \rangle = \delta_I^J$$

 $\mathcal{N} = 1$ Kähler potential and dilaton are given as

$$K = K_{+} + 4\varphi$$
$$e^{-2\varphi} = \frac{|\mathcal{C}|^{2}}{16|a|^{2}|b|^{2}}e^{-K_{-}} = \frac{i}{16|a|^{2}|b|^{2}}\int_{\mathcal{M}} \langle \mathcal{C}\Phi_{-}, \overline{\mathcal{C}\Phi_{-}} \rangle$$

 $\mathcal{N} = 1$ SUSY variations yield superpotential and D-term:

$$\delta \psi_{\mu} = \nabla_{\mu} \varepsilon - \overline{n}^{\mathcal{A}} S_{\mathcal{A}\mathcal{B}} n^{*\mathcal{B}} \gamma_{\mu} \varepsilon^{c} \equiv \nabla_{\mu} \varepsilon - e^{\frac{K}{2}} \mathcal{W} \gamma_{\mu} \varepsilon^{c}$$
$$\delta \chi^{A} = \operatorname{Im} F^{A}_{\mu\nu} \gamma^{\mu\nu} \varepsilon + \mathrm{i} D^{A} \varepsilon$$

$$\mathcal{W} = \frac{\mathrm{i}}{4\overline{a}b} \Big[4\mathrm{i}\,\mathrm{e}^{\frac{K_{-}}{2} - \varphi} \int_{\mathcal{M}} \left\langle \Phi_{+}, \mathcal{D}\mathrm{Im}(ab\Phi_{-}) \right\rangle + \frac{1}{\sqrt{2}} \int_{\mathcal{M}} \left\langle \Phi_{+}, \mathbf{G} \right\rangle \Big]$$
$$D^{A} = 2\,\mathrm{e}^{K_{+}}(K_{+})^{c\overline{d}} D_{c} X^{A} \overline{D_{d}} \overline{X^{B}} \Big[\overline{n}^{\mathcal{C}}(\sigma_{x})_{\mathcal{C}} \mathcal{B}n_{\mathcal{B}} \Big] \Big(\mathcal{P}^{x}_{B} - \mathcal{N}_{BC} \widetilde{\mathcal{P}}^{xC} \Big)$$

On the SU(3) geometry $(\eta^1_+ = \eta^2_+)$:

$$d_{H}\omega_{A} = m_{A}{}^{I}\alpha_{I} - e_{IA}\beta^{I} \qquad d_{H}\widetilde{\omega}^{A} = 0$$

$$d_{H}\alpha_{I} = e_{IA}\widetilde{\omega}^{A} \qquad d_{H}\beta^{I} = m_{A}{}^{I}\widetilde{\omega}^{A}$$

where NS three-form H deforms the differential operator:

$$dH = 0, \qquad H = H^{\mathsf{fl}} + dB, \qquad H^{\mathsf{fl}} = m_0^{I} \alpha_I - e_{I0} \beta^I$$
$$d_H \equiv d - H^{\mathsf{fl}} \wedge$$

background	charges	
NS-flux charges	$e_{I0} m_0^I$	
torsion	$e_{Ia} m_a{}^I$	

On the $SU(3) \times SU(3)$ geometry $(\eta^1_+ \neq \eta^2_+ \text{ at some points})$:

Extend to the generalized differential operator \mathcal{D} :

$$d_H = d - H^{\mathsf{fl}} \land \longrightarrow \mathcal{D} \equiv d - H^{\mathsf{fl}} \land -Q \cdot -R \sqcup$$

$$\mathcal{D}\omega_A \sim m_A{}^I \alpha_I - e_{IA}\beta^I \qquad \mathcal{D}\widetilde{\omega}^A \sim -q^{IA}\alpha_I + p_I{}^A\beta^I$$
$$\mathcal{D}\alpha_I \sim p_I{}^A \omega_A + e_{IA}\widetilde{\omega}^A \qquad \mathcal{D}\beta^I \sim q^{IA}\omega_A + m_A{}^I\widetilde{\omega}^A$$

Necessary to introduce new fluxes Q and R to make a consistent algebra...

On the $SU(3) \times SU(3)$ geometry $(\eta^1_+ \neq \eta^2_+ \text{ at some points})$:

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Necessary to introduce new fluxes Q and R to make a consistent algebra...

But the compactified geometry becomes nongeometric:

 $\begin{array}{lll} (Q \cdot C)_{m_1 \cdots m_{k-1}} & \equiv & Q^{ab}_{[m_1} C_{|ab|m_2 \cdots m_{k-1}]} & \mbox{feature of T-fold} \\ (R \llcorner C)_{m_1 \cdots m_{k-3}} & \equiv & R^{abc} C_{abcm_1 \cdots m_{k-3}} & \mbox{locally nongeometric background} \end{array}$

Structure group contains Diffeo. + Duality trsf. \rightarrow Doubled formalism³

3: C. Albertsson, R.A. Reid-Edwards, TK "D-branes and doubled geometry," arXiv:0806.1783

RR-fluxes $F^{\text{even}} = e^B G$ without localized sources on the SU(3) geometry:

$$G = G_0 + G_2 + G_4 + G_6 = G^{\mathsf{fl}} + d_H A$$

$$F_n^{\mathsf{even}} = dC_{n-1} - H \wedge C_{n-3}, \quad C = e^B A$$

$$d_H F^{\mathsf{even}} = 0$$

Extension of RR-fluxes on the $SU(3) \times SU(3)$ geometry:

$$G = G^{\mathsf{fl}} + \mathcal{D}A, \qquad \mathcal{D}G = 0$$

$$G^{\mathsf{fl}} = \sqrt{2} \left(m_{\mathsf{RR}}^A \,\omega_A - e_{\mathsf{RR}A} \,\widetilde{\omega}^A \right), \qquad A = \sqrt{2} \left(\xi^I \,\alpha_I - \widetilde{\xi}_I \,\beta^I \right)$$

$$\Downarrow$$

$$G \sim G^{A} \omega_{A} - \widetilde{G}_{A} \widetilde{\omega}^{A}$$
$$G^{A} \sim \sqrt{2} \left(m_{\mathsf{RR}}^{A} + \xi^{I} p_{I}^{A} - \widetilde{\xi}_{I} q^{IA} \right), \qquad \widetilde{G}_{A} \sim \sqrt{2} \left(e_{\mathsf{RR}A} - \xi^{I} e_{IA} + \widetilde{\xi}_{I} m_{A}^{I} \right)$$

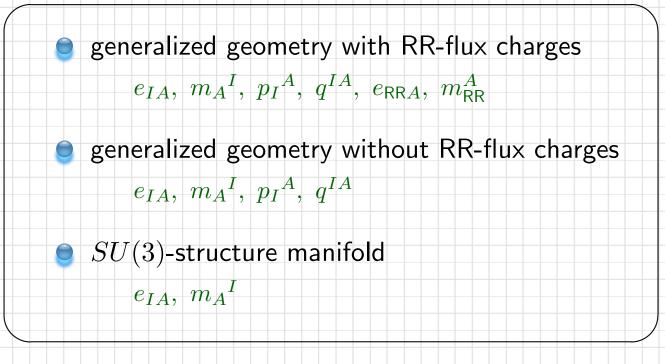
fluxes	charges	
NS three-form H	e_{I0}	$m_0{}^I$
torsion	e_{Ia}	$m_a{}^I$
nongeometric fluxes	p_I^A	q^{IA}
RR-fluxes	e_{RRA}	$m^A_{\sf RR}$

backgrounds	flux charges	
Calabi-Yau		
Calabi-Yau with H	e_{I0} m_0^I	
SU(3) geometry	e_{IA} $m_A{}^I$	
SU(3) imes SU(3) geometry	e_{IA} $m_A{}^I$ $p_I{}^A$	q^{IA}

Note: SU(3) generalized geometry without RR-fluxes ~ SU(3)-structure manifold

We are ready to search SUSY vacua in 4D $\mathcal{N} = 1$ supergravity.

Consider three typical situations given by



Notice: 4D physics given by Calabi-Yau three-fold with RR-fluxes is forbidden.

RR-fluxes induce the non-zero NS-fluxes as well as torsion classes in SUSY solutions.

D. Lüst, D. Tsimpis hep-th/0412250

Search of SUSY vacua: flux vacua attractors

$$V = e^{K} \left(K^{\mathcal{M}\overline{\mathcal{N}}} D_{\mathcal{M}} \mathcal{W} \overline{D_{\mathcal{N}}} \mathcal{W} - 3|\mathcal{W}|^{2} \right) + \frac{1}{2} |D^{a}|^{2}$$
$$\equiv V_{\mathcal{W}} + V_{D}$$

Search of vacua $\partial_{\mathcal{P}} V \big|_* = 0$

 $V_* > 0$: de Sitter space $V_* = 0$: Minkowski space $V_* < 0$: Anti-de Sitter space

$$0 = \partial_{\mathcal{P}} V_{\mathcal{W}} = e^{K} \left\{ K^{\mathcal{M}\overline{\mathcal{N}}} D_{\mathcal{P}} D_{\mathcal{M}} \mathcal{W} \overline{D_{\mathcal{N}}} \mathcal{W} + \partial_{\mathcal{P}} K^{\mathcal{M}\overline{\mathcal{N}}} D_{\mathcal{M}} \mathcal{W} \overline{D_{\mathcal{N}}} \mathcal{W} - 2 \overline{\mathcal{W}} D_{\mathcal{P}} \mathcal{W} \right\}$$
$$0 = \partial_{\mathcal{P}} V_{D} \quad \dashrightarrow \quad D^{a} = 0$$

Consider the SUSY condition $D_{\mathcal{P}}\mathcal{W} \equiv (\partial_{\mathcal{P}} + \partial_{\mathcal{P}}K)\mathcal{W} = 0$ in various cases.

- 1. Set a simple prepotential: $\mathcal{F} = D_{abc} \frac{X^a X^b X^c}{X^0}$
- 2. Consider the simplest model: single modulus t of Φ_+ (and U of Φ_-)

Derivatives of the Kähler potential are

$$\partial_t K = -\frac{3}{t - \overline{t}} \qquad \qquad \partial_U K = -\frac{2}{U - \overline{U}}$$

The superpotential is reduced to

$$\mathcal{W} = \mathcal{W}^{\mathsf{R}\mathsf{R}} + U \mathcal{W}^{\mathbb{Q}}$$
$$\mathcal{W}^{\mathsf{R}\mathsf{R}} = m_{\mathsf{R}\mathsf{R}}^{0} t^{3} - 3 m_{\mathsf{R}\mathsf{R}} t^{2} + e_{\mathsf{R}\mathsf{R}} t + e_{\mathsf{R}\mathsf{R}0}$$
$$\mathcal{W}^{\mathbb{Q}} = p_{0}^{0} t^{3} - 3 p_{0} t^{2} - e_{0} t - e_{00}$$

Consider the SUSY condition $D_{\mathcal{P}}\mathcal{W} \equiv (\partial_{\mathcal{P}} + \partial_{\mathcal{P}}K)\mathcal{W} = 0$:

$$D_t \mathcal{W} = 0 \dashrightarrow 0 = D_t \mathcal{W}^{\mathsf{R}\mathsf{R}} + U D_t \mathcal{W}^{\mathbb{Q}}$$
$$D_U \mathcal{W} = 0 \dashrightarrow 0 = \frac{\mathrm{i}}{\mathrm{Im}U} \Big(\mathcal{W}^{\mathsf{R}\mathsf{R}} + \mathrm{Re}U \mathcal{W}^{\mathbb{Q}} \Big)$$

Note: $ImU \neq 0$ to avoid curvature singularity

The discriminant of the superpotential $\mathcal{W}^{\mathsf{RR}}$ (and $\mathcal{W}^{\mathbb{Q}})$ governs the SUSY solutions.

$$\Delta(\mathcal{W}^{\mathsf{RR}}) = -27 \, (m_{\mathsf{RR}}^0 \, e_{\mathsf{RR0}})^2 - 54 \, m_{\mathsf{RR}}^0 \, e_{\mathsf{RR0}} \, m_{\mathsf{RR}} \, e_{\mathsf{RR}} + 9 \, (m_{\mathsf{RR}} \, e_{\mathsf{RR}})^2 + 108 \, (m_{\mathsf{RR}})^3 e_{\mathsf{RR0}} - 4 \, m_{\mathsf{RR}}^0 (e_{\mathsf{RR}})^3$$

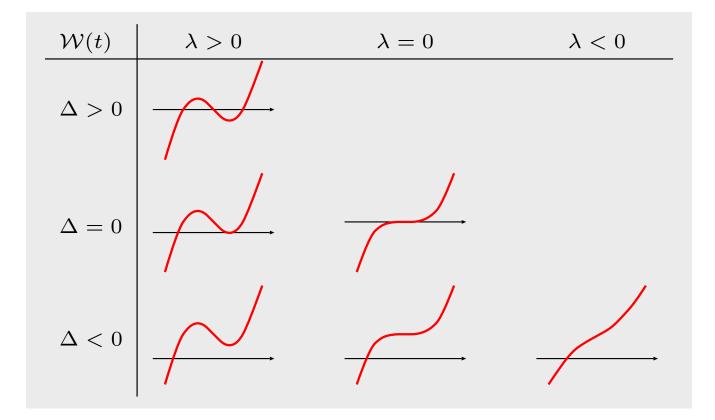
Discriminant of cubic equation

Consider a cubic function and its derivative:

$$\begin{cases} \mathcal{W}(t) = a t^3 + b t^2 + c t + d \\ \partial_t \mathcal{W}(t) = 3a t^2 + 2b t + c \end{cases}$$

Discriminants $\Delta(\mathcal{W})$ and $\Delta(\partial_t \mathcal{W})$ are

$$\Delta(\mathcal{W}) \equiv \Delta = -4b^3d + b^2c^2 - 4ac^3 + 18abcd - 27a^2d^2$$
$$\Delta(\partial_t \mathcal{W}) \equiv \lambda = 4(b^2 - 3ac)$$



 $\Delta^{\text{RR}} > 0$ case: always $\lambda^{\text{RR}} > 0$, and exists a zero point: $D_t \mathcal{W}^{\text{RR}} = 0$

$$D_{t}\mathcal{W}^{\mathsf{R}\mathsf{R}}|_{*} = 0$$

$$t_{*}^{\mathsf{R}\mathsf{R}} = \frac{6(3\,m_{\mathsf{R}\mathsf{R}}^{0}\,e_{\mathsf{R}\mathsf{R}0} + m_{\mathsf{R}\mathsf{R}}\,e_{\mathsf{R}\mathsf{R}})}{\lambda^{\mathsf{R}\mathsf{R}}} - 2\mathrm{i}\frac{\sqrt{3\,\Delta^{\mathsf{R}\mathsf{R}}}}{\lambda^{\mathsf{R}\mathsf{R}}}$$

$$\mathcal{W}_{*}^{\mathsf{R}\mathsf{R}} = -\frac{24\,\Delta^{\mathsf{R}\mathsf{R}}}{(\lambda^{\mathsf{R}\mathsf{R}})^{3}} \Big(36\,(m_{\mathsf{R}\mathsf{R}})^{3} + 36\,(m_{\mathsf{R}\mathsf{R}}^{0})^{2}e_{\mathsf{R}\mathsf{R}0} - 3\,m_{\mathsf{R}\mathsf{R}}\lambda^{\mathsf{R}\mathsf{R}} - 4\mathrm{i}\,m_{\mathsf{R}\mathsf{R}}^{0}\sqrt{3\,\Delta^{\mathsf{R}\mathsf{R}}}\Big)$$

 $\Delta^{\text{RR}} > 0$ case: always $\lambda^{\text{RR}} > 0$, and exists a zero point: $D_t \mathcal{W}^{\text{RR}} = 0$

$$\begin{split} D_t \mathcal{W}^{\mathsf{RR}}|_* &= 0 \\ t_*^{\mathsf{RR}} &= \frac{6 \left(3 \, m_{\mathsf{RR}}^0 \, e_{\mathsf{RR0}} + m_{\mathsf{RR}} \, e_{\mathsf{RR}}\right)}{\lambda^{\mathsf{RR}}} - 2\mathrm{i} \frac{\sqrt{3 \, \Delta^{\mathsf{RR}}}}{\lambda^{\mathsf{RR}}} \\ \mathcal{W}^{\mathsf{RR}}_* &= -\frac{24 \, \Delta^{\mathsf{RR}}}{(\lambda^{\mathsf{RR}})^3} \left(36 \, (m_{\mathsf{RR}})^3 + 36 \, (m_{\mathsf{RR}}^0)^2 e_{\mathsf{RR0}} - 3 \, m_{\mathsf{RR}} \lambda^{\mathsf{RR}} - 4\mathrm{i} \, m_{\mathsf{RR}}^0 \sqrt{3 \, \Delta^{\mathsf{RR}}}\right) \end{split}$$

 $\Delta^{RR} < 0$ case: only $\lambda^{RR} < 0$ is (physically) allowed, and exists a zero point: $W^{RR} = 0$

$$\begin{split} \mathcal{W}_{*}^{\mathsf{RR}} &= m_{\mathsf{RR}}^{0}(t_{*}-e)(t_{*}-\alpha)(t_{*}-\overline{\alpha}) = 0, \quad t_{*} = \alpha^{\mathsf{RR}} = \alpha_{1} + \mathrm{i}\,\alpha_{2} \\ \alpha_{1} &= \frac{\lambda^{\mathsf{RR}} + F^{2/3} + 12\,m_{\mathsf{RR}}\,F^{1/3}}{12\,m_{\mathsf{RR}}^{0}\,F^{1/3}} \\ (\alpha_{2})^{2} &= \frac{1}{m_{\mathsf{RR}}^{0}} \Big(e_{\mathsf{RR}} - 6\,m_{\mathsf{RR}}\,\alpha_{1} + 3\,m_{\mathsf{RR}}^{0}\,(\alpha_{1})^{2} \Big) \\ e &= -\frac{1}{m_{\mathsf{RR}}^{0}} \Big(-3\,m_{\mathsf{RR}} + 2\,m_{\mathsf{RR}}^{0}\,\alpha_{1} \Big) \\ F &= 108\,(m_{\mathsf{RR}}^{0})^{2}e_{\mathsf{RR0}} + 12\,m_{\mathsf{RR}}^{0}\sqrt{-3\Delta^{\mathsf{RR}}} + 108\,(m_{\mathsf{RR}})^{3} - 9\,m_{\mathsf{RR}}\,\lambda^{\mathsf{RR}} \\ D_{t}\mathcal{W}^{\mathsf{RR}}|_{*} &= 2\mathrm{i}\,m_{\mathsf{RR}}^{0}(e - \alpha^{\mathsf{RR}})\alpha_{2} \end{split}$$

... Analysis of $\mathcal{W}^{\mathbb{Q}}$ is also discussed.

Three types of solutions to satisfy $0 = D_t \mathcal{W}^{\mathsf{RR}} + U D_t \mathcal{W}^{\mathbb{Q}}$ and $0 = \mathcal{W}^{\mathsf{RR}} + \operatorname{Re}U \mathcal{W}^{\mathbb{Q}}$:

SUSY AdS vacuum: attractor point

$$\Delta^{\mathsf{RR}} > 0, \quad \Delta^{\mathbb{Q}} > 0; \quad D_t \mathcal{W}^{\mathsf{RR}}|_* = 0 = D_t \mathcal{W}^{\mathbb{Q}}|_*$$
$$t_*^{\mathsf{RR}} = t_*^{\mathbb{Q}}, \quad \operatorname{Re} U_* = -\frac{\mathcal{W}_*^{\mathsf{RR}}}{\mathcal{W}_*^{\mathbb{Q}}}$$
$$V_* = -3 \operatorname{e}^K |\mathcal{W}_*|^2 = -\frac{4}{[\operatorname{Re}(\mathcal{CG}_0)]^2} \sqrt{\frac{\Delta^{\mathbb{Q}}}{3}}$$

SUSY Minkowski vacuum: attractor point

$$\Delta^{\mathsf{RR}} < 0, \qquad \Delta^{\mathbb{Q}} < 0; \qquad \mathcal{W}_*^{\mathsf{RR}} = 0 = \mathcal{W}_*^{\mathbb{Q}}$$
$$\alpha^{\mathsf{RR}} = \alpha^{\mathbb{Q}}, \qquad U_* = -\frac{D_t \mathcal{W}^{\mathsf{RR}}|_*}{D_t \mathcal{W}^{\mathbb{Q}}|_*} \neq 0$$
$$V_* = 0$$

SUSY AdS vacua, but moduli t and U are not fixed: non attractor point

$$U = -\frac{D_t \mathcal{W}^{\mathsf{RR}}(t)}{D_t \mathcal{W}^{\mathbb{Q}}(t)}, \qquad \operatorname{Re} U = -\frac{\mathcal{W}^{\mathsf{RR}}(t)}{\mathcal{W}^{\mathbb{Q}}(t)}$$

- 1. Set $e_{RRA} = 0 = m_{RR}^A$
- 2. Set a simple prepotential: $\mathcal{F} = D_{abc} \frac{X^a X^b X^c}{X^0}$
- **3**. Consider the simplest model: single modulus t of Φ_+ (and U of Φ_-)

The SUSY conditions on $\mathcal{W} = U \mathcal{W}^{\mathbb{Q}}$ are

 $D_t \mathcal{W} = 0 \quad \dashrightarrow \quad 0 = D_t \mathcal{W}^{\mathbb{Q}}$ $D_U \mathcal{W} = 0 \quad \dashrightarrow \quad 0 = \operatorname{Re} U \mathcal{W}^{\mathbb{Q}}$

- 1. Set $e_{RRA} = 0 = m_{RR}^A$
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$$\mathcal{W} = U \mathcal{W}^{\mathbb{Q}}$$
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 $D_t \mathcal{W} = 0 \quad \dashrightarrow \quad 0 = D_t \mathcal{W}^{\mathbb{Q}}$
 $D_U \mathcal{W} = 0 \quad \dashrightarrow \quad 0 = \operatorname{Re} U \mathcal{W}^{\mathbb{Q}}$

The solution is given only when $\Delta^{\mathbb{Q}} > 0$, and the AdS vacuum emerges:

$$t_{*}^{\mathbb{Q}} = -\frac{6 (3 p_{0}^{0} e_{00} + p_{0} e_{0})}{\lambda^{\mathbb{Q}}} - 2i \frac{\sqrt{3 \Delta^{\mathbb{Q}}}}{\lambda^{\mathbb{Q}}}, \quad \text{Re} U_{*} = 0$$
$$V_{*} = -3 e^{K} |\mathcal{W}_{*}|^{2} = -\frac{4}{[\text{Re}(\mathcal{CG}_{0})]^{2}} \sqrt{\frac{\Delta^{\mathbb{Q}}}{3}}$$

Example 3: SU(3)-structure manifold

- 1. Set $e_{RRA} = 0 = m_{RR}^A$ and $p_I^A = 0 = q^{IA}$
- 2. Set a simple prepotential: $\mathcal{F} = D_{abc} \frac{X^a X^b X^c}{X^0}$
- **3.** Consider the simplest model: single modulus t of Φ_+ (and U of Φ_-)

Functions are reduced to

$$D_t \mathcal{W} = \frac{U}{t - \overline{t}} \left(e_0 (2t + \overline{t}) + 3 e_{00} \right), \qquad D_U \mathcal{W} = \mathrm{i} \frac{\mathrm{Re}U}{\mathrm{Im}U} \mathcal{W}^{\mathbb{Q}}$$

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There are neither SUSY solutions under the conditions $D_t \mathcal{W} = 0 = D_U \mathcal{W}$ nor non-SUSY solutions satisfying $\partial_{\mathcal{P}} V = 0$!

Ansatz 2. "Neglecting all α' corrections on the compactified space" is too strong!

2'. Set a deformed prepotential: $\mathcal{F} = \frac{(X^t)^3}{X^0} + \sum_n N_n \frac{(X^t)^{n+3}}{(X^0)^{n+1}}$

Consider a simple case as $N_1 \neq 0$, otherwise $N_n = 0$:

$$D_{t} \mathcal{W}^{\mathbb{Q}} = -e_{00} + \frac{3(t-\bar{t})^{2} - \partial_{t} P}{(t-\bar{t})^{3} - P} \left(e_{00} + e_{0} t\right)$$
$$P \equiv -2 \left(N_{1} t^{4} - \overline{N_{1}} \bar{t}^{4} - 2N_{1} t^{3} \bar{t} + 2\overline{N_{1}} t \bar{t}^{3}\right)$$

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SUSY AdS solution appears under the conditions $D_t \mathcal{W} = 0$ and $D_U \mathcal{W} = 0$:

$$t^{\mathbb{Q}}_{*} = -\frac{2 e_{00}}{e_{0}}, \quad \operatorname{Re} U_{*} = 0$$
$$\mathcal{W}^{\mathbb{Q}}_{*} = e_{00}, \quad \operatorname{Im} N_{1} < 0$$
$$V_{*} = -3 e^{K} |\mathcal{W}_{*}|^{2} = \frac{1}{[\operatorname{Re}(\mathcal{CG}_{0})]^{2}} \frac{3 (e_{0})^{4}}{16 (e_{00})^{2} \operatorname{Im} N_{1}}$$

This is also given by the heterotic string compactifications on SU(3)-structure manifolds with torsion, which carries α' corrections.

Summary and Discussions

Summary

- We obtained SUSY AdS or Minkowski vacuum on an attractor point
- We found a powerful rule to evaluate the attractor points: Discriminants
- \blacksquare We confirmed that α' corrections are included in reduced configurations

Discussions

- Complete stabilization via nonperturbative corrections
- Duality transformations
- Understanding the physical interpretation of nongeometric fluxes
- Connection to doubled formalism

de Sitter vacua?

In order to build (stable) de Sitter vacua perturbatively in type IIA,

in addition to the usual RR and NSNS fluxes and O6/D6 sources,

one must minimally have geometric fluxes and non-zero Romans' mass parameter.

S.S. Haque, G. Shiu, B. Underwood, T. Van Riet arXiv:0810.5328

Romans' mass parameter $\sim G_0$

Search a (meta)stable de Sitter vacuum in this formulation

Appendix: compactifications in type II strings

Moduli spaces in $\mathcal{N}=2$ supergravity are

vector multiplets: Hodge-Kähler geometry hypermultiplets: quaternionic geometry

We look for the origin of the moduli spaces in 10D string theories

Decomposition of vector bundle on 10D spacetime:

$$T\mathcal{M}_{1,9} = T_{1,3} \oplus F$$

- $\begin{cases} T_{1,3}: \text{ a real } SO(1,3) \text{ vector bundle} \\ F: \text{ an } SO(6) \text{ vector bundle which admits a pair of } SU(3) \text{ structures} \end{cases}$

10D spacetime itself is not decomposed yet, i.e., do not yet consider truncation of modes.

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Decomposition of Lorentz symmetry:

$$Spin(1,9) \rightarrow Spin(1,3) \times Spin(6) = SL(2,\mathbb{C}) \times SU(4)$$
$$\mathbf{16} = (\mathbf{2},\mathbf{4}) \oplus (\overline{\mathbf{2}},\overline{\mathbf{4}}) \qquad \mathbf{16} = (\mathbf{2},\overline{\mathbf{4}}) \oplus (\overline{\mathbf{2}},\mathbf{4})$$

Decomposition of supersymmetry parameters (with $a, b \in \mathbb{C}$):

 $\begin{cases} \epsilon_{\mathrm{IIA}}^{1} = \varepsilon_{1} \otimes (a\eta_{+}^{1}) + \varepsilon_{1}^{c} \otimes (\overline{a}\eta_{-}^{1}) \\ \epsilon_{\mathrm{IIA}}^{2} = \varepsilon_{2} \otimes (\overline{b}\eta_{-}^{2}) + \varepsilon_{2}^{c} \otimes (b\eta_{+}^{2}) \end{cases} \begin{cases} \epsilon_{\mathrm{IIB}}^{1} = \varepsilon_{1} \otimes (a\eta_{+}^{1}) + \varepsilon_{1}^{c} \otimes (\overline{a}\eta_{-}^{1}) \\ \epsilon_{\mathrm{IIB}}^{2} = \varepsilon_{2} \otimes (b\eta_{+}^{2}) + \varepsilon_{2}^{c} \otimes (\overline{b}\eta_{-}^{2}) \end{cases}$

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Set SU(3) invariant spinor $\eta_+^{\mathcal{A}}$ s.t. $\nabla^{(T)}\eta_+^{\mathcal{A}} = 0$ $(\mathcal{A} = 1, 2)$

a pair of SU(3) on $F(\eta^1_+, \eta^2_+) \quad \longleftrightarrow$ a single SU(3) on $F(\eta^1_+ = \eta^2_+ = \eta_+)$

Requirement that we have a pair of SU(3) structures means there is a sub-supermanifold

$$\mathcal{N}^{1,9|4+4} \subset \mathcal{M}^{1,9|16+16}$$

 $\begin{pmatrix} (1,9): & \text{bosonic degrees} \\ 4+4: & \text{eight Grassmann variables as spinors of } Spin(1,3) & \text{and singlet of } SU(3) \\ \end{pmatrix}$

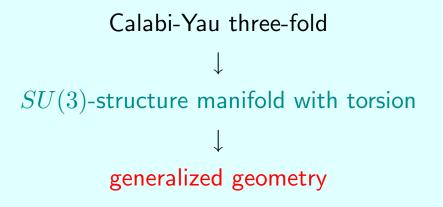
Equivalence such as

eight SUSY theory reformulation of type II supergravity a pair of SU(3) structures on vector bundle F an SU(3) imes SU(3) structure on extended $F \oplus F^*$

10D spinors in type IIA in Einstein frame

$$\begin{split} \delta\Psi_{M}^{\mathcal{A}} &= \nabla_{M}\epsilon^{\mathcal{A}} - \frac{1}{96}\mathrm{e}^{-\phi} \Big(\Gamma_{M}{}^{PQR}H_{PQR} - 9\Gamma^{PQ}H_{MPQ}\Big)\Gamma_{(11)}\epsilon^{\mathcal{A}} \\ &- \sum_{n=0,2,\dots,8} \frac{1}{64n!}\mathrm{e}^{\frac{5-n}{4}\phi} \Big[(n-1)\Gamma_{M}{}^{N_{1}\cdots N_{n}} - n(9-n)\delta_{M}{}^{N_{1}}\Gamma^{N_{2}\cdots N_{n}}\Big]F_{N_{1}\cdots N_{n}}(\Gamma_{(11)})^{n/2}(\sigma^{1}\epsilon)^{\mathcal{A}} \\ &\epsilon^{1} &= \varepsilon_{1}\otimes(a\eta_{+}^{1}) + \varepsilon_{1}^{c}\otimes(\overline{a}\eta_{-}^{1}) \qquad \epsilon^{2} &= \varepsilon_{2}\otimes(\overline{b}\eta_{-}^{2}) + \varepsilon_{2}^{c}\otimes(b\eta_{+}^{2}) \\ &0 &\equiv \delta\psi_{m}^{\mathcal{A}} &= \nabla_{m}\eta_{+}^{\mathcal{A}} + (\mathsf{NS-fluxes}\cdot\eta)^{\mathcal{A}} + (\mathsf{RR-fluxes}\cdot\eta)^{\mathcal{A}} \end{split}$$

Information of 6D SU(3) Killing spinors $\eta_{+}^{\mathcal{A}}$



• on a single
$$SU(3)$$
:
• a real two-form
• a complex three-form
• a complex three-form
• on a pair of $SU(3)$:
• (J^A, Ω^A)
• $(J^A,$

If $\eta_+^1 \neq \eta_+^2$ globally, a pair of SU(3) is reduced to global single SU(2) w/ (j, w, v, v')If $\eta_+^1 = \eta_+^2$ globally, a pair of SU(3) is reduced to a single global SU(3) w/ (J, Ω)

$$\eta_{+}^{2} = c_{\parallel} \eta_{+}^{1} + c_{\perp} (v + iv')^{m} \gamma_{m} \eta_{-}^{1}, \qquad |c_{\parallel}|^{2} + |c_{\perp}|^{2} = 1$$

a pair of SU(3) on $T\mathcal{M} \sim \text{an } SU(3) \times SU(3)$ on $T\mathcal{M} \oplus T^*\mathcal{M}$

Appendix: Calabi-Yau compactifications

One can embed 4D $\mathcal{N}=2$ theory into 10D type II theory compactified on Calabi-Yau three-fold

Т

	vector multiplets	hypermultiplets
generic	coord. of Hodge-Kähler	coord. of quaternionic
IIA on Calabi-Yau	Kähler moduli	complex moduli + RR
IIB on Calabi-Yau	complex moduli	Kähler moduli + RR

NS-NS fields in ten-dimensional spacetime are expanded as

$$\phi(x,y) = \varphi(x)$$

$$G_{m\overline{n}}(x,y) = i v^{a}(x)(\omega_{a})_{m\overline{n}}(y), \quad G_{mn}(x,y) = i \overline{z}^{k}(x) \left(\frac{(\overline{\chi}_{k})_{m\overline{pq}}\Omega^{\overline{pq}}}{||\Omega||^{2}}\right)(y)$$

$$B_{2}(x,y) = B_{2}(x) + b^{a}(x)\omega_{a}(y)$$

RR fields in type IIA are

$$C_1(x,y) = C_1^0(x)$$

$$C_3(x,y) = C_1^a(x)\omega_a(y) + \xi^K(x)\alpha_K(y) - \widetilde{\xi}_K(x)\beta^K(y)$$

RR fields in type IIB are

$$C_0(x,y) = C_0(x)$$

$$C_2(x,y) = C_2(x) + c^a(x)\omega_a(y)$$

$$C_4(x,y) = V_1^K(x)\alpha_K(y) + \rho_a(x)\widetilde{\omega}^a(y)$$

cohomology class	basis	
$H^{(1,1)}$	ω_a	$a = 1, \dots, h^{(1,1)}$
$H^{(0)}\oplus H^{(1,1)}$	$\omega_A = (1, \omega_a)$	$A = 0, 1, \dots, h^{(1,1)}$
$H^{(2,2)}$	$\widetilde{\omega}^a$	$a = 1, \dots, h^{(1,1)}$
$H^{(2,1)}$	χ_k	$k = 1, \dots, h^{(2,1)}$
$H^{(3)}$	(α_K, β^K)	$K = 0, 1, \dots, h^{(2,1)}$

4D type IIA $\mathcal{N} = 2$ ungauged supergravity action of bosonic fields is

$$\begin{split} S_{\text{IIA}}^{(4)} &= \int_{\mathcal{M}_{1,3}} \left(-\frac{1}{2}R * \mathbf{1} + \frac{1}{2} \operatorname{Re}\mathcal{N}_{AB}F^A \wedge F^B + \frac{1}{2} \operatorname{Im}\mathcal{N}_{AB}F^A \wedge *F^B \\ &- G_{a\overline{b}} \, \mathrm{d}t^a \wedge * \mathrm{d}\overline{t}^{\overline{b}} - h_{uv} \, \mathrm{d}q^u \wedge * \mathrm{d}q^v \right) \\ \hline \\ & \\ \hline \\ & \\ \text{gravity multiplet} \qquad g_{\mu\nu}, \ C_1^0 & 1 \\ & \\ & \\ \text{vector multiplet} \qquad C_1^a, \ v^a, \ b^a & a = 1, \dots, h^{(1,1)} \\ & \\ & \\ & \\ \end{matrix}$$

hypermultiplet	$z^k,\;\xi^k,\;\widetilde{\xi}_k$	$k=1,\ldots,h^{(2,1)}$
tensor multiplet	$B_2,\;arphi,\;\xi^0,\;\widetilde{\xi_0}$	1

Various functions in the actions:

$$\begin{split} B + \mathrm{i}J &= (b^a + \mathrm{i}v^a)\,\omega_a \ = \ t^a\omega_a & K^{\mathsf{KS}} = -\log\left(\frac{4}{3}\int_{\mathcal{M}_6}J\wedge J\wedge J\right) \\ \mathcal{K}_{abc} &= \int_{\mathcal{M}_6}\omega_a\wedge\omega_b\wedge\omega_c & \mathcal{K}_{ab} = \int_{\mathcal{M}_6}\omega_a\wedge\omega_b\wedge J \ = \ \mathcal{K}_{abc}v^c \\ \mathcal{K}_a &= \int_{\mathcal{M}_6}\omega_a\wedge J\wedge J \ = \ \mathcal{K}_{abc}v^bv^c & \mathcal{K} \ = \int_{\mathcal{M}_6}J\wedge J\wedge J \ = \ \mathcal{K}_{abc}v^av^bv^c \\ \mathrm{Re}\mathcal{N}_{AB} &= \left(\begin{array}{c} -\frac{1}{3}\mathcal{K}_{abc}b^ab^bc \ \frac{1}{2}\mathcal{K}_{abc}b^bb^c \\ \frac{1}{2}\mathcal{K}_{abc}b^bb^c \ -\mathcal{K}_{abc}b^c \end{array}\right) & \mathrm{Im}\mathcal{N}_{AB} = -\frac{\mathcal{K}}{6}\left(\begin{array}{c} 1+4G_{ab}b^ab^b \ -4G_{ab}b^b \\ -4G_{ab}b^b \ 4G_{ab} \end{array}\right) \\ G_{a\overline{b}} \ = \ \frac{3}{2}\frac{\int\omega_a\wedge\ast\omega_b}{\int J\wedge J\wedge J} \ = \ \partial_{t^a}\overline{\partial}_{\overline{t}\overline{b}}K^{\mathsf{KS}} \end{split}$$

4D type IIB $\mathcal{N} = 2$ ungauged supergravity action of bosonic fields is

$$S_{\text{IIB}}^{(4)} = \int_{\mathcal{M}_{1,3}} \left(-\frac{1}{2}R * \mathbf{1} + \frac{1}{2} \operatorname{Re}\mathcal{M}_{KL}F^{K} \wedge F^{L} + \frac{1}{2} \operatorname{Im}\mathcal{M}_{KL}F^{K} \wedge *F^{L} - G_{k\bar{l}} \, \mathrm{d}z^{k} \wedge *\mathrm{d}\overline{z}^{\bar{l}} - h_{pq} \, \mathrm{d}\widehat{q}^{p} \wedge *\mathrm{d}\widehat{q}^{q} \right)$$

gravity multiplet	$g_{\mu u},V^0_1$	1
vector multiplet	V_1^k,z^k	$k=1,\ldots,h^{(2,1)}$
hypermultiplet	$v^a,\ b^a,\ c^a,\ ho_a$	$a = 1, \dots, h^{(1,1)}$
tensor multiplet	$B_2,\ C_2,\ arphi,\ C_0$	1

Various functions in the actions:

$$\Omega = Z^{K} \alpha_{K} - \mathcal{G}_{K} \beta^{K} \qquad z^{k} = Z^{K} / Z^{0} \qquad \mathcal{G}_{KL} = \partial_{L} \mathcal{G}_{K}$$
$$K^{\mathsf{CS}} = -\log\left(\mathrm{i} \int_{\mathcal{M}_{6}} \Omega \wedge \overline{\Omega}\right) \qquad G_{k\overline{l}} = -\frac{\int \chi_{k} \wedge \overline{\chi}_{\overline{l}}}{\int \Omega \wedge \overline{\Omega}} = \partial_{z^{k}} \overline{\partial}_{\overline{z}^{\overline{l}}} K^{\mathsf{CS}}$$

$$\mathcal{M}_{KL} = \overline{\mathcal{G}}_{KL} + 2i \frac{(\mathrm{Im}\mathcal{G})_{KM} Z^M (\mathrm{Im}\mathcal{G})_{LN} Z^N}{Z^N (\mathrm{Im}\mathcal{G})_{NM} Z^M}$$

Appendix: SU(3)-structure manifold with torsion

1 Information from Killing spinor eqs. with torsion $D^{(T)}\eta_{\pm} = 0$ (³complex Weyl η_{\pm})

• Invariant *p*-forms on SU(3)-structure manifold:

a real two-form
$$J_{mn} = \mp 2i \eta_{\pm}^{\dagger} \gamma_{mn} \eta_{\pm}$$

a holomorphic three-form $\Omega_{mnp} = -2i \eta_{-}^{\dagger} \gamma_{mnp} \eta_{+}$
 $dJ = \frac{3}{2} \operatorname{Im}(\overline{W}_{1}\Omega) + W_{4} \wedge J + W_{3} \quad d\Omega = W_{1}J \wedge J + W_{2} \wedge J + \overline{W}_{5} \wedge \Omega$

	C	/• . • • \			
Five classes	ot	(intrinsic)	torsion	are given	as
		()		0	

components	interpretations	SU(3)-representations
\mathcal{W}_1	$J\wedge \mathrm{d}\Omega$ or $\Omega\wedge \mathrm{d}J$	${f 1}\oplus{f 1}$
\mathcal{W}_2	$(\mathrm{d}\Omega)^{2,2}_0$	${\bf 8}\oplus {\bf 8}$
\mathcal{W}_3	$(\mathrm{d}J)_0^{2,1} + (\mathrm{d}J)_0^{1,2}$	${\bf 6}\oplus\overline{\bf 6}$
${\mathfrak W}_4$	$J\wedge \mathrm{d}J$	${\bf 3}\oplus \overline{\bf 3}$
${\mathcal W}_5$	$(\mathrm{d}\Omega)^{3,1}$	${\bf 3}\oplus \overline{\bf 3}$

> Vanishing torsion classes in SU(3)-structure manifolds:

complex	hermitian	$\mathcal{W}_1 = \mathcal{W}_2 = 0$
	balanced	$\mathcal{W}_1 = \mathcal{W}_2 = \mathcal{W}_4 = 0$
	special hermitian	$\mathcal{W}_1 = \mathcal{W}_2 = \mathcal{W}_4 = \mathcal{W}_5 = 0$
	Kähler	$\mathcal{W}_1 = \mathcal{W}_2 = \mathcal{W}_3 = \mathcal{W}_4 = 0$
	Calabi-Yau	$\mathcal{W}_1 = \mathcal{W}_2 = \mathcal{W}_3 = \mathcal{W}_4 = \mathcal{W}_5 = 0$
	conformally Calabi-Yau	$\mathcal{W}_1 = \mathcal{W}_2 = \mathcal{W}_3 = 3\mathcal{W}_4 + 2\mathcal{W}_5 = 0$
almost complex	symplectic	$\mathcal{W}_1=\mathcal{W}_3=\mathcal{W}_4=0$
	nearly Kähler	$\mathcal{W}_2 = \mathcal{W}_3 = \mathcal{W}_4 = \mathcal{W}_5 = 0$
	almost Kähler	$\mathcal{W}_1 = \mathcal{W}_3 = \mathcal{W}_4 = \mathcal{W}_5 = 0$
	quasi Kähler	$\mathcal{W}_3 = \mathcal{W}_4 = \mathcal{W}_5 = 0$
	semi Kähler	$\mathcal{W}_4 = \mathcal{W}_5 = 0$
	half-flat	$\mathrm{Im}\mathcal{W}_1 = \mathrm{Im}\mathcal{W}_2 = \mathcal{W}_4 = \mathcal{W}_5 = 0$

Appendix: generalized geometry

Introduce a generalized almost complex structure \mathcal{J} on $T\mathcal{M}_d \oplus T^*\mathcal{M}_d$ s.t.

$$\mathcal{J}: T\mathcal{M}_d \oplus T^*\mathcal{M}_d \longrightarrow T\mathcal{M}_d \oplus T^*\mathcal{M}_d$$
$$\mathcal{J}^2 = -\mathbb{1}_{2d}$$

$$\exists O(d,d)$$
 invariant metric L , s.t. $\mathcal{J}^T L \mathcal{J} = L$

Structure group on $T\mathcal{M}_d \oplus T^*\mathcal{M}_d$:

$\exists L$	GL(2d)	>	O(d,d)
$\mathcal{J}^2 = -\mathbb{1}_{2d}$	O(d,d)	>	U(d/2,d/2)
$\mathcal{J}_1,\mathcal{J}_2$	$U_1(d/2, d/2) \cap U_2(d/2, d/2)$	>	$U(d/2) \times U(d/2)$
integrable $\mathcal{J}_{1,2}$	U(d/2) imes U(d/2)	>	SU(d/2) imes SU(d/2)

▶ Integrability is discussed by "(0,1)" part of the complexified $T\mathcal{M}_d \oplus T^*\mathcal{M}_d$:

$$\Pi \equiv \frac{1}{2}(\mathbb{1}_{2d} - \mathrm{i}\mathcal{J})$$

 $\Pi A = A$ where $A = v + \zeta$ is a section of $T\mathcal{M}_d \oplus T^*\mathcal{M}_d$

We call this A i-eigenbundle $L_{\mathcal{J}}$, whose dimension is $\dim L_{\mathcal{J}} = d$. Integrability condition of \mathcal{J} is

$$\overline{\Pi} \big[\Pi(v+\zeta), \Pi(w+\eta) \big]_{\mathcal{C}} = 0 \qquad v, w \in T \mathcal{M}_d \qquad \zeta, \eta \in T^* \mathcal{M}_d$$
$$[v+\zeta, w+\eta]_{\mathcal{C}} = [v,w] + \mathcal{L}_v \eta - \mathcal{L}_w \zeta - \frac{1}{2} \mathrm{d}(\iota_v \eta - \iota_w \zeta): \text{ Courant bracket}$$

► Two typical examples of generalized almost complex structures:

$$\begin{aligned} \mathcal{J}_{-} &= \left(\begin{array}{cc} I & \mathbf{0} \\ \mathbf{0} & -I^{T} \end{array} \right) & \qquad \text{w} / \ I^{2} = -\mathbbm{1}_{d} \text{: almost complex structure} \\ \mathcal{J}_{+} &= \left(\begin{array}{cc} \mathbf{0} & -J^{-1} \\ J & \mathbf{0} \end{array} \right) & \qquad \text{w} / \ J \text{: almost symplectic form} \end{aligned}$$

integrable $\mathcal{J}_{-} \leftrightarrow$ integrable Iintegrable $\mathcal{J}_{+} \leftrightarrow$ integrable J

On a usual geometry, $J_{mn} = I_m{}^p g_{pn}$ is given by an SU(3) invariant (pure) spinor η_+ as $J_{mn} = -2i \eta^{\dagger}_+ \gamma_{mn} \eta_+ \qquad \gamma^i \eta_+ = 0 \qquad \gamma^{\overline{\iota}} \eta_+ \neq 0$

In a similar analogy, we want to find Cliff(6, 6) pure spinor(s) Φ .

::) Compared to almost complex structures, (pure) spinors can be easily utilized in supergravity framework.

On $T\mathcal{M}_6 \oplus T^*\mathcal{M}_6$, we can define $\mathsf{Cliff}(6,6)$ algebra and Spin(6,6) spinor Φ :

$$\{\Gamma^m, \Gamma^n\} = 0 \qquad \{\Gamma^m, \widetilde{\Gamma}_n\} = \delta_n^m \qquad \{\widetilde{\Gamma}_m, \widetilde{\Gamma}_n\} = 0$$

Irreducible repr. of Spin(6, 6) spinor is a Majorana-Weyl

 \rightarrow a generic Spin(6,6) spinor bundle S splits to S^{\pm} (Weyl)

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Weyl spinor bundles S^{\pm} are isomorphic to bundles of forms on $T^*\mathcal{M}_6$:

$$S^+ \text{ on } T\mathfrak{M}_6 \oplus T^*\mathfrak{M}_6 \quad \sim \quad \wedge^{\mathsf{even}} T^*\mathfrak{M}_6$$

 $S^- \text{ on } T\mathfrak{M}_6 \oplus T^*\mathfrak{M}_6 \quad \sim \quad \wedge^{\mathsf{odd}} T^*\mathfrak{M}_6$

Thus we often regard a $\mathsf{Cliff}(6,6)$ spinor as a form on $\wedge^{\mathsf{even/odd}} T^*\mathfrak{M}_6$

A form-valued representation of the algebra

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A form-valued representation of the algebra

$$\Gamma^m = \mathrm{d} x^m \wedge , \qquad \qquad \widetilde{\Gamma}_n = \iota_n$$

IF Φ is annihilated by half numbers of the Cliff(6, 6) generators:

 $\rightarrow \Phi$ is called a pure spinor

cf.) SU(3) invariant spinor η_+ is a Cliff(6) pure spinor: $\gamma^i \eta_+ = 0$

An equivalent definition of a Cliff(6, 6) pure spinor is given by "Clifford action":

$$(v+\zeta) \cdot \Phi = v^m \iota_{\partial_m} \Phi + \zeta_n \, \mathrm{d} x^n \wedge \Phi \quad \text{w/} v: \text{vector} \quad \zeta: \text{ one-form}$$

Define the annihilator of a spinor as

$$L_{\Phi} \equiv \left\{ v + \zeta \in T\mathcal{M}_{6} \oplus T^{*}\mathcal{M}_{6} \, \middle| \, (v + \zeta) \cdot \Phi = 0 \right\}$$
$$\dim L_{\Phi} \leq d$$

If dim $L_{\Phi} = 6$ (maximally isotropic) $\rightarrow \Phi$ is a pure spinor

Correspondence between pure spinors and generalized almost complex structures:

$$\mathcal{J} \leftrightarrow \Phi$$
 if $L_{\mathcal{J}} = L_{\Phi}$ with $\dim L_{\Phi} = 6$

More precisely: $\mathcal{J} \leftrightarrow$ a line bundle of pure spinor Φ

 \therefore) rescaling Φ does not change its annihilator L_{Φ}

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Then, we can rewrite the generalized almost complex structure as

$$\mathcal{J}_{\pm\Pi\Sigma} = \left\langle \mathrm{Re}\Phi_{\pm}, \Gamma_{\Pi\Sigma}\,\mathrm{Re}\Phi_{\pm} \right\rangle$$

w/ Mukai pairing:

even forms: $\langle \Psi_+, \Phi_+ \rangle = \Psi_6 \wedge \Phi_0 - \Psi_4 \wedge \Phi_2 + \Psi_2 \wedge \Phi_4 - \Psi_0 \wedge \Phi_6$ odd forms: $\langle \Psi_-, \Phi_- \rangle = \Psi_5 \wedge \Phi_1 - \Psi_3 \wedge \Phi_3 + \Psi_1 \wedge \Phi_5$ Correspondence between pure spinors and generalized almost complex structures:

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 \mathcal{J} is integrable \longleftrightarrow \exists vector v and one-form ζ s.t. $d\Phi = (v \downarrow + \zeta \land) \Phi$ generalized CY \longleftrightarrow $\exists \Phi$ is pure s.t. $d\Phi = 0$ "twisted" GCY \longleftrightarrow $\exists \Phi$ is pure, and H is closed s.t. $(d - H \land) \Phi = 0$ A Cliff(6, 6) spinor can also be mapped to a bispinor:

$$C \equiv \sum_{k} \frac{1}{k!} C_{m_1 \cdots m_k}^{(k)} \, \mathrm{d}x^{m_1} \wedge \cdots \wedge \mathrm{d}x^{m_k} \quad \longleftrightarrow \quad \mathcal{Q} \equiv \sum_{k} \frac{1}{k!} C_{m_1 \cdots m_k}^{(k)} \, \gamma_{\alpha\beta}^{m_1 \cdots m_k}$$

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On a geometry of a single SU(3)-structure, the following two SU(3,3) spinors:

$$\Phi_{0+} = \eta_{+} \otimes \eta_{+}^{\dagger} = \frac{1}{4} \sum_{k=0}^{6} \frac{1}{k!} \eta_{+}^{\dagger} \gamma_{m_{k} \cdots m_{1}} \eta_{+} \gamma^{m_{1} \cdots m_{k}} = \frac{1}{8} e^{-iJ}$$

$$\Phi_{0-} = \eta_{+} \otimes \eta_{-}^{\dagger} = \frac{1}{4} \sum_{k=0}^{6} \frac{1}{k!} \eta_{-}^{\dagger} \gamma_{m_{k} \cdots m_{1}} \eta_{+} \gamma^{m_{1} \cdots m_{k}} = -\frac{i}{8} \Omega$$

Check purity: $(\delta + iJ)_m{}^n \gamma_n \eta_+ \otimes \eta_{\pm}^{\dagger} = 0 = \eta_+ \otimes \eta_{\pm}^{\dagger} \gamma_n (\delta \mp iJ)^n{}_m$

One-to-one correspondence: $\Phi_{0-} \leftrightarrow \mathcal{J}_1, \quad \Phi_{0+} \leftrightarrow \mathcal{J}_2$

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On a generic geometry of a pair of SU(3)-structure defined by (η^1_+, η^2_+)

$$\Phi_{0+} = \eta_{+}^{1} \otimes \eta_{+}^{2\dagger} = \frac{1}{8} (\bar{c}_{\parallel} e^{-ij} - i\bar{c}_{\perp} w) \wedge e^{-iv \wedge v'} |c_{\parallel}|^{2} + |c_{\perp}|^{2} = 1$$

$$\Phi_{0-} = \eta_{+}^{1} \otimes \eta_{-}^{2\dagger} = -\frac{1}{8} (c_{\perp} e^{-ij} + ic_{\parallel} w) \wedge (v + iv')$$

$$\Phi_{\pm} = e^{-B} \Phi_{0\pm}$$

Each Φ_{\pm} defines an SU(3,3) structure on E. Common structure is $SU(3) \times SU(3)$. (F is extended to E by including e^{-B})

Compatibility requires

$$\begin{split} \left\langle \Phi_{+}, V \cdot \Phi_{-} \right\rangle \; = \; \left\langle \overline{\Phi}_{+}, V \cdot \Phi_{-} \right\rangle \; = \; 0 \quad \text{ for } \forall V = x + \xi \\ \left\langle \Phi_{+}, \overline{\Phi}_{+} \right\rangle \; = \; \left\langle \Phi_{-}, \overline{\Phi}_{-} \right\rangle \end{split}$$

Start with a real form $\chi_f \in \wedge^{\text{even/odd}} F^*$ (associated with a real Spin(6, 6) spinor χ_s) Regard χ_f as a stable form satisfying

$$q(\chi_f) = -\frac{1}{4} \langle \chi_f, \Gamma_{\Pi\Sigma} \chi_f \rangle \langle \chi_f, \Gamma^{\Pi\Sigma} \chi_f \rangle \in \wedge^6 F^* \otimes \wedge^6 F^*$$
$$U = \{ \chi_f \in \wedge^{\mathsf{even/odd}} F^* : q(\chi_f) < 0 \}$$

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Define a Hitchin function

$$H(\chi_f) \equiv \sqrt{-\frac{1}{3}q(\chi_f)} \in \wedge^6 F^*$$

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Define a Hitchin function

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which gives an integrable complex structure on U

Then we can get another real form $\hat{\chi}_f$ and a complex form Φ_f by Mukai pairing

$$\langle \hat{\chi}_f, \chi_f \rangle = -dH(\chi_f)$$
 i.e., $\hat{\chi}_f = -\frac{\partial H(\chi_f)}{\partial \chi_f}$
--> $\Phi_f \equiv \frac{1}{2}(\chi_f + i\hat{\chi}_f)$ $H(\Phi_f) = i\langle \Phi_f, \overline{\Phi}_f \rangle$

Hitchin showed: Φ_f is a (form corresponding to) pure spinor!

N.J. Hitchin math/0010054, math/0107101, math/0209099

Consider the space of pure spinors Φ ...

Mukai pairing $\langle *, * \rangle \longrightarrow$ symplectic structure Hitchin function $H(*) \longrightarrow$ complex structure \Downarrow

The space of pure spinor is Kähler

Consider the space of pure spinors Φ ...

 $\begin{array}{cccc} \text{Mukai pairing } \left< \ast, \ast \right> & \longrightarrow & \text{symplectic structure} \\ \text{Hitchin function } H(\ast) & \longrightarrow & \text{complex structure} \\ & & \downarrow \end{array}$

The space of pure spinor is Kähler

Quotienting this space by the \mathbb{C}^* action $\Phi \to \lambda \Phi$ for $\lambda \mathbb{C}^*$

 \rightarrow The space becomes a local special Kähler geometry with Kähler potential K:

$$e^{-K} = H(\Phi) = i\langle \Phi, \overline{\Phi} \rangle = i(\overline{X}^A \mathcal{F}_A - X^A \overline{\mathcal{F}}_A) \in \wedge^6 F^*$$

 X^A : holomorphic projective coordinates

 \mathcal{F}_A : derivative of prepotential \mathcal{F} , i.e., $\mathcal{F}_A = \partial \mathcal{F} / \partial X^A$

These are nothing but objects which we want to introduce in $\mathcal{N} = 2$ supergravity!

Spaces of pure spinors Φ_{\pm} on $F \oplus F^*$ with $SU(3) \times SU(3)$ structure $\|$ special Kähler geometries of local type = Hodge-Kähler geometries

For a single SU(3)-structure case:

$$\Phi_{+} = \frac{1}{8} e^{-B - iJ} \qquad K_{+} = -\log\left(\frac{1}{48}J \wedge J \wedge J\right)$$
$$\Phi_{-} = -\frac{i}{8} e^{-B}\Omega \qquad K_{-} = -\log\left(\frac{i}{64}\Omega \wedge \overline{\Omega}\right)$$

Structure of forms is exactly same as the one in the case of Calabi-Yau compactification! We should truncate Kaluza-Klein massive modes from these forms to obtain 4D supergravity.

Appendix: puzzle on conventional differential forms

M. Graña, J. Louis, D. Waldram hep-th/0612237

Recall that Φ_{\pm} are expanded in terms of truncation bases Σ_{+} and Σ_{-} .

Whenever $c_{\parallel} \neq 0$, the structure Φ_+ contains a scalar. This implies that at least one of the forms in the basis Σ_+ contains a scalar. Let us call this element Σ_+^1 , and take the simple case where the only non-zero elements of \mathbb{Q} are those of the form $\mathbb{Q}_{\hat{I}}^{-1}$ (where $\hat{I} = 1, \ldots, 2b^- + 2$). Thus $d(\Sigma_-)_{\hat{I}} = \mathbb{Q}_{\hat{I}}^{-1}\Sigma_+^1$ and so if $\mathbb{Q}_{\hat{I}}^{-1} \neq 0$ then $(d\Sigma_-)_{\hat{I}}$ contains a scalar.

But this is not possible if d is an honest exterior derivative, acting as $d: \Lambda^p \to \Lambda^{p+1}$.

The same is true if c_{\parallel} is zero. In this case, there may be no scalars in any of the even forms Σ_{-} , and for an "honest" d operator, there should be then no one-forms in $d\Sigma_{-}$. But we again see from that Φ_{-} contains a one-form, and as a consequence so do some of the elements in Σ_{-} .

One way to generate a completely general charge matrix \mathbb{Q} in this picture is to consider a modified operator d which is now a generic map $d: U^+ \to U^-$ which satisfies $d^2 = 0$ but does not transform the degree of a form properly.

In particular, the operator ${f d}$ can map a p-form to a (p-1)-form.

Of course, this d does not act this way in conventional geometrical compactifications.

One is thus led to conjecture that to obtain a generic \mathbb{Q} we must consider non-geometrical compactifications. One can still use the structures

$$\mathrm{d}\Sigma_{-} \sim \mathbb{Q}\Sigma_{+}, \quad \mathrm{d}\Sigma_{+} \sim \mathcal{S}_{+}\mathbb{Q}^{T}(\mathcal{S}_{-})^{-1}\Sigma_{-}$$

to derive sensible effective actions, expanding in bases Σ_+ and Σ_- with a generalised d operator, but there is of course now no interpretation in terms of differential forms and the exterior derivative.

--→ introduce generalized fluxes

(not only geometrical H- and f-fluxes, but also Q- and R-fluxes)

For a geometrical background it is natural to consider forms of the type

$$\omega = e^{-B} \omega_{m_1 \cdots m_p} e^{m_1} \wedge \cdots \wedge e^{m_p} \quad w/ \omega_{m_1 \cdots m_p} \text{ constant}$$

Action of d on ω is

$$\mathrm{d}\omega = -H^{\mathsf{fl}} \wedge \omega + f \cdot \omega, \qquad (f \cdot \omega)_{m_1 \cdots m_{p+1}} = f^a{}_{[m_1 m_2]} \omega_{a|m_3 \cdots m_{p+1}]}$$

The natural nongeometrical extension is then to an operator \mathcal{D} such that

$$\mathcal{D} := \mathrm{d} - H^{\mathsf{fl}} \wedge -f \cdot -Q \cdot -R \sqcup$$

$$(Q \cdot \omega)_{m_1 \cdots m_{p-1}} = Q^{ab}_{[m_1} \omega_{|ab|m_2 \cdots m_{p-1}]}, \qquad (R \sqcup \omega)_{m_1 \cdots m_{p-3}} = R^{abc} \omega_{abcm_1 \cdots m_{p-3}}$$

Requiring $D^2 = 0$ implies that same conditions on fluxes as arose from the Jacobi identities for the extended Lie algebra

$$[Z_a, Z_b] = f_{ab}{}^c Z_c + H_{abc} X^c$$

$$[X^a, X^b] = Q^{ab}{}_c X^c + R^{abc} Z_c$$

$$[X^a, Z_b] = f^a{}_{bc} X^c - Q^{ac}{}_b Z_c$$

We can see nongeometrical information in \mathbb{Q} as contribution from Q and R.

Setup in $\mathcal{N}=1$ theory

Functionals are given by two Kähler potentials on two Hodge-Kähler geometries of Φ_{\pm} :

$$K = K_{+} + 4\varphi$$

$$K_{+} = -\log i \int_{\mathcal{M}} \langle \Phi_{+}, \overline{\Phi}_{+} \rangle = -\log i (\overline{X}^{A} \mathcal{F}_{A} - X^{A} \overline{\mathcal{F}}_{A})$$

$$K_{-} = -\log i \int_{\mathcal{M}} \langle \Phi_{-}, \overline{\Phi}_{-} \rangle = -\log i (\overline{Z}^{I} \mathcal{G}_{I} - Z^{I} \overline{\mathcal{G}}_{I})$$

$$\int_{\mathcal{M}} \operatorname{vol}_{6} = \frac{1}{8} e^{-K_{\pm}} = e^{-2\varphi + 2\phi^{(10)}}$$

Introduce $C = \sqrt{2}ab e^{-\phi^{(10)}} = 4ab e^{\frac{K_{-}}{2}-\varphi}$

$$\therefore \quad e^{-2\varphi} = \frac{|\mathcal{C}|^2}{16|a|^2|b|^2} e^{-K_-} = \frac{i}{16|a|^2|b|^2} \int_{\mathcal{M}} \langle \mathcal{C}\Phi_-, \overline{\mathcal{C}\Phi}_- \rangle$$
$$= \frac{1}{8|a|^2|b|^2} \Big[\operatorname{Im}(\mathcal{C}Z^I) \operatorname{Re}(\mathcal{C}\mathcal{G}_I) - \operatorname{Re}(\mathcal{C}Z^I) \operatorname{Im}(\mathcal{C}\mathcal{G}_I) \Big]$$

See the SUSY variation of 4D $\mathcal{N} = 2$ gravitinos:

$$\delta \psi_{\mathcal{A}\mu} = \nabla_{\mu} \varepsilon_{\mathcal{A}} - S_{\mathcal{A}\mathcal{B}} \gamma_{\mu} \varepsilon^{\mathcal{B}} + \dots$$
$$S_{\mathcal{A}\mathcal{B}} = \frac{\mathrm{i}}{2} \mathrm{e}^{\frac{K_{+}}{2}} \begin{pmatrix} \mathcal{P}^{1} - \mathrm{i}\mathcal{P}^{2} & -\mathcal{P}^{3} \\ -\mathcal{P}^{3} & -\mathcal{P}^{1} - \mathrm{i}\mathcal{P}^{2} \end{pmatrix}_{\mathcal{A}\mathcal{B}}$$

The components are also written by Φ_{\pm} :

$$\mathcal{P}^{1} - \mathrm{i}\mathcal{P}^{2} = 2 \mathrm{e}^{\frac{K_{-}}{2} + \varphi} \int_{\mathcal{M}} \left\langle \Phi_{+}, \mathcal{D}\Phi_{-} \right\rangle, \qquad \mathcal{P}^{1} + \mathrm{i}\mathcal{P}^{2} = 2 \mathrm{e}^{\frac{K_{-}}{2} + \varphi} \int_{\mathcal{M}} \left\langle \Phi_{+}, \mathcal{D}\overline{\Phi}_{-} \right\rangle$$
$$\mathcal{P}^{3} = -\frac{1}{\sqrt{2}} \mathrm{e}^{2\varphi} \int_{\mathcal{M}} \left\langle \Phi_{+}, G \right\rangle$$

Note: $\hat{\Psi}_{\mathcal{A}\mu} = \Psi_{\mathcal{A}\mu} + \frac{1}{2}\Gamma_{\mu}{}^{m}\Psi_{m}^{\mathcal{A}} = \psi_{\mathcal{A}\mu\pm} \otimes \eta_{+} + \psi_{\mathcal{A}\mu\mp} \otimes \eta_{-} + \dots$

4D $\mathcal{N} = 1$ fermions given by the SUSY truncation from 4D $\mathcal{N} = 2$ system:

SUSY parameter : $\varepsilon \equiv \overline{n}^{\mathcal{A}} \varepsilon_{\mathcal{A}} = a \varepsilon_1 + \overline{b} \varepsilon_2$

gravitino: $\psi_{\mu} \equiv \overline{n}^{\mathcal{A}} \psi_{\mathcal{A}\mu} = a \psi_{1\mu} + \overline{b} \psi_{2\mu}, \qquad \widetilde{\psi}_{\mu} \equiv \left(b \psi_{1\mu} - \overline{a} \psi_{2\mu} \right)$ gauginos: $\chi^{\mathcal{A}} \equiv -2 e^{\frac{K_{+}}{2}} D_{b} X^{\mathcal{A}} \left(\overline{n}^{\mathcal{C}} \epsilon_{\mathcal{C}\mathcal{E}} \chi^{\mathcal{E}b} \right)$ where $\overline{n}^{\mathcal{A}} = \left(a, \overline{b} \right), \quad \epsilon_{\mathcal{A}\mathcal{B}} = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}$ SUSY variations yield the superpotential and the D-term:

$$\delta \psi_{\mu} = \nabla_{\mu} \varepsilon - \overline{n}^{\mathcal{A}} S_{\mathcal{A}\mathcal{B}} n^{*\mathcal{B}} \gamma_{\mu} \varepsilon^{c} \equiv \nabla_{\mu} \varepsilon - e^{\frac{K}{2}} \mathcal{W} \gamma_{\mu} \varepsilon^{c}$$
$$\delta \widetilde{\psi}_{\mu} = 0$$
$$\delta \chi^{A} = \operatorname{Im} F^{A}_{\mu\nu} \gamma^{\mu\nu} \varepsilon + i D^{A} \varepsilon$$

$$\mathcal{W} = \frac{\mathrm{i}}{4\overline{a}b} \Big[4\mathrm{i} \, \mathrm{e}^{\frac{K_{-}}{2} - \varphi} \int_{\mathcal{M}} \left\langle \Phi_{+}, \mathcal{D}\mathrm{Im}(ab\Phi_{-}) \right\rangle + \frac{1}{\sqrt{2}} \int_{\mathcal{M}} \left\langle \Phi_{+}, G \right\rangle \Big]$$

$$\equiv \mathcal{W}^{\mathsf{R}\mathsf{R}} + U^{I} \, \mathcal{W}^{\mathbb{Q}}_{I} + \widetilde{U}_{I} \, \widetilde{\mathcal{W}}^{I}_{\mathbb{Q}}$$

$$\mathcal{W}^{\mathsf{R}\mathsf{R}} = -\frac{\mathrm{i}}{4\overline{a}b} \Big[X^{A} \, e_{\mathsf{R}\mathsf{R}A} - \mathcal{F}_{A} \, m_{\mathsf{R}\mathsf{R}}^{A} \Big]$$

$$\mathcal{W}^{\mathbb{Q}}_{I} = \frac{\mathrm{i}}{4\overline{a}b} \Big[X^{A} \, e_{IA} + \mathcal{F}_{A} \, p_{I}^{A} \Big], \qquad \widetilde{\mathcal{W}}^{I}_{\mathbb{Q}} = -\frac{\mathrm{i}}{4\overline{a}b} \Big[X^{A} \, m_{A}^{I} + \mathcal{F}_{A} \, q^{IA} \Big]$$

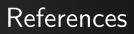
$$D^{A} = 2 e^{K_{+}} (K_{+})^{c\overline{d}} D_{c} X^{A} \overline{D_{d}} \overline{X^{B}} [\overline{n}^{\mathcal{C}} (\sigma_{x})_{\mathcal{C}} {}^{\mathcal{B}} n_{\mathcal{B}}] \left(\mathcal{P}_{B}^{x} - \mathcal{N}_{BC} \widetilde{\mathcal{P}}^{xC} \right)$$

$\mathcal{N}=2$ multiplets: $(t^a=X^a/X^0,\ z^i=Z^i/Z^0)$		gravity multiplet	$g_{\mu u},\;A^0_\mu$		
		vector multiplets	$A^a_\mu, t^a = b^a + \mathrm{i} v^a$	$a = 1, \ldots, b^+$	
		hypermultiplets	$z^i,\;\xi^i,\;\widetilde{\xi_i}$	$i=1,\ldots,b^-$	
		tensor multiplet	$B_{\mu u}, \; arphi, \; \xi^0, \; \widetilde{\xi}_0$		
		orientifold projection: $\mathcal{O} \equiv \Omega_{\rm WS} (-1)^{F_L} \sigma$			
$\mathcal{N}=1$ multiplets:	gravity multiplet		$g_{\mu u}$		
	vector multiplets		$A^{\hat{a}}_{\mu}$	$\hat{a} = 1, \dots, \hat{n}_v = b^+ - n_{ch}$	
	chiral multiplets		$t^{\check{a}} = b^{\check{a}} + \mathrm{i} v^{\check{a}}$	$\check{a} = 1, \dots, n_{ch}$	
	chiral/linear multiplets		$U^{\check{I}} = \xi^{\check{I}} + \mathrm{i}\mathrm{Im}(\mathcal{C}Z^{\check{I}})$	$I = (\check{I}, \hat{I}) = 0, 1, \dots, b^{-}$	
			$\widetilde{U}_{\hat{I}} = \widetilde{\xi}_{\hat{I}} + \mathrm{i}\mathrm{Im}(\mathcal{C}\mathcal{G}_{\hat{I}})$		
	(projected out)		$B_{\mu\nu}, \ A^0_{\mu}, \ A^{\check{a}}_{\mu}, \ t^{\hat{a}}, \ U$	$T^{\hat{I}}, \ \widetilde{U}_{\check{I}}$	

 \mathcal{N}

Parameters are restricted as $a = \overline{b} e^{i\theta}$ and $|a|^2 = |b|^2 = \frac{1}{2}$

T.W. Grimm hep-th/0507153



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and more...