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COLUMBIA UNIVERSITY
BROADWAY AND 120TH STREET, NEW YORK 27, NEW YORK

PHYSICS COLLOQUIUM
7:30 P.M. ROOM 301 PUPIN PHYSICS LABORATORIES

Friday, April 14, 1950

Speaker: Dr. J. R. Haynes
Bell Telephone Laboratories

Subject: The mobility and lifetime of injected
holes and electrons in germanium

Friday, April 21, 1950

Speaker: Dr. H. Yukawa

Subject: Present status of meson theory

Friday, April 28, 1950

There will be no Physics Colloquium

- (i) General structure of Rel. Theory of Quantized Field.
- (ii) Our knowledge about mesons about a year ago (H. Yukawa, Rev. Mod. Phys. 21 (1949), 474)
- (iii) Nuclear forces (low energy region, high energy region)
- (iv) Neutral mesons
- (v) meson scattering.
- (vi) Creation of mesons
- (vii) Remarks on heavier mesons

Present Status of Meson Theory

02

By Hideki Yukawa

(1)

Physics Colloquium, Columbia University
Room 301, Pupin, April 21, 1950

The development of theoretical physics can be regarded in a sense as the development of the concept of "field" ^{contrasted} ~~side by side~~ ^{in a constant} ~~side~~ ^{close correlation} with that of "matter". We find the best example in the development of electrodynamics. So was the development of meson theory. It started from the extension of the concept of the field of force so as to include the "nuclear forces" in addition to the electromagnetic and gravitational forces.

However, the situation in meson theory was somewhat different from the earlier that in electrodynamics. ^{Relativity theory} and quantum theory were there already ^{it was because} and they were mixed ^{mixed together into} into the theory of "quantized field". In this theory, the ~~most fundamental~~ ^{most} fundamental point ~~was~~ ^{not} there was not so much as field ^{and} ~~versus~~ ^{plus} matter as free fields and interaction between them.

In order to make clear this point, I would like to mention very briefly the whole structure of the present theory of quantized field.

- (I) Free field
field quantities \rightarrow (operators)
field equations \neq commutation relations
spin \leftrightarrow statistics; mass.
- (II) Interaction
Schrödinger function or functional Schrödinger equation
Hamiltonian or Lagrangian \neq (operator)
Coupling constants e, g etc.

So far as the free field is concerned, the present theory seems to be complete by itself. The requirements of relativity theory and quantum theory are sufficient to deduce all possible type of field equations and commutation relations, provided if ~~that~~ we assume further

- (i) field equations are linear.
- (ii) ~~of~~ particles ^{associated with} accompanying the field ~~to~~ are dimensionless particles and obey either the B. E. statistics or F. D. statistics.

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However, there is a great arbitrariness as to which ~~of the~~ types of fields are realized in nature. In particular, the values of masses of existing particles must be ~~considered~~ ^{accepted} only as such without any fundamental theoretical reason. As for the interaction, the situation is still worse. ~~It is true that as~~ ~~is~~ well known, the Schrödinger equation for the system consisting of two or more fields interacting with each other has no finite solution, so that the ^{present} theory of interaction of elementary particles is inconsistent from the beginning. It is true that recent development of ~~the~~ so called many-time-formalism together with the renormal procedure of renormalization of mass and charge (more generally ~~the~~ ^{makes and} coupling constants) threw new light on this difficulty. ~~problem~~, but it is certain that ~~it~~ we are still far from the satisfactory solution. Apart from this fundamental difficulty, there is further a great deal of arbitrariness in the choice of the interaction between quantized fields. The only criterion, which was useful, is that the interaction

operator must be as simple as possible, as just we postulated that the field equations ~~must be as simple as possible~~ ^{and commutation relations}.

Under these circumstances, we must admit that the whole theory of elementary particles, particularly meson theory, is provisional at the present stage.

The ~~best~~ ~~that I can~~ ~~tell~~ we can do today is to discriminate between

- (i) what we understand about mesons or more generally about elementary particles
- (ii) and what we do not understand as clearly as possible.

It is generally accepted
Today ~~we know now~~ that there are several groups of elementary particles:

Group I
nucleon

II

III[±]

IV

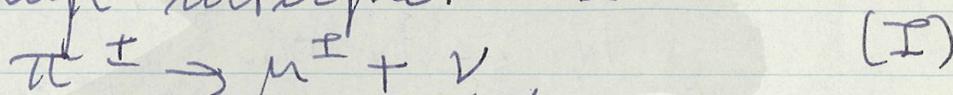
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(N, P); (e[±], ν); μ[±]; π[±], π⁰; γ

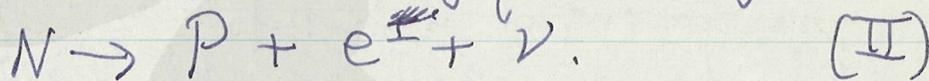
~~So far as~~ ~~the phenomena up to about~~ ~~neutrons and protons constitute~~ a distinct group I, because a nucleon can be created, only if another nucleon is annihilated, or vice versa.

There is no indication of ^{the creation or annihilation} one nucleon, which is not accompanied by the annihilation or creation of another.

As for the group II and III, existing experimental results could be understood more easily by assuming regarding them as one single group, because, for example, π - μ -decay can be usually interpreted as



just as the ^{prototype of the} ordinary β -decay is

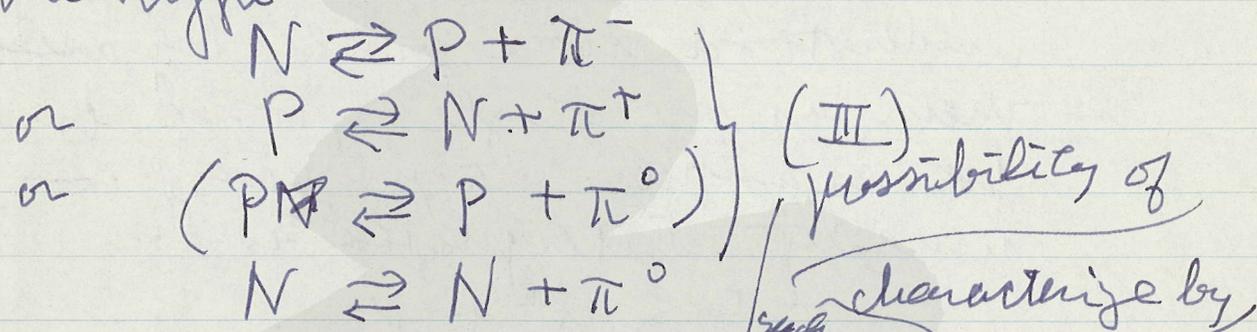


In other words, the creation or annihilation of a μ -meson is associated with that of a neutrino, just as the creation or annihilation of an electron is associated with that of a neutrino in the case of β -decay. Alternatively, you may say that μ , e , ν are different states of ^{one} same type of elementary particles just as N , P are those of another type.

As for π^\pm and π^0 , among ^{as for the} group IV, we must discuss ^{a little} more in detail, because the existence and nature of ~~the~~ π^\pm , π^0 , particularly of the latter

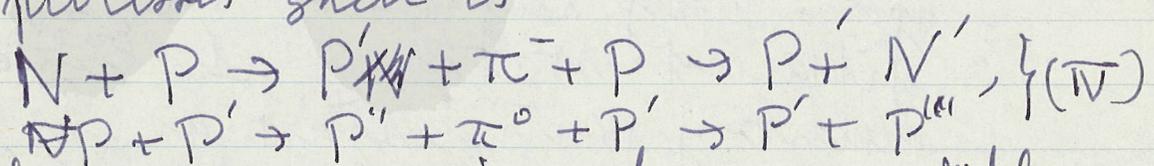
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were confirmed experimentally only very recently. The most important property of π meson group is the ~~possibility~~ possibility of the processes of the type



As well known, ~~that~~ this process with a large coupling constant g was the ~~fundamental assumption of meson~~ has been regarded as the main support of the meson theory.

Of course, such processes cannot occur be observed directly because the laws of conservation of energy and momentum are not satisfied. These processes are combined together into the processes such as



which are supposed to be responsible for the nuclear forces* manifested in the ~~in the~~ however, it should be noticed that the notion of nuclear forces is not as clear as the notion of

* Discussions of Deuteron problem and N-P, P-P scattering

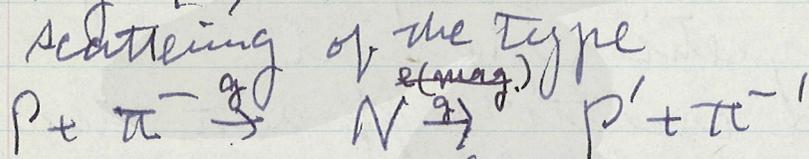
photo electrostatic force (Coulomb force), because the simple notion of nuclear static potential cannot be applied to a very short distance of the order of $\frac{h}{mc}$ between nucleons (and ^{hence} nucleon-nucleon scattering at high energy), whereas Coulomb potential has always the definite meaning, because it appears as the inevitable consequence of the elimination of longitudinal and scalar time components of the electromagnetic potential.

~~We~~ the creation of a π -meson by collision of a high energy proton or a neutron with the nucleus ^{certainly} was regarded as a more direct evidence for (IV). However, ~~there remains still various~~ it does not follow necessarily that (IV) is the main contribution to the nuclear forces, as long as the collided nucleus is a complex nucleus such as C carbon, because the interaction between the nucleon and the carbon nucleus might be more complicated and the emission of a π -meson might be ^{the final result} a consequence of such various complicated competent ~~pro~~ possible processes.

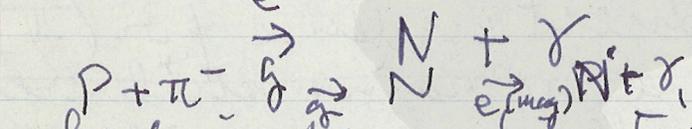
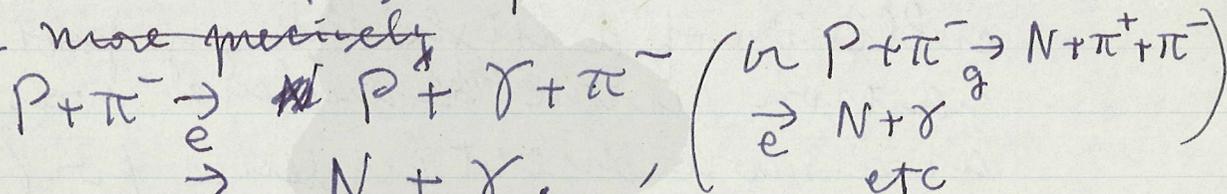
In this connection, recent experiments in Berkeley concerning the interaction of π -mesons with protons is of decisive importance. In this case, ~~another~~ ^{well known} type of processes such as

$$P \rightleftharpoons P' + \gamma \quad (IV)$$

can be supposed to take part together with the process (III). Namely, when a ~~negative~~ π^- is absorbed by P , the result must be a simple scattering of the type



or more precisely



which is the combination of (IV). It was known that the specific nuclear interaction has a contribution appreciably to the collision between π^- and P , only when π^- is slowed down and captured into a low Bohr orbit in a time short compared with the mean life-time $\tau_\pi \sim 10^{-8}$ sec. (at least for liquid hydrogen). * Moreover, the subsequent

* R. E. Marshak and A. S. Wightman, Phys. Rev. 76 (1949), 114. Wightman, ibid 77 (1950), 521. See also G. Araki, Sci. Paper 29 (1941), 14.

probability of subsequent absorption of π^- resulting in the emission of γ is much larger than ^{that of} the spontaneous decay of π^- . Recently, calculations of this process were made by Marshak and Wightman* for the scalar and pseudo-scalar mesons. They are

$$\tau_S = 8.0 \times 10^{-15} \text{ sec.} \quad *$$

$$\tau_P = 1.1 \times 10^{-15} \text{ sec.}$$

for the absorption from the bound state in liquid hydrogen with $g^2/\hbar c = 0.25$ $f^2/\hbar c = 0.25$.

At any rate thus, we expect that ~~a~~ monochromatic γ -rays of energy around mc^2 .

Recent experiment by Panofsky** indicates on the contrary the ~~appearance~~ spectrum of γ -rays with the band around 65 MeV and a line at 130 MeV. This could only be understood, only if ^{we} assume the existence of π^0 with the mass $m_{\pi^0} < m_{\pi} - (M_N - M_P)$,

$$* \quad 1/\tau_S = 2 (e^2/\hbar c)^4 (g^2/\hbar c) (\Gamma \mu/M)^2 (\bar{\mu}/\mu)^2 (\mu c^2/\hbar)$$

Γ : difference of magnetic moments, $|\Gamma| = 4.71$

$$1/\tau_P = 8 (e^2/\hbar c)^4 (f^2/\hbar c) (\bar{\mu}/\mu)^2 (\mu c^2/\hbar)$$

$\bar{\mu} = \mu M / (m + M)$

** New York meeting & Private Conversation

'07

(6)

very rapidly, when the energy of the proton becomes larger than the threshold for the production of π^\pm -mesons. Moreover ~~the intensity of~~ For 290 MeV and 340 MeV, the intensity is of the same order of magnitude as the intensity of π^\pm -mesons ~~with~~ ^{and} the energy distribution has a maximum around 70 MeV. ~~70 MeV~~

This could be ~~so~~ accounted for consistently, by assuming the 2γ -decay of π^0 -mesons, which were produced side by side with π^\pm . They estimated the mass ~~around~~ ^{about} 300 me, but their results are not inconsistent with the assumption that the mass m_{π^0} is a little smaller than that of the $\cdot m_\pi$. Moreover, Panofsky and Steinberger (Appendix) carried out the experiment, which confirming the coincidence of two γ -rays due to the 2γ -decay.

Thus the existence of neutral π -mesons, which have been expected from the for long time theoretically, seems to be finally confirmed. Theoretical arguments for postulating the neutron mesons were (sometimes called neutrons)

- (i) Approximate charge independence of Nuclear forces
- (ii) Interpretation of P-P scattering

Now the γ -decay of neutral mesons, which was also predicted theoretically years ago, is an interesting subject by itself. It has been known that a scalar or a pseudoscalar mesons can decay into ~~three mesons~~ two photons and the life-time may be of the order of 10^{-14} sec or smaller* whereas a vector meson or a p.v. meson can decay only into three photons with much long life-time.

Recent some evidence, which support this theoretical argument, was obtain in Rochester in connection with the so-called "R-shower" ** in cosmic rays. This is a high energy shower consisting of a ^{narrow} dense core of relativistic particles and (2352 particles) and ~~surrounding~~ surrounded by a more diffuse shower of 33 relativistic particles. In addition 18 non-relativistic particles carrying at 23 units of charge and a total energy of ~ 3 Bev emerge from the origin of the shower.

** M.F. Kaplan, B. Peters and H.L. Bradt, Phys. Rev. 76 (1949), 1735; R.E. Marshak, *ibid.* 1736.

* According to recent calculations based on the many-time formalism, the life-time turns out to be 10^{-10} sec for ps. scalar mesons.

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hand an evidence for the existence of π^0
meson of ~ 300 me and life-time less than
 10^{-13} sec.

Apart from the problem of neutral mesons,
existing experimental material indicate
in general meson interaction between
 π^\pm -mesons and nucleons, which is consistent
with the assumption $g^2/\hbar c = 0.1 \sim 1$,
However although it is not clear whether
 π^\pm, π^0 -meson are mainly responsible
for nuclear forces. The uncertainty comes
from the complicated ion of the interaction
between nucleons at very short distance
~~between \hbar/mc and \hbar/c of the order of~~
 ~~\hbar/mc , and a little~~ This might be a
deeper question, which is intimately
connected with the question of divergence
in general field theory.

However, there is another problem, which
is not ^{yet} settled down, ~~that~~ which does not
seem so deep a question. Namely, the
scattering of π^\pm -mesons by nucleons has
been dealt with theoretically from the
beginning of meson theory, since it is one
of the simplest possible processes.
According to conventional perturbation
calculations, the cross-section for

scalar mesons is *

$$d\sigma_s = \frac{1}{4} \left(\frac{g^2}{Mc^2} \right)^2 \left(\frac{E}{E_0} \right)^2 \left\{ 1 + \left(\frac{2Mc^2}{E_0} \right)^2 + \frac{E}{E_0} \right\} d\Omega,$$

where E_0, E are the energies of incident and scattered mesons. The total cross-section

$$\sigma_s = \pi \left(\frac{g^2}{Mc^2} \right)^2 \left\{ \frac{(2+\gamma)^2}{\gamma^2(1+2\gamma)} + \frac{1+\gamma}{(1+2\gamma)^2} \right\},$$

with $\gamma = E_0/Mc^2$. For $\gamma \gg 1$, σ_s is of the order of 10^{-29} cm^2

so that we can expect that no appreciable contribution to the absorption of π^\pm in matter. For $\gamma \ll 1$ (but not too small),

$$\begin{aligned} \sigma_s &= 4\pi \left(\frac{g^2}{E_0} \right)^2 = 4\pi \left(\frac{g^2}{\hbar c} \right)^2 \left(\frac{\hbar c}{E_0} \right)^2 \\ &= 4\pi \left(\frac{g^2}{\hbar c} \right)^2 \left(\frac{\hbar}{\mu c} \right)^2 \left(\frac{\mu c^2}{E_0} \right)^2 \\ &\sim 10^{-26} \sim 10^{-27} \text{ cm}^2. \end{aligned}$$

For pseudoscalar mesons, **

$$d\sigma_{ps} = \left(\frac{\hbar}{\mu c} \right)^2 \left(\frac{f^2}{\hbar c} \right)^2 \left(\frac{p_0 c}{E_0} \right)^2 \left(\frac{p_0}{\mu c} \right)^2 d\Omega$$

$$\text{or } \sigma_s = 4\pi \left(\frac{\hbar}{\mu c} \right)^2 \left(\frac{f^2}{\hbar c} \right)^2 \left(\frac{p_0 c}{E_0} \right)^2 \left(\frac{p_0}{\mu c} \right)^2$$

for $p \sim \mu c$. The first factor $4\pi \left(\frac{\hbar}{\mu c} \right)^2 \left(\frac{f^2}{\hbar c} \right)^2$ again has a value $10^{-26} \sim 10^{-27} \text{ cm}^2$.

* H. Yukawa and S. Sakata, Proc. Phys.-Math. Soc. Japan 19 (1937), 1084.

** Y. Tanikawa and H. Yukawa, ibid #23 (1941), 445.

The difference is between ~~ps~~ scalar and pseudoscalar mesons is that the cross-section decreases with the energy E_0 , proportional to E_0^{-2} (first) and then E_0^{-1} , whereas in the former case, whereas it increases with E_0 proportional to E_0 . ~~first~~ In any case, the cross-section is always of the order of $10^{-26} \sim 10^{-27} \text{ cm}^2$ for moderate energies.*

On the contrary, recent experiments by Piccioni** and Fretter*** gave much smaller values for the cross-section.

Namely, according to Piccioni for $E_0 \sim 300 \text{ MeV}$
 $\sigma < 2 \times 10^{-27} \text{ cm}^2 (\lambda > 1000 \text{ g/cm}^2)$

According to Fretter
 $\sigma < 3 \times 10^{-28} \text{ cm}^2 (\lambda > 150 \text{ g/cm}^2)$
 for lead for scattering angle $> 15^\circ$

According to
 * Recent calculation by R. E. Marshak, Cosmic Rays. Proc. Echo Lake Symposium, 1949

	$E_0 = 1$	$E_0 = 3$
σ_s	$2 \cdot 10^{-26} \text{ cm}^2$	$1 \cdot 10^{-27}$
σ_{ps}	$3 \cdot 10^{-26} \text{ cm}^2$	$6 \cdot 10^{-27}$
for $g_{\pi K}^2 \sim 1/4$		$f_{\pi K}^2 \sim 4$

** O. Piccioni, Phys. Rev. 72 (1950), 1, 6.

*** W. B. Fretter, 76 (1949), 511.

On the other hand, the process
 $\gamma + p \rightarrow \pi^+ + n$

for γ -energy up to ^{max.} 330 MeV was investigated recently*, which seems to be very simple and important for meson theory.

(Paraffin - Carbon) The calculations are performed years ago and recently by Feshbach.**

According to Araki*** the cross-section of production of a ₁ meson by γ -p-reaction is

$$d\Phi_{p_1} = 2\Phi_0 \left(\frac{Mc^2}{E_0} \right) \left(\frac{E}{E_0} - \frac{1}{2} \right) dE$$

$$\text{where } \frac{1}{2} = 2 \left\{ \frac{mc^2 \cdot p \cdot c \cdot \sin\theta}{\sqrt{[(p_0 - p_1)^2 + m^2c^4] + (E_0 - E)^2}} \right\}^2 + \frac{(mc^2)^2}{2Mc^2 \cdot E_0} : \text{negligible}$$

$$\Phi_0 = \pi \left(\frac{1}{2} \frac{e t_1}{Mc^2} + \frac{e t_2}{mc^2} \right)^2$$

$$\Phi_{p_1} = 4\pi \left(\frac{1}{2} \frac{e t_1}{Mc^2} + \frac{e t_2}{mc^2} \right)^2 \frac{(E_0 + Mc^2) \{ Mc^2 E_0 + \frac{1}{2}(mc^2)^2 \}}{E_0 (2E_0 + Mc^2)^2} \times \sqrt{\left(1 - \frac{m}{2M} \cdot \frac{mc^2}{E_0} \right)^2 - \left(\frac{mc^2}{E_0} \right)^2}$$

*** G. Araki, Sci. Pap. 39 (1941), 14.
 See Appendix.

* J. Steinberger and A. S. Bishop,
 Berkeley Report.

** H. Feshbach and M. Lax, Phys. Rev. 26 (1949), 134;

For $\mu c^2 < E_0 < Mc^2$

$$\begin{aligned}\Phi_{ps} &\sim 4\Phi_0 \\ &= 4\pi \left(\frac{1}{2} \frac{ef_1}{Mc^2} + \frac{ef_2}{\mu c^2} \right)^2 \\ &\sim 10^{-28} \sim 10^{-29} \text{ cm}^2\end{aligned}$$

The corresponding cross-section for scalar mesons is

$$\begin{aligned}d\Phi_s &= 2\pi \left(\frac{ef_1}{\mu c^2} \right)^2 \frac{Mc^2}{E_0^2} \left\{ \left(1 - \left(\frac{\mu}{2M} \right)^2 \right) \right. \\ &\quad \left. + \left(\frac{\mu}{2M} \right)^2 \left(\frac{E}{E_0} - \frac{2Mc^2}{E_0} \right) \right\} dE\end{aligned}$$

and

$$\Phi_s = 4\pi \left(\frac{ef_1}{E_0} \right)^2 \log \left(\frac{E_0}{\mu c^2} \right) \text{ for } \mu c^2 < E_0 < Mc^2$$

which is a little smaller than Φ_{ps} and decreases with energy E_0 .

In both cases, Φ_{ps} and Φ_s decrease with E_0^{-1} for $E_0 \gg Mc^2$.

Another difference between two cases is that the $d\Phi_{ps}$ gives ~~an~~ ~~angular~~ a rather flat angular distribution with the maximum for an angle larger than 90° , whereas $d\Phi_s$ gives a maximum for ~~the~~ an angle less than 90° . The experiment is in favour of ~~the~~ the pseudoscalar mesons. (See, however, Appendix)

Finally I would like to add a few ~~conclusion~~ remarks on the problem of heavier mesons.

Theoretically, we can easily conceive that if a type of ~~meson~~ heavier mesons heavier than π -meson is created in cosmic rays ~~in sufficiently large number~~ frequently, this implies that the life-time of such mesons will be very short.

On the contrary, if the life-time is sufficiently long to be observed, they may not be created abundantly.

This is because the probability of creation is proportional to g_{π}^2 , whereas the life-time is proportional to g_{π}^{-2} .

Under these circumstances, we can say not expect a large percentage of heavier mesons in cosmic rays, but still we cannot exclude the existence of them. Recent experiment by Anderson might be (and other previous experiments on π -mesons) could be understood in this way.

Appendix to
 "Present Status of Meson Theory"

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On Neutral Mesons.

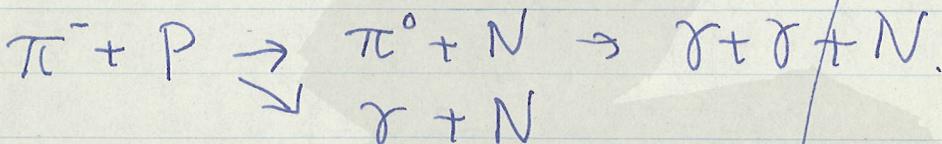
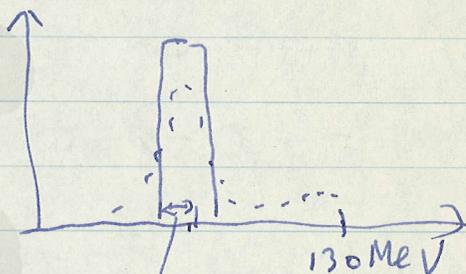
April 11, 1950

W. Panofsky, π^- -P-reaction

(New York Meeting, Private talk)

L. Aamodt,
 H. York

130	1.35
85	4.
60	13.
40	6



($\pi^- + He$: no γ -ray)

$$\frac{A_\pi}{A_\gamma} \cong \frac{g^2}{e^2} \cdot \frac{p_{\pi^0}}{mc}$$

? $\pm 5 \text{ MeV}$ $0.1 < g^2/mc$

$$\Delta M < \Delta \mu < 2.9 \text{ MeV}$$

$$1.3 \text{ MeV} < \Delta \mu < 4.7 \text{ MeV}$$

$$M_N - M_p < m_\pi - m_{\pi^0}$$

S. Hayakawa, Letter Priv. Comm.

π, π^0 : scalar, neutral theory:

$$A_\pi \gg A_\gamma$$

pseudoscalar, symmetrical theory:

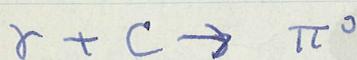
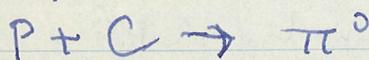
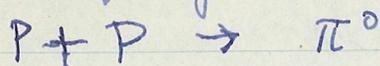
$$A_\pi \ll A_\gamma$$

pseudoscalar symmetrical

$$A_\pi \sim A_\gamma.$$

Marshak, Washington Meeting

(Panofsky?)

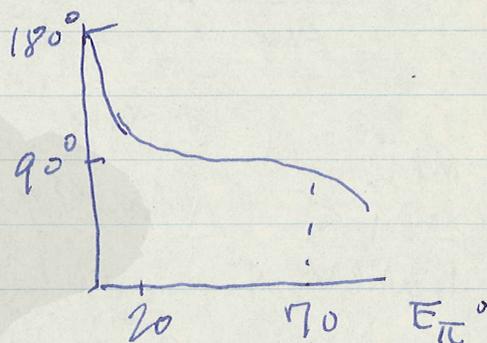
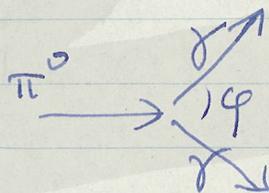
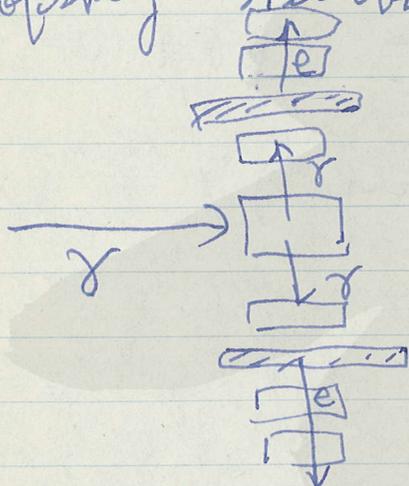


$$\frac{\sigma_H}{\sigma_C} < 0.02$$

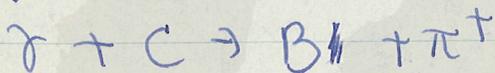
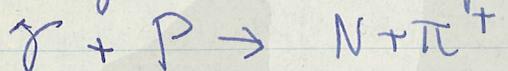
energy threshold

$$\frac{\sigma_H}{\sigma_C} < 0.1$$

Panofsky - Steinberger



Steinberger and Bishop, New York Meeting, 1950



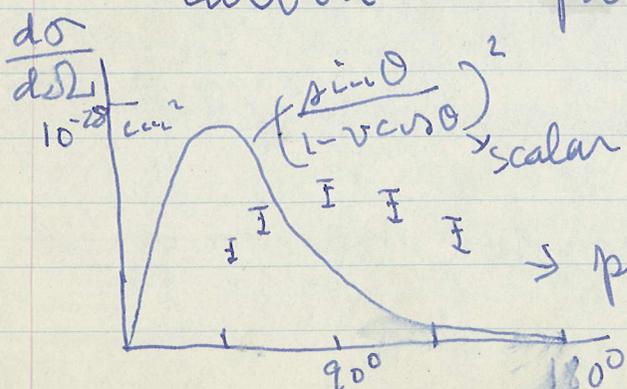
$$\frac{\sigma_H}{\sigma_C} > \frac{1}{12}$$

Physical Rev. 78 (1950),

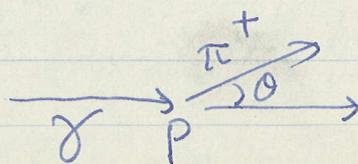
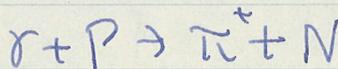
494.

carbon

paraffin

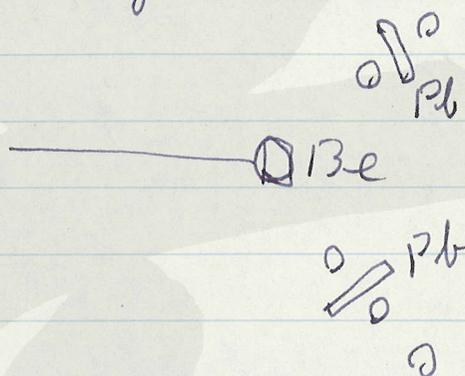


$$h\nu = 250 \text{ MeV}$$



Evidence for the Production of
Neutral Mesons by Photons
by J. Steinberger, W.K.H. Panofsky
and J. Steller.

Synchrotron X-rays of 330 MeV maximum
energy strike a target of Be. Simultaneous
pulses (resolving time 10^{-7} sec.) are recorded
in the outer four crystals, but none in
two near the target.



$$\sigma_{H, \pi^0} / \sigma_{C, \pi^0} = 0.12 \pm 0.03$$

$$\sigma_{H\pi^+} / \sigma_{C\pi^+} \cong 0.55$$

partly

This might be due to the fact that only
protons contribute to π^+ production and that
partly to the exclusion principle, which
is more effective for $\sigma_{C\pi^+}$ than for $\sigma_{C\pi^0}$

$$\sigma_{Be} = 7.5 \times 10^{-28} \text{ cm}^2$$

$$\sigma_C = 10 \times 10^{-28}$$

$$\sigma_H = 1.3 \times 10^{-28}$$

Hydrogen cross-section is approximately
the same as that for π^+ production, these

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for Carbon and Beryllium are somewhat larger.

γ -decay of neutral mesons

$$\pi^0 \rightarrow P + (P') \xrightarrow{g} P'' + \gamma + (P') \xrightarrow{e} \gamma + \gamma'$$

$$\left(\frac{mc^2}{\hbar}\right) \left(\frac{g^2}{\hbar c}\right) \left(\frac{e^2}{\hbar c}\right)^2 \xrightarrow{e} P + \gamma + (P'') \xrightarrow{e} \gamma + \gamma'$$

$$\sim 10^{+23} 10^{-1} \cdot 10^{-4} \sim 10^{+18}$$

Selection rule:

C.N. Yang, Phys. Rev. 77 (1950), 242, 722

Relativistic Calculations on
 γ - $P \rightarrow N$ - π^+ -reaction

Z. Koba, T. Kotani and S. Nakai,
 Prog. Theor. Phys. ~~10~~ 5 (1950), 137.
 ps. scalar theory:

$$\sigma_- = 2.0 \left(\frac{e^2}{\hbar c}\right) \left(\frac{f^2}{\hbar c}\right) \left(\frac{\hbar}{mc}\right)^2 \times 10^{-2}$$

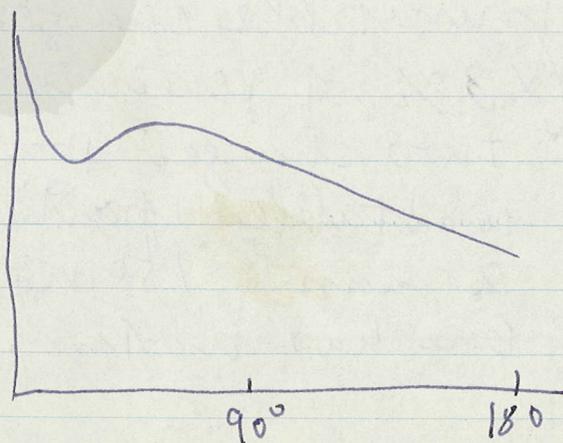
$$\sigma_+ \sim 1.5 \text{ "}$$

For energy of γ -ray = $2mc^2$

$$\sigma_- / \sigma_+ = 1.4$$

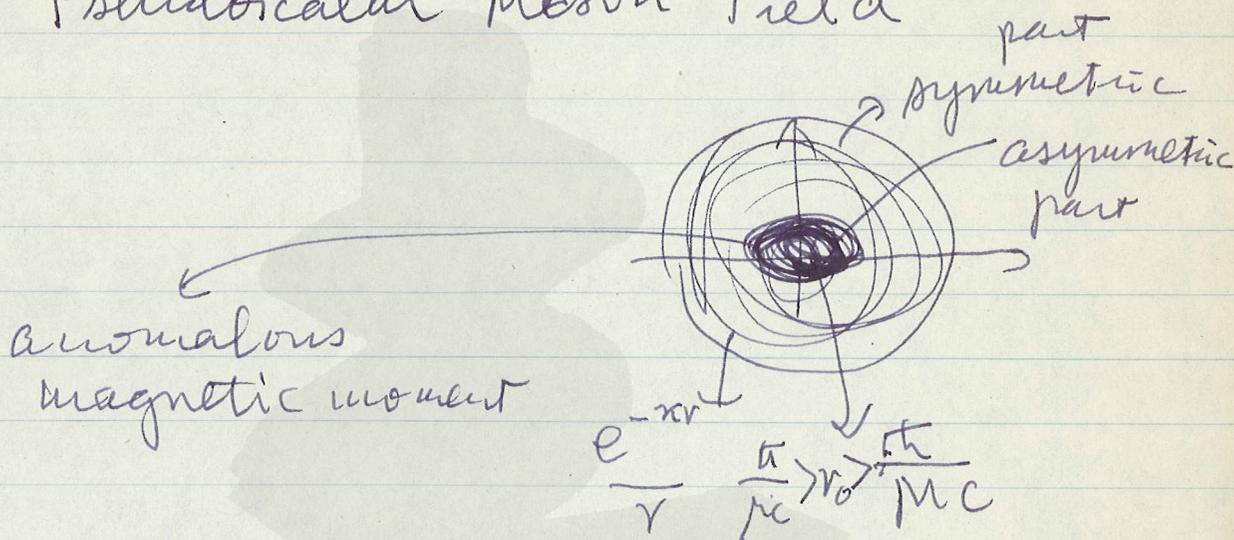
For $f^2 / \hbar c \sim 0.1$ $\sigma \sim 10^{-29} \text{ cm}^2$

Scalar theory
 $\sin^2 \theta$ -distribution
 near threshold



Nuclear Potential in Pseudoscalar Meson Field

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K. A. Brueckner
Production of Mesons by Photons
(UCRL-597)

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Feshbach and Lax, Phys. Rev. 76 (1949), 134.

Foldy, *ibid.* 372

Brueckner and Goldberger, *ibid.* 72 1725.

$$\frac{\sigma_{p+}}{\sigma_{-}} = \left(\frac{e \frac{\vec{v}_p \cdot \vec{A}}{1 - v_p/c \cos \theta_p}}{-e \frac{\vec{v}_p \cdot \vec{A}}{1 - v_p/c \cos \theta_p} + e \frac{\vec{v}_p \cdot \vec{A}}{1 - v_p/c \cos \theta_p}} \right)$$

$$= \left[1 - \frac{g_0}{Mc^2} \left(1 - \frac{v_p}{c} \cos \theta_p \right) \right]^2$$

$$g_0: \text{meson energy} = \frac{mc^2}{\sqrt{1 - v_p^2/c^2}}$$

θ_p : angle between meson and photon

effect of magnetic moments.

$$\frac{\sigma_{+}}{\sigma_{-}} = \left(1 - \frac{\gamma_p - \gamma_N}{\gamma_p + \gamma_N} \frac{g_0}{Mc^2} \left(1 - \frac{v_p}{c} \cos \theta_p \right) \right)^2$$

$$= \left[1 - .20 \frac{g_0}{Mc^2} \left(1 - \frac{v_p}{c} \cos \theta_p \right) \right]^2$$

Angular distribution

$\frac{1}{(1 - \beta \cos \theta)^2}$ is inclined to the forward

direction for $\beta \sim \beta_p$ and near symmetric with respect to 90° for $\beta \sim \beta_M$.

A. Scalar Meson

predominance of $\mu \cdot \gamma$ (electric dipole) interaction.

B. Pseudoscalar Meson

Roughly isotropic distribution due to magnetic moment of the nucleon coupled to the photon. This is because the electric dipole interaction between the meson and the photon is small, the mesons are strongly bound to a distance small compared with $\frac{h}{mc}$.

Preliminary experimental results indicate that the spectrum of mesons observed is roughly isotropic.

Vector and Pseudovector mesons show a strong asymmetry in angular distribution

Stationary state of two Nucleon
 System in Rel. Theory.

$$H \Psi = E \Psi \quad H = H_1 + H_2 + H_{12}$$

$$\Psi (\mathbf{r}_1^{(1)} \sigma_3^{(1)} \rho_3^{(1)} ; \mathbf{r}_2^{(2)} \sigma_3^{(2)} \rho_3^{(2)} ; n_R)$$

$$H_1 = -(\alpha p + Mc^2 \beta)$$

$$H_2 = \sum_R n_R k_0 \tau c$$

$$H_{12} = \sqrt{4\pi} f_1 (p_2^{(1)} U(\mathbf{r}_1) + p_2^{(2)} U(\mathbf{r}_2))$$

$$+ \sqrt{4\pi} f_2 \frac{1}{\kappa} \left(\sigma^{(1)} \cdot \text{grad } U(\mathbf{r}_1) \right)$$

$$+ \sigma^{(2)} \cdot \text{grad } U(\mathbf{r}_2) + c \rho_1^{(1)} U^+(\mathbf{r}_1) + c \rho_2^{(2)} U^+(\mathbf{r}_2)$$

$$+ \frac{1}{2} \left(\frac{f_2}{\kappa} \right)^2 \rho_1^{(1)} \rho_2^{(2)} \delta(\mathbf{r}_1 - \mathbf{r}_2)$$

$$\gamma_1 \gamma_2 \gamma_3 \gamma_4 \begin{pmatrix} \gamma_1 \\ \gamma_2 \\ \gamma_3 \\ \gamma_4 \end{pmatrix} = \begin{pmatrix} -\gamma_2 \gamma_3 \gamma_4 \\ +\gamma_1 \gamma_3 \gamma_4 \\ -\gamma_1 \gamma_2 \gamma_4 \\ \gamma_1 \gamma_2 \gamma_3 \end{pmatrix}$$

$$= - \begin{pmatrix} \gamma_2 \gamma_3 \gamma_4 \\ -\gamma_1 \gamma_3 \gamma_4 \\ \gamma_1 \gamma_2 \gamma_4 \\ -\gamma_1 \gamma_2 \gamma_3 \end{pmatrix}$$

Nuclear Forces, continued.
Effect of High Order Processes
Strong Coupling Theory
Extended source

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Scattering of Mesons
by Nucleons

Production of Mesons by
Nucleon-Nucleon Collision

Nuclear β -Decay and μ - β -Decay