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II. Non-local Field
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Nonlocal Field Theory

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N.L.F. (1)

Last time I discussed the S-matrix theory from a more general standpoint than usual according to Heisenberg. As I mentioned, however, Heisenberg's approach was not ~~different~~ much different from the orthodox field theory that many of the important features of present field theory which was ~~successful~~ enabled as successfully to account for great variety of phenomena in terms of the virtual creation ~~and~~ annihilation of particles such as photons, neutrons, pairs of electrons, pairs of nucleons and so on. Moreover, only the formal general considerations of the structure of S-matrix alone give us no hint to the solution of the second important problem of mass-spectrum of elementary particles, and its relation to ~~mass~~ ^{mass} ~~is~~ ^{is} ~~not~~ ^{not} ~~clear~~ ^{clear} ~~at~~ ^{at} ~~all~~ ^{all} ~~in~~ ⁱⁿ ~~the~~ ^{the} ~~present~~ ^{present} ~~theory~~ ^{theory}.

the first problem of convergence.
~~In connection with~~ ~~for~~ the second problem,

I discussed the non-local inter-
between local field. It is clear,
however, that such a theory cannot
claim to be a fundamental field
theory because ~~it~~ there is no
principle in it to ~~determine~~
~~the~~ form function ^{whose} from all
other except for the convergence
criterion. Nor does it suggest
any connection to the second
problem of mass spectrum.

It seems to me that ^{the} only
conceivable way of ~~dealing~~ coping
with the two fundamental problems
as mutually related is to change
the concept of field itself.

Problem, Mass Spectrum as Eigenvalue Problem
~~in Lagrangian formalism~~
 in N. h. F. T.

Since the ~~ordinary~~^A local field, say a scalar or pseudoscalar field, $\varphi(x_\mu)$ is ordinarily supposed to satisfy a field equation of Klein-Gordon type in which

$$\left(\frac{\partial^2}{\partial x_\mu \partial x_\mu} - m^2 \right) \varphi(x_\mu) = 0$$

in which m is an arbitrary parameter, ~~and there is no rule~~ which give such a field is ~~too~~^{whatsoever} simple to give any information on further restriction ~~to~~ the possible (discrete) values of the mass m . For that reason, ~~we~~

we had better start from a more complicated ~~of~~ classical field which depends on two sets of space-time points parameters

$x_\mu, x_{\mu'} : \varphi(x_\mu, x_{\mu'})$
 or $\varphi(x_{\mu'}, x_{\mu''}) = \varphi(x_\mu, x_{\mu'})$
 where $x_\mu = \frac{1}{2}(\alpha_{\mu'} + \alpha_{\mu''})$ η
 $x_{\mu'} = \alpha_{\mu'} - \alpha_{\mu''}$

This field is supposed to satisfy the field equation:

$$F(x'_\mu, x''_\mu, \frac{\partial}{\partial x'_\mu}, \frac{\partial}{\partial x''_\mu}) \varphi(x'_\mu, x''_\mu) = 0$$

where F is an invariant function of $x'_\mu, x''_\mu, \frac{\partial}{\partial x'_\mu}, \frac{\partial}{\partial x''_\mu}$.

In order that the field eq. is invariant with respect to ^{an arb.} Lorentz transformation, F must be independent of x_μ , so that we may write

$$F(\frac{\partial}{\partial x_\mu}, x_\mu, \frac{\partial}{\partial x_\mu}) \varphi(x_\mu, x_\mu) = 0$$

In general, one may also write

~~φ can be~~

$$\varphi(x_\mu, x_\mu) = \int u(k_\mu, x_\mu) e^{ik_\mu x_\mu} (dK)^4$$

$$F(ik_\mu, x_\mu, \frac{\partial}{\partial x_\mu}) u(k_\mu, x_\mu) = 0$$

Now F is a function of $k_\mu k_\mu, x_\mu x_\mu, \frac{\partial^2}{\partial x_\mu \partial x_\mu}, k_\mu x_\mu, k_\mu \frac{\partial}{\partial x_\mu}, x_\mu \frac{\partial}{\partial x_\mu}$

let us take the simplest case in which Q03

suppose that one ^{may} write

$$F = k_\mu k_\mu + f(r_\mu r_\mu, \frac{\partial^2}{\partial x_\mu \partial x_\mu}, k_\mu x_\mu, k_\mu \frac{\partial}{\partial x_\mu}, r_\mu \frac{\partial}{\partial r_\mu})$$

so that the eigenvalue problem

$$f(r_\mu r_\mu, \frac{\partial^2}{\partial x_\mu \partial x_\mu}, k_\mu x_\mu, k_\mu \frac{\partial}{\partial x_\mu}, r_\mu \frac{\partial}{\partial r_\mu}) \times u(k_\mu, r_\mu) = m^2 u(k_\mu, r_\mu)$$

for the operator f can be solved with a certain eigenvalue

m_{n, k_n}^2
 and eigenfunction $u_n(k_\mu, r_\mu)$.

Then we have

$$(k_\mu k_\mu + m_{n, k_n}^2) u_n(k_\mu, r_\mu) = 0$$

Thus m_{n, k_n} can be interpreted as the mass of the particle, ^{due to} being determined by its internal motion, provided that m_{n, k_n}^2 is independent of k_μ . ^{the direction of} In other words, either if F has the form

$$F = k_\mu k_\mu + f(r_\mu r_\mu, \frac{\partial^2}{\partial x_\mu \partial x_\mu}, r_\mu \frac{\partial}{\partial r_\mu})$$

or m_n^2 is actually independent of k_μ the direction of k_μ due to some supplementary imposed on u_n .

The above argument suggests us that

$$F^{(int)} = F^{(ext)} + F^{(int)}$$

$$F^{(ext)} \left(-\frac{\partial}{\partial x_\mu} \right) = F^{(int)} \left(\frac{\partial^2}{\partial x_\mu \partial x_\mu} \right)$$

$$F^{(int)} \left(r_\mu r_\mu, \frac{\partial^2}{\partial v_\mu \partial v_\mu}, \frac{\partial^2}{\partial x_\mu \partial x_\mu} \right)$$

$$\left(\frac{\partial}{\partial x_\mu}, r_\mu \frac{\partial}{\partial v_\mu} \right)$$

$$\varphi(x_\mu, r_\mu) = \sum_n \varphi_n^{(ext)}(x_\mu) \varphi_n^{(int)}(r_\mu)$$

where

$$F^{(int)} \varphi_n^{(int)}(r_\mu) = m_n^2 \varphi_n^{(int)}(r_\mu) \quad (1)$$

$$(F^{(ext)} + m_n^2) \varphi_n^{(ext)}(x_\mu) = 0$$

Thus, if we take

$$F^{(ext)} = -\frac{\partial^2}{\partial x_\mu \partial x_\mu}$$

m_n becomes the mass of the particle.
 by solving the eigenvalue problem (1),
 we have the mass spectrum for scalar
 elementary particles with ~~or pseudoscalar~~
 0 or integer spin.

Example 1. (Oscillator Model)

$$F^{(int)} = -\frac{\partial^2}{\partial x_\mu \partial x_\mu} + \frac{1}{\lambda^2} \gamma_\mu \gamma_\mu$$

$$H = -p^2 + \frac{1}{\lambda^2} q^2 = \frac{1}{\lambda^2} (P^2 + Q^2)$$

$$[q, p] = -1. \quad P = \lambda p, \quad Q = \frac{q}{\lambda}$$

$$[Q, P] = -1.$$

$$E_n = \frac{1}{\lambda^2} (2n+1)$$

$$[P + Q, P - Q] =$$

$$\frac{1}{\sqrt{2}}(Q + P), \frac{1}{\sqrt{2}}(Q - P) = -\frac{1}{2}[Q, P] + \frac{1}{2}[P, Q]$$

$$= 1$$

$$b = \frac{1}{\sqrt{2}}(Q + P) \quad b^* = \frac{1}{\sqrt{2}}(Q - P)$$

$$b^* b = \frac{1}{2}(Q^2 - P^2) + \frac{1}{2}[Q, P]$$

$$b b^* = \frac{1}{2}(Q^2 - P^2) - \frac{1}{2}[Q, P]$$

$$Q^2 - P^2 = b^* b + b b^* = 2n + 1$$

$$\frac{\partial}{\partial x'_\mu} = \frac{1}{2} \frac{\partial}{\partial x_\mu} + \frac{\partial}{\partial r_\mu}$$

$$\frac{\partial}{\partial x''_\mu} = \frac{1}{2} \frac{\partial}{\partial x_\mu} - \frac{\partial}{\partial r_\mu}$$

$$\frac{\partial^2}{\partial x'_\mu \partial x'_\nu} = \frac{1}{4} \frac{\partial^2}{\partial x_\mu \partial x_\nu} + \frac{\partial^2}{\partial r_\mu \partial r_\nu} + \frac{\partial}{\partial x_\mu \partial r_\nu}$$

$$\frac{\partial^2}{\partial x''_\mu \partial x''_\nu} = \frac{1}{4} \frac{\partial^2}{\partial x_\mu \partial x_\nu} + \frac{\partial^2}{\partial r_\mu \partial r_\nu} - \frac{\partial}{\partial x_\mu \partial r_\nu}$$

$$\frac{\partial^2}{\partial x'_\mu \partial x'_\nu} + \frac{\partial^2}{\partial x''_\mu \partial x''_\nu} = \frac{1}{2} \frac{\partial^2}{\partial x_\mu \partial x_\nu} + 2 \frac{\partial^2}{\partial r_\mu \partial r_\nu}$$

$$F^{(ex)} + F^{(in)} = -\frac{\partial^2}{2 \partial x'_\mu \partial x'_\mu} - \frac{\partial^2}{2 \partial x''_\mu \partial x''_\mu} + \frac{2}{\lambda^2} (x'_\mu - x''_\mu)(x'_\mu - x''_\mu)$$

$$= -\frac{1}{2} \frac{\partial^2}{\partial x_\mu \partial x_\mu} + 2 \left(\frac{\partial^2}{\partial r_\mu \partial r_\mu} + \frac{1}{\lambda^2} r_\mu r_\mu \right)$$

$$\frac{2}{\lambda^2} \left\{ (2n_1 + 1) + (2n_2 + 1) + (2n_3 + 1) - (2n_0 + 1) \right\}$$

$$m_{n_1, n_2, n_3, n_0}^2 = \frac{8}{\lambda^2} \{ n_1 + n_2 + n_3 - n_0 + 1 \}$$

$$n_1 = n_2 = n_3 = n_0 = 0$$

$$m_0^2 = \frac{8}{\lambda^2}$$

$$\frac{hc}{\lambda} = \frac{\sqrt{8}}{\lambda} \mu c$$

$$\lambda = \frac{\sqrt{8} \hbar}{\mu c}$$

$$\mu = 2 \frac{e^2}{e^2} m_e$$

$$\lambda = \sqrt{8} \cdot \frac{1}{2} \frac{e^2}{m_e c^2} = \sqrt{2} \frac{e^2}{m_e c^2}$$

$$\lambda' = \frac{\lambda}{\sqrt{2}} = \frac{e^2}{m_e c^2}$$

$$\Gamma^{(e^+)} + \Gamma^{(e^-)} = \frac{1}{2} \frac{\partial^2}{\partial x_\mu \partial x_\mu} + 2 \frac{\partial^2}{r r_\mu \partial r_\mu}$$

$$+ \frac{1}{\lambda^4} \gamma_\mu \gamma_\mu$$

$$\lambda' = \frac{e^2}{m_e c^2}$$

$$m_1 = \sqrt{2} m_0$$

$$m_2 = \sqrt{3} m_0$$

$$m_3 = 2 m_0$$

$$\vdots$$

$$\left. \begin{array}{l} m_1 = \sqrt{2} m_0 \\ m_2 = \sqrt{3} m_0 \\ m_3 = 2 m_0 \\ \vdots \end{array} \right\} e^2$$

The trouble with this ~~and any other~~ model ~~is~~ is that m_{eff}^2 may not be positive, but restricted to positive eigenvalues, but can be negative, if $n_0 > n_1 + n_2 + n_3 + 1$.

It is obviously possible to restrict m_{eff}^2 to positive values by taking for example:

$$F^{(\text{in})} = \left(\frac{\partial^2}{\partial x_{\mu} \partial x_{\mu}} + \frac{1}{\lambda^2} r_{\mu} r_{\mu} \right)^2,$$

but this does not remove the more serious difficulty of degeneracy. Namely, for any $n = n_1 + n_2 + n_3$ (not)

there are infinitely many combinations of n_1, n_2, n_3, n_0 .

In order to avoid such an ~~unpleasant~~ undesirable complication, one has to assume a certain supplement.

condition of the type:

$$\frac{\partial^2}{\partial x_{\mu} \partial x_{\mu}} \varphi(x_{\mu}, r_{\mu}) = 0$$

N.L.F. (5)

In particular, if we take a solution of the form

$$\varphi = e^{ik_\mu x_\mu} \varphi^{(in)}(r_\mu)$$

$$k_\mu \frac{\partial}{\partial r_\mu} \varphi^{(in)} = 0$$

or another simple case:

$$\left(\frac{\partial^2}{\partial x_\mu \partial x_\mu} + \frac{\partial}{\partial r_\mu} \right) \varphi = 0$$

$$\text{or } \frac{\partial}{\partial x_\mu} \left(\frac{\partial}{\partial r_\mu} + \frac{1}{r_\mu} \right) \varphi = 0$$

$$\varphi = e^{ik_\mu x_\mu} \varphi^{(in)}(r_\mu)$$

$$k_\mu \left(\frac{\partial}{\partial r_\mu} + \frac{1}{r_\mu} \right) \varphi^{(in)} = 0$$

$$\text{or } k_\mu l_\mu \varphi^{(in)} = 0$$

If we carry out a ~~trans~~ coordinate transformation

with the coefficients

$$l'_\mu = a_{\nu\lambda} b_\nu$$

$$(a_{\mu\nu}) = \begin{pmatrix} 1 + \left(\frac{k_1}{\kappa}\right)^2 & \frac{k_1 k_2}{\kappa^2} & \frac{k_1 k_3}{\kappa^2} & i k_1 / \kappa \\ \frac{k_1 k_2}{\kappa^2} & 1 + \left(\frac{k_2}{\kappa}\right)^2 & \frac{k_2 k_3}{\kappa^2} & i k_2 / \kappa \\ k_1 k_3 / \kappa^2 & k_2 k_3 / \kappa^2 & 1 + \left(\frac{k_3}{\kappa}\right)^2 & i k_3 / \kappa \\ i k_1 / \kappa & i k_2 / \kappa & i k_3 / \kappa & i k_4 / \kappa \end{pmatrix}$$

$$\kappa = \sqrt{\kappa^2 + k_4^2}$$

$$\begin{aligned} & \left(1 + \left(\frac{k_1}{\kappa}\right)^2\right) \frac{k_1 k_2}{\kappa^2} + \left(\frac{k_1 k_2}{\kappa^2}\right) \left(1 + \left(\frac{k_2}{\kappa}\right)^2\right) \\ & + \frac{k_1 k_3}{\kappa^2} \frac{k_2 k_3}{\kappa^2} - \frac{k_1 k_2}{\kappa^2} \\ & = \frac{k_1 k_2}{\kappa^2} \left\{ 2 + \frac{k_1^2 + k_2^2 + k_3^2}{\kappa^2} \right\} \\ & \quad - \frac{k_1 k_2}{\kappa^2} \left\{ + k_1^2 + k_2^2 + k_3^2 \right\} \\ & = \frac{k_1 k_2 \left(2\kappa^2 + 2\kappa(k_1^2 + k_2^2 + k_3^2 + \kappa^2) - k_1^2 k_2^2 \right)}{\kappa^4} \\ & = \frac{k_1 k_2 \left(\kappa + \sqrt{k_1^2 + k_2^2 + k_3^2 + \kappa^2} \right)^2}{\kappa^4} \end{aligned}$$

$$\begin{aligned} & - \frac{k_1 k_2}{\kappa^2} \\ & = \frac{k_1 k_2}{\kappa^2} - \frac{k_1 k_2}{\kappa^2} = 0 \end{aligned}$$

N. h. F. ①

$$\left(1 - \left(\frac{k_1}{\kappa}\right)^2\right)^2 + \left(\frac{k_1 k_2}{\kappa^2}\right)^2 + \left(\frac{k_1 k_3}{\kappa^2}\right)^2$$

$$- \frac{k_1^2}{\kappa^2} = 1 + 2 \frac{k_1^2}{\kappa^2} + \frac{k_1^2 (k_1^2 + k_2^2 + k_3^2)}{\kappa^4}$$

$$- \frac{k_1^2}{\kappa^2} = \frac{1}{\kappa^4} \left\{ \kappa^4 + \cancel{2k_1^2 \kappa^2} + \cancel{k_1^2 k_2^2} + \cancel{k_1^2 k_3^2} \right.$$

$$\left. + k_1^2 (k_1^2 + k_2^2 + k_3^2) - k_1^2 (\kappa + \sqrt{\quad})^2 \right\}$$

$$= \frac{1}{\kappa^4} \left\{ \kappa^4 + 2\kappa^2 k_1^2 - k_1^2 \kappa^2 \right.$$

$$\left. - 2k_1^2 \kappa \sqrt{k_1^2 + k_2^2 + k_3^2 + \kappa^2} - k_1^2 \kappa^2 \right\}$$

$$= \frac{1}{\kappa^4} \left\{ \kappa^2 (\kappa + \sqrt{\quad})^2 + 2k_1^2 \kappa (\kappa + \sqrt{\quad}) \right.$$

$$\left. - 2k_1^2 \kappa^2 - 2k_1^2 \kappa \sqrt{\quad} \right\}$$

$$= 1$$

One may alternatively subtract from $\frac{\partial^2}{\partial x_\mu \partial x_\mu} F^{(in)}$, the term $1 + \left(\frac{\partial^2}{\partial x_\mu \partial x_\mu}\right)^2 \frac{1}{\kappa^2} F^{(in)}$ in F after multiplying the first terms by $-\frac{\partial^2}{\partial x_\mu \partial x_\mu}$ pair second

Then, we have

$$F^{(in)} = \frac{1}{\lambda^2} (2b_{\mu}^{*'} b_{\mu}' + 1)$$

$$\frac{1}{\lambda^2} (2b_{\mu}^{*'} b_{\mu}' + 2) \varphi^{(in)} = 0$$

with the supplementary cond.

$$b_4' \varphi^{(in)} = 0,$$

$$E_n = \frac{Q}{\lambda^2} (n_1' + n_2' + n_3' + 1)$$

$$= (137 m_e)^2 (n_1' + n_2' + n_3' + 1)$$

$$m_n^2 = (294 m_e)^2 (n_1' + n_2' + n_3' + 1)$$

Thus, in the rest system, the eigenfn

$$\varphi_{0, k_{\mu}}^{(in)} \text{ has the form } \varphi_{0, k_{\mu}}^{(in)} = \frac{1}{\sqrt{\pi \cdot \lambda}} e^{-\sqrt{q_1'^2 + q_2'^2 + q_3'^2} / 2\lambda^2}$$

$$= \frac{1}{\pi \lambda^2} e^{-\sqrt{q_1'^2 + q_2'^2 + q_3'^2} / 2\lambda^2}$$

$$\therefore \frac{d^2}{dq^2} e^{-aq^2} = \frac{d}{dq} (-2aq e^{-aq^2})$$

$$= -2a + 4a^2 q^2 e^{-aq^2}$$

$$\left(-\frac{d^2}{dq^2} + 4a^2 q^2\right) e^{-aq^2} = 2a e^{-aq^2}$$

$$a = \frac{1}{2}. \quad \left(-\frac{d^2}{dq^2} + q^2\right) e^{-q^2/2} = e^{-q^2/2}$$

N.L.F. (8)

The corresponding total Schrödinger fn
is

$$\varphi_0(x_\mu, r_\mu) = \int \frac{1}{\sqrt{(2\pi)^3}} a_{\mathbf{k}} e^{i\mathbf{k}\cdot\mathbf{x}_\mu} \frac{1}{\pi\lambda^2} e^{-\frac{(r_1'^2 + r_2'^2 + r_3'^2 + r_0'^2)}{2\lambda^2}} (d\mathbf{k})^3$$

$$= \iint \frac{1}{\sqrt{(2\pi_0(2\pi))^3}} (a_{\mathbf{k}} e^{i\mathbf{k}\cdot\mathbf{x}_\mu} + a_{\mathbf{k}}^* e^{-i\mathbf{k}\cdot\mathbf{x}_\mu})$$

$$k_4 = i k_0 \quad k_0 = +\sqrt{k^2 + k_0^2}$$

$$\left[\frac{1}{\pi\lambda^2} e^{-\frac{(r_1'^2 + r_2'^2 + r_3'^2 + r_0'^2)}{2\lambda^2}} (d\mathbf{k})^3 \right]$$

$$[a_{\mathbf{k}}, a_{\mathbf{k}'}^*] = \delta(\mathbf{k}, \mathbf{k}')$$

$$\int (\varphi_0(x_\mu, r_\mu))^2 = \int \varphi_0(x_\mu, r_\mu) \varphi_0(x_\mu, -r_\mu) \frac{1}{\pi\lambda^2} (d\mathbf{x})^3 (d\mathbf{r})^4$$

$$= \int (a_{\mathbf{k}}^* a_{\mathbf{k}} + a_{\mathbf{k}} a_{\mathbf{k}}^*) (d\mathbf{k})^3 + \text{fluctuation}$$

$$\int_{-\infty}^{+\infty} e^{-2aq^2} dq = \sqrt{\frac{\pi}{2a}} = \sqrt{\frac{\pi \cdot 2\lambda^2}{2}} = \sqrt{\pi} \cdot \lambda$$

$$\iint e^{-2a(x^2+y^2)} dx dy = 2\pi \int_0^\infty e^{-2ar^2} r dr$$

$$= \pi \int_0^\infty e^{-2ax} dx = \frac{\pi}{2a}$$

There are other excited states
which correspond to

$$n = n_1 + n_2 + n_3 = 1 \quad n_0 = 0$$

$$= 2 \quad n_0 = 0$$

$$= 3 \quad n_0 = 0$$

etc.

All of them are degenerate

$n=1$. triple ~~q~~ (P)

$n=2$. $(2,0,0)$ $(0,2,0)$ $(0,0,2)$

$(1,1,0)$ $(1,0,1)$ $(0,1,1)$

6-fold

(D, S)

$n=3$,

$n=4$

The total Schrödinger ψ is

$$\psi(x, r) = \sum_n \varphi_n(x, r),$$

the procedure of
the quantization being similar
to the one for $n=0$.

General eigenvalue problem

$$F \left(\frac{\partial^2}{\partial x_\mu \partial x_\mu}, \frac{\partial^2}{\partial r_\mu \partial r_\mu}, \gamma_\mu \gamma_\mu, \frac{\partial^2}{\partial x_\mu \partial r_\mu}, \gamma_\mu \frac{\partial}{\partial x_\mu}, \gamma_\mu \frac{\partial}{\partial r_\mu} \right) \varphi(x_\mu, r_\mu) = 0$$

Among 6 operators $\frac{\partial^2}{\partial x_\mu \partial x_\mu}$ commutes with all others, so that one can take the representation in which $\frac{\partial^2}{\partial x_\mu \partial x_\mu}$ expand φ into series of eigenfunctions of satisfying

$$\frac{\partial^2 \varphi_n(x_\mu)}{\partial x_\mu \partial x_\mu} = m_n^2 \varphi_n(x_\mu)$$

then are coefficients $\varphi_{n,k}(r_\mu)$ of expansion

$$\varphi(x_\mu, r_\mu) = \sum \varphi_{n,k}(x_\mu) \varphi_{n,k}(r_\mu)$$

where must satisfy

$$F(m_n^2, \frac{\partial^2}{\partial r_\mu \partial r_\mu}, \gamma_\mu \gamma_\mu, \frac{\partial}{\partial r_\mu} \gamma_\mu \frac{\partial}{\partial r_\mu}, \gamma_\mu \frac{\partial}{\partial r_\mu}) \varphi_{n,k}(r_\mu) = 0$$

Example:

$$m_n^2 \left(-\frac{\partial^2}{\partial r_\mu \partial r_\mu} + \frac{1}{\lambda^2} \gamma_\mu \gamma_\mu \right) + \left(\gamma_\mu \frac{\partial}{\partial r_\mu} \right)^2 + \frac{1}{\lambda^2} (\gamma_\mu \gamma_\mu)^2$$

$$F = m_n^4 -$$

$$= m_n^2 \left\{ m_n^2 - \frac{4}{\lambda^2} (n_1 + n_2 + n_3 + 1) \right\}$$

However, this method is unsatisfactory in that the $\varphi^{(in)}(r_\mu)$ is not sufficiently restricted in $k_\mu r_\mu$ -direction. *

On the other hand, the supplementary condition

$$\frac{\partial}{\partial r_\mu} \left(\frac{\partial}{\partial r_\mu} + \frac{1}{\lambda^2} r_\mu \right) \varphi = 0$$

is compatible with the field eq. because

$$\begin{aligned} & \frac{\partial^2}{\partial r_\mu \partial r_\mu} \left(\frac{1}{\lambda^2} r_\mu \right) \varphi \\ &= \frac{\partial}{\partial r_\mu} \left\{ \frac{1}{\lambda^2} r_\nu \frac{\partial}{\partial r_\mu} \right\} \varphi + \frac{\partial}{\partial r_\mu} \frac{1}{\lambda^2} \delta_{\mu\nu} \varphi \\ &= \frac{1}{\lambda^2} r_\nu \cdot \frac{\partial^2 \varphi}{\partial r_\mu \partial r_\mu} + 2 \frac{1}{\lambda^2} \delta_{\mu\nu} \cdot \frac{\partial \varphi}{\partial r_\mu} \\ & \quad + \frac{1}{\lambda^4} r_\mu r_\mu \left(\frac{\partial}{\partial r_\nu} \right) \varphi \\ &= \frac{\partial}{\partial r_\nu} \frac{1}{\lambda^4} r_\mu r_\mu \varphi \\ &= \frac{2}{\lambda^4} \delta_{\mu\nu} r_\nu \varphi \end{aligned}$$

$$\begin{aligned} & \left(-\frac{\partial^2}{\partial r_\mu \partial r_\mu} + \frac{1}{\lambda^4} r_\mu r_\mu \right) \left(\frac{\partial}{\partial r_\nu} + \frac{1}{\lambda^2} r_\nu \right) \varphi \\ &= \left(\frac{\partial}{\partial r_\nu} + \frac{1}{\lambda^2} r_\nu \right) \left(-\frac{\partial^2}{\partial r_\mu \partial r_\mu} + \frac{1}{\lambda^4} r_\mu r_\mu \right) \varphi - 2 \frac{1}{\lambda^2} \delta_{\mu\nu} \left(\frac{\partial}{\partial r_\nu} + \frac{1}{\lambda^2} r_\nu \right) \varphi \end{aligned}$$

* In order to get rid of this difficulty
one can further impose a S.C.

$$k_{\mu} r_{\mu} \varphi = 0.$$

or

$$r_{\mu} \frac{\partial}{\partial x_{\mu}} \varphi = 0$$

This condition is compatible with
the field equation, because Γ can
be written in terms of

$$r_1, r_2, r_3 \quad \frac{\partial}{\partial r_1}, \frac{\partial}{\partial r_2}, \frac{\partial}{\partial r_3}$$

alone and independent of

$$r_1' = -\frac{\hbar r_{\mu}}{m} \quad \text{and} \quad \frac{\partial}{\partial r_0}$$

This gives rise to the factor

$$\varphi \propto \delta(k_{\mu} r_{\mu})$$

Another possibility is to assume
a S.C.

$$\left\{ \left(\frac{\partial^2}{\partial x_{\mu} \partial x_{\mu}} \right)^2 - \frac{1}{\lambda^2} \left(r_{\mu} \frac{\partial}{\partial x_{\mu}} \right)^2 \right\} \varphi = 0$$

$$\frac{\partial^2}{\lambda^2} \frac{\partial^2 \varphi}{\partial x_{\mu} \partial x_{\mu}} \quad ?$$

N. h. F. (9)

hagen formalism

$$\begin{aligned}
 & \int \left(\frac{\partial \varphi(x', x'')}{\partial x'_\mu} \frac{\partial \varphi(x', x'')}{\partial x''_\mu} + \frac{\partial \varphi(x', x'')}{\partial x''_\mu} \frac{\partial \varphi(x', x'')}{\partial x'_\mu} \right) (dx')^4 (dx'')^4 \\
 & + \int \left(\frac{1}{\lambda^4} \frac{\partial \varphi(x', x'')}{\partial x'_\mu} \frac{\partial \varphi(x', x'')}{\partial x''_\mu} \right) (dx')^4 (dx'')^4
 \end{aligned}$$

$\delta \bar{L} = 0$

$$- \frac{\partial^2 \varphi}{\partial x'_\mu \partial x'_\mu} - \frac{\partial^2 \varphi}{\partial x''_\mu \partial x''_\mu} + \frac{1}{\lambda^4} (x'_\mu - x''_\mu)^2 \varphi = 0$$

$$\left(-\frac{1}{2} \frac{\partial^2}{\partial x'_\mu \partial x'_\mu} - 2 \frac{\partial^2}{\partial x''_\mu \partial x''_\mu} + \frac{2}{\lambda^4} r_\mu r_\mu \right) \varphi = 0$$

$$\left(-\frac{\partial^2}{\partial x'_\mu \partial x'_\mu} + m_n^2 \right) \varphi_n^{out}(x'_\mu) = 0$$

$$\left(-\frac{\partial^2}{\partial x''_\mu \partial x''_\mu} + \frac{1}{\lambda^4} r_\mu r_\mu \right) \varphi_n^{in}(x''_\mu) = m_n^2 \varphi_n^{in}$$

$$m_n^2 = \frac{8}{\lambda^2} \{ n_1 + n_2 + n_3 + n_0 + 1 \}$$

$$= \frac{\mu c}{\hbar} \{ n_1 + n_2 + n_3 + n_0 + 1 \} \quad \text{for } n_1 = n_2 = n_3 = n_0 = 0$$

$$\text{for } \lambda = \sqrt{2} \frac{e^2}{m c^2} \quad \mu = 274 m_e$$

or $\varphi(x, y) = \varphi(x, -y)$,
 one may write \bar{L} as the trace
 of a non-local operator.

Namely, if we introduce the
 non-local operators

$$(x' | \varphi | x'') \equiv \varphi(x', x'')$$

$$(x' | x_\mu | x'') = x'_\mu \delta(x' - x'')$$

$$(x' | p_\mu | x'') = i \delta'_\mu(x' - x'')$$

$$(x' | [x_\mu, p_\nu] | x'') = \int x'_\mu \delta(x' - x''') i \delta'_\nu(x'' - x''') dx'''$$

$$= \int i \delta'_\nu(x' - x''') x''_\mu \delta(x'' - x''') dx'''$$

$$= -i \int x'_\mu \frac{\partial}{\partial x''_\nu} \delta(x' - x''') \delta(x'' - x''') dx'''$$

$$= i \int \delta_{\mu\nu} \delta(x' - x''') \delta(x'' - x''') dx'''$$

$$= i \int \delta(x' - x''') x''_\mu \frac{\partial}{\partial x''_\nu} \delta(x'' - x''') dx'''$$

$$= -i \delta_{\mu\nu} \delta(x' - x'')$$

$$\int f(x') \delta'_\mu(x' - x'') dx'$$

$$= - \frac{\partial f(x'')}{\partial x''_\mu}$$

$$\int \delta'_\mu(x'-x'') f(x'') dx''$$

$$= \frac{\partial f(x')}{\partial x'_\mu}$$

$$(x'| [p_\mu, \varphi] | x'') = \int i \delta'_\mu(x'-x''') \varphi(x''', x'') dx'''$$

$$- \int i \varphi(x', x''') \delta'_\mu(x''', x'') dx'''$$

$$= i \frac{\partial \varphi(x', x'')}{\partial x'_\mu} + i \frac{\partial \varphi(x', x'')}{\partial x''_\mu}$$

$$= i \frac{\partial \varphi(x, r)}{\partial x_\mu}$$

$$(x'| [p_\mu, \varphi]_+ | x'') = i \frac{\partial \varphi(x', x'')}{\partial x'_\mu} - i \frac{\partial \varphi(x', x'')}{\partial x''_\mu}$$

$$= 2i \frac{\partial \varphi(x, r)}{\partial r_\mu}$$

$$\text{Trace} \{ [p_\mu, \varphi]_+ [p_\mu, \varphi] \} = - \frac{\int \left(\frac{\partial \varphi(x, r)}{\partial x_\mu} \frac{\partial \varphi(x, r)}{\partial x_\mu} \right)}{(dx)(dr)}$$

$$\left. \begin{aligned} x'_\mu &= x_\mu + \frac{1}{2} r_\mu \\ x''_\mu &= x_\mu - \frac{1}{2} r_\mu \end{aligned} \right\}$$

$$\frac{\partial}{\partial x'_\mu} = \frac{\partial}{\partial x_\mu} + \frac{\partial}{\partial x''_\mu}, \quad \frac{\partial}{\partial r_\mu} = \frac{1}{2} \left(\frac{\partial}{\partial x'_\mu} - \frac{\partial}{\partial x''_\mu} \right)$$

$$\text{Trace } [\rho_\mu, \varphi]_+ [\rho_\mu, \varphi]_+$$

$$= -4 \int \int \frac{\partial \varphi(x, r)}{\partial x_\mu} \frac{\partial \varphi(x, -r)}{\partial x_\mu} dx dr$$

$$\text{Trace } [\alpha_\mu, \varphi] [\alpha_\mu, \varphi]$$

$$= \int \int \alpha_\mu \gamma_\mu \varphi(x, r) \varphi(x, -r) dx dr$$

$$\bar{L} = \text{Trace } \{ [\rho_\mu, \varphi] [\rho_\mu, \varphi]$$

$$+ [\rho_\mu, \varphi]_+ [\rho_\mu, \varphi]_+ + \frac{4}{\lambda^4} [\alpha_\mu, \varphi] [\alpha_\mu, \varphi] \}$$

$$\lambda' = \frac{\lambda}{\sqrt{2}} = \frac{e^2}{m_e c^2}$$

$$\lambda'^4 = \frac{\lambda^4}{4}$$

$$\bar{L} = \text{Trace } \{ [\rho_\mu, \varphi] [\rho_\mu, \varphi]$$

$$+ [\rho_\mu, \varphi]_+ [\rho_\mu, \varphi]_+ + \frac{1}{\lambda^4} [\alpha_\mu, \varphi] [\alpha_\mu, \varphi] \}$$

N. h. T. (11)

If one adopts the unit in which the length is measured in units of $\lambda' = \frac{\hbar}{m_0 c}$, one obtains simply

$$\bar{L} = \text{Trace} \{ [\gamma_\mu, \varphi] [\gamma_\mu, \varphi] + [\gamma'_\mu, \varphi]_+ [\gamma_\mu, \varphi]_+ + [\alpha_\mu, \varphi] [\alpha_\mu, \varphi] \}$$

by multiplying the original \bar{L} by a factor λ'^2 and replacing

$$\begin{aligned} \lambda' \gamma_\mu &\rightarrow \gamma_\mu \\ \frac{\alpha_\mu}{\lambda'} &\rightarrow \alpha_\mu \end{aligned}$$

The supplementary condition will be

$$[\gamma_\mu, \frac{1}{2i} [\gamma_\mu, \varphi]_+ + [\alpha_\mu, \varphi]] = 0$$

is commutative with ρ

The notion of non-local fields is modified a little bit.

Interaction of Non-local Fields

Suppose that two kinds of fields ϕ and ψ interact with each other. Let us take a simple example of interaction between two scalar fields, ~~because~~ ^{since} non-local spinor fields are more difficult to deal with, we postpone the discussion of a more important case of interaction of a spinor and scalar fields.

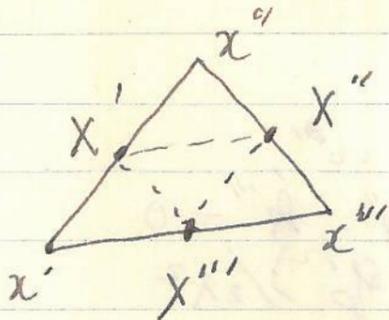
Suppose that the interaction is represented by an additional term

$$L_{int} = \text{Trace } g \cdot \psi^* \phi \psi$$

in the lagrangian, where ϕ is a real scalar field and ψ is a complex scalar field.

$$\text{Trace } g \psi^* \phi \psi = \int (x' | \psi^* | x'') \times (x'' | \phi | x''') (x''' | \psi | x') dx' dx'' dx'''$$

N.L.F (12)



$$\left. \begin{aligned} X' &= \frac{1}{2}(x' + x'') \\ X'' &= \frac{1}{2}(x'' + x''') \\ X''' &= \frac{1}{2}(x''' + x') \end{aligned} \right\}$$

$$\left. \begin{aligned} \gamma' &= x' - x'' = 2(X''' - X'') \\ \gamma'' &= x'' - x''' = 2(X' - X''') \\ \gamma''' &= x''' - x' = 2(X'' - X') \end{aligned} \right\}$$

$$\varphi(x) = \sum_{n, k} \varphi_{n, k}^{ex}(x) \varphi_{n, k}^{in}(r)$$

$$\sum_{\substack{n', n'', n''' \\ k', k'', k'''}} \int \varphi_{n', k'}^{ex}(x') \varphi_{n'', k''}(x'') \varphi_{n''', k'''}(x''') dx' dx'' dx'''$$

$$\Phi_{n', k'; n'', k''; n''', k'''}(x''' - x'', x' - x'', x' - x''')$$

$$\Phi = \int \varphi_{n', k'}^{ex}(2x''' - x'') \varphi_{n'', k''}^{(in)}(2x'' - x''') \varphi_{n''', k'''}^{(in)}(x' - x'')$$

$$= \int \varphi_{n', k'}^{in}(x') \varphi_{n'', k''}^{in}(x'') \varphi_{n''', k'''}^{(in)}(x''')$$

or some of
 If all the Φ 's reduce to a certain multiple of $\delta(x''' - x'') \delta(x'' - x')$, we have the ordinary local field theory.

Actually, for

$$r_i = r_i' = r_i''$$

we have $\rho_i' = \rho_i'' = \rho_i''' = 0$

$$e^{-(q_1'^2 + q_2'^2 + q_3'^2 + q_0'^2)/2\lambda^2}$$

$$\times e^{-(q_1''^2 + q_2''^2 + q_3''^2 + q_0''^2)/2\lambda^2}$$

$$\times e^{-(q_1'''^2 + q_2'''^2 + q_3'''^2 + q_0'''^2)/2\lambda^2}$$

$$= e^{-(r_1'^2 + \dots)/2\lambda^2}$$

$$\times e^{-(r_1'' + r_1''')^2 + \dots)/2\lambda^2}$$

$$\times e^{-(r_1'''^2 + \dots)/2\lambda^2}$$

$$= e^{-r_1'^2/\lambda^2} e^{-r_2'''^2/\lambda^2} e^{-\lambda r_1' r_1''/\lambda^2}$$

$$\times \dots$$

$$= e^{-3(r_1' + r_1''')^2/4\lambda^2} e^{-(r_1' - r_1''')^2/4\lambda^2}$$

$$\times \dots$$

$$\times \dots$$

$$\times \dots$$

$$\int_{-\infty}^{+\infty} e^{-2aq^2} dq = \sqrt{\frac{\pi}{2a}}$$

$$a = \frac{3}{\lambda^2}$$

$$\int e^{-3X^2/\lambda^2} dX = \sqrt{\frac{\pi}{6}} \cdot \lambda$$

$$a = \frac{1}{4\lambda^2}$$

$$\int e^{-r^2/4\lambda^2} dr = \sqrt{\frac{\pi}{8}} \cdot \lambda$$

$$\left(\frac{\pi}{4\sqrt{3}}\lambda^2\right)^4 \delta\left(\frac{r'_\mu + r''_\mu}{2}\right) \delta(r'_\mu - r''_\mu) \\ \delta(X'' - X') \delta(2X' + X'' - 2X''')$$

Normalization factor.

$$\left(\frac{1}{\pi\lambda^2}\right)^3$$

Plus the coupling C_p must be proportional to

$$C_p \sim \text{const. } g/\lambda^2$$

More generally, if ^{some of} k', k'', k''' are not zero, we must investigate the Fourier transform of we have

$$q_1'^2 + q_2'^2 + q_3'^2 + q_0'^2 = f_{\mu\nu} r_\mu' r_\nu'$$

$$q_1''^2 + q_2''^2 + q_3''^2 + q_0''^2 = g_{\mu\nu} r_\mu'' r_\nu''$$

$$q_1'''^2 + q_2'''^2 + q_3'''^2 + q_0'''^2 = h_{\mu\nu} r_\mu''' r_\nu'''$$

where $f_{\mu\nu}$, $g_{\mu\nu}$, $h_{\mu\nu}$ are different symmetric coefficients.

$$\sum_{\mu} (q_{\mu}'^2 + q_{\mu}''^2 + q_{\mu}'''^2) = f_{\mu\nu} r_\mu' r_\nu' + h_{\mu\nu} r_\mu''' r_\nu''' + g_{\mu\nu} (r_\mu' + r_\mu''') (r_\nu' + r_\nu''')$$

$$= f_{\mu\nu} r_\mu' r_\nu' + h_{\mu\nu} r_\mu''' r_\nu'''$$

$$+ g_{\mu\nu} (r_\mu' r_\nu''' + r_\mu''' r_\nu')$$

$$e^{-\sum_{\mu} (q_{\mu}'^2 + q_{\mu}''^2 + q_{\mu}'''^2) / 2\lambda^2} \equiv f(r_\mu', r_\mu''')$$

converges to a multiple of
 in the limit $\delta(r_\mu') \delta(r_\mu''')$
 with $\lambda \rightarrow 0$.

Only question is whether the
 normalization numerical factor is
 independent of λ 's or not.

$$q_1'^2 + q_2'^2 + q_3'^2 + q_0'^2 = q_1''^2 + q_2''^2 + q_3''^2 + q_0''^2 - q_0''^2 + 2q_0''^2$$

$$\begin{aligned}
 &= r_1'^2 + r_2'^2 + r_3'^2 - r_0'^2 + 2 \left(\frac{k_1' r_1'}{m} \right)^2 \\
 &= r_1'^2 + r_2'^2 + r_3'^2 - r_0'^2 + \frac{2}{m^2} (k_1'^2 r_1'^2 + k_2'^2 r_2'^2 \\
 &\quad + k_3'^2 r_3'^2 + k_0'^2 r_0'^2 + 2 k_1' k_2' r_1' r_2' + \dots \\
 &\quad + \dots - 2 k_1' k_0' r_1' r_0' - \dots) \\
 &= \left(1 + \frac{2 k_1'^2}{m^2} \right) r_1'^2 + \left(1 + \frac{2 k_2'^2}{m^2} \right) r_2'^2 + \dots \\
 &\quad + \left(-1 + \frac{2 k_0'^2}{m^2} \right) r_0'^2 + \frac{4 k_1' k_2'}{m^2} r_1' r_2' + \dots \\
 &\quad - \frac{4 k_1' k_0'}{m^2} r_1' r_0' - \dots
 \end{aligned}$$

$$\begin{aligned}
 &(q_1''')^2 + (q_2''')^2 + (q_3''')^2 + (q_0''')^2 \\
 &= \left(1 + \frac{2 k_1'''^2}{m^2} \right) r_1'''^2 + \left(1 + \frac{2 k_2'''^2}{m^2} \right) r_2'''^2 + \dots \\
 &\quad + \left(-1 + \frac{2 k_0'''^2}{m^2} \right) r_0'''^2 + \frac{4 k_1''' k_2'''}{m^2} r_1''' r_2''' + \dots \\
 &\quad - \frac{4 k_1''' k_0'''}{m^2} r_1''' r_0''' - \dots
 \end{aligned}$$

$$\begin{aligned}
 &(q_1'')^2 + (q_2'')^2 + (q_3'')^2 + (q_0'')^2 \\
 &= \left(1 + \frac{2 k_1''^2}{m^2} \right) (r_1' + r_1'')^2 + \dots + (r_2' + r_2'')^2 \\
 &\quad + \left(-1 + \frac{2 k_0''^2}{m^2} \right) (r_0' + r_0'')^2 + \frac{4 k_1'' k_2''}{m^2} (r_1' + r_1'') \\
 &\quad + \dots - \frac{4 k_1'' k_0''}{m^2} (r_1' + r_1'') (r_0' + r_0'') - \dots
 \end{aligned}$$

Interaction between Non-local N.L.T. (15)

Let us take a simpler example of interaction between a non-local scalar (or pseudoscalar) field and a local scalar or spinor field. The interaction Lagrangian will be

$$\begin{aligned} \bar{L}_{int} &= G \text{Trace } \rho \varphi = G \text{Trace } \varphi \rho \\ &= G \int \cdot \cdot (x' | \varphi | x'') \psi^\dagger(x'') \psi(x') dx' dx'' \end{aligned}$$

or

$$= G \int \cdot \cdot \psi^\dagger(x') (x' | \varphi | x'') \psi(x'') dx' dx''$$

provided that

$$(x' | \varphi | x'') = (x'' | \varphi | x')$$

$$\text{or } \varphi(x, r) = \varphi(x, -r)$$

$$\bar{L}_{int} = G \int \cdot \cdot \int dx dr \varphi(x, r) \psi^\dagger(x - \frac{r}{2}) \psi(x + \frac{r}{2})$$

$$\left. \begin{aligned} \varphi(x, r) &= \sum_n \int v_{e,n} e^{il''x} dl'' \varphi_{e,n}(r) \\ \psi(x + \frac{r}{2}) &= \int u_{e''} e^{il''x} e^{i\frac{l''}{2}r} dl'' \\ \psi(x - \frac{r}{2}) &= \int u_{e'}^+ e^{il'x} e^{-i\frac{l'}{2}r} dl' \end{aligned} \right\}$$

$$\overline{h}_{int} = \sum_n \int \cdot \int (2\pi)^4 \delta(l' + l'' + l''') d l' d l'' d l'''$$

$$u_{-l'}^+ v_{l'',n} u_{l'''} \int \phi_{l'',n}(r) e^{-i \frac{l' - l''}{2} r} d r$$

$$\begin{aligned} l' + l'' + l''' &= K = 0 \\ \frac{l' - l''}{2} &= l & -l'' &= l \\ l' + l''' &= L \end{aligned}$$

$$\overline{L}_{int} = \sum_n \int \cdot \int (2\pi)^4 d l d l d L$$

(with)

$$\times u_{-\frac{l}{2} + l}^+ v_{-L,n} u_{\frac{l}{2} - l} \Phi(\quad)$$

$$\Phi(\quad) = \int e^{-\frac{1}{2\lambda^2} (r_1^2 + r_2^2 + r_3^2 + r_0^2)} e^{-i \frac{(l') - (l'')}{2} r'} d r'$$

$$= \text{const} \cdot \exp\left[-\text{const} \frac{((l') - (l''))^2}{2}\right]$$

$$+ \frac{((l_2') - (l_2''))^2}{2} + \frac{((l_3') - (l_3''))^2}{2} + \frac{((l_0') - (l_0''))^2}{2}$$

For a time-like L, we have

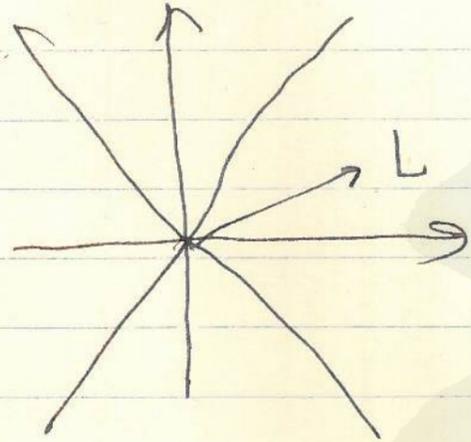
$$= \text{const} \exp\left[-\text{const} \left\{ l^2 - \frac{2(lh)^2}{L^2} \right\}\right]$$

See: Katayama, *Sobyn. Ken.* 4 (1952),
 Nr. 9, 97

N.L.F. (16)

If the above formula is true
 for a ~~space~~-like h , but, if
 h is space-like,

~~$l^2 + 2l_0$~~



~~$l_1^2 +$~~

$\vec{l}'^2 + l_0'^2$

$= 2\vec{l}'^2 + (l_0'^2 - \vec{l}'^2)$

$= 2 \frac{(lL)^2}{h^2} - l^2 > 0$

$l' = \frac{l_1 h_1 + l_2 h_2 + l_3 h_3 - l_0 h_0}{L_1^2 + L_2^2 + L_3^2 - L_0^2}$

$(h_1 = l_1, h_2 = l_2, h_3 = l_3, h_0 = 0) \quad l_1^2 + l_2^2 + l_3^2 + l_0^2$

Anyway, ~~independent~~ the Fourier transform
 of the form function decreases
 exponentially in any direction

for large values of any one of

$(l'_\mu - l''_\mu)^2 \quad (\mu = 1, 2, 3, 0)$ or ~~l'_μ~~

~~$(l'_\mu - l''_\mu)(l'_\mu - l''_\mu)$~~

$l^2 - \frac{2(lL)^2}{h^2}$

for any time-like h

ans
$$2 \frac{(lh)^2}{c^2} - l^2$$

for any space-like L .

Field Equations N.H.F. (19)

$$\begin{aligned} \bar{L} = & \text{Trace} \frac{1}{2} \{ - [\gamma_\mu, \varphi] [\gamma_\mu, \varphi] \\ & + [\gamma_\mu, \varphi]_+ [\gamma_\mu, \varphi]_+ + \frac{1}{\lambda^2} [\alpha_\mu, \varphi] [\alpha_\mu, \varphi] \\ & + \# [\gamma_\mu, \rho] + \# M, \rho \} + \# \frac{1}{\lambda^2} \text{Trace} \{ g \varphi \rho \}^* \\ (\alpha' | \rho_{\alpha\beta} | \alpha'') = & \bar{\Psi}_\alpha(x') \Psi_\beta(x'') \end{aligned}$$

for scalar meson.

$g \varphi \rho$ is to be replaced by $g \varphi \tilde{\sigma}_5 \rho$ where

$$(\tilde{\sigma}_5)_{\alpha\beta} = (\sigma_5)_{\beta\alpha}$$

more precisely,

$$(\alpha' | (\tilde{\sigma}_5)_{\alpha\beta} | \alpha'') = (\sigma_5)_{\beta\alpha} \delta(\alpha' - \alpha'')$$

$$\begin{aligned} \bar{L} = & \frac{1}{2} \int \left\{ \frac{\partial \varphi(x, r)}{\partial x_\mu} \frac{\partial \varphi(x, -r)}{\partial x_\mu} \right. \\ & + 4 \frac{\partial \varphi(x, r)}{\partial r_\mu} \frac{\partial \varphi(x, -r)}{\partial r_\mu} + \frac{4}{\lambda^2} \gamma_\mu \gamma_\mu \varphi(x, r) \varphi(x, -r) \\ & + \left. \left[-\gamma_\mu \frac{\partial \bar{\Psi}(x')}{\partial x'_\mu} + \# M \bar{\Psi}(x') \right] \Psi(x'') \delta(x' - x'') \right\} \\ & + \bar{\Psi}(x') \left[\sigma_\mu \frac{\partial \Psi(x'')}{\partial x''_\mu} + \# M \Psi(x'') \right] \delta(x' - x'') \end{aligned}$$

$$\begin{aligned} * (\alpha' | [\gamma_\mu, \varphi]_+ | \alpha') & = - \left(\frac{\partial}{\partial x'_\mu} - \frac{\partial}{\partial x''_\mu} \right) \varphi(x', x'') \\ * \left(\frac{\partial}{\partial x''_\mu} - \frac{\partial}{\partial x'_\mu} \right) \varphi(x'', x') & = 4 \frac{\partial \varphi(x, r)}{\partial r_\mu} \frac{\partial \varphi(x, -r)}{\partial r_\mu} \end{aligned}$$

$\varphi_n(x) \varphi_n^{\dagger}$

$$\int \int g(x' | \varphi(x'')) \bar{\psi}_\alpha(x'') \psi_\alpha(x') \frac{d^4x' d^4x''}{\xi}$$
$$\text{or } \int \int g(x, r) \bar{\psi}_\alpha(x - \frac{1}{2}r) \psi_\alpha(x + \frac{1}{2}r) \frac{d^4x dr}{\xi}$$

Field equations:

$$-\frac{\partial^2 \varphi(x, r)}{\partial x_\mu \partial x_\mu} - 4 \frac{\partial^2 \varphi(x, r)}{\partial r_\mu \partial r_\mu}$$
$$+ \frac{4}{\lambda^4} r_\mu r_\mu \varphi(x, r) - g \bar{\psi}_\alpha(x + \frac{1}{2}r)$$
$$\times \psi_\alpha(x - \frac{1}{2}r) = 0$$

$$\text{or } \frac{\partial^2 \varphi(x, r)}{\partial x_\mu \partial x_\mu} + 4 \frac{\partial^2 \varphi(x, r)}{\partial r_\mu \partial r_\mu}$$

$$- \frac{4}{\lambda^4} r_\mu r_\mu \varphi(x, r) = -g \bar{\psi}_\alpha(x + \frac{1}{2}r) \psi_\alpha(x - \frac{1}{2}r)$$

$$\left(\gamma_\mu \frac{\partial \psi(x')}{\partial x'_\mu} + M \psi(x') \right) = g \int \frac{\psi(x'') (x'' | \varphi | x')}{dx''}$$

$$\left(\gamma_\mu \frac{\partial \bar{\psi}(x')}{\partial x'_\mu} - M \bar{\psi}(x') \right) = -g \int \frac{(x' | \varphi | x'') \bar{\psi}(x'')}{dx''}$$

$\varphi_n^{\text{in}}(r)$: complete set of orthogonal
 eigenfunctions of satisfying

$$-4 \frac{\partial^2 \varphi_n(X, r)}{\partial r_\mu \partial r_\mu} + \frac{4}{\lambda^4} r_\mu r_\mu \varphi_n(X, r) = m_n^2 \varphi_n(X, r).$$

$$\varphi(X, r) = \sum_n \varphi_n^{\text{ex}}(X) \varphi_n^{\text{in}}(r)$$

$$\frac{\partial^2 \varphi_n(X)}{\partial x_\mu \partial x_\mu} - m_n^2 = -g \int \varphi_n(r) \bar{\varphi}_n(X + \frac{1}{2}r)$$

$$= -g \int \varphi_n^{\text{in}}(x-x'') \bar{\varphi}_n(x) \varphi_n^{\text{ex}}(x''') \delta(\frac{x'+x''}{2} - x''') dx'''$$

provided that φ_n is normalized:

$$\int (\varphi_n^{\text{in}}(r))^2 dr = 1.$$

$$\left(r_\mu \frac{\partial \psi(x')}{\partial x'_\mu} + M \psi(x') \right) = g \int$$

$$x \sum_n \left(\psi(x'') \varphi_n^{\text{ex}}(\frac{x'+x''}{2}) \varphi_n^{\text{in}}(x''-x') dx'' \right)$$

$$= g \sum_n \int \psi(x'') \varphi_n^{\text{ex}}(x''') \varphi_n^{\text{in}}(x''-x') \times \delta(\frac{x'+x''}{2} - x''') dx'' dx'''$$

$$\left(\frac{\partial^2 \varphi_n^{ea}(x'')}{\partial x_\mu \partial x_\mu} - m_n^2 \varphi_n^{ea}(x'') \right)$$

$$= -g \int \Phi_n(x', x'', x''') \bar{\psi}_\alpha(x') \psi_\alpha(x'') dx' dx''$$

with

$$\Phi_n(x', x'', x''') = \varphi_n^{in}(x' - x''') \delta\left(\frac{x' + x''}{2} - x'''\right)$$

$$\left(\gamma_\mu \frac{\partial \psi(x')}{\partial x'_\mu} + M \psi(x') \right)$$

$$= g \sum_n \int \Phi'_n(x', x'', x''') \varphi_n^{ea}(x'') \psi(x''') dx'' dx'''$$

$$\Phi'_n(x', x'', x''') = \varphi_n^{in}(x'' - x') \times \delta\left(\frac{x' + x''}{2} - x'''\right)$$

$\Phi'_n = \pm \bar{\Phi}_n$
 according as n is even or odd.

$$\begin{aligned}
 & \gamma_\mu \frac{\partial \bar{\Psi}(x')}{\partial x'_\mu} - M \bar{\Psi}(x') \\
 & = g \sum_n \int \bar{\Psi}(x'') \varphi_n^{\downarrow \lambda} \left(\frac{x'+x''}{2} \right) \varphi_n^{\uparrow \lambda}(x'-x'') dx'' \\
 & = g \sum_n \int \bar{\Psi}(x''') \varphi_n^{\downarrow \lambda}(x''') \Phi_n^{\uparrow \lambda}(x', x'', x''') dx''
 \end{aligned}$$

$$\Phi_n^{\uparrow \lambda}(x', x'', x''') = \varphi_n^{\uparrow \lambda}(x'-x''') \delta\left(\frac{x'+x'''}{2} - x''\right)$$

If we confine our attention to any one of the even state
 (~~with~~ a set of n_1, n_2, n_3, n_0
 with $n = n_1 + n_2 + n_3 - n_0 = \text{even}$)
 one can proceed just as shown in
 Moller paper.

$$\bar{\Phi}(x', x'', x''') = (2\pi)^{-8} \int G(l', l''')$$

$$\exp i \int d^4x \left\{ \bar{\psi} \left(\frac{x'+x'''}{2} - x'' \right) + \bar{\psi}(x'-x''') \right\}$$

$$\bar{\psi} = \int d^4x \left\{ l' x' + l'' x'' - (l' + l'') x' \right\}$$

$$= (2\pi)^{-8} \int G(l', l''') \exp i \left[l \left(\frac{x' + x'''}{2} - x'' \right) + l' (x' - x''') \right] \times dL dl$$

with

$$L = l' + l'''$$

$$l = \frac{l' - l'''}{2}$$

$$G(l, L) = G(l, L)$$

in our case.

In particular, for

$$n_1 = n_2 = n_3 = n_0 = 0,$$

we ~~have~~ may take for our coord. sys.

$$G(l) = \exp - \frac{\lambda^2}{2} (l_1^2 + l_2^2 + l_3^2 + l_0^2)$$

independent of the external motion.

The form factors appearing in the self-energies are

$$|G(k, -k - P')|^2 \delta(k^2 + M^2)$$

$$l = k + P'/2$$

$$|G(P', k - P')|^2 \delta(k^2 + m^2)$$

$$l = P' - k/2$$

$$|G''(k, -P')|^2 \delta(k^2 + M^2)$$

$$l = k + P'$$

N. h. F. (20)

plus, as long as one can take up a definite ~~internal~~ set of even ~~internal~~ state, the ~~consequence~~ of self-excitations are secured.

However, the situation is more complicated due to the ~~for~~ supplement condition.

Let us assume that, even in the presence of interaction, the S.C.,

$$\frac{\partial}{\partial x_\mu} \left(\frac{\partial}{\partial \gamma_\mu} + \frac{1}{\pi} \gamma_\mu \right) \varphi = 0$$

is true.

whether this is (I shall discuss the legitimacy or not later on.)

Once we assume S.C., the form factor for

$$v_1 = v_2 = v_3 = 0$$

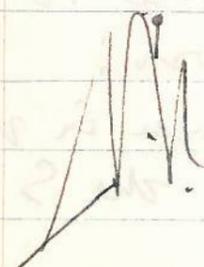
will be

$$G(l) = \exp -\frac{\lambda^2}{2} (l_1^2 + l_2^2 + l_3^2 + l_0^2)$$

in the rest system for the meson in question; or in the original system

$$G(l) = \exp -\frac{\lambda^2}{2} \left(l^2 - \frac{2(l \cdot k)^2}{k^2} \right)$$

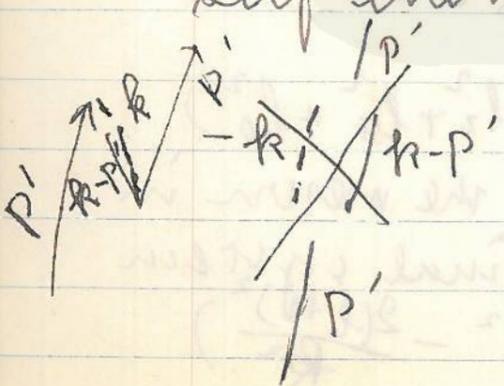
Now, the k for the nucleon in the calculation of self-energy of the nucleon is that of the nucleon in the initial and final states.



So that it is simply ($k = -L = P'$)
 $k_1 = k_2 = k_3$ $k_0 = m$
 for a nucleon at rest.

In other words, the form factor is
 i.e. $G(l) = \exp\left(\frac{i l_0}{2} \left(\vec{k}^2 + \left(k_0 + \frac{m}{2}\right)^2 \right)\right)$
~~independent~~ with l_1, l_2, l_3, l_0
 for the original system.
 (rest)

The ~~correct~~ k for the nucleon self-energy is for the first term:



$$-k = L$$

$$\frac{(P' - k)^2 - 2\left(P' - \frac{k}{2}\right)k}{k^2}$$

For the nucleon at rest
 $P_1 = P_2 = P_3 = 0$ $P_0 = M$

N. L. F. (21)

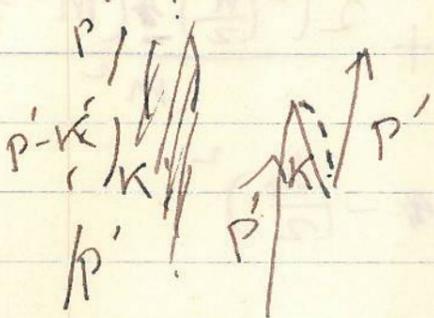
$$\begin{aligned}
 & \vec{k}^2 = -m^2 \\
 & \left(\frac{\vec{k}}{2}\right)^2 - \left(M - \frac{k_0}{2}\right)^2 + \frac{2\left(-\frac{k^2}{2}\right)\left(M - \frac{k_0}{2}\right)k_0}{m^2} \\
 & = \left(\frac{\vec{k}}{2}\right)^2 - M^2 + M k_0 - \left(\frac{k_0}{2}\right)^2 \\
 & \quad + \frac{2(m^2 - M k_0)^2}{m^2} \quad (+ M^2 k_0^2) \\
 & = -m^2 - M^2 + M k_0 + \frac{2(m^4 - 2m^2 M k_0)}{m^2} \\
 & = +m^2 - M^2 - 3M k_0 + \frac{2M^2 k_0^2}{m^2} \rightarrow 0 \\
 & = \frac{2M^2 k_0^2}{m^2} - 3M k_0 - M^2 + m^2 > 0
 \end{aligned}$$

In any case, for a very large value of

$$k_0 \gg M,$$

the argument in the Gauss law is always very large as k_0^2 .

For the second term, the meson
 in the intermediate state



$$k = -L = p' - p,$$

which is space-like

~~If the supplementary condition
 is added to the ^{virtual} meson in the
 virtual intermediate state~~

~~$$k_\mu \gamma_\mu \psi = 0,$$~~

~~we have no restriction on
 the internal motion of the
 meson:~~

~~$$n'_0 = 0$$~~

~~if we take the ~~int~~ as new
 x-axis the direction of k_μ .
 In that case, n'_0 is no longer
 restricted to 0 and the infinite
 degeneracy reappears.~~

However, ~~if~~ this came from
 the (incoming) free meson field
 for which k_μ is always

time-like. The factor G in
 $|G(\mathbf{k}, -P')|^2 \delta(\mathbf{k}^2 + M^2)$

becomes thus

$$\exp - \frac{\lambda^2}{2} \left((\mathbf{k} - \mathbf{P}')_1^2 + (\mathbf{k} - \mathbf{P}')_2^2 + (\mathbf{k} - \mathbf{P}')_3^2 + (\mathbf{k} - \mathbf{P}')_0^2 \right)$$

in the rest system for the meson

$$\vec{k} = \mathbf{k} - \mathbf{P}'$$

$$k_0 = \sqrt{(\mathbf{k} - \mathbf{P}')^2 + m^2}$$

$$\text{rest system: } \frac{(\vec{k} - \vec{P}')^2 - (k_0 - P'_0)^2 + 2 \{ (\mathbf{k} - \mathbf{P}')(\mathbf{k} - \mathbf{P}') - k_0 (k_0 - P'_0) \}^2}{m^2}$$

$$\frac{(\vec{k} - \vec{P}')^2 - (k_0 - P'_0)^2}{m^2} \quad k_0 = \sqrt{\mathbf{k}^2 + m^2}$$

$$\frac{\mathbf{k}^2 - (k_0 - M)^2}{m^2}$$

$$+ \frac{2 \{ \mathbf{k}^2 - k_0 (k_0 - M) \}^2}{m^2} \quad \underline{2M(k_0 - M)}$$

$$= \frac{\mathbf{k}^2 - k_0^2 + 2k_0 M - M^2}{m^2}$$

$$+ \frac{2}{m^2} \left\{ \mathbf{k}^2 - \sqrt{\mathbf{k}^2 + m^2} (k_0 - M) \right\}^2 > 0$$

$$\frac{(\mathbf{k}^2)^2 - 2\sqrt{\mathbf{k}^2 + m^2} (k_0 - M) \mathbf{k}^2 + (\mathbf{k}^2 + m^2) (k_0 - M)^2}{m^4}$$

$$\varphi(x, r) = \int v_{l', l''} e^{i(l'x' + l''x'')} (dl') (dl'')$$

$$= \int v(L, l) e^{i(Lx + lr)} (dL) (dl)$$

$$\left. \begin{aligned} L &= l' + l'' \\ l &= \frac{l' - l''}{2} \end{aligned} \right\} \left. \begin{aligned} l' &= \frac{L}{2} + l \\ l'' &= \frac{L}{2} - l \end{aligned} \right\}$$

$$\left\{ -L^2 - 4\left(l^2 - \frac{1}{\lambda^2} \frac{\partial^2}{\partial l_\mu \partial l_\mu}\right) \right\} v(L, l) e^{i(Lx + lr)}$$

$$= -g \int \bar{u}_\alpha\left(-\frac{L}{2} - l\right) u_\alpha\left(\frac{L}{2} - l\right) e^{i(Lx + lr)} (dL) (dl)$$

$$\Psi_\alpha(x'') = \int u_\alpha(l'') e^{il''x''} (dl'')$$

$$\bar{\Psi}_\alpha(x') = \int \bar{u}_\alpha(l') e^{-il'x'} (dl')$$

$$= \int \bar{u}_\alpha(-l') e^{il'x} (dl')$$

$$\bar{\Psi}_\alpha\left(x + \frac{1}{2}r\right) \Psi_\alpha\left(x - \frac{1}{2}r\right)$$

$$= \int \bar{u}_\alpha(-l') u_\alpha(l'') e^{i(l'x' + l''x'')} (dl') (dl'')$$

$$= \int \bar{u}_\alpha\left(-\frac{L}{2} - l\right) u_\alpha\left(\frac{L}{2} - l\right) e^{i(Lx + lr)} (dL) (dl)$$

$$\left\{ -L^2 - 4 \left(l^2 - \frac{1}{\lambda^2} \frac{\partial^2}{\partial l_\mu \partial l_\mu} \right) \right\} v(L, l)$$

$$= -g \bar{u}_\alpha \left(-\frac{L}{2} - l \right) u_\alpha \left(\frac{L}{2} - l \right)$$

$$\int (i \gamma_\mu l'_\mu + M) u_\alpha(l') e^{i l' x'} (d l')$$

$$= g \int u_\alpha(l'') e^{i l'' x''} W_{\alpha\beta}(l'', l') e^{i(l'' x'' + l' x')} dx'' d l'' d l' d l''$$

$$= g \int u_\alpha(-l''') W(l''', l') \cdot (2\pi)^4 d l''' e^{i l' x'} d l'$$

$$(i \gamma_\mu l'_\mu + M) u(l') = \int u(-l''') W(l''', l') \times d l'''$$

Particular solution:

$$v(L, l) = N_{\#0} e^{-\frac{\Delta^2}{2} (l_1^2 + l_2^2 + l_3^2 + l_0^2)}$$

$$v(L, l) = b_L v_{0L} \int (v_{0L}(l))^2 d l = 1$$

$$(-L^2 - m_0^2) b_L v_{0L}(l)$$

$$= -g \bar{u}_\alpha \left(-\frac{L}{2} - l \right) u_\alpha \left(\frac{L}{2} - l \right)$$

$$(-L^2 - m_0^2) b_L = -g \int v_{0L}(l) \bar{u}_\alpha \left(-\frac{L}{2} - l \right) u_\alpha \left(\frac{L}{2} - l \right) d l$$

N.L.F. (24)

$$\int (-L^2 - m_0^2) b_L e^{iLx} dL$$

$$= -g (2\pi)^4 \int v_{0L}(l) \bar{u}_a(-\frac{L}{2}-l) e^{i(\frac{L}{2}+l)(x+\frac{1}{2}r)} u_a(\frac{L}{2}-l) e^{-i(\frac{L}{2}-l)(x-\frac{1}{2}r)} (dL)(dl) e^{-i\epsilon r}$$

$$= -g (2\pi)^4 \int \bar{u}_a(-l') e^{il'x'} u_a(l'') e^{il''x''} v_{0L}(l) e^{-i\epsilon r} dL dl$$

$$= -g (2\pi)^4 \int \int \bar{u}_a(-l') e^{il'x'} u_a(l'') e^{il''x''} v_0(l', l'') e^{-i\frac{l'-l''}{2}(x'-x'')} dl' dl''$$

~~$$= -g (2\pi)^4 \int \int \bar{u}_a(-l') e^{il'x'} u_a(l'') e^{il''x''} v_0(l', l'') e^{-i\frac{l'-l''}{2}(x'-x'')} dl' dl'' dl''' dl'''' \delta(l'-l''') \delta(l''-l''')$$~~

~~$$= -g (2\pi)^4 \bar{u}_a(x') u_a(x'')$$~~

$$= -g (2\pi)^4 \int \int \Phi(x', x'') \bar{u}_0(x') u_a(x'')$$

Interaction between
local spinor field
and N.L. scalar field

$$\bar{L}_{\text{total}} = \bar{L}_{\text{sc}} + \bar{L}_{\text{sp}} + \bar{L}_{\text{int}}$$

$$\bar{L}_{\text{sc}} = \frac{1}{2} \text{Trace} \{ [\gamma_{\mu}, \varphi] [\gamma_{\mu}, \varphi]$$

$$- [\gamma_{\mu}, \varphi]_+ [\gamma_{\mu}, \varphi]_+ + \frac{4}{\lambda^4} [\alpha_{\mu}, \varphi] [\alpha_{\mu}, \varphi] \}$$

$$\bar{L}_{\text{sp}} = \frac{1}{2} \text{Trace} \{ +i \cancel{p}_{\mu} \cancel{p}_{\mu}, p_{\mu} \}_+ - [M, p]_+ \}$$

$$\bar{L}_{\text{int}} = g \text{Trace} \varphi \rho \quad \begin{array}{l} (x) \rho_{\mu}^{\alpha\beta} = \bar{\psi}(x) \sigma_{\mu} \psi(x) \\ (x' | \rho | x'') \\ = \bar{\psi}(x') \psi(x'') \end{array}$$

$$(x' | \rho_{\alpha\beta} | x'') = \bar{\psi}_{\alpha}(x') \psi_{\beta}(x'')$$

$$(x'' | \rho_{\beta\alpha} | x') = \bar{\psi}_{\beta}(x'') \psi_{\alpha}(x')$$

$$= (x' | \rho_{\alpha\beta}^* | x'') = \bar{\psi}_{\beta}(x'') \psi_{\alpha}(x')$$

hermitian

$$(x') \tilde{\varphi} | x'' = (x'' | \varphi | x')$$

$$\tilde{\varphi}(x, y) = \varphi(x, -y)$$

$$\bar{L}_{\text{sc}} = -\frac{1}{2} \int d^4x \left\{ \frac{\partial \tilde{\varphi}(x, y)}{\partial x_{\mu}} \frac{\partial \varphi(x, y)}{\partial x_{\mu}} \right\} dx dy$$

$$+ 4 \frac{\partial \tilde{\varphi}(x, y)}{\partial y_{\mu}} \frac{\partial \varphi(x, y)}{\partial y_{\mu}} + \frac{4}{\lambda^4} \gamma_{\mu} \gamma_{\mu} \tilde{\varphi}(x, y) \varphi(x, y)$$

$$\begin{aligned} \bar{h}_{op} &= + \frac{1}{2} \int \left\{ \cancel{\gamma}_\mu \frac{\partial \bar{\psi}(x')}{\partial x'_\mu} - M \bar{\psi}(x') \right\} \psi(x') dx' \\ &= \int \bar{\psi}(x) \left\{ \gamma_\mu \frac{\partial \psi(x)}{\partial x_\mu} + M \psi(x) \right\} dx \end{aligned}$$

$$\bar{h}_{int} = g \int \int x' \phi(x, r) \bar{\psi}_a(x - \frac{1}{2}r) \psi_a(x + \frac{1}{2}r) dx dr$$

$$g \int \int (x' | \phi | x'') \bar{\psi}_a(x') \psi_a(x'') dx' dx''$$

Field equations:

$$\frac{\partial^2 \phi(x, r)}{\partial x_\mu \partial x_\mu} + 4 \frac{\partial^2 \phi(x, r)}{\partial r_\mu \partial r_\mu} - \frac{4}{x^4} r_\mu r_\mu \phi(x, r)$$

$$+ g \bar{\psi}_a(x + \frac{1}{2}r) \psi_a(x - \frac{1}{2}r) = 0$$

$$\gamma_\mu \frac{\partial \psi(x')}{\partial x'_\mu} + M \psi(x')$$

$$- g \int (x'' | \phi | x') \bar{\psi}_a(x'') dx'' = 0$$

$$\cancel{\gamma}_\mu \frac{\partial \bar{\psi}(x')}{\partial x'_\mu} - M \bar{\psi}(x')$$

$$+ g \int (x' | \phi | x'') \bar{\psi}(x'') dx'' = 0$$

N.L.F. (26)

First, ignore all about supplementary condition and expand $\varphi(x, r)$ into series of orthogonal function defined by

$$\left\{ \begin{aligned} \frac{\partial^2 \varphi}{\partial x \partial x} + \frac{4}{x^2} r \mu \varphi + m_n^2 \varphi &= \gamma \varphi_n^{(in)}(r) \\ m_n^2 &\geq 0 \end{aligned} \right.$$

$\varphi_n^{(e)}(r)$: even - real

$\varphi_n^{(o)}(r)$: odd - pure imag.
 $\varphi(r) = -\varphi(-r)$

$$\varphi(x, r) = \sum_n \varphi_n^{(e)}(x) \varphi_n^{(e)}(r)$$

normalization

$$\int \tilde{\varphi}_n(r) \varphi_n(r) dr = 1$$

$$\frac{\partial^2 \varphi_n^{(e)}(x)}{\partial x \partial x} - m_n^2 \varphi_n^{(e)}(x) = 0$$

$x' = x - x''$
 $x'' = x + x'$

$$= -g \int \tilde{\varphi}_n(x, r) \varphi_n(x + \frac{1}{2}r) \varphi_n(x - \frac{1}{2}r) dr$$

$$\gamma \frac{\partial \varphi(x')}{\partial x'} + M \varphi(x') \quad \begin{aligned} r &= 2(x' - x'') \\ &= 2(x'' - x') \end{aligned}$$

$$= g \sum_n \int \tilde{\varphi}_n(x') \varphi_n(x'') dx''$$

$$\gamma \frac{\partial \varphi(x')}{\partial x'} - M \varphi(x') \quad \varphi_n^{(e)}(x) \varphi_n(-r)$$

$$= -g \sum_n \int \tilde{\varphi}_n(x'') \varphi_n(x') dx''$$

$$\frac{\partial^2 \varphi_{n\mu}^{(x)}(x'')}{\partial x''_{\mu} \partial x''_{\mu}} - m_{n\mu}^2 \varphi_{n\mu}^{(x)}(x'')$$

$$= -g \int \Phi_{n\mu}(x', x'', x''') \varphi_{n\mu}^{(x)}(x') \varphi_{n\mu}^{(x)}(x''') dx' dx''$$

$$\Phi_{n\mu}(x', x'', x''') = \frac{1}{\chi_{n\mu}} \varphi_{n\mu}^{(x)}(x' - x'') \delta\left(\frac{x' + x''}{2} - x'''\right)$$

$$\gamma_{n\mu} \frac{\partial \psi(x')}{\partial x'_{\mu}} + M \psi(x')$$

$$= g \sum_n \int \Phi_{n\mu}(x', x'', x''') \varphi_{n\mu}^{(x)}(x') \psi(x''')$$

etc.

Thus, if a form function, for example

$$\Phi_0(x', x'', x''')$$

was chosen as to give convergent approximately chosen, one can expect the same convergence as in Moller's theory.

N.L.F. (27)

Unfortunately, ~~the~~ this such a function does not seem to exist for the following reasons:

An eigenfunction of the type

$$e^{-\frac{1}{2\lambda^2}(x_1^2 + x_2^2 + x_3^2 + x_0^2)}$$

is not invariant. One might write

$$e^{-\frac{1}{2\lambda^2}(x_\mu x_\mu - \frac{2(k_\mu x_\mu)^2}{k_\mu k_\mu})}$$

which is formally invariant, where k_μ is an arbitrary fixed time-like vector. However, one will be asked how the direction of k_μ is determined. Only, if k_μ is determined in a reasonable way, above considerations really make sense.

This difficulty is intimately connected with the fact that hermitian invariant operators are, in general, not positive definite. Even, if one makes it positive definite by taking the square, for instance, the infinite degeneracy remains

(Born,

Schrödinger

In order to ~~suppress~~ remove the degeneracy, one need further restriction on the internal motion.

In the case of oscillator model, one may assume a S.C., however, the S.C. was so chaoticⁱⁿ to connect the external and internal motion. ~~that~~ For that reason, the further develop of field equations does not have the simple form which is equivalent to those reduced to those of n.l. interaction theory.

In other words, the internal motion is affected by the change in the external motion, which makes our picture really very complicated.

Supplementary Conditions, I, N. L. F. (28)

For the reasons above mentioned, the coupling between external and internal motion is to be considered more seriously:

$$\frac{\partial}{\partial x_\mu} \left(-\frac{\partial}{\partial r_\mu} + \frac{1}{\lambda^2} r_\mu \right) \varphi = 0$$

or

$$\left[\hat{p}_\mu \left\{ -i \left(\frac{\partial}{\partial r_\mu} + \frac{1}{\lambda^2} r_\mu \right) \varphi \right\} + \frac{1}{\lambda^2} \left[x_\mu, \varphi \right] \right] = 0$$

$$\sum_{j=1,2,3} \frac{\partial}{\partial x_j} \left(-\frac{\partial}{\partial r_j} + \frac{1}{\lambda^2} r_j \right) \varphi(x, r)$$

$$- (i) \frac{\partial}{\partial x_0} \left((i) \frac{\partial}{\partial r_0} + \frac{1}{\lambda^2} (i) r_0 \right) \varphi = 0$$

If φ is independent of x_j and only depends on x_0

$$\left(\frac{\partial}{\partial r_0} + \frac{1}{\lambda^2} r_0 \right) \varphi(x_0, r) = 0$$

$$\varphi \propto e^{-\frac{1}{2\lambda^2} r_0^2}$$

Further if φ is a function of $k_\mu x_\mu$, where k_μ is a time-like vector, then φ must be of the form have

the factor

$$e^{-\frac{1}{2\lambda^2}(r_0')^2}$$

$$r_0' = * \frac{k_\mu r_\mu}{\sqrt{-k_\mu k_\mu}}$$

$$\left[\lambda \frac{\partial}{\partial r_0} + \frac{1}{\lambda} r_0, -\lambda \frac{\partial}{\partial r_0} + \frac{1}{\lambda} r_0 \right] \varphi$$

$$= 2\varphi$$

$$b_0 = \frac{1}{\sqrt{2}} \left(\lambda \frac{\partial}{\partial r_0} + \frac{1}{\lambda} r_0 \right) \quad [b_0, b_0^*] = 1$$

$$b_0^* = \frac{1}{\sqrt{2}} \left(-\lambda \frac{\partial}{\partial r_0} + \frac{1}{\lambda} r_0 \right)$$

Similarly,

$$b_j = \frac{1}{\sqrt{2}} \left(\lambda \frac{\partial}{\partial r_j} + \frac{1}{\lambda} r_j \right) \quad \}$$

$$b_j^* = \frac{1}{\sqrt{2}} \left(-\lambda \frac{\partial}{\partial r_j} + \frac{1}{\lambda} r_j \right)$$

$$\left\{ \frac{\partial}{\partial x_j} \cdot b_j^* + \frac{\partial}{\partial x_0} b_0 \right\} \varphi = 0$$

$$b_j^* \chi_n = \sqrt{n+1} \chi_{n+1} \quad \}$$

$$b_j \chi_n = \sqrt{n} \chi_{n-1}$$

$$(n+1)(b_j^* | n) = \sqrt{n+1}, \quad (n-1)(b_j | n) = \sqrt{n}.$$

χ_n : normalized chc. eigen fn

$$\chi_n = N_n H_n^{(\alpha)} e^{-\frac{x^2}{2}} \quad x = \frac{r}{\lambda}$$

$$\varphi(X, r) = \sum_n \varphi_{n, n_2, n_3, n_0}^{(ex)}(X) \chi_{n, n_2, n_3, n_0}^{(v)}(r)$$

$$\left\{ \frac{\partial}{\partial X_j} b_j^* + \frac{\partial}{\partial X_0} b_0 \right\} \varphi$$

$$= \left\{ \sum_j \frac{\partial \varphi_{n, n_2, n_3, n_0}^{(ex)}}{\partial X_j} \sqrt{n_j+1} \chi_{n_j, n_2, n_3, n_0}^{(v)} + \frac{\partial \varphi_{n, n_2, n_3, n_0}^{(ex)}}{\partial X_0} \sqrt{n_0} \chi_{n, n_2, n_3, n_0-1}^{(v)} \right\} = 0$$

$$= \sum \chi_{n, n_2, n_3, n_0} \left\{ \sqrt{n_1} \frac{\partial \varphi_{n, -1, n_2, n_3, n_0}^{(ex)}(X)}{\partial X_1} + \sqrt{n_2} \frac{\partial \varphi_{n, n_2-1, n_3, n_0}^{(ex)}(X)}{\partial X_2} + \sqrt{n_3} \frac{\partial \varphi_{n, n_2, n_3-1, n_0}^{(ex)}(X)}{\partial X_3} + \sqrt{n_0+1} \frac{\partial \varphi_{n, n_2, n_3, n_0+1}^{(ex)}(X)}{\partial X_0} \right\} = 0$$

$n_1 = n_2 = n_3 = 0$:

$$\frac{\partial \varphi_{n, n_2, n_3, 0, 0, 0, n_0+1}}{\partial X_0} = 0$$

$$n_1 = 1, n_2 = n_3 = 0 \quad \frac{\partial \varphi_{0, 0, 0, n_0}}{\partial X_1} + \sqrt{n_0+1} \frac{\partial \varphi_{1, 0, 0, n_0+1}}{\partial X_0} = 0$$

$$\varphi_{n_1, n_2, n_3, n_0}^{(ex)}(X) = \int \varphi_{n_1, n_2, n_3, n_0}^{(ex)}(k) e^{i(kX - k_0 X_0)} d^4 k$$

$$\begin{aligned} & \sqrt{n_1} k_1 \varphi_{n_1-1, n_2, n_3, n_0}^{(ex)}(k) \\ & + \sqrt{n_2} k_2 \varphi_{n_1, n_2-1, n_3, n_0}^{(ex)}(k) + \sqrt{n_3} k_3 \varphi_{n_1, n_2, n_3-1, n_0}^{(ex)}(k) \\ & - \sqrt{n_0+1} k_0 \varphi_{n_1, n_2, n_3, n_0+1}^{(ex)}(k) = 0 \end{aligned}$$

$$\begin{aligned} n_1 = n_2 = n_3 = 0 \\ \sqrt{n_0+1} k_0 \varphi_{0,0,0,n_0}^{(ex)}(k) = 0 \\ \varphi_{0,0,0,1}^{(ex)}(k) = \varphi_{0,0,0,2}^{(ex)}(k) = \dots = 0 \\ \text{if } k_0 \neq 0. \end{aligned}$$

$$n_1 = 1, n_2 = n_3 = 0. \quad \varphi_{0,0,0,n_0}^{(ex)}(k) = \sqrt{n_0+1} k_0 \varphi_{1,0,0,n_0+1}^{(ex)}(k)$$

$$k_1 \varphi_{0,0,0,0}^{(ex)}(k) = k_0 \varphi_{1,0,0,1}^{(ex)}(k)$$

$$\varphi_{0,0,0,2}^{(ex)}(k) = \varphi_{1,0,0,3}^{(ex)}(k) = \dots = 0 \\ \text{if } k_1 \neq 0.$$

If further $k_2 = k_3 = 0.$

$$\begin{aligned} & \sqrt{n_1} k_1 \varphi_{n_1-1, n_2, n_3, n_0}^{(ex)} \\ & = \sqrt{n_0+1} k_0 \varphi_{n_1, n_2, n_3, n_0+1}^{(ex)} \end{aligned}$$

$$\begin{aligned} (0, 0, 0, 0) & \rightarrow (1, 0, 0, 1) \rightarrow (2, 0, 0, 2) \\ & \rightarrow \dots \rightarrow (n, 0, 0, n) \end{aligned}$$

$$\sqrt{n+1} k_1 \varphi_{n,0,0,n} = \sqrt{n+1} k_0 \varphi_{n+1,0,0,n+1}$$

$$\frac{\varphi_{n+1,0,0,n+1}}{\varphi_{n,0,0,n}} = \frac{k_1}{k_0}$$

Thus, if $|k_1/k_0| < 1$

$$\sum_n \int \varphi_{n,0,0,n}^{(ex)}(k) \chi_{n,0,0,n}(r) \cdot e^{i(kx - k_0 x_0)} dk$$

$$= \int f(k, r) e^{i(kx - k_0 x_0)} dk$$

$$f(k, r) = \sum_n \varphi_{n,0,0,n}^{(ex)}(k) \chi_{n,0,0,n}(r)$$

$$= \sum_n \left(\frac{k_1}{k_0}\right)^n \chi_{n,0,0,n}(r) e^{i(kx - k_0 x_0)} dk$$

$$\sum_n \int \left(\frac{k_1}{k_0}\right)^{2n} \chi_{n,0,0,n}^2(r) dr$$

$$= \frac{1}{1 - \left(\frac{k_1}{k_0}\right)^2} = \frac{k_0^2}{k_0^2 - k_1^2}$$

For any sp-time-like direction k

$f(k, r)$

is quadratically integral \int_n in r -space.

On the contrary, for any space-like k ,
i.e. $|\frac{k}{k_0}| > 1$, $f(k, v)$ is not a
~~bounded~~ quadratically integral fns of
 v . Thus, in order that $\varphi(x, v)$
it is a quadratically integral fns
of x and v , φ must be the superposition
of plane waves with ~~space-like~~ ^{time-like} k .
This is certainly true for a free
particle with a real mass, but for
the field quantity φ , which ~~is~~
~~satisfies~~ the field equation with
the interaction, contains, in general,
the superposition of plane waves
with space-like k as well as
time-like k . For instance, the
field equation with the static source
is the superposition of ^{wave functions} ~~space waves~~
with ~~the space-like k~~ ^{$k_0=0$} and of
those with the free plane-waves.

In any perturbation calculation,
Møller has suggested a form function
which decreases for $|\pi^2| > M$
for both signs of $\pi^2 = l^2 - \frac{(lh)^2}{L^2}$,

, which may be negative for a space-like
 L . (For instance $L_1=l_1, L_2=l_2, L_3=l_3$
 $L_0=0, \quad \Pi^2 = l_1^2 + l_2^2 + l_3^2 - l_0^2 = \frac{(l_1^2 + l_2^2 + l_3^2)^2}{l_1^2 + l_2^2 + l_3^2}$
 $= -l_0^2 < 0$)

Another suggestion may be to ~~start~~ introduce
 a system of four unit vectors $e_\mu^{(\nu)}$ orthog
 to each other and assume $(\nu=1,2,3,4)$

$$F^{\mu\nu} = \sum_{\nu} \left\{ - (e_\mu^{(\nu)} \frac{\partial}{\partial x^\nu})^2 + \frac{1}{\lambda^2} (e_\mu^{(\nu)} \gamma_\nu)^2 \right\}$$

which is formally invariant*. However,
 the objection immediately arises
 as for the arbitrariness of $e_\mu^{(\nu)}$.
 Only if all the ~~case~~ physically meaningful
 results are independent of the choice
 of $e_\mu^{(\nu)}$ (or the ~~arbitrary~~ internal
 coordinate system), the above
 idea makes sense. Unfortunately,
 there is no guarantee for that,
 because for example, if $e_4^{(1)}$ winds

* In this case, S.C. is no longer
 necessary, because the eigenvalues
 are of the form $\frac{1}{\lambda^2} (n_1 + n_2 + n_3 + n_0 + 1)$

the direction
with ~~the~~ wave vector k_n , the external motion with ^{this} k_n may behave in such particular ~~way~~ ^{ways} ~~property~~ ^{to} which may ~~well~~ violate the relativistic principle.

Thus, ^{it is} ~~we are~~ very likely that we have to give up the method of direction integration of field equations. We had better invent a method of ~~solving~~ ~~expressing~~ the outgoing field $\phi^{(out)}$ in terms of the incoming field $\phi^{(in)}$. This amounts to the same thing as to construct directly the S-matrix without taking recourse to Y. F. K.'s method of approach.

The supplementary condition may or may not be satisfied by the actual fields ϕ .

In other words, ^{is there} ~~is there~~ the interaction ~~exists~~ ^{due to} ~~the relation~~ ^{between} the external and internal motion is so different from that for free field that S.C. may not work at all.

Supplementary Conditions II, N. L. F. (32)

In the presence of the interaction, the S.C. cannot be

$$\frac{\partial}{\partial x_\mu} \left(-\frac{\partial}{\partial r_\mu} + \frac{1}{\lambda} r_\mu \right) \varphi = 0$$

is not adequate, because we have to take into account both the time-like k and the space-like k . Let us assume the new supplementary conditions:

$$k_\mu \left(-\frac{\partial}{\partial r_\mu} + \frac{1}{\lambda} r_\mu \right) \varphi = 0 \quad \text{for } k_\mu k_\mu < 0$$

$$k_\mu \left(\frac{\partial}{\partial r_\mu} + \frac{1}{\lambda} r_\mu \right) \varphi = 0 \quad \text{for } k_\mu k_\mu > 0$$

In the limit $k_\mu k_\mu \neq 0$, both of these conditions can be satisfied, because they ~~do~~ the two operators commute with each other for $k_\mu k_\mu = 0$.

The ^{orthogonal} ~~orthogonal~~ eigenfunctions of the 8-dimensional eigenvalue problem

$$\frac{\delta^2 \varphi}{\delta x_\mu \delta x_\mu} + 4 \frac{\delta^2 \varphi}{\delta r_\mu \delta r_\mu} + \frac{4}{\lambda^2} r_\mu r_\mu \varphi = \mu^2 \varphi$$

which satisfies the above S.C., has, in general, the form

$$e^{i k_\mu r_\mu} \chi_{n, k_\mu}(r)$$

where, $n = (n_1, n_2, n_3, n_0)$, $n \in n_0 + n_1 + n_2 + n_3$, $\chi_{n, k_\mu}(r)$ is normalized.

$$\int \chi_{n, k_\mu}(r) \chi_{n, k_\mu}(-r) (dr)^4 = 1$$

k_μ and m_n is connected by the relation

$$-k_\mu k_\mu - m_n^2 = \mu^2$$

or $k_\mu k_\mu + m_n^2 + \mu^2 = 0$,

$$m_n^2 = \frac{8}{\lambda^2} (n_1^2 + n_2^2 + n_3^2 - n_0^2).$$

So k_μ can be time-like or space-like according as

$$m_n^2 + \mu^2 \gtrless 0.$$

For a time-like k_μ , one can choose a coordinate system, in which $n_0 = 0$. For a space-like k_μ , n may not simply a linear combination of n_i 's, but one can define because of the degeneracy m_n^2 is always degenerate, but n can still be

$$n = n_1^2 + n_2^2 + n_3^2 - n_0^2$$

in an appropriate coordinate system, in which one of the space axes coincide with k_μ .

Suppose, then, that

$$\varphi(X, r) = \sum_n \int V_{n, k_\mu} e^{i k_\mu X_\mu} \chi_{n, k_\mu}(r) (dk)^\mu,$$

where k_μ can take arbitrary vector. The field equations

$$\frac{\partial^2 \varphi(X, r)}{\partial X_\mu \partial X_\mu} + 4 \frac{\partial^2 \varphi}{\partial r_\mu \partial r_\mu} + \frac{4}{\lambda^4} r_\mu r_\mu \varphi + g \bar{\Psi}(X + \frac{1}{2}r) \Psi(X - \frac{1}{2}r) = 0$$

becomes

$$-k_\mu k_\mu V_{n, k_\mu} + m_n^2 V_{n, k_\mu}$$

$$+ \frac{g}{(4\pi)^4} \int \underbrace{e^{-i k_\mu X_\mu}}_{\varphi(-r)} \bar{\Psi}(X + \frac{1}{2}r) \Psi(X - \frac{1}{2}r) dX dr = 0$$

Then, let us introduce an equivalent local scalar field

$$V_n(x'') = \int V_{n, k_\mu} e^{i k_\mu x''_\mu} (dk)^\mu$$

$$\frac{\partial^2 V_n(x'')}{\partial x''_\mu \partial x''_\mu} = m_n^2 V_n(x'')$$

$$+ \frac{g}{(4\pi)^4} \int \Phi_n \left(\begin{matrix} x'' - X \\ X - x'' \end{matrix}, -r \right) \bar{\Psi}(X + \frac{1}{2}r) \Psi(X - \frac{1}{2}r) dX dr$$

where

$$\Phi_n(x'' - X, +r) = \frac{1}{(2\pi)^4} \int_{\text{space } k} e^{-i k_\mu (X - x'')} \tilde{\Phi}(-r)(dk)^4$$

$$g \int \Phi_n(x'' - X, +r) \bar{\psi}(x + \frac{1}{2}r) \psi(x - \frac{1}{2}r) dx dr$$

$$= g \int \Phi_n(\frac{x' + x''}{2} - X, x' - x'') \bar{\psi}(x') \psi(x'') dx' dx''$$

$$\left. \begin{aligned} x' &= x + \frac{1}{2}r \\ x'' &= x - \frac{1}{2}r \end{aligned} \right\}$$

$$\frac{\partial^2 V_n(x'')}{\partial x''^\mu \partial x''^\mu} - m_n^2 V_n(x'')$$

$$= -g \int \Phi_n(\frac{x' + x''}{2} - X, x' - x'') \bar{\psi}(x') \psi(x'') dx' dx''$$

This is equationivalent to the field equation with the form function Φ_n in the non-local interaction theory.

$$\begin{aligned}
 & \gamma_{\mu} \frac{\partial \psi(x')}{\partial x'_{\mu}} + M \psi(x') \\
 &= g \int (x'' | \varphi | x') \psi(x'') dx'' \\
 &= g \sum_n \left(v_{n, k_{\mu}} e^{i k_{\mu} x''} \chi_{n, k_{\mu}}(r) \psi(x'' - \frac{1}{2} r) \right. \\
 & \quad \left. x' = X + \frac{1}{2} r \right. \\
 & \quad \left. x'' = X - \frac{1}{2} r \right) \times dx'' \\
 & \quad x'' = X - \frac{1}{2} r
 \end{aligned}$$

Then, the general form for $\varphi(x, r)$ is

$$\begin{aligned} \varphi(x, r) &= \sum_{n, k, \mu} \overline{c_{n, k, \mu}} e^{i k r} \chi_{n, k, \mu}(x) \\ &= \sum_{n, k, \mu} \int e^{i k r} c_{n, k, \mu} \chi_{n, k, \mu}(r) \frac{1}{x} (dk)^3 \\ &= \sum_{n, j} \left(v_{n, k, \mu}^{(j)} e^{i k r} \chi_{n, k, \mu}^{(j)}(r) \right) (dk)^3 \end{aligned}$$

where j refers to different orthogonal states with the same $n = n_1 + n_2 + n_3 = n_0$.

$$\sum_{n, j} v_{n, k, \mu}^{(j)} \chi_{n, k, \mu}^{(j)}(r) = \sum_{n, k, \mu} c_{n, k, \mu} \chi_{n, k, \mu}(r)$$

$$v_{n, k, \mu}^{(j)} \chi_{n, k, \mu}^{(j)}(r) = \sum_{n, k, \mu} c_{n, k, \mu} \chi_{n, k, \mu}(r)$$

$$v_{n, k, \mu}^{(j)} = \sum_{n, k, \mu} c_{n, k, \mu} f(n, k, \mu) \quad (n_1 + n_2 + n_3 = n_0 = j)$$

$$\frac{\partial^2 \varphi_{n, k, \mu}(x)}{\partial x_\mu^2} - m_n^2 \varphi_{n, k, \mu}(x)$$

$$= -g \int \overline{\varphi}_{n, k, \mu}(x', x'', x''') \overline{\psi}(x') \psi(x'') dx' dx'' dx'''$$

$$\varphi_{n, k, \mu}(x) = \int e^{i k r} c_{n, k, \mu} (dk)^3$$

$(-k_\mu k_\mu + m_n^2) \varphi_{n, k, \mu}(x) = c_{n, k, \mu}$

$$= -g \int \frac{e^{-i k_\mu x''}}{(2\pi)^4} \Phi_{n\mu}(x', x'', x''') \bar{\psi}(x') \psi(x''') dx' dx'' dx'''$$

$$(-k_\mu k_\mu - m_n^2) v_{n, k_\mu}^{(j)}$$

$$= -g \sum_{n\mu}^{(j)} \int \frac{e^{-i k_\mu x''}}{(2\pi)^4} \Phi_{n\mu}(x', x'', x''') \bar{\psi}(x') \psi(x''') dx' dx'' dx'''$$

$$\left(\frac{\partial^2}{\partial x_\mu \partial x_\mu} - m_n^2 \right) \Phi_n(x, y)$$

$$= -g \sum_{n\mu}^{(j)} \int \frac{\Phi_{n\mu}(x', x'', x''') e^{i k_\mu (x'' - x''')} dk dx'' dx'''}{\bar{\psi}(x') \psi(x''')}$$

$$\gamma_\mu \frac{\partial \psi(x')}{\partial x'_\mu} + M \psi(x')$$

$$= g \sum_{n\mu}^{(j)} \Phi_{n\mu}(x', x'', x''') \int C_{n\mu, k_\mu} e^{i k_\mu x''} dk dx'' dx'''$$

$$v_{n, k_\mu}^{(j)} \chi_{n, k_\mu}^{(j)}(r) = \sum_{\substack{n=n_1, n_2, n_3, \dots, n_6 \\ l=0, 1, 2, \dots, l_0}} C_{n\mu, k_\mu} \chi_{n\mu}(r)$$

$$v_{n, k_\mu}^{(j)} \int_{n, n_\mu, k_\mu}^{(j)} = \int \chi_{n, k_\mu}^{(j)}(r) \chi_{n_\mu, k_\mu}^{(j)}(r) dr$$

$$C_{n\mu, k_\mu} = \int v_{n, k_\mu}^{(j)} \int_{n, n_\mu, k_\mu}^{(j)}$$

$$\begin{aligned}
& \gamma_{\mu} \frac{\partial \psi(x')}{\partial x'_{\mu}} + M \psi(x') \\
& = g \sum_{\nu, \mu} \int \mathbb{P}_{\nu\mu}(x', x'', x''') f_{\nu, \mu, k_{\nu}}^{(j)} v_{\nu, k_{\nu}}^{(j)} dt \\
& \quad e^{i k_{\nu} x''} \psi(x''') dx'' dx'''
\end{aligned}$$

= g

Direct Integration of
Field Equations

N.L.F. (27)

$$\frac{\partial^2 \varphi_{n\mu}(x'')}{\partial x''^\mu \partial x''^\mu} - m_{n\mu}^2 \varphi_{n\mu}(x'')$$

$$= -g \int \Phi_{n\mu}(x', x'', x''') \varphi(x') \psi(x''')$$

$$\times dx' dx'''$$

$$\Phi_{n\mu}(x', x'', x''') = \tilde{\chi}_{n\mu}(x' - x'')$$

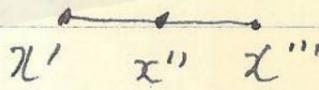
$$\times \delta\left(\frac{x' + x''}{2} - x'''\right)$$

$$\tilde{\chi}_{n\mu} \frac{\partial \psi(x')}{\partial x'^\mu} + M \psi(x')$$

$$= g \sum_{n\mu} \int \Phi_{n\mu}(x', x'', x''') \varphi_{n\mu}(x'') \psi(x''')$$

$$\times dx'' dx'''$$

etc



Since all functions $\Phi_{n\mu}(x', x'', x''')$ decrease rapidly as the distance, between x'' and x''' (and x'' and x') increases, in any direction, one can formally integrate the field equations as follows, according to Y.F.K. method:

$$\varphi_{n\mu}(x) = \varphi_{n\mu}^{(in)}(x) + g \int \Delta_{n\mu}^{(ret)}(x-x'') \times \bar{\Phi}_{n\mu}(x', x'', x''') \bar{\Psi}(x') \Psi(x''') dx' dx'' dx'''$$

$$\psi(x) = \psi^{(in)}(x) - g \int S_M^{(ret)}(x-x') \times \Phi_{n\mu}(x', x'', x''') \varphi_{n\mu}(x'') \psi(x''') dx' dx'' dx'''$$

$$\varphi_{n\mu}(x) = \varphi_{n\mu}^{(out)}(x) + g \int \Delta_{n\mu}^{(adv)}(x-x'') \times \bar{\Phi}_{n\mu}(x', x'', x''') \bar{\Psi}(x') \Psi(x''') dx' dx'' dx'''$$

$$\psi(x) = \psi^{(out)}(x) - g \sum_{n\mu} \int S_M^{(adv)}(x-x') \times \bar{\Phi}_{n\mu}(x', x'', x''') \varphi_{n\mu}(x'') \psi(x''') dx' dx'' dx'''$$

$$(\mathcal{D} - m_{n\mu}^2) \begin{pmatrix} \Delta_{n\mu}^{(ret)}(x-x') \\ \Delta_{n\mu}^{(adv)}(x-x') \end{pmatrix} = -\delta(x-x')$$

$$\left. \begin{aligned} \lim_{x_0 \rightarrow \mp \infty} \Delta_{n\mu}^{ret}(x-x') &= 0 \\ \lim_{x_0 \rightarrow \mp \infty} \frac{\partial \Delta_{n\mu}^{ret}(x-x')}{\partial x_0} &= 0 \end{aligned} \right\}$$

For $m_{np}^2 \geq 0$, the properties of $\Delta_{\kappa}^{\text{ret, adv}}$ are well known

$$\Delta_{\kappa}^{\text{ret}} = \bar{\Delta}_{\kappa} - \frac{1}{2} \Delta_{\kappa}$$

$$\Delta_{\kappa}^{\text{adv}} = \bar{\Delta}_{\kappa} + \frac{1}{2} \Delta_{\kappa}$$

with

$$\bar{\Delta}_{\kappa} = \frac{1}{(2\pi)^4} \int \frac{1}{k^2 + \kappa^2} e^{ikx} dk$$

$$\Delta_{\kappa} = \frac{1}{(2\pi)^3 i} \int \varepsilon(k) \delta(k^2 + \kappa^2) e^{ikx} dk$$

where

$$\kappa = m_{np}$$

For $m_{np}^2 < 0$, $\kappa = +\sqrt{-m_{np}^2}$,

$$\bar{\Delta}_{i\kappa} = \frac{1}{(2\pi)^4} \int \frac{1}{k^2 - \kappa^2} e^{ikx} dk$$

still satisfies ^{formally} the equation, ^{however,} if we have ^{to} first carry out the integral over angular variables of \vec{k} , and then take the principal value for the integration over $|\vec{k}|$, because the integrand has poles at $\pm\sqrt{k_0^2 + \kappa^2}$. Finally one integrates over k_0 . Similarly, for $\Delta_{i\kappa}$, one has to start from the

space interactions. Δ_{in} is no longer zero outside the light-cone. These and other things will give rise to further complications.

For that reason we had better restrict the eigenvalues of m_{in}^{-2} to non-negative values by choosing a suitable eigenvalue problem for the mass. For instance,

if $F^{(in)}$ has the form

$$\left(\frac{\partial}{\partial x^\mu \partial x^\mu} \right)^2 (\chi + \theta)^2$$

$$\left(\frac{\partial^2}{\partial x^\mu \partial x^\mu} - \frac{1}{\lambda^2} r_\mu r_\mu \right) \chi = \frac{m^2}{\lambda^2} \chi$$

or

$$\left(\frac{\partial^2}{\partial x^\mu \partial x^\mu} - \frac{1}{\lambda^2} r_\mu r_\mu \right) \chi = \frac{m + \frac{m}{\lambda}}{\lambda} \theta$$

$$\left(\frac{\partial^2}{\partial x^\mu \partial x^\mu} - \frac{1}{\lambda^2} r_\mu r_\mu \right) \theta = \mu \chi$$

$$\begin{aligned} & (\chi + \theta) = \mu (\chi + \theta) \\ & (\chi - \theta) = -\mu (\chi - \theta) \end{aligned}$$

N.L.F. (39)

Thus, if we take a complete set of eigenfunctions for the original oscillator problem, it is also the eigenfunctions for the new eigenvalue problem, which has, however, only positive eigenvalues.
(The infinite degeneracy still remains)

Now

$$m_{n_1} = I \frac{8}{\lambda} (n_1 + n_2 + n_3 - n_0 + 1)$$

If we restrict, in this way, the eigenvalues m_n to real values, we have further

$$\left. \begin{aligned} \varphi_{m_n}^{(out)}(x) &= \varphi_{m_n}^{(in)}(x) - g \int \Delta_{m_n}(x-x') \bar{\varphi}_{m_n}(x', x'', x''') \bar{\psi}(x') \\ &\quad \Delta_{m_n}(x-x') \psi(x''') dx' dx'' dx''' \end{aligned} \right\}$$

$$\left. \begin{aligned} \psi^{(out)}(x) &= \psi^{(in)}(x) + g \int \bar{\varphi}_{m_n}(x', x'', x''') \\ &\quad S_{m_n}(x-x') \varphi_{m_n}(x') \psi(x''') dx' dx'' dx''' \end{aligned} \right\}$$

As shown by Bloch, the incoming and outgoing fields are connected by a unitary real S-matrix by the

relation, ~~if~~

$$\varphi_0^{(out)}(x) = S^{-1} \varphi_0^{(in)}(x) S$$

if ~~we~~ we restrict our attention to $n_n=0$ and ignore all other states and ~~assume~~ quantize the fields $\varphi_0^{(out)}(x)$ as usual:

$$[\varphi_0^{(in)}(x), \varphi_0^{(in)}(x')] = i \Delta_{m_0}(x-x')$$

However, such a restriction to $n_n=0$ will inevitably violate the relativistic invariance.

We have at least to take into account all the states belonging to the same lowest eigenvalue $m = m_0$ in order to be ~~in~~ conform to relativity principle.

Then the question of convergence ~~could not~~ ^{could not} ~~be~~ ^{be} ~~again~~ ^{again} ~~be~~ ^{be} ~~easy~~ ^{easy} ~~to~~ ^{to} ~~be~~ ^{be} ~~answered~~ ^{answered}.

Perhaps From here on, we have to take up the restriction due to the S.C. again seriously.

Quantization of N.H.S.F. N.L.F. (40)

Now our problem is as follows:

Take an incoming field $\varphi_{\mu\nu}^{(in)}(x)$ which is so chosen as to satisfy the S.C. ~~Is it possible that is the~~ relation between ~~(out)~~ outgoing field $\varphi_{\mu\nu}^{(out)}(x)$, which would be obtained by solving the integral equations, related to ~~the~~ $\varphi_{\mu\nu}^{(in)}(x)$ by a unitary matrix

$$\varphi_{\mu\nu}^{(out)}(x) = S^{-1} \varphi_{\mu\nu}^{(in)}(x) S,$$

so that $\varphi_{\mu\nu}^{(out)}(x)$ satisfies ^{not only} the same field equation as $\varphi_{\mu\nu}^{(in)}(x)$, but also it satisfies the same commutation relations and the same supplementary condition?

To answer this question is not easy, because the commutation relations for $\varphi_{\mu\nu}^{(in)}(x)$ is to be modified ~~to~~ due to S.C.

In any case, we have to investigate the C.R. for $\varphi_{\mu\nu}^{(out)}(x)$ in detail.

First of all, the general form for the field $\phi(x, r)$ which satisfies both the supplementary condition and the field equation is:

$$\phi(x, r) = \sum_{n,j} v_{k_n}^{(n,j)} e^{i k_n x} \chi_{k_n}^{(n,j)}(r) \times \delta(k_n^2 + m_n^2) d^4 k$$

where n is the quantum which characterizes the mass m_n and j is that what discriminate orthogonal ~~states~~ internal eigenfunction belong to the same n or same m_n . $v_{k_n}^{(n,j)}$ are is an arbitrary coefficient.

Although $\chi_{k_n}^{(n,j)}(r)$ can be normalized in r -space, two such functions with the same n, j and different k_n are not orthogonal to each other. ~~And~~ let us take the case $n_1 = n_2 = n_3 = 0$ ($n_0 = 0$) without k_n . In this case, the extra

quantum number j is suppressed,

$$\chi_{k\mu}^{(0)}(r) = N_{k\mu} e^{-\frac{1}{2\lambda^2} \left\{ r_\mu r_\mu + \frac{2(k_\mu r_\mu)^2}{m_0^2} \right\}}$$

where $N_{k\mu}$ is the normalization factor,

$$\int |\chi_{k\mu}^{(0)}(r)|^2 (dr)^4 = |N_{k\mu}|^2 \int e^{-\frac{1}{\lambda^2} \left\{ r_\mu r_\mu + \frac{2(k_\mu r_\mu)^2}{m_0^2} \right\}} \times (dr)^4$$

$$= |N_{k\mu}|^2 \int e^{-\frac{1}{\lambda^2} (r_1'^2 + r_2'^2 + r_3'^2 + r_0'^2)} \times dr_1' dr_2' dr_3' dr_0'$$

$$= |N_{k\mu}|^2 \left(\sqrt{\frac{\pi \lambda^2}{2}} \right)^4 = |N_{k\mu}|^2 \left(\frac{\pi \lambda^2}{2} \right)^2$$

or $= 1$

or $N_{k\mu} = \frac{1}{\sqrt{\pi \lambda^2}}$

$$\int \chi_{k\mu}^{(0)}(r) \chi_{k\mu}^{(0)}(r) (dr)^4 = \left(\frac{1}{\pi \lambda^2} \right)^2 \int e^{-\frac{1}{\lambda^2} \left\{ r_\mu r_\mu + \frac{(k_\mu r_\mu)^2 + (k'_\mu r_\mu)^2}{m_0^2} \right\}} (dr)^4$$

If we go over to the centre of mass

system, in which

$$k'_1 = -l'_1, \quad k'_0 = l'_0$$

$$\frac{(k'_1 r'_1 + k'_0 r'_0)^2 + (-k'_1 r'_0 - k'_0 r'_1)^2}{m_0^2}$$

$$= 2(k_1'^2 r_1'^2 + k_0'^2 r_0'^2) / m_0^2$$

$$r'_\mu r'_\mu + \frac{2}{m_0^2} (k_1'^2 r_1'^2 + k_0'^2 r_0'^2)$$

$$= \left(1 + \frac{2k_1'^2}{m_0^2}\right) r_1'^2 + r_2'^2 + r_3'^2 + \left(1 + \frac{2k_0'^2}{m_0^2}\right) r_0'^2$$

Thus

$$\int \chi_{k'_\mu}^{(\omega)}(r') \chi_{l'_\mu}^{(\omega)}(r') d^3r'$$

$$= \frac{1}{\sqrt{\left(1 + \frac{2k_1'^2}{m_0^2}\right) \left(1 + \frac{2k_0'^2}{m_0^2}\right)}}$$

$$= \frac{1}{\sqrt{\left(1 + \frac{2k_1'^2}{m_0^2}\right) \left(1 + \frac{2k_0'^2}{m_0^2}\right)}} = \frac{1}{1 + \frac{2k_1'^2}{m_0^2}}$$

$$= \frac{1}{\frac{2k_0'^2}{m_0^2} - 1} = \frac{m_0^2}{k_0 l_0 - k_j l_j} = \frac{m_0^2}{-k_j l_j}$$

N.L.F. (42)

Prop: i

in the centre of mass system

$$\frac{k_0' - l_0'}{2} = \frac{k_0}{m_0} \left[(k_1 + l_1)^2 + (k_2 + l_2)^2 + l^2 \right] - (k_0 + l_0)^2 = \kappa^2$$

$$k_0' = - \frac{k_1 + l_1}{\kappa} k_1 - \frac{k_2 + l_2}{\kappa} k_2$$

$$- \frac{k_3 + l_3}{\kappa} k_3 + \frac{k_0 + l_0}{\kappa} k_0$$

$$= \frac{m_0^2 + (k_1 l_1 + k_2 l_2 + k_3 l_3 - k_0 l_0)}{\kappa}$$

$$= l_0' \neq$$

$$(k_1 + l_1)^2 + (k_2 + l_2)^2 + (k_3 + l_3)^2 - (k_0 + l_0)^2$$

$$\Rightarrow 2m_0^2 + 2(k_1 l_1 + k_2 l_2 + k_3 l_3 - k_0 l_0)$$

$$= \kappa^2$$

$$\kappa^2 = 2m_0^2 - 2(k_1 l_1 + k_2 l_2 + k_3 l_3 - k_0 l_0)$$

$$k_0' = \frac{1}{\sqrt{2}} \sqrt{m_0^2 + (k_0 l_0 - k_j l_j)}$$

$$k_\mu = l_\mu: \quad k_0' = m_0$$

$$\frac{2k_0'^2}{m_0^2} - 1 = \frac{m_0^2 + (k_0 l_0 - k_j l_j)}{m_0^2} - 1$$

$$= (k_0 l_0 - k_j l_j) / m_0^2$$

On the other hand, φ can be expanded in the form

$$\varphi(X, r) = \sum_{n, \mu} \varphi_{n, \mu}(X) \chi_{n, \mu}(r)$$

where

$$\varphi_{n, \mu}(X) = \sum_{n_1, j_1} \dots \sum_{n_j, j_j} v_{n, j}^{(n, j)} e^{i k_{n, j} X} \frac{\delta(k_{n, j} + m_n)}{(d k)^4}$$

$$\times \int \chi_{n, \mu}(r) \chi_{k, \mu}^{(n, j)}(r) (dr)^4$$

Now the last $\chi_{n, \mu}(r)$ and $\chi_{k, \mu}^{(n, j)}$ are orthogonal to each other, if

$$n \neq n_1 + n_2 + n_3 - n_0,$$

because they belong to different eigenvalues of the same oscillator problem with the same eigenvalues.

$$C_{n, \mu, k, \mu} = \int \chi_{n, \mu}(r) \chi_{k, \mu}^{(n, j)}(r) (dr)^4$$

~~Let us take again the case $n=0$. Then the eigenfunction $\chi_{n, \mu}$ which gives non-vanishing~~

N.L.F. (43)

For $C_{n\mu, k\mu}^{(n,j)}$, we have

$$\sum_{n,j} |C_{n\mu, k\mu}^{(n,j)}|^2 = 1$$

provided that $\chi_{n\mu}$ is normalized,

because

$$\chi_{k\mu}^{(n,j)}(r) = \sum_{n\mu} C_{n\mu, k\mu}^{(n,j)} \chi_{n\mu}(r)$$

$$1 = \sum_{n\mu} C_{n\mu, k\mu}^{(n,j)} \int \chi_{k\mu}^{(n,j)}(r) \chi_{n\mu}(r) dr$$

$$\varphi_{n\mu}(x) = \sum_{n,j} \int_{k\mu} e^{ik_{\mu}x} C_{n\mu, k\mu}^{(n,j)} d(k_{\mu} + \mu_n) (dk)^\mu$$

also we have

$$\sum_{n,j} |C_{n\mu, k\mu}^{(n,j)}|^2 \leq 1$$

because $\{\chi_{k\mu}^{(n,j)}(r)\}$ is ~~not~~ ^{may} not a complete set of orthogonal functions for a fixed k_{μ} , because we have chosen

~~$\chi_{n\mu}(r)$~~ ~~$\chi_{k\mu}(r)$~~

only these functions which satisfy S.C.

$$\chi_{n\mu}(r) = \sum_{(n,j)} C_{n\mu, k_{n,j}} \chi_{k_{n,j}}^{(n,j)}(r) + \chi'_{n\mu, k_{n,j}}(r)$$

where $\chi'_{n\mu, k_{n,j}}(r)$ is the residual which does not satisfy S.C., and which is orthogonal to all of $\chi_{k_{n,j}}^{(n,j)}$.

In this sense, $\phi_{n\mu}(x)$ if $v_{k_{n,j}}^{(n,j)}$ satisfies the usual cond. R. of the type

$$[v_{k_{n,j}}^{(n,j)}, v_{k'_{n,j}}^{(n',j')}] = \text{const.} \cdot \delta_{nn'} \delta_{jj'} \delta(k_{n,j} + k'_{n',j'})$$

$$[\phi_{n\mu}(x), \phi_{n'\mu'}(x)] = \int e^{i k_{n,j} x} \cdot e^{i k'_{n',j'} x'} dk dk'$$

$$\sum_{(n,j)} [v_{k_{n,j}}^{(n,j)}, v_{k'_{n',j'}}^{(n',j')}] |C_{n\mu, k_{n,j}}|^2 \delta(k_{n,j} + k'_{n',j'})$$

$$= \int_{\text{const.}} e^{i k_{\mu} (x_{\mu} - x'_{\mu})} \sum_{n,j} |c_{n,j} k_{\mu}|^2 dk$$

$$\frac{1}{2(k_{\mu}) \delta(k_{\mu} k_{\mu} + m_n^2)}$$

Thus, the commutation relations for $\varphi_{n\mu}(x)$, $\varphi_{n\mu}(x')$ will take the usual form only if

$$\sum_{n,j} |c_{n,j} k_{\mu}|^2 = 1$$

independent of k_{μ} .

In other words, $\varphi_{n\mu}(x)$ and $\varphi_{n\mu}(x')$ may not be commutative for two x, x' which are space-like to each other.

Also, $\varphi_{n\mu}(x)$, $\varphi_{n'\mu}(x')$ are not in general commutative even for different n, n' and n, n' .

On the contrary, if we assume usual commutation relations for $\varphi_{n\mu}(x)$, in order to obtain a unitary S-matrix,

then the C. R. for $\psi_{k_{\mu}}^{(n,j)}$ will be different:

$$\varphi_{n_{\mu}}(X) = \sum_{n,j} \int \psi_{k_{\mu}}^{(n,j)} C_{n_{\mu}, k_{\mu}}^{(n,j)} e^{i k_{\mu} X_{\mu}} \delta(k_{\mu} k_{\mu} + m_n^2) (d k)^4$$

$$\rightarrow k^2 + k_0^2 - m_n^2 = X$$

$$k_0 = \pm \sqrt{X + k^2 + m_n^2}$$

$$2 k_0 dk_0 = dX$$

$$\int_{-\infty}^{+\infty} f(k_0) \delta(k_{\mu} k_{\mu} + m_n^2) d k_0 = \int_0^{\infty} f(\sqrt{X}) \frac{1}{2\sqrt{X}} \delta(X) dX$$

$$+ \int_0^{\infty} f(-\sqrt{X}) \frac{1}{2\sqrt{X}} \delta(X) dX$$

$$+ \int_0^{\infty} f(-\sqrt{X}) \frac{1}{2\sqrt{X}} \delta(X) dX$$

$$\varphi_{n_{\mu}}(X) = \sum_{n,j} \int \psi_{k_{\mu}, k_0}^{(n,j)} C_{n_{\mu}, k_{\mu}, k_0}^{(n,j)} e^{i k_{\mu} X_{\mu}} \delta(k_{\mu} k_{\mu} + m_n^2) (d k)^3$$

$$+ \psi_{k_{\mu}, k_0}^{(n,j)} C_{n_{\mu}, k_{\mu}, k_0}^{(n,j)} e^{i k_{\mu} X_{\mu}} \delta(k_{\mu} k_{\mu} + m_n^2) (d k)^3$$

$$k_0 = +\sqrt{k^2 + m_n^2}$$

From the commutation relation
 $[\varphi_{n_{\mu}}(X), \varphi_{n_{\mu}}(X')] = i \Delta_{m_n}(X-X')$

we obtain

$$\left\{ \sum_{(n,j)} v_{\mu, k, +k_0}^{(n,j)} + k_0 C_{\mu, k, +k_0}^{(n,j)}, \right. \\ \left. \sum_{(n',j')} v_{\mu, k', -k'_0}^{(n',j')} - k'_0 C_{\mu, k', -k'_0}^{(n',j')} \right\} \\ = \text{const. } \delta(k, k')$$

If, $\varphi_{n,j}(x)$ is real,

$$C_{\mu, k, +k_0}^{(n,j)} = C_{\mu, -k, -k_0}^{(n,j)}$$

~~$$\int C_{\mu, k, +k_0}^{(n,j)} v_{\mu, k, +k_0}^{(n,j)} C_{\mu, k, +k_0}^{(n,j)}$$~~

$$\sum_{\substack{(n,j) \\ (n',j')}} C_{\mu, k, +k_0}^{(n,j)} C_{\mu, -k', -k'_0}^{(n',j')} \left[v_{k, +k_0}^{(n,j)}, v_{k', -k'_0}^{(n',j')} \right] \\ = \text{const. } \delta(k, k')$$

$$\sum_{\substack{(n,j) \\ (n',j')}} f(n, j, k, -k') \left[v_{k, +k_0}^{(n,j)}, v_{-k', -k'_0}^{(n',j')} \right] \\ = \text{const. } \delta(k, k')$$

~~$$\left[v_{k, +k_0}^{(n)}, v_{-k', -k'_0}^{(n)} \right]$$~~

Thus, the C.R. for $V_{ij}^{(ij)}$ will
~~be singular, does not~~ not leave
such a form as to allow the
the interpretation:

usual $V + \kappa_0$: annihil. operator

$V - \kappa_0$: creat. operator

~~The real annihilation~~

In any case, if we started from
the integral equations for
the fields with the initial
incoming field satisfying
the S.C., the outgoing field
which would be obtained by
as the result of successive
integration, would not satisfy
the S.C. in other words

words, the unitary S-matrix
which connect the in- and
out- fields must be obtained
by some other way other than
the direct integration of field
equations.

thought

N. h. I. (46)

It might be possible to modify the S.C. However, the mere change in S.C. does not help addition of an interaction term much, unless the spinor field is restricted at the same time by another S.C.

Or else, one can say that the field equations themselves are already so much modified by S.C. that there remains a little similarity between them and the field equations for N. h. I. ~~to~~

As for the method of constructing S.M., ~~independently of the~~ ^{by giving up deducing the} ~~field equations~~ ^{from variation principle} was suggested ~~by~~ in my earlier papers.

One may also go over first to the moment space as rep. of Lagrangian Action Operator.

Summary of the course
of A. Q. M. for Winter
Semester Sept. 1952
~ Jan. 1953

A. General formulation
of Relativistic Field
Theories of

~~local~~ local fields

- (a) Variation principle.
- (b) Direct integration
of Field Equations.
- (c) Structure of S-matrix
- (d) Divergence - Renormalization
- (e) A. F. D. (Supp. Cond.)
- (f) General Rules of Constructing
S-matrix

B. ~~local~~ Non-local Field Interaction
Field Theories

- (a) N. L. Interaction
- (b) N. L. Fields
mass-spectrum versus
form factor.

There are a number of problems which were not discussed in detail: Among other things (for collision process)

- (i) Computation of physically observable quantities
- (ii) Problems of bound states

However, I think I could cover the most fundamental problems of the present field theories and, in particular, their most serious difficulties.

Program of Spring Semester

Application of Group Theory to Phys. Quantum Mechanics

- (i) permutation group
- (ii) rotation group
- (iii) Lorentz group
- (iv) possible other groups of possible interest.

S.Z

In the meantime, I hope I shall be able to discuss further developments in field theories.

In any case, I think I better change the subject

Another possibility
More elementary approach to Meson-Nucleon Problem without bothering too much about relativistic formulation of the theory.

Weyl
van der Waerden
Wigner

Carson, Perturbation Methods in the Q. M. of N. Electron Systems, 1951

Murugan, Theory of G. Repr.

Variation Principle

with constraint

$$F\left(\frac{\partial}{\partial x_\mu}, r_\mu, \frac{\partial}{\partial r_\mu}\right)\varphi(x, r) = 0$$

$$C\left(\frac{\partial}{\partial x_\mu}, r_\mu, \frac{\partial}{\partial r_\mu}\right)\varphi(x, r) = 0$$

F, C; self-adjoint operators

$$\delta \int \varphi \cdot F \varphi dx dr = 0$$

$$\delta \int \varphi \cdot C \varphi dx dr = 0$$

μ : Lagrange multiplier

$$\delta \left[\int \varphi F \varphi + \mu \int \varphi \cdot C \varphi \right] = \delta I = 0$$

for arbitrary change of φ and μ .
 φ must satisfy the eq.

$$F\varphi + \mu C\varphi = 0$$

μ : certain positive constant.

~~where μ is to be so chosen as to~~

~~give φ where μ is to be so chosen as to give~~

$$\int \varphi \cdot C \varphi = 0$$

positive definite operator
 for tunneling term

$$\varphi(x, r) = \int_{k_\mu} C_{k_\mu, \mu} e^{i k_\mu x_\mu} \chi_{k_\mu}(x) (dk)^\mu$$

$$(-\hbar^2 k_\mu^2 - m_n^2) C_{k_\mu, \mu} = \mu \left(\frac{\hbar^2 k_\mu^2 + m_n^2}{\sqrt{2\pi} \hbar k_\mu} \right) C_{k_\mu, \mu} = 0$$

For $k_1 = k_2 = k_3 = 0$:

$$(k_0^2 - m_n^2) C_{0,0,0,k_0; n_1, n_2, n_3, n_0}$$

$$\mu m_n C_{0,0,0,k_0; n_1, n_2, n_3, n_0} = 0$$

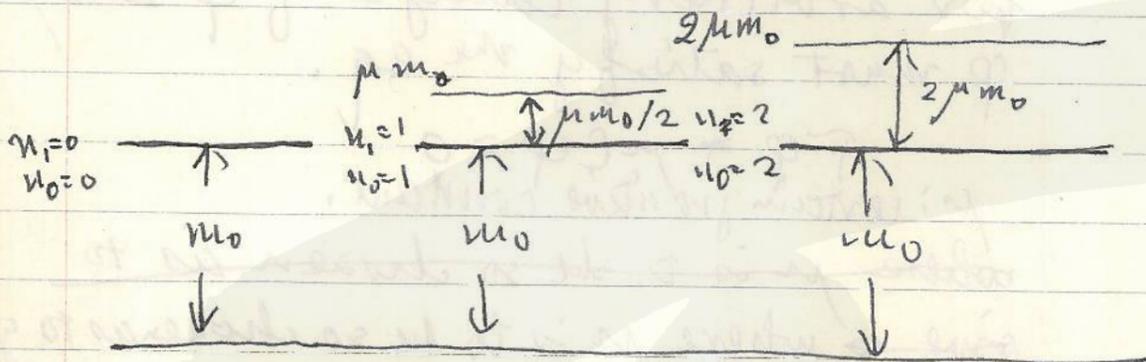
If $k_0^2 = m_n^2$: $n_0 = 0$

or $k_0^2 = 0$ otherwise

$$(k_0^2 - m_n^2 - \mu m_n^2 n_0) C_{0,0,0,k_0; n_1, n_2, n_3, n_0} = 0$$

$$k_0 = \pm \sqrt{m_n^2 + \mu m_n^2 n_0} = m_0 \sqrt{1 + \mu n_0}$$

$n_1 = 1$
 $n_0 = 0$



$$k_0 = +\sqrt{m_0^2 + \mu m_0 n_0}$$

$$= m_0 \sqrt{1 + \frac{\mu n_0}{m_0}} \approx m_0 + \frac{\mu m_0 n_0}{2}$$

If μ is positive, we have a non-degenerate lowest mass state $n_0^* = n_1 = n_2 = n_3$ for the particle at rest.

N.L.F. (48)

$$a(n_1 + n_2 + n_3 - n_0)^2 + b(n_1 + n_2 + n_3)n_0$$

$$= a(n_1 + n_2 + n_3)^2 - 2af$$

$$= an^2 - 2an n_0 + an_0^2$$

$$+ b n n_0 - b n_0^2$$

$$a=1, \quad b=\mu$$

$$n^2 - (2-b)nn_0 + n_0^2$$

$$= \left(n - \left(1 - \frac{b}{2}\right)n_0 \right)^2$$

$$+ \left(1 - \frac{b}{2}\right)^2 n_0^2$$

$$-1 \leq 1 - \frac{b}{2} \leq 1$$

$$\frac{b}{2} \leq 2 \quad b \leq 4$$

$$\frac{b}{2} \geq 0 \quad b \geq 0$$

$$0 \leq \frac{b}{2} \leq 2$$

$$0 \leq \frac{b}{4} \leq 1$$

$$\boxed{0 \leq b \leq 4}$$

||
μ

$$\mu = 4, \quad (n + n_0)^2 = (u_1 + u_2 + u_3 + u_0)^2$$

$$u_1 = 2, \quad \underline{u_1 = 1, u_2 = 1}$$

$$u_1 = 1, \quad \underline{\underline{\hspace{2cm}}}$$

$$u_m = 0, \quad \underline{\hspace{2cm}}$$

$$(n_1 + n_2 + n_3 - n_0)^2$$

$$+ \mu (n_1 + n_2 + n_3 - n_0) n_0$$

$$= (n - n_0)^2 + \mu (n n_0 - n_0^2)$$

$$= n^2 - (2 - \mu) n n_0 + \mu n_0^2$$

$$= (n - (1 - \frac{\mu}{2}) n_0)^2 + (1 - \frac{\mu}{2})^2 n_0^2$$

$$1 + \mu - (1 - \frac{\mu}{2})^2 > 0$$

$$(\frac{\mu}{2})^2 < 0 \quad \times$$

N.L.F. (49)

$$-\left(\frac{\partial^2}{\partial x_\mu \partial x_\mu} + \frac{1}{\lambda^2} r_\mu\right) \left(-\frac{\partial^2}{\partial x_\nu \partial x_\nu} + \frac{1}{\lambda^2} r_\nu\right) \varphi = 0$$

$$\varphi = e^{i k_\mu x_\mu} \chi_{k_\mu}(r)$$

$$-i k_\nu \left(\frac{\partial}{\partial x_\nu} + \frac{1}{\lambda} r_\nu\right) \cdot i k_\mu \left(\frac{\partial}{\partial x_\mu} + \frac{1}{\lambda} r_\mu\right)$$

" " " "

$i k_0$ " " $i r_0$

k_μ : time-like

$$-i k'_0 b'_0{}^* \cdot i k'_0 k'_0 = k'_0{}^2 b'_0{}^* b'_0 \quad (\text{pos. op.})$$

k_μ : space-like

$$-i k'_j b'_j{}^* \cdot i k'_j k'_j = k'_j{}^2 b'_j{}^* b'_j \quad (\text{pos. op.})$$

Quantization:
 Interaction:

Free field can be expanded into series of eigenfunctions of the e.v.p.

$$(F - \mu C) \varphi = \kappa \varphi$$

Thus, we have an eigenvalue problem of κ in general, with two numerical

for constants in addition to the fundamental length λ ,

$$\varphi(x, r) = \sum_{n, n_0, j} \int_{k_n, n}^{v_0, j} e^{i k_n x + i \omega_n t} \chi_{k_n, n}^{n_0, j}(r)$$

For the free field, for which $(\square - \mu^2)\varphi = 0$

is satisfied, all k_n must be ~~the~~ time-like with the factor $\delta(k_n^2 + m_n^2)$.

If there is the interaction with another field, k_n can no longer be restricted to the ~~wave~~ time-like vector. For a space-like vector k_n , $\chi_{k_n, n}^{n_0, j}(r)$ still is a function which decreases with $v_1^2, v_2^2, v_3^2, v_0^2$
 increasing

$$k_n k_n + m_n^2 + \mu m_n \left(n_j + \frac{1}{2} \right) = \kappa.$$

$$\begin{aligned} & (n-n_0)^2 + \mu(n-n_0)n_0 \\ &= n^2 - 2nn_0 + \underbrace{\mu nn_0}_{+n_0^2} - \mu n_0^2 \\ &= n^2 - (2-\mu)nn_0 + (1-\mu)n_0^2 \\ &= \left(n - \left(1 - \frac{\mu}{2}\right)n_0\right)^2 \\ &+ \left\{ (1-\mu) - \left(1 - \frac{\mu}{2}\right)^2 \right\} n_0^2 \\ & \quad 1-\mu - 1 + \mu - \frac{\mu^2}{4} \end{aligned}$$

$$\begin{aligned} & (n_1 + n_2 + n_3 - 2n_0)^2 (1 + \mu n_0) \\ & \quad \left(1 + \mu \left(n_j + \frac{1}{2}\right)\right) \end{aligned}$$

Direct Integration of F. E.

05c

$$\left\{ \frac{\partial^2}{\partial x_\mu \partial x_\mu} - F\left(\frac{\partial}{\partial x_\mu}, r_\mu, \frac{\partial}{\partial r_\mu}\right) \right\} \varphi(x, r)$$

$$= -g \bar{\Psi}\left(x + \frac{1}{2}r\right) \Psi\left(x - \frac{1}{2}r\right)$$

$$\gamma_\mu \frac{\partial \Psi(x')}{\partial x'_\mu} + M \Psi(x')$$

$$= g \int (x'' | \varphi | x') \Psi(x'') dx''$$

$$\left\{ \frac{\partial^2}{\partial x_\mu \partial x_\mu} - F(\dots) \right\} \chi_{k_\mu}^{(n'_\mu)}(x, r)$$

$$= \gamma \chi_{k_\mu}^{(n'_\mu)}(x, r)$$

$$\varphi(x, r) = g \int G(x, r; x', r') dx' dr'$$

$$\underbrace{\varphi^{(0)}(x, r)} + \bar{\Psi}\left(x' + \frac{1}{2}r'\right) \Psi\left(x' - \frac{1}{2}r'\right)$$

$$\} G(x, r; x', r') = -\delta(x-x') \delta(r-r')$$

$$G(x, r; x', r') = \sum_{n'_\mu} \int \frac{\chi_{k_\mu}^{(n'_\mu)}(x, r) \chi_{k_\mu}^{(n'_\mu)}(x', r')}{\gamma} dk$$

$$+ M \sum_{(P)} \int \chi_{k_\mu}^{n'_\mu}(x, r) \chi_{k_\mu}^{n'_\mu}(x', r') \delta(\alpha) dk$$

$$\gamma = \gamma_1 k_p k_m + \gamma_2$$

$$\gamma_2 \geq 0, \quad \gamma_1 \leq 0$$

$(\gamma_1 - 1) k_p k_m$ according as
 \geq odd always.

or $\gamma_1 < 1$ for $k_p k_m < 0$

$\gamma_1 > 1$ for $k_p k_m > 0$

$$\chi_{k_p}^{(n)}(X, r) = \int e^{i k_p X} \chi_{k_p}^{(n)}(r) \times d k_p$$

$$G(X, r; X', r') = \int_P d k \frac{\chi_{k_p}^{(n)}(X, r) \chi_{k_p}^{(n)}(X', r')}{\gamma_1 k_p k_m + \gamma_2}$$

$$= \int d k_0 e^{-i k_0 (X_0 - X'_0)} e^{i k (X - X')} \times \frac{1}{\gamma_1 (k^2 - k_0^2) + \gamma_2} \times \chi_{k_p}^{(n)}(r) \chi_{k_p}^{(n)}(r') \times d k$$

$$\left(\frac{1}{\gamma_1 (k^2 - k_0^2) + \gamma_2} * \mu \delta(\gamma_1 (k^2 - k_0^2) + \gamma_2) \right)$$

Quantum Mechanical

V.P. (1)

Variation Principle

$$I = \iint L(\varphi(x', x'')) dx' dx''$$

$$\varphi(x', x'') = \sum a_n \chi_n(x', x'')$$

χ_n : x -dimensional eigenfunctions
 $\delta I = 0$ of a certain eigenvalue problem.

$$\frac{\partial}{\partial a_n} I = \iint \frac{\partial}{\partial a_n} L(a) dx' dx''$$

$$= 0$$

$$[\hat{a}_n, I] = 0 \quad (b_n \equiv -\frac{\partial}{\partial a_n})$$

$$[\hat{a}_n, \hat{a}_n] = \delta_{nn}$$

classical variation principle is equivalent to the operator relation

$$[\hat{a}_n, I] = 0,$$

if the a_n in I are all commutative with each other.

Generalization to q.m. would be as follows:

$$[\hat{a}_n, I(a_n, \hat{b}_n)] = 0 \quad \}$$

$$[a_n, I(a_n, \hat{b}_n)] = 0$$

where a_n, \hat{b}_n could be identified with annihilation and creation

operators ~~from~~ if we choose

$$a_n^* = b_n$$

then we obtain, from

$$L = \sum_{n, n'} G_{nn'} a_n^* a_{n'} \quad \tilde{G}_{nn'} = G_{n'n}$$

$$[a_n^*, L] = \sum_{n'} G_{nn'} a_{n'} = 0$$

$$[a_n, L] = - \sum_{n'} a_n^* G_{n'n} = 0$$

$$G_{nn'} = \int \tilde{c}_n c_{n'} \tilde{\chi}_n(x', x'') \chi_{n'}(x', x'') dx' dx''$$

$\tilde{F}(\frac{\partial}{\partial x'}, \frac{\partial}{\partial x''}, x' - x'')$

$$\tilde{F} \chi_n = f_n \chi_n$$

$$G_{nn'} = |c_n|^2 f_n \delta_{nn'}$$

$$|c_n|^2 f_n a_n = 0$$

$$a_n^* |c_n|^2 f_n = 0$$

These equations are satisfied only for such c_n :

$$c_n = 0 \quad \text{for } f_n \neq 0$$

$$c_n \neq 0 \quad \text{for } f_n = 0$$

$$c_n = c_n' \delta_{fn, 0}$$

v. P. ②

There is an essential difference between the field equations and commutation relations as operator relations ~~are~~ the ordinary formulation as the above field eq. ~~and c. r. for the b-eq~~ and ~~c. r.~~ for a_n, a_n^* , which are to be regarded as the conditions satisfied by the Schrödinger function Ψ

$$[a_n, L]\Psi = [c_n]^2 f_n a_n \cdot \Psi = 0$$

$$[a_n^*, L]\Psi = a_n^* [c_n]^2 f_n \cdot \Psi = 0$$

while ~~the~~ c. r.

$$[a_n, a_n^*] = \delta_{nn'}$$

are regarded as operator relations.

~~For that reason, the relations like~~

$$[a_n^* [a_n^*, L]] = 0$$

~~is not true~~

Interaction

$$L = \sum_{n, n'} G_{n, n'} a_n^\dagger a_{n'} + \sum_{m, m'} H_{m, m'} b_m^* b_{m'}$$

$$+ \sum_{n, m, m'} J_{n, m, m'} a_n \cdot b_m^* b_{m'} + \sum_{n, m, m'} \tilde{J}_{n, m, m'}^* a_n^* b_m b_{m'}$$

$$\sum_{m'} G_{n,n'} a_{n'} + \sum_{m,m'} \tilde{J}_{n,m',m} b_m^* b_{m'} = 0$$

$$\sum_{n'} G_{n',n} a_{n'}^* + \sum_{m,m'} J_{n,m,m'} b_m b_{m'}^* = 0$$

$$[b_m^*, b_{m'}^*]_+ = \delta_{m,m'}$$

$$[b_m, b_{m'}^*]$$

$$[b_m^*, b_m b_{m'}^*]_+ = [b_m^*, b_{m'}^*]_+ b_m + b_m^* [b_m^*, b_{m'}]_+ = \delta_{m,m'} b_{m'}$$

$$\sum_{m'} H_{mm'} b_{m'} + \sum_{n,m,m'} J_{n,m,m'} a_n b_{m'} + \sum_{n,m'} \tilde{J}_{n,m',m} a_n^* b_{m'} = 0$$

Green's functions:

$$\sum_{n'} K_{n',n} G_{n,n'} = \delta_{n',n}$$

$$a_n + \sum_{m,m'} K_{n',n} \tilde{J}_{n,m',m} b_m^* b_{m'} = 0$$

$$|c_n|^2 f_n a_n + \sum_{m, m'} \tilde{J}_{n, m, m'} b_m^* b_{m'} = 0$$

$$|c_n|^2 f_n a_n^* + \sum_{m, m'} \tilde{J}_{n, m, m'} b_m^* b_{m'}^* = 0$$

$$\begin{aligned} & [b_m^* b_{m'}, b_{m''}^* b_{m''}] \\ &= [b_m^* b_{m'}, b_{m''}^*] b_{m''} \\ &+ b_{m''}^* [b_m^* b_{m'}, b_{m''}] \\ &= \{ [b_m^*, b_{m''}^*] + b_{m''}^* \\ &+ b_m^* [b_{m'}, b_{m''}^*] \} b_{m''} \\ &+ b_{m''}^* \{ b_m^* [b_{m'}, b_{m''}] + \\ &- [b_m^*, b_{m''}] b_{m'} \} \\ &= + b_m^* b_{m''} \delta_{m' m''} - b_{m''}^* b_{m'} \delta_{m m''} \end{aligned}$$

$$(|c_n|^2 f_n)^2 \delta_{nn'}$$

$$\begin{aligned} & + \sum_{m, m'} \tilde{J}_{n, m, m'} \tilde{J}_{n', m', m''} (b_m^* b_{m''} \delta_{m' m''} \\ & - b_{m''}^* b_{m'} \delta_{m m''}) b_m^* b_{m'} \\ & \equiv \sum_{m, m'} \tilde{J}_{n, m, m'} \tilde{J}_{n', m', m''} b_m^* b_{m'} - \tilde{J}_{n, m, m'} \tilde{J}_{n', m', m''} \end{aligned}$$

$$\sum_m L_{m''m'} H_{m''m'} = \delta_{m'',m'}$$

$$b_{m''} + \sum J_{n, m''} L_{m'', m'} J_{n, m', m''} a_n b_{m'} \\
 + \sum L_{m'', m'} \tilde{J}_{n, m', m''} a_n^* b_{m'} = 0$$

$$K_{n'', n'} = \tilde{L}_n$$

$$G_{n, n'} = \tilde{L}_n \tilde{L}_{n'} \tilde{X}_n$$

$$\sum \tilde{J}_{n, m', m''} \tilde{J}_{n', m'', m'''} b_m^* b_{m'''}$$

$$- \sum \tilde{J}_{n, m'', m'''} \tilde{J}_{n', m', m''} b_m^* b_{m'''}$$

General Formulation

Free non-local field

$$\varphi(x, y) \equiv (x' | \varphi | x'')$$

$$\bar{I} = \int d^4x d^4y \left(\frac{1}{2} \frac{\partial \tilde{\varphi}(x, y)}{\partial x_\mu} \frac{\partial \varphi(x, y)}{\partial x_\mu} + \text{pos. definite quadratic form of } \varphi \right)$$

$$\frac{\partial}{\partial x_\mu} \left(-\lambda \frac{\partial}{\partial y_\mu} + \frac{1}{\lambda} \gamma_\mu \right) \varphi$$

$$F_2 = \frac{\partial}{\partial x_\mu} \left(-\lambda \frac{\partial}{\partial y_\mu} + \frac{1}{\lambda} \gamma_\mu \right) \tilde{\varphi}$$

$$\left(-\lambda^2 \frac{\partial^2}{\partial y_\mu \partial y_\mu} + \frac{1}{\lambda^2} \gamma_\mu \gamma_\mu \right) \varphi(x, y)$$

$$F_1 = \frac{dk}{\lambda^2} \left(-\lambda^2 \frac{\partial^2}{\partial y_\mu \partial y_\mu} + \frac{1}{\lambda^2} \gamma_\mu \gamma_\mu \right) \tilde{\varphi}(x, y)$$

$$\varphi = e^{ik_\mu x_\mu} F_0$$

$$\bar{I} = \int \left[\frac{1}{2} \frac{\partial \tilde{\varphi}(x, y)}{\partial x_\mu} \frac{\partial \varphi(x, y)}{\partial x_\mu} + \alpha F_1 + \beta F_2 \right] d^4x d^4y$$

α, β : positive constant.

$$(F_0 + \alpha F_1 + \beta F_2) \varphi = \alpha \varphi$$

$$\varphi = e^{ik_\mu x_\mu} \chi_{k_{\mu\nu}}(y)$$

F_1', F_2' have both discrete positive eigenvalues m_1^2, m_2^2

$$k_\mu k_\mu + \alpha m_1^2 + \beta m_2^2 (n_\mu, k_\mu) = 0$$

where m_2^2 depends both on n_μ and k_μ , while m_1^2 depends only on n_μ .
 In general

$$\varphi(x, r) = \sum_{n_\mu} \left(c_{k_\mu, n_\mu} e^{i k_\mu x_\mu} \chi_{n_\mu}(r_\mu) (dk) \right)^4$$

Free field: $\underline{r=0}$

$$k_\mu k_\mu < 0, \quad k_\mu: \text{time-like}$$

General field:

$$k_\mu k_\mu - \gamma < 0$$

$$\text{or } k_\mu k_\mu < \gamma - \alpha m_1^2 - \beta m_2^2$$

$$\gamma = \alpha m_1^2 + \beta m_2^2;$$

$$k_\mu k_\mu = 0.$$

$$\gamma > \alpha m_1^2 + \beta m_2^2$$

$$k_\mu k_\mu > 0 \quad k_\mu: \text{space-like}$$

Feb 1953
 Coupling between External
 and Internal Motions (C.1)

$$F \equiv -\frac{\partial^2}{\partial x_\mu \partial x_\mu} + \alpha^2 \lambda^2 \left(-\frac{\partial^2}{\partial r_\mu \partial r_\mu} + \frac{1}{\lambda^4} r_\mu r_\mu \right)^2$$

$$- \beta^2 \lambda^2 \left\{ -\left(\frac{\partial}{\partial x_\mu \partial r_\mu} \right)^2 + \frac{1}{\lambda^4} \left(r_\mu \frac{\partial}{\partial x_\mu} \right)^2 \right\}$$

$$\varphi(x, r) = e^{i k_\mu x_\mu} \chi(k_\mu, r_\mu)$$

$$\left[k_\mu k_\mu + \alpha^2 \lambda^2 \left(-\frac{\partial^2}{\partial r_\mu \partial r_\mu} + \frac{1}{\lambda^4} r_\mu r_\mu \right)^2 \right.$$

$$\left. + \beta^2 \lambda^2 \left\{ -\left(k_\mu \frac{\partial}{\partial r_\mu} \right)^2 + \frac{1}{\lambda^4} (k_\mu r_\mu)^2 \right\} \right]$$

$$\times \chi(k_\mu, r_\mu) = 0$$

$k_\mu k_\mu > 0$: space-like

$$k_1 = \kappa, \quad k_2 = k_3 = k_0 = 0$$

$$\left[\kappa^2 + \alpha^2 \lambda^2 \left(-\frac{\partial^2}{\partial r_\mu \partial r_\mu} + \frac{1}{\lambda^4} r_\mu r_\mu \right)^2 \right.$$

$$\left. + \beta^2 \lambda^2 \kappa^2 \left\{ -\left(\frac{\partial}{\partial r_1} \right)^2 + \frac{1}{\lambda^4} r_1^2 \right\} \right]$$

$$\times \chi(\kappa, 0, 0, 0; r_\mu) = 0$$

$$\chi = \chi_{n_1, n_2, n_3, n_0}$$

$$\left[\kappa^2 + 2\alpha^2 \lambda^2 (n_1 + n_2 + n_3 - n_0 + 1)^2 \right.$$

$$\left. + 2\beta^2 \kappa^2 (n_1 + \frac{1}{2}) \right] \chi_{n_1, n_2, n_3, n_0} \neq 0$$

Thus, there's no solution for $k_n k_n > 0$,
 because any solution ~~can be expanded~~ ^{should}
 into series of $\chi_{n_1, n_2, n_3, n_0}$ with non-
 zero coefficient, ~~which is impossible~~
 (but that)

for $k_n < 0$ $k_1 = k_2 = k_3 = 0$, $k_0 = \kappa$

$$\left[-\kappa^2 + \alpha^2 \left(-\frac{\partial^2}{\partial r_\mu \partial r_\mu} + \frac{1}{\lambda^2} \gamma_\mu \gamma_\mu \right) + \beta^2 \lambda^2 \kappa^2 \right]$$

$$\left(-\frac{\partial^2}{\partial r_0^2} + \frac{1}{\lambda^2} r_0^2 \right) \chi(0, 0, 0, \kappa; r_n)$$

$$= 0$$

$$\left\{ -\kappa^2 + \frac{4\alpha^2}{\lambda^2} (n_1 + n_2 + n_3 - n_0 + 1) \right. \\ \left. + 2\beta^2 \kappa^2 (n_0 + \frac{1}{2}) \right\} \chi_{n_1 n_2 n_3 n_0} = 0$$

This is satisfied only if

$$2\beta^2 (n_0 + \frac{1}{2}) < 1,$$

$$\text{or } n_0 < \left(\frac{1}{2\beta^2} - \frac{1}{2} \right)$$

$$\text{or } n_0 < \frac{1}{2} \left(\frac{1}{\beta^2} - 1 \right)$$

~~n_1, n_2, n_3, n_0~~

Thus, among infinitely many states belonging to the same value of

$$n \equiv n_1 + n_2 + n_3 - n_0,$$

only those states with $n_0 < \frac{1}{2}(\frac{1}{s^2} - 1)$ can survive, if we introduce the coupling between external and internal motions/degrees of freedom.

On the contrary, if ~~introduce~~ we adopt the perturbation theory, ~~only~~ all of the states with arbitrary n_1, n_2, n_3, n_0 ~~can~~ survive and the mass values ~~to~~ are changed a little bit, so that the degeneracy is removed partly.

Field Equations with Interaction. ~~for~~

(C.3)

$$\left\{ -\frac{\partial^2}{\partial x_\mu \partial x_\mu} + F' \right\} \varphi(X, r) = F \varphi(X, r)$$

$$= -g \sum_a \bar{\psi}_a(x + \frac{1}{2}r) \psi_a(x - \frac{1}{2}r)$$

$$\gamma_\mu \frac{\partial \psi(x')}{\partial x'_\mu} + M \psi(x') = -g \int \psi(x'') \varphi(x')$$

$$\frac{\partial \bar{\psi}(x')}{\partial x'_\mu} \gamma_\mu - M \bar{\psi}(x') = g \int \bar{\psi}(x'') \varphi(x'')$$

8-dimensional Green's function

$$F\left(\frac{\partial}{\partial x_\mu}, r_\mu, \frac{\partial}{\partial x'_\mu}\right) \varphi_0(X, r)$$

$$\text{suppose that } = \mu \varphi(X, r)$$

This eigenvalue problem can be solved and, in particular, let us assume that

$$F = -\frac{\partial^2}{\partial x_\mu \partial x_\mu} + F' \left(r_\mu r_\mu, \frac{\partial^2}{\partial x_\mu \partial x_\mu}, r_\mu \frac{\partial}{\partial x'_\mu}, r_\mu \frac{\partial}{\partial x_\mu}, \frac{\partial^2}{\partial x_\mu \partial x_\mu} \right)$$

~~any~~
 The eigenfunctions ~~ϕ_s~~ belonging to
 and eigenvalue μ_s be ϕ_s , one can
 expand an arbitrary ϕ

$$\phi(x, r) = \sum c_s \phi_s(x, r).$$

Eigenvalues μ_s are, in general,
 degenerate and consist constitute
 (both discrete and) continuous
 spectra. One had better write
 them as

$$\phi_{k_n, n_\mu}(x, r) \equiv \int e^{i k_n x} \phi_{n_\mu}(k_n, r_\mu) \frac{1}{(2\pi)^4}$$

where n_μ are discrete and k_n are
 continuous and as eigenvalues
 μ_{k_n, n_μ}

Then one can define

$$\bar{G}(x_\mu, r_\mu; x'_\mu, r'_\mu) \equiv \int_{n_\mu} \frac{e^{-i k_n x_\mu} \phi_{n_\mu}(k_n, r_\mu) e^{i k_n x'_\mu}}{\mu(k_n, n_\mu)} \times \phi_{n_\mu}(k_n, r'_\mu) (dk)^4$$

(C, 4)

Then

$$F\left(\frac{\partial}{\partial x_\mu}, r_\mu \frac{\partial}{\partial r_\mu}\right) \bar{G}(x_\mu, r_\mu; x'_\mu, r'_\mu) \\ = -\delta(x_\mu - x'_\mu) \delta(r_\mu - r'_\mu)$$

provided that F is bilinear
 even with respect to $\frac{\partial}{\partial x_\mu}$. (The
 particular examples as given above
 all satisfy this ~~axip~~ condition
 requirement)

Then

$$\int \bar{G}(x_\mu, r_\mu; x'_\mu, r'_\mu) \varphi(x'_\mu, r'_\mu) dx'_\mu dr'_\mu \\ \times g \sum \bar{\psi}_\alpha(x'_\mu + \frac{1}{2}r'_\mu) \psi_\alpha(x'_\mu - \frac{1}{2}r'_\mu)$$

satisfies the field equation and the
 general solution becomes for $\varphi(x, r)$

$$\varphi(x, r) = \varphi^{(f)}(x, r)$$

$$+ \int \bar{G}(x_\mu, r_\mu; x'_\mu, r'_\mu)$$

$$\times g \sum \bar{\psi}_\alpha(x'_\mu + \frac{1}{2}r'_\mu) \psi_\alpha(x'_\mu - \frac{1}{2}r'_\mu) dx'_\mu dr'_\mu$$

where $\varphi^{(f)}$ is any solution of the
 free field equation for $\varphi(x, r)$.

\bar{G} corresponds to $\bar{\Delta}$ in the ordinary
 theory.

Nature of Green's Δ .
 Our next question is, whether it is possible to construct ^{other} a Green's Δ which corresponds to retarded and advanced Δ -functions in the local field theory.

If we take

$$\varphi_0(k_\mu, r_\mu) \equiv \varphi_{0,0,0,0}(k_\mu, r_\mu)$$

alone,

$$\equiv \frac{1}{\pi \lambda^2} \cdot e^{-\frac{1}{2\lambda^2} \left\{ r_\mu r_\mu + \frac{2(k_\mu r_\mu)^2}{\kappa^2} \right\}}$$

$$\tilde{\varphi}_0(k_\mu, r_\mu) \varphi_0(k_\mu, r'_\mu)$$

$$= \frac{1}{\pi^2 \lambda^4} \cdot e^{-\frac{1}{2\lambda^2} \left\{ r_\mu r_\mu + r'_\mu r'_\mu + \frac{2(k_\mu r_\mu)^2}{\kappa^2} + \frac{2(k_\mu r'_\mu)^2}{\kappa^2} \right\}}$$

$$= \frac{1}{\pi^2 \lambda^4} e^{-\frac{1}{2\lambda^2} \left[2z_\mu^2 + \frac{1}{2} \eta_\mu^2 + \frac{2}{\kappa^2} (k_\mu z_\mu)^2 + \frac{1}{2} (k_\mu \eta_\mu)^2 \right]}$$

$$\frac{1}{2}(r_\mu + r'_\mu) = z_\mu$$

$$r_\mu - r'_\mu = \eta_\mu$$

$$2z_\mu^2 + \frac{1}{2}\eta_\mu^2 = r_\mu^2 + r'^2_\mu$$

$$(k_\mu r_\mu)^2 + (k_\mu r'_\mu)^2$$

$$= 2(k_\mu z_\mu)^2$$

$$+ \frac{1}{2}(k_\mu \eta_\mu)^2$$

if we may regard λ very small,

Thus φ_0 is small, except for very small values of r_μ, r'_μ

or r_μ, r'_μ , or, since

$$x'_\mu = x_\mu + \frac{1}{2} r_\mu$$

$$x''_\mu = x_\mu - \frac{1}{2} r_\mu$$

$$x'''_\mu = x'_\mu + \frac{1}{2} r'_\mu$$

$$\text{except for } x''_\mu = x'_\mu - \frac{1}{2} r'_\mu$$

$$x'_\mu \approx x''_\mu, \quad x'''_\mu \approx x''_\mu$$

We have then approximately

$$G(x_\mu, r_\mu; x'_\mu, r'_\mu) = - \frac{1}{\pi^2 \lambda^4} \int \frac{e^{-ik_\mu x_\mu} e^{ik'_\mu x'_\mu}}{k_\mu^2 + m^2} \delta(r_\mu) \delta(r'_\mu) \frac{d^3 k}{(2\pi)^3}$$

$$= \frac{\lambda^4}{\pi^2} \cdot \bar{\Delta}_{m_0}(x_\mu - x'_\mu) \delta(r_\mu) \delta(r'_\mu)$$

Now one can construct

$$G(x_\mu, r_\mu; x'_\mu, r'_\mu) = - \frac{1}{(2\pi)^4} \sum_{\mu} \int \frac{e^{-ik_\mu x_\mu} \Phi_{\mu}(k_\mu, r_\mu) e^{ik'_\mu x'_\mu}}{\Phi_{\mu}(k_\mu, r_\mu)} \times \Phi_{\mu}(k_\mu, r'_\mu) \cdot \epsilon(k_\mu) \delta(\mu) (d^3 k)$$

which satisfies ^{relative to Δ - μ} the free field equation and approximately satisfies the requirements for Δ - μ .

The commutation relation for the free field would then be

$$[\varphi^{(H)}(x_\mu, r_\mu), \varphi^{(H)}(x'_\mu, r'_\mu)]$$

$$= i G(x_\mu, r_\mu; x'_\mu, r'_\mu)$$

Now the ^{remaining} question is whether we can construct the Green fun. corresp. to $\Delta^{(ret)}$, $\Delta^{(adv)}$.

今のところでは

~~素粒子~~ 素粒子に何か局所場では表現しきれない未知の自由度を盛えなければ、質量スペクトル~~の~~自身の理解がいきなればかりでなく、これ等の電磁粒子がこれと比較的に長い寿命を持つていこう現象を説明するための^{必要}選択規則も導き出せぬ

新しい

~~い~~ ~~素粒子~~ ~~の~~ ~~局所場~~ ~~の~~ ~~問題~~ ~~を~~ ~~扱~~ ~~う~~ ~~こと~~ ~~が~~、はつきりしてきた^色に見える。これと同時に最近、~~素粒子~~ ~~論~~ ~~の~~ ~~相~~ ~~互~~ ~~作~~ ~~用~~ ~~を~~ ~~局~~ ~~所~~ ~~的~~ ~~相~~ ~~互~~ ~~作~~ ~~用~~ ~~を~~ ~~く~~ ~~り~~ ~~に~~ ~~み~~ ~~理~~ ~~論~~ ~~的~~ ~~立~~ ~~場~~ ~~か~~ ~~ら~~ ~~分~~ ~~析~~ ~~す~~ ~~こ~~ ~~と~~ ~~が~~ ~~先~~ ~~立~~ ~~屋~~ ~~グ~~ ~~ル~~ ~~ー~~ ~~プ~~ ~~の~~ ~~こ~~ ~~と~~ ~~あ~~ ~~つ~~ ~~て~~ ~~試~~ ~~み~~ ~~ら~~ ~~れ~~、⁽²⁾ ~~その~~ ~~結~~ ~~果~~ ~~と~~ ~~し~~ ~~て~~ ~~あ~~ ~~る~~ ~~種~~ ~~の~~ ~~局~~ ~~所~~ ~~的~~ ~~相~~ ~~互~~ ~~作~~ ~~用~~ ~~に~~ ~~く~~ ~~り~~ ~~に~~ ~~み~~ ~~理~~ ~~論~~ ~~を~~ ~~ど~~ ~~こ~~ ~~で~~ ~~と~~ ~~適~~ ~~用~~ ~~し~~ ~~て~~ ~~ゆ~~ ~~け~~ ~~は~~ ~~結~~ ~~局~~ ~~非~~ ~~局~~ ~~所~~ ~~的~~ ~~相~~ ~~互~~ ~~作~~ ~~用~~ ~~に~~ ~~な~~ ~~つ~~ ~~て~~ ~~し~~ ~~ま~~ ~~う~~ ~~こ~~ ~~と~~ ~~が~~ ~~明~~ ~~ら~~ ~~か~~ ~~と~~ ~~な~~ ~~つ~~ ~~た~~。一方において局所場の間の非局所的相互作用の問題を Peierls, Rayski, Bloch, ~~Her~~ 等によって詳しく検討され⁽³⁾、場の理論における散逸と、非局所的相互作用にあらわれる形状因子の間の関係が大方わかつてきた。特に Moller 及び Kristensen⁽⁴⁾ は、核子と中間子の間の非局所的相互作用を⁽⁴⁾ ~~適~~ ~~用~~ ~~に~~ ~~加~~ ~~へ~~、第一近似においては核子及び中間子の自己エネルギーを有限にできることを示した。~~か~~ ~~ら~~ ~~い~~ ~~へ~~ ~~ば~~ ~~単~~ ~~に~~ ~~非~~ ~~局~~ ~~所~~ ~~的~~ ~~相~~ ~~互~~ ~~作~~ ~~用~~ ~~に~~ ~~関~~ ~~する~~ ~~問~~ ~~題~~ ~~に~~ ~~対~~ ~~し~~ ~~て~~、形状因子の任意性が除去できないばかりでなく、質量スペクトルの問題とは何れも関連も見

せいのである。

2. 一般的考察

~~この~~所が非局所場という立場からこの系の問題を
 新しく見直すと、~~次の系に興味ある~~ 質量スペクトルの問題
 と ~~収斂因子の~~ 固有値の問題の密接な
 関係が有り得ることから明かになるのである。
 最初の着眼点はこの質量スペクトルの問題を非局所場
 によって記述される粒子の内部運動に於ける固有
 値問題と解釈することにあり、前者のために
 スカラー場 $(x' | \varphi | x'')$ を考え、これを ^{内部座標} 座標
 (外部座標)

$$X_\mu = \frac{1}{2}(x'_\mu + x''_\mu) \quad \mu = 1, 2, 3, 4$$

及び相対座標 (内部座標)

$$r_\mu = x'_\mu - x''_\mu$$

の函数

$$(x' | \varphi | x'') \equiv \varphi(X, r)$$

と見直すと、自由場の方程式は一般に

$$F\left(\frac{\partial}{\partial X_\mu}, r_\mu, \frac{\partial}{\partial r_\mu}\right) \varphi(X, r) = 0 \quad (1)$$

という形式に書けるものと推定される。Fは $\frac{\partial}{\partial X_\mu}, r_\mu,$

$\frac{\partial}{\partial r_\mu}$ の相対論的に不変な函数で、特に四次元座
 標の原点の運動に於て不変であるため、

X_μ を含む必要はないことが必要である。... かくれば F

$$は \frac{\partial^2}{\partial X_\mu \partial X_\mu}, r_\mu r_\mu, \frac{\partial^2}{\partial r_\mu \partial r_\mu}, r_\mu \frac{\partial}{\partial X_\mu}, \frac{\partial^2}{\partial X_\mu \partial r_\mu}, r_\mu \frac{\partial}{\partial r_\mu}$$

の固有函数の集である。

今特に前章は場論として、外部運動と内部運動とが分離できるものとして

$$F \equiv -\frac{\partial^2}{\partial x_\mu \partial x_\mu} + F^{(r)}(r_\mu r_\mu, \frac{\partial^2}{\partial r_\mu \partial r_\mu}, r_\mu \frac{\partial}{\partial r_\mu}) \quad (2)$$

場という形に書けるとすると、~~場~~ 場方程式(1)は

$$\varphi(X, r) = u(X) \chi(r)$$

と置くことにより分離できる。u及び χ はそれぞれ分離常数を μ とすると

$$\left(\frac{\partial^2}{\partial x_\mu \partial x_\mu} - \mu\right) u(X) = 0 \quad (3)$$

$$\left(F^{(r)} - \mu\right) \chi(r) = 0 \quad (4)$$

の解であることは先に示した。従って $F^{(r)}$ なる演算子の固有値0又は正の離散固有値しか持たないという選んでおけば、方程式(4)で定義された四次元的固有値問題の解として得られる μ の固有値 $\mu_n (n=1, 2, \dots)$ から質量スペクトル $m_n = \sqrt{\mu_n}$ (5) $\mu = c=1$ 単位系での

が決定されることになる。一般には固有値 μ_n は複素数であるので、そのために問題はもっと複雑化するが、ここでは先ず実数値であるものとして、 μ_n は

次元は9次元
 固有規格化された固有函数 $\chi_n(r)$ とすると、場自由

場は一般に

$$\varphi(X, r) = \sum_n u_n(X) \chi_n(r) \quad (6)$$

と書けることになる。ここで u_n は

$$\left(\frac{\partial^2}{\partial x_\mu \partial x_\mu} - m_n^2\right) u_n(X) = 0 \quad (7)$$

を満足するから、これを m_n なる質量の粒子の外殻運動を記述する局所場と見出すことができる。

すると非局所場 $\phi(X, r)$ の量子化の問題は、局所場 $u_n(X)$ の量子化という問題に還元される。従って $u_n(X)$ を

$$\langle u_n(X), u_{n'}(X') \rangle = i \Delta_n \quad (18)$$

$$[u_n(X), u_{n'}(X')] = i \delta_{nn'} \Delta_n (X - X')$$

なる交換関係を満足する Hermite 型演算子と見出せば、非局所スカラー場 $\phi(X, r)$ はある意味で種々の質量を持つた ~~伸縮可能な~~ 粒子の集りと同義になる。但し Δ_n は質量 m_n なる場合の不変行列関数である。

3. 収縮因子の問題

次の問題は内部運動の固有函数 $\chi_n(r)$ がどういう役割を演ずるかという点にある。それを明らかにするには、相互作用のある場合の非局所場の方程式を考察する必要がある。順序として先ず自由場に対する方程式 (1) が 変分原理

$$\delta \int \tilde{\varphi} \cdot F \left(\frac{\partial}{\partial x_\mu}, r_\mu, \frac{\partial}{\partial r_\mu} \right) \varphi (dX)^4 (dr)^4 = 0 \quad (9)$$

から導き出されることから、自由場に対する Lagrange 密役を

$$L_{sc} = \frac{1}{2} \tilde{\varphi}(X, r) \cdot F \left(\frac{\partial}{\partial x_\mu}, r_\mu, \frac{\partial}{\partial r_\mu} \right) \varphi(X, r) \quad (10)$$

の後の $u_n(X)$ は n 量子化する前は n 変関数であり、従って量子化後は n 変 Hermite 演算子となるが、 n 変後は

$$[u_n(X), u_{n'}(X')] = i \delta_{nn'} \Delta_n(X-X') \quad (10)$$

なる交換関係を満足する Hermite 演算子となる。但し Δ_n は n 変関数 u_n がある場合の不変デルタ関数である。この Δ_n は n 変で相局所スカラー場 $\varphi(X, Y)$ は種々の質量を持つたボース粒子の集りと同等であると考えられる。特に上の例の Δ_2 は $\varphi(X, Y) = \frac{1}{2} \varphi(X) \varphi(Y)$ として Hermite 型演算子を取れば、中性のボース粒子の集りと同等になる。

$$x'_j = x_j + v_j \left[\frac{x_j v_j}{v_j v_j} \left(\frac{1}{\sqrt{1-v^2}} - 1 \right) - \frac{v x_0}{\sqrt{1-v^2}} \right]$$

$$x'_0 = x_0 - \frac{v_j x_j}{\sqrt{1-v^2}} = \frac{1}{m} (k_0 x_0 - k_j x_j)$$

$$k_j x_j - k_0 x_0 \quad x'_\mu = i \frac{-k_\mu x_\mu}{\sqrt{-k_\mu k_\mu}}$$

$$k_j = \frac{m v_j}{\sqrt{1-v^2}} \quad k_0 = \frac{m}{\sqrt{1-v^2}}$$

$$\left\{ \begin{aligned} x'_j &= x_j + \frac{k_j}{k_0} \left[\frac{k_0^2}{(k_j k_j)} \left(\frac{k_j}{k_0} x_j \right) \left(\frac{k_0}{m} - 1 \right) - \frac{k_0 x_0}{m} \right] \\ &= x_j + k_j \left[\frac{k_0 x_0}{k_j k_0} \left(\frac{k_0}{m} - 1 \right) - \frac{x_0}{m} \right] \\ x'_0 &= \frac{k_0}{m} \left(x_0 - \frac{k_j x_j}{k_0} \right) = \frac{-k_\mu x_\mu}{\sqrt{-k_\mu k_\mu}} \end{aligned} \right.$$

$$\langle x' | \psi | x'' \rangle \equiv \psi(x', x'')$$

又はこれと同様の量で定義しておく。次に ~~もう~~ もう

一つ別の非局所場を φ とし、これに対する通常の Lagrange 密度を加えればよい。この二つの非局所場の間の相互作用は例えば ~~Lagrangian 密度~~ 相互作用積分

$$\int g(x^{(1)} | \psi^* | x^{(2)})(x^{(2)} | \varphi | x^{(3)})(x^{(3)} | \psi | x^{(4)}) dx^{(1)} dx^{(2)} dx^{(3)} dx^{(4)}$$

(この項を g に加えることに g として記述せよ)

$$(11)$$

~~これは~~ g は結合定数で、今の例では ψ は ~~一般~~ 一般の Hermite 型でない非局所スカラー場と見なす。しかし下の参考文献は Møller-Kristensen が詳しく考察した局所スカラー場と局所スピノル場の間の非局所的相互作用の場合との対応を明らかにしていると思うので、 ψ として局所スピノル場 $\psi_\alpha(x')$ を取ることにする。従って相互作用の Lagrange 密度も例えば

$$L_{int} = g(x' | \varphi | x'') \sum_\alpha \bar{\psi}_\alpha(x'') \psi_\alpha(x') \quad (12)$$

* 非局所場 $\langle x' | \varphi | x'' \rangle \equiv \varphi(x', x'')$ が Hermite 型演算子であるという制限を加えれば、 $\langle x' | \varphi | x'' \rangle$ と $\langle x'' | \varphi | x' \rangle$ は ~~$\langle x' | \varphi | x'' \rangle$~~ 複素共転であることになり、従って

$$\tilde{\varphi}(x, y) = \varphi(x, -y) \quad (13)$$

なる関係が満たされるような関数に限らねばならぬことになる。所で固有函数 $\chi_n(y)$ は一般に偶函数か奇函数のいずれかに制限できるから、そのいずれかがあることに従って

$$\tilde{\chi}_n(y) = \chi_n(-y) = \pm \chi_n(y)$$

の複素共転のどちらかを取らねばならぬ。これ等の詳細は別の機会に譲る。

と取ることをできる** すると変分原理から導き出される場の方程式は、下として(2)の表式を取れば、

$$\left\{ -\frac{\partial^2}{\partial x_\mu \partial x_\mu} + F^{(n)}(\gamma_\mu \gamma_\mu, \frac{\partial^2}{\partial r_\mu \partial r_\mu}, \gamma_\mu \frac{\partial}{\partial r_\mu}) \right\} \varphi(X, r) = -g \sum_a \bar{\Psi}_a(X + \frac{1}{2}r) \Psi_a(X - \frac{1}{2}r) \quad (13)$$

$$\gamma_\mu \frac{\partial \Psi(x')}{\partial x'_\mu} + M \Psi(x') = -g \int \varphi(x'') (x'' | \varphi | x') dx'' \quad (14)$$

$$\left(\gamma_\mu \frac{\partial \bar{\Psi}(x')}{\partial x'_\mu} - M \bar{\Psi}(x') \right) = g \int (x' | \varphi | x'') \bar{\Psi}(x'') dx'' \quad (15)$$

と取る。

所で一般の非局所場 $\varphi(X, r)$ は固有函数 $\tilde{\chi}_n(r)$ を使つて(6)なる形に展開できるから、これを(13)の左辺に代入し、(14)を考慮して、両辺に $\tilde{\chi}_n(r)$ を乗じて ~~積分~~ 四次元の r 空間で積分すると

$$\left(\frac{\partial^2}{\partial x_\mu \partial x_\mu} - m_n^2 \right) u_n(X) = g \int \tilde{\chi}_n(r) \sum_a \bar{\Psi}_a(X + \frac{1}{2}r) \Psi_a(X - \frac{1}{2}r) dr \quad (16)$$

が得られる。同様にして(14)の右邊を変形して、脚註の関係式(†)を考慮すると、(14)は

** スピノル場を非局所化するとは、スピノルが半整数の粒子のフェルミ粒子の質量スペクトルを導き出す上に必要であるが、問題を単純化するためにスピノル場場の方には通常の局所場を取ることにしておく。これに関しては高片山泰久：素粒子論研究 (Vol. 4, No. 9 (1952), 97) を参照されたい。

$$\gamma_\mu \frac{\partial \psi(x')}{\partial x'_\mu} + M \psi(x') = -g \sum_n \int \tilde{\chi}_n(x'-x'') u_n\left(\frac{x'+x''}{2}\right) \psi(x'') dx'' \quad (17)$$

となる。これは等の方程式を Møller-Kristensen の得た方程式

$$\left(\frac{\partial^2}{\partial x''_\mu \partial x''_\mu} - m^2\right) u(x'') = \int \Phi(x', x'', x''') \underbrace{\sum_\alpha \bar{\psi}_\alpha(x') \psi_\alpha(x''')}_{dx' dx'''} \quad (16)'$$

$$\gamma_\mu \frac{\partial \psi(x')}{\partial x'_\mu} + M \psi(x') = - \int \underbrace{\Phi(x', x'', x''') u(x'')}_{\psi(x'') dx'''} \quad (17)'$$

両者の間の

と比較して見ると、若しい類似性と同様に、本質的な相違が明瞭になる。即ち非局所スカラー場で記述される色々の質量のボース粒子の中で、特定の統一的一つ、例えば質量 m_n なる粒子だけを取り出し、他はすべて無視してしまおうと、(16) 及び (17) はそれぞれ (16)' 及び (17)' において

$$m = m_n, \quad \Phi(x', x'', x''') = \underbrace{\tilde{\chi}_n(x'-x'')}_{\delta\left(\frac{x'+x''}{2} - x''\right)} \quad (18)$$

とおいたのと同等になる。従って内部運動の固有函数 $\chi_n(x)$ が非局所的相互作用の理論における形状函数ないし収斂因子の役割をなすことになる。

しかしある算量の場ではやり易いこと、精々低エネルギー領域における近似解を求めるのに役立つという以上の意味を持つ得ない。非局所場理論では算量スペクトルの問題が収斂因子の問題と切り離せられ、これが局所場の非局所的相互作用の理論と本質的に異なるとしてある。

4. 振動子模型

以上の一般化が一つの具体的な場合においてどの程度正統化されるか、^{また}どの様な複雑な問題が伏在しているかを明らかにする^{ため}に、最も簡単らしく見える振動子模型を算例として取り上げることにしよう。即ち(2)の形に分解できる演算子 F の中で、特に

$$F = -\frac{\partial^2}{\partial x_{\mu} \partial x_{\mu}} + \alpha \left(-\frac{\partial^2}{\partial r_{\mu} \partial r_{\mu}} + \frac{1}{\lambda^2} r_{\mu} r_{\mu} \right) \quad (19)$$

なる形を假定すると、内部運動の固有函数

$$\chi_{n_1, n_2, n_3, n_0}(r) = \frac{H_{n_1}(r_1/\lambda) H_{n_2}(r_2/\lambda) H_{n_3}(r_3/\lambda) H_{n_0}(r_0/\lambda)}{\sqrt{\exp\{-\frac{1}{2\lambda^2}(r_1^2 + r_2^2 + r_3^2 + r_0^2)\}}} \quad (20)$$

に対応する算量の固有値 m_{n_1, n_2, n_3, n_0} は

$$(m_{n_1, n_2, n_3, n_0})^2 = \frac{2\alpha}{\lambda^2} (n_1 + n_2 + n_3 - n_0 + 1) \quad (21)$$

となる。他し α は次元なしの定数、 λ は長さの次元を持つ定数、 r_0 は ~~実数~~ ~~実数~~ 相対時間 に比例する変数で、虚数座標 r_4 との ~~関係は~~ ~~関係は~~ ~~関係は~~

• $r_4 = i r_0$

である。 n_1, n_2, n_3, n_0 はそれぞれ 0 ~~より~~ 正の整数値に等しい取り得る量子数を表わし、 $H_n(x)$ は変数 x に関する n 次の Hermite 多項式を表わす。

振動子模型の特徴の一つは、固有函数 $\chi_{n_1, n_2, n_3, n_0}$ が量子数 n_1, n_2, n_3, n_0 の如何にかかわらず、 $|r_1|, |r_2|, |r_3|, |r_0|$ の大きな値に対して充分早く減りずるばかりでなく、これ等の函数の Fourier 変換を作ると、全く同じ形を持つてゐるから、運動量・エネルギー空間でも充分よく収斂してくれることである。* ここで直ぐ気がつくのは、(20) なる形の固有函数が相対論的に不変でないことであるが、その点自身は別に理論の真偽と関係ない。ただこれに関連して次の二つの重要な困難を無視できない。
(i) 第一は (21) の右辺が負なる場合以外、質量が虚数となり、物理的意味づけが困難になることである。

~~この点の改良は容易であつて~~、例えは演算子 F として (19) の代りに

$$F \equiv -\frac{\partial^2}{\partial x_\mu \partial x_\mu} + \alpha \left(-\frac{\partial^2}{\partial r_\mu \partial r_\mu} + \frac{1}{\lambda^2} r_\mu r_\mu \right) \quad (19)'$$

を取れば、内部運動の固有函数 χ の代りに (20) を取つてよく、質量の固有値は

$$(m_{n_1, n_2, n_3, n_0})^2 = \frac{4\alpha^2}{\lambda^2} (n_1 + n_2 + n_3 - n_0 + 1)^2 \quad (22)'$$

* 振動子模型は Born 達の提唱した相互性の立場から見ると、自己相反性を満足する場合に當つており、従つて通常の四次元空間での表現と、運動量・エネルギー空間での表現とが當然同形になつてくる。しかし Born-Green の理論は (5) 素粒子の内部自由度を考慮したことに依つていないから、場の理論の収斂性の問題とは関係がつかはつた。

(19)
しかし非局所場理論ではある状態の場を他から切り出して取り出すことは、一種の近似としてしか許されない。高エネルギーの過程においては一般に近似は悪くなり、少なくとも幾つかの低い状態の場を同時に考える必要がある。局所場の非局所的相互作用の理

$$\beta^2 \lambda^2 \left\{ - \left(\frac{\partial}{\partial x_\mu \partial r_\mu} \right)^2 + \frac{1}{\lambda^2} \left(r_\mu \frac{\partial}{\partial x_\mu} \right)^2 \right\}$$

$$k_\mu k_\mu$$

$$\left\{ - \left(k_\mu \frac{\partial}{\partial r_\mu} \right)^2 + \frac{1}{\lambda^2} \left(r_\mu k_\mu \right)^2 \right\}$$

$$- (k_j r_j - k_0 r_0)^2 = - k_0'^2 r_0'^2$$

$$- \left(k_j \frac{\partial}{\partial r_j} + k_0 \frac{\partial}{\partial r_0} \right)^2 = - k_0'^2 \frac{\partial^2}{\partial r_0'^2}$$

$$(+ k_\mu k_\mu)$$

$$\beta^2 \lambda^2 \left(\frac{m^2}{\lambda^2} \right) \frac{(2n_0' + 1)}{\lambda^2}$$

$$= -m^2$$

$$-m^2 + \left(\quad \right)^2 + \beta^2 m^2 (2n_0' + 1)$$

$$m^2 = \frac{\left(\quad \right)^2}{1 - \beta^2 (2n_0' + 1)} \quad n_0' = 0, 1, 2, \dots$$

である。 H_n は n 次の Hermite 多項式を表現する。従って
 固有函数 χ_{n_1, n_2, n_3}

$$(m_{n_1, n_2, n_3, n_0})^2 = \frac{\frac{4d^2}{\lambda^4} (n_1 + n_2 + n_3 - 2n_0 + 1)^2}{1 - 2\beta^2 (n_0 + \frac{1}{2})} \quad (27)$$

2 番より n_0 と

これは、但し n_0 の場合は

$$1 - 2\beta^2 (n_0 + \frac{1}{2}) \neq 0$$

$$n_0 + \frac{1}{2} \leq \frac{1}{2\beta^2}$$

$$n_0 < \frac{1}{2} (\frac{1}{\beta^2} - 1)$$

は満足する 0 又は 2 の整数に限られる。

例えば $\beta^2 = \frac{1}{2}$ とすると、 $n_0 = 0$ 又は 2 のみが可能
 である。 (27) は $n_0 = 0$ の場合

$$m_{n_1, n_2, n_3, 0} = \frac{\sqrt{8}d}{\lambda^2} (n_1 + n_2 + n_3 + 1) \quad (28)$$

2 番より、従って

となるから、常数の正負に依りて α の値を正負に
 しか現われないこととなる。

(ii) 第二の困難は α の値を $\alpha = 0$ 又は
 正の他と限つて差支ない) に制限しておいても、常に
 固有値が無限に縮退していることである。例として
 (21)' の場合について、最小の値 $m_0 = 0$

を取ると、この固有値に属する固有函数 ψ は

$$n_1 + n_2 + n_3 - n_0 = -1$$

を満す n_1, n_2, n_3, n_0 の組合せの α は取つてお
 いから、無限に縮退していることとなる。

この第二の困難は第一の困難より 是るかに重大
 であつて、恐らく振動子模型自体の特性でなく、相對
 論的固有値問題に關する α の問題らしく
 思われる。* 此の難點を除くには α の値を $\alpha = 0$ に
 内部運動と外部運動の間の組合せを考慮する必要がある
 ことである。その一例として、(19)' なる演算子に

$$-\beta^2 \lambda^2 \left\{ - \left(\frac{\partial}{\partial x_\mu \partial x_\mu} \right)^2 + \frac{1}{\lambda^2} \left(\gamma_\mu \frac{\partial}{\partial x_\mu} \right)^2 \right\} \quad (25)$$

なる項を附加したとすると、自由場の方程式 (1) は

$$\varphi(X, Y) = e^{ik_\mu x_\mu} \chi(k_\mu, \gamma_\mu) \quad (26)$$

なる形の解を有することが直ちにわかる。但し β は実常数とする。

* Schrödinger が Born-Green の振動子模型に
 對する難點としてあげている諸點もこの問題と
 密接に關係している。(6)

$$\left[k_\mu k_\mu + \alpha^2 \left(-\frac{\partial^2}{\partial r_\mu \partial r_\mu} + \frac{1}{\lambda^4} r_\mu r_\mu \right)^2 + \beta^2 \lambda^2 \left(-k_\mu \frac{\partial}{\partial r_\mu} \right)^2 + \frac{1}{\lambda^4} (k_\mu r_\mu)^2 \right] \chi(k_\mu, r_\mu) = 0 \quad (25)$$

の解 ~~は~~ (20) の固有函数の完全系で展開することによって ~~得られる~~ ~~一般の場合には~~ ~~複雑な~~ ~~形式を~~ ~~とる~~、しかしこの手順では次の事情を考慮することによって、非常に簡単化される、即ち波数ベクトル k_μ が時間的である場合、 $k_\mu k_\mu < 0$ なる場合にのみ (25) が解を有することを考慮して、*座標系の変換を行い、新しい座標系では $k_1 = k_2 = k_3 = 0$ 、 $k_0 = \kappa$ とおくとすると、(25) は

$$\left[-\kappa^2 + \alpha^2 \left(-\frac{\partial^2}{\partial r_0^2} + \frac{1}{\lambda^4} r_0^2 \right)^2 + \beta^2 \lambda^2 \kappa^2 \left(-\frac{\partial^2}{\partial r_0^2} + \frac{1}{\lambda^4} r_0^2 \right) \right] \chi(0,0,0,\kappa; r_\mu) = 0, \quad (26)$$

となり、~~この解が (20) が~~ ~~従って~~ κ が ~~固有値~~ ~~である~~

~~$$\chi(0,0,0,\kappa; r_\mu)$$~~

$$-\kappa^2 + \frac{4\alpha^2}{\lambda^4} (n_1 + n_2 + n_3 - n_0 + 1)^2 + 2\beta^2 \lambda^2 (n_0 + \frac{1}{2}) = 0 \quad (28)$$

なる関係が満たされ、~~が満たされる~~ 場合には限って (20) が (25) を満足することになる。従って固有値スペクトルは (27) の代りに

* k_μ が空間的ベクトル、即ち $k_\mu k_\mu > 0$ の場合には、 $|r_{11}|, |r_{21}|, |r_{31}|, |r_{01}|$ の大きさを他に対して ~~減少~~ ~~減少~~ するならば χ が存在しないことが容易に証明できる。

$$m_{n_1, n_2, n_3, n_0} = \frac{2\alpha}{\lambda^{3/2}} \frac{|n_1 + n_2 + n_3 - n_0 + 1|}{\sqrt{1 - 2\beta^2(n_0 + \frac{1}{2})}} \quad (29)$$

となる。但し α は正の整数に限り、
 n_0 の値は

$$n_0 < \frac{1}{2} \left(\frac{1}{\beta^2} - 1 \right) \quad (30)$$

を満足する 0 または正の整数に限られる。例えば
 $\beta = \frac{1}{\sqrt{2}}$ と取ると、 $n_0 = 0$ だけが許されることになる。
 質量スペクトルは

$$m_{n_1, n_2, n_3, 0} = \frac{2\sqrt{2}\alpha}{\lambda^{3/2}} (n_1 + n_2 + n_3 + 1) \quad (29)'$$

となる。従って固有値の一般に有限の縮退しか持たぬ
 ことになる。特に最小の固有値

$$m_{0,0,0,0} = \frac{2\sqrt{2}\alpha}{\lambda^{3/2}} \quad (31)$$

は縮退しておらず、これに対応する固有函数は

$$\chi_{0,0,0,0}(r) = \exp \left\{ -\frac{1}{2\lambda^2} (r_1^2 + r_2^2 + r_3^2 + r_0^2) \right\} \quad (32)$$

に限られることになる。

以上は外部運動のない重心運動に関する静止系、いわ
 ゆる重心静止系から見た時の話であるが、元の座標
 系に於いては、質量スペクトルは変化なく、固有
 函数の中の変数が座標変換を要するだけである。
 例えば (30) の固有函数は、元の座標系では

$$\chi_{0,0,0,0}(k_\mu, r_\mu) = \exp \left\{ -\frac{1}{2\lambda^2} (r_\mu r_\mu + 2 \frac{(k_\mu r_\mu)^2}{k_\mu k_\mu}) \right\} \quad (32)'$$

となる。
 而 (29) がパイ中間子と同じ位の質量を備えるための
 には、 α として1の程度の整数、長さの次元を持つ
 定数~~を~~入としてパイ中間子の Compton 波長
 (通常の単位で $\hbar/m_\pi c$) の程度の値を取らねば~~い~~。
 すると (27) からその整数倍の質量の粒子がすべて
 同時に存在しなくてはならぬことになるが、それは
 パイ中間子よりもっと複雑な内部構造を持ち、従って
 色に違つた性質を持つたボース粒子と考へられる。
 この点の結果が正しいかどうか今の所判定できない。
 もし入として核子の Compton 波長に近い程小さい値
 を取るならば、 α は1よりずっと小さくなければなら
 ぬことになる。~~これは少し自然の順序に見えぬ。~~

5. ~~結論~~ 結論を述べ

以上 核子模型という特定の場合に関連して得られ
た結果が、どの程度の一般性を持つかの推定は容易で

* 特に例えに $\alpha = 1/\sqrt{2}$ と仮定し、~~を~~入か
 古典電子半径 (通常の単位で $e^2/m_e c^2$, m_e は
 電子質量) である場合に
 $m_{0,0,0} = 274 m_e$

となる。
 ** 今までに知られている不安定粒子の質量は近似的に
 整数法則に従うように見えるが (7) 自己エネルギーの~~計算~~修正
 の補正がどの位か推定できない。 (27) の正否の判定は
 不可能である。

はいが、~~非局所~~ 望ましくない 端退を 取除く ため に 外部運
動 と 内部運動 を 結合 せしめる こと が 一般 に 重要 に 関
定 すると、 第三節 で 内部運動 を 平均化 して 非局所場 を
非局所的 相互作用 に 歸着 せしめる 手續 は、 その 中 で
は 使 用 され ない こと に なる。 従つて 非局所場 理論 の 收
斂 性 の 問題 を 簡潔 に 見 透 し かつ せ にく くなる か、
この 問題 の 詳細 な 検討 は 次 の 機 會 に 譲 る こと
にして、 一先 手 摺 筆 する。(8)

也、内部場自由なと関係して新しい選擇規則の導
出が ~~...~~

文 献

- (1) H. Yukawa, Phys. Rev. 77 (1950), 219; 80 (1950),
~~1047~~ 1047; D. R. Yennie, Phys. Rev. 80 (1950), 1053.

$\{ (\alpha' | \phi | \alpha) \}$ $\gamma' = \alpha' - \alpha$

$= \left[\sum_n u_n(\alpha) \chi_n(\alpha), \sum_n u_n(\alpha') \chi_n(\alpha') \right]$

$= i \sum_n \chi_n(\alpha) \chi_n(\alpha') \Delta_n(\alpha - \alpha')$

~~$\int \frac{e^{i k \mu \alpha} \chi_n(\alpha) e^{-i k \mu \alpha'} \chi_n(\alpha')}{\epsilon(k) \delta(k \mu + m_n^2)} (d k)^4$~~

$\mathbb{F} \cdot e^{i k \mu \alpha} \chi_n(\alpha) = \mu e^{i k \mu \alpha} \chi_n(\alpha)$

$\int \phi_{k,n}(\alpha, r) \phi_{k,n}(\alpha', r') \delta(\mu) (d k)^4$

[Faint handwritten notes and scribbles, including the word "Number" and various mathematical symbols.]

Structure of and Mass Spectrum of Elementary Particles (1)

By Hideki Yukawa
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if it ^{into} exists at all,

A unified field theory of elementary particles is expected to account for the mass spectrum of a ^{great} variety of stable and unstable particles on the one hand and to be free from the pathological divergences on the other. The non-local field was ^{supposed} introduced in order to describe an ~~entity~~ ^{in a relativistically} which could be elementary ^{in a} system with its own internal ^{way} structure a system an object which ~~is~~ ~~was~~ ~~not~~ ~~a~~ ~~con~~ ~~did~~ ~~not~~ ~~was~~ elementary in the sense that it could not be decomposed into more elementary ~~not~~ objects, but was complex enough to have ~~take~~ be in various kinds of states with different masses, spins and other intrinsic properties. In the

It seems to the author that ~~was~~ disappointing
consequences are ~~intrinsic to the~~ ~~field~~ ~~in general~~ but are
intrinsic to the ~~field~~ ~~in general~~ but are
heretofore in ~~field~~ ~~in general~~ but are

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previous papers, the author tried to
decompose a general non-local field
into irreducible parts, by imposing
a number of ~~compatible~~ ~~relations~~
as ~~many~~ ~~restrictions~~ ~~as~~ ~~possible~~
upon this ~~it~~. However, the conclusions
reached were very unsatisfactory
for a number of reasons. (2) Among
other things, the ~~possible~~ ~~masses~~ ~~of~~
the particles associated with the
irreducible fields remained completely
arbitrary and ~~reasonable~~ ~~assumptions~~
on the interaction between these fields
did not result in the expected

convergence of self-energies. Since
this may well be due to the
imposition of too many ~~restrictions~~ ~~upon~~
a single ~~non-local~~ ~~field~~. Instead, one could
have started from one field equation
of the following type, ~~without~~ ~~laying~~
aside the question of reduction of
of paying too much attention to the
reduction of fields a field into
irreducible parts.

In order to avoid too much ~~trouble~~

(2)

let us take, just for simplicity, a scalar (or pseudoscalar) non-local field

$$(x'_\mu | \varphi | x''_\mu) \equiv \varphi(X_\mu, r_\mu),$$

where x'_μ, x''_μ ($\mu=1, 2, 3, 4$) stand for two sets of space-time parameters with imaginary time and

$$X_\mu = \frac{1}{2} (x'_\mu + x''_\mu)$$

$$\text{free, } r_\mu = x'_\mu - x''_\mu. *$$

The field equation is supposed to have a general form

$$\underbrace{F}_{\text{the operator}} \left(\frac{\partial}{\partial x_\mu}, r_\mu, \frac{\partial}{\partial r_\mu} \right) \varphi(X_\mu, r_\mu) = 0, \quad (1)$$

where F is a certain invariant function of $\frac{\partial}{\partial x_\mu}, r_\mu$ and $\frac{\partial}{\partial r_\mu}$. It ^{must be in} cannot depend on x_μ of X_μ in order to be invariant under the translation of the coord. origin of the coordinate system. In particular, if ^{we} assume that F is linear in $\frac{\partial^2}{\partial x_\mu \partial r_\mu}$ and separable, i.e.

$$F \equiv -\frac{\partial^2}{\partial x_\mu \partial r_\mu} + f\left(\frac{\partial}{\partial r_\mu}, r_\mu, \frac{\partial}{\partial r_\mu}\right), \quad (2)$$

* In contradistinction to the previous papers, ^{$\mu=4$ corresponds to an} ~~we use~~ imaginary time ~~parameter~~ ^{parameter}, for instance, $r_4 = i r_0$, where r_0 is the real relative time.

we have ~~an~~ i -eigen-solutions of
the form $\varphi(x_\mu, r_\mu) = u(x_\mu) \chi(r_\mu)$
where u and χ satisfy

$$\left(\frac{\partial^2}{\partial x_\mu \partial x_\mu} - \mu\right) u = 0 \quad (3)$$

$$\mathcal{L}\left(f\left(r_\mu, \frac{\partial}{\partial r_\mu}\right) \chi\right) = 0, \quad (4)$$

μ being the separation constant.
Thus, the masses of the particles associated
with the non-local field φ are
given as the eigenvalues of μ of
the equation (4) for χ . The
interval eigenvalue χ . If it is
possible to choose the operator f
such that the eigenvalues are
all discrete, ~~as say~~, then ~~the~~
free field φ can be expanded
arbitrary non-local fields φ can
be expanded

Novi-local
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A Field Theory of Particles (with Internal Structure) I.
By Michio YUKAWA*

Abstract

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Abstract

Introduction

* On leave of absence from Columbia University, New York.

implication
of the introduction of the concept
of non-local fields, implies
that some new insight was twofold.

08.6

1. Introduction

The original intention of the author, which is ~~an~~ ^{an} ~~more~~ ^{simple} way of describing a system ^{relativistically} ~~with the~~ ^{internal} ~~introduction~~ ^{structure} of field, in such a way that the elementary particles with internal structure was to extend the field theory twofold. On the one hand, by ~~introducing~~ ^{possible internal} taking into account the structure of elementary particles and formulating it ~~relativistically~~ in terms of the non-local field, the mass-spectrum of elementary particles was expected to come out as the result by solving a certain eigen-value problem. On the other hand, it was hoped that the eigenfunction which was associated with ^{for the internal motion} the eigenvalues of the mass might ~~serve~~ ^{each of} play the role of the convergence factor in the interaction between elementary particles.

In the previous papers⁽¹⁾, however, these points were not discussed. On the contrary, the main effort was concentrated on restricting the possible forms of the

a notion

non-local field, which was too vague and too broad to be dealt with effectively. For that purpose, the author ~~took the advantage~~ ^{adopted} of the concept of reciprocity in a certain particular way ~~and unfortunately~~, decomposed the global non-local field into irreducible parts.

It ~~was found~~ ^{turned out}, however, that the restrictions thus imposed on the non-local field in order to obtain irreducible fields were so severe that an irreducible ~~for~~ non-local field reduced to the local field became equivalent to a certain type of local field as first pointed out by Fierz (2). For that reason, it was felt necessary to give up the ~~red~~ complete reduction of ϕ . It is inappropriate to try to find a correspondence between ~~an~~ ^{each of} irreducible non-local fields and ~~the~~ ^{each} ~~each one~~ type of elementary particles. This means

(3)

that we have to drop, at least, one of the field equations, for the non-local field, ~~or~~ or replace it so that we ^{can} ~~admit~~ admit the ~~irreducible~~ reducible fields as the ~~model~~ possible models for the particles. We can start from the following general equation or one can. For instance, one can assume only one field eq. for the non-local field instead of a set of three simultaneous equations. From this, one can formulate the general eigenvalue problem for the mass as shown in the next section.

In the mean-time non-local interaction between local fields has been investigated by a number of authors⁽³⁾. Among them, a recent paper by Moller and Rosenfeld & Kristensen⁽⁴⁾ is of particular interest in that the self-energies of both of the meson and nucleon could be made ~~over~~ finite, at least ^{up to} the lowest order, in the coupling constant by introducing a suitable form function, into the interaction for the interaction in certain type of meson theory.

In this connection, it
will be shown in the third
section, ^{that} the eigenfunctions for the
internal motion, if one ignores
the possible degeneracy of eigenvalues
for the mass, the field equations for
the non-local field interacting
with some other field can be
brought to the form, in which
the eigenfunction for the internal
motion plays the ^{role} ~~form~~ similarly
to the form function in the
theory of non-local interact. between
local fields. However, there is an
important difference between these
two theories. Namely, ^{states} non-local
field is ^{shown} ~~equivalent~~ to a mixture
of local-fields with different
infinite number of masses
and the form functions are uniquely
defined in connection with the
eigenvalue problem for the mass spectrum.

This has its own ~~at~~ disadvantage
as well as advantage compared with

the theory of non-local interaction.
It is more satisfactory in that we
~~for~~ have the means of determining the
form function, whereas on the
other hand, however, the problem
of convergence of self-energy
~~and~~ becomes. The quantities like
much more ~~complicated~~ ^{involved} due to the
appearance of infinitely many
terms already in the lowest
approximation!

~~The complication~~ There is a
further complication which we
~~has~~ ignored in the ^{above} in order to ~~state~~ ^{show}
the general feature of the more clearly
the essential points of our arguments.
Namely, the four-dimensional
in the case

eigenvalue problem ^{with} which we ^{are going} to
deal with, the degeneracy is ~~so~~ very
serious. The difficulty of the infinite
series. In order to make clear
the situation, we take up the

(5)

simple case of the four-dimensional harmonic oscillation in §4. This is the one which were investigated by Born and Green⁽⁵⁾ in connection with the problem of mass-spectrum, although their interpretation was ~~essent~~ quite different from that of the present paper.

In §5, ~~the~~ a tentative method of removing the infinite degeneracy by introducing the coupling between external and internal motion will be discussed. The introduction of ~~the~~ complex ϕ this, however, leads us to the new situation that the field equations for the interacting ^{non-local} fields cannot no longer be brought to the form which is similar to the theory of int non-local fields interaction. ^{Phase} In order to carry out the formal integration of field equations, we must adopt a new method which will be discussed in the next paper.

(6)

2. General considerations on the four dimensional eigenvalue problem

An arbitrary scalar non-local field, which is characterized a scalar $(\phi | \phi | x'_\mu, x''_\mu)$ function of two sets of space-time parameters x'_μ, x''_μ ($\mu = 1, 2, 3, 4$), can be written as in the previous paper as a function

$$(\phi | \phi | x'_\mu, x''_\mu) \equiv \phi(X_\mu, r_\mu)$$

of two sets of ~~vari~~ real variables

$$X_\mu = \frac{1}{2}(x'_\mu + x''_\mu)$$

$$r_\mu = x'_\mu - x''_\mu.$$

We adopt the usual ~~notations~~ convention and the fourth component is always pure imaginary and related to the real time variables which is characterized by the index 0. For instance,

$$x'_4 = ix'_0, \quad X_0 = i r_0 = i r_0$$

where r_0 denotes the ^{internal} relative time in natural units in which $c = \hbar = 1$.

The field equation for the free non-local
For convenience, let us call them external and internal variables respectively.

field $\varphi(X, r)$ is supposed to have the general form

$$F\left(\frac{\partial}{\partial x_\mu}, r_\mu, \frac{\partial}{\partial r_\mu}\right)\varphi(X, r) = 0, \quad (1)$$

where F is a certain invariant function of $\frac{\partial}{\partial x_\mu}, r_\mu, \frac{\partial}{\partial r_\mu}$. Obviously F ~~cannot~~ ^{should be} independent of x_μ in order that it is invariant with respect to the translation of the origin of the Lorentz frame-work.

The ~~explicit~~ ^{explicit} dependence of F both on r_μ and $\frac{\partial}{\partial r_\mu}$ is not only admissible, but also essential for the derivation of the discrete mass-spectrum. ~~It is~~ ^{the dependence} on r_μ

is ~~admissible~~ ^{admissible}, because r_μ denotes ~~the~~ ^{internal or} relative coordinates, ~~while Born and Green's theory is to be reinterpreted~~

It seems to the author that this is the only reasonable re-interpretation of the ~~for~~ the eigenvalue problem proposed by Born and Green ⁽¹⁾.

Thus, in general, F ^{can be} is a function of $\frac{\partial^2}{\partial x_\mu \partial x_\nu}, r_\mu r_\nu, \frac{\partial^2}{\partial r_\mu \partial r_\nu}, r_\mu \frac{\partial}{\partial r_\nu}, r_\mu \frac{\partial}{\partial x_\nu},$

Since $\frac{\partial^2}{\partial x_\mu \partial x_\mu}$. The last two operators refer to the coupling between external and internal motions, let us ignore them for the moment and ^{discuss} assume the simple case $F \equiv$ in which

$$F \equiv -\frac{\partial^2}{\partial x_\mu \partial x_\mu} + F^{(r)}(r_\mu, r_\mu, \frac{\partial^2}{\partial r_\mu \partial r_\mu}, r_\mu \frac{\partial^2}{\partial r_\mu}) \quad (2)$$

so that the field equation (1) can be readily separated by putting $\psi(x, r) = u(x) \chi(r)$.

The external and internal wave functions $u(x)$ and $\chi(r)$ satisfy the equations

$$\left(\frac{\partial^2}{\partial x_\mu \partial x_\mu} - \mu \right) u(x) = 0 \quad (3)$$

$$\left(F^{(r)} - \mu \right) \chi(r) = 0 \quad (4)$$

respectively, where μ is the separation constant,

$\varphi(x_\mu)$: spin mass: arbitrary, continuous
 $\varphi(x'_\mu, x''_\mu)$ spin mass: discrete

$$L(x'_\mu, x''_\mu, \frac{\partial}{\partial x'_\mu}, \frac{\partial}{\partial x''_\mu}) \varphi(x'_\mu, x''_\mu) = 0$$

$$L(x_\mu, \frac{\partial}{\partial x_\mu}, \frac{\partial}{\partial r_\mu}) \varphi(x_\mu, r_\mu) = 0$$

$$x_\mu = \frac{1}{2}(x'_\mu + x''_\mu), \quad r_\mu = x'_\mu - x''_\mu$$

$$h = h^{(ext)} \left(\frac{\partial}{\partial x_\mu} \right) + L^{(int)} \left(r_\mu, \frac{\partial}{\partial r_\mu} \right)$$

$$\varphi(x_\mu, r_\mu) = \sum_n \varphi_n^{(ex)}(x_\mu) \varphi_n^{(in)}(r_\mu)$$

$$h^{(int)} \left(r_\mu, \frac{\partial}{\partial r_\mu} \right) \varphi_n^{(in)}(r_\mu) = -\frac{1}{2} m_n^2 \varphi_n^{(in)}(r_\mu)$$

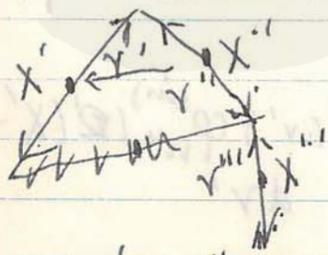
$$h^{(int)} \left(\frac{\partial}{\partial x_\mu} \right) = \frac{1}{2} \frac{\partial^2}{\partial x_\mu \partial x_\mu}$$

$m_n^2 \geq 0$: $h^{(int)}$ has positive eigenvalues.

Interaction

$$L_I = \int \bar{\psi}(x) \varphi(x) \psi(x) = \int \int \int \bar{\psi}_n^{(ex)}(x) \varphi_n^{(ex)}(x) \psi_n^{(ex)}(x) \Phi_{n'n''n'''}(x, x', x'') dx' dx'' dx'''$$

$$\Phi_{n'n''n'''} = \int \bar{\psi}_n^{(in)}(r) \varphi_n^{(in)}(r) \psi_n^{(in)}(r) x dr'$$



$$= \int \bar{\psi}_n(r') \varphi_n(2(x'' - x') - r')$$

$$\times \psi_n(r' + 2(x'' - x') - 2(x'' - x')) dr' = \Phi_{n'n''n'''}(x'' - x', x'' - x''')$$

$$r' + r'' = 2(x'' - x')$$

$$r'' + r''' = 2(x'' - x''')$$

$$r''' - r' = 2(x''' + x' - 2x'')$$

$$r' + r''' = 2(x''' - x')$$

happening formalism:

$$\bar{L}_0 = \int \varphi(x', x'') \left(L^{(ex)} + L^{(in)} \right) \varphi(x''', x''') dx' dx'' dx'''$$

$$\left(\cancel{p_\mu \varphi} + \varphi p_\mu \right) + L^{(in)} \left(\cancel{p_\mu \varphi} + \varphi p_\mu \right)$$

$$\varphi_{n'}^{(ex)}(x') \varphi_{n''}^{(in)}(r') \quad \varphi_{n'''}^{(ex)}(x''') \varphi_{n''''}^{(in)}(r''')$$

$$\bar{L}_0 = \text{Tr} \left[\frac{1}{2} [p_\mu \varphi] [p_\mu \varphi] + \frac{1}{2} [p_\mu \varphi] + [p_\mu \varphi] + \frac{1}{2} [x_\mu \varphi] [x_\mu \varphi] \right]$$

$$= \int \left(\frac{1}{2} \sum_{n'n''} \frac{\partial \varphi_{n'}^{(ex)}(x')}{\partial x'} \frac{\partial \varphi_{n''}^{(ex)}(x'')}{\partial x''} \right) \left(\varphi_{n'}^{(in)}(r') \varphi_{n''}^{(in)}(r'') \right) dx' dx''$$

$\gamma' + \gamma'' = 2(x' - x'')$
 $\gamma'' = 2(x' - x'') - \gamma'$

$$\bar{L}_0 = \frac{1}{2} \int \left(\frac{\partial \varphi_{n'}^{(ex)}(x')}{\partial x'} \frac{\partial \varphi_{n''}^{(ex)}(x'')}{\partial x''} \right) + \frac{1}{2} \sum_{n'n''} \varphi_{n'}^{(ex)}(x') \varphi_{n''}^{(ex)}(x'')$$

$\kappa_{n'n''} = \varphi_{n'}^{(in)}(r') \varphi_{n''}^{(in)}(r'')$

$$\bar{\Gamma}_{n'n''}(x' - x'') = \int \varphi_{n'}^{(in)}(r') \varphi_{n''}^{(in)}(R(x' - x'') - r')$$

Main Factor:

Normalization factor: $\bar{\Gamma}_{n'n''}(x' - x'')$

Form factor: $\bar{\Phi}_{n'n''n'''}(x' - x'', x'' - x''')$

$$\int \left\{ \frac{\partial \varphi(r')}{\partial x'_\mu} \frac{\partial \varphi(r'')}{\partial x''_\mu} + r'_\mu \varphi(r') r''_\mu \varphi(r'') \right\} dr'$$

$$r'' = 2(X' - X'') - r'$$

$$\bar{h}_0 = \iint (x' | h_0 | x'') d(x' - x'')$$

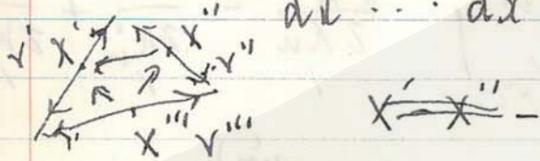
$$= \frac{1}{2} \int \left(\frac{\partial \varphi_n(x')}{\partial x'} \right)^2 + m^2 \varphi_n^2(x') \frac{\pm \delta_{n'n''}}{r_{n'n''}} dx'$$

$$\bar{r}_{n'n''} = \int \varphi_n^{(in)}(r') \varphi_n^{(in)}(-r'') dr'$$

$$\varphi_n^{(in)}(r') = \pm \varphi_n^{(in)}(r'')$$

$$= \pm \delta_{n'n''}$$

$$\bar{L}_1 = g \int \bar{\varphi}(x'x'') \varphi(x''x''') \psi(x''', x''') d(x', x''')$$



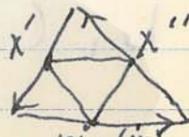
$\Rightarrow 2f$

$$r' + r'' + r''' = 0$$

$$r' = 2(X'' - X' + 2X''' + X' - 2X'')$$

$$\bar{\varphi}_{n'n''n'''} = \int \bar{\varphi}_n(r') \varphi_n(2(X' - X'') - r') \psi(2(X''' - X'') - r')$$

$$r' = 2(X''' - X'')$$



$$\bar{\varphi}_{n'n''n'''} \propto \int \varphi_n^{(in)}(X''' - X'') \varphi_n^{(in)}(2(X' - X'')) \psi_n(2(X'' - X'))$$

$$\varphi_n^{(in)} \propto \delta(2(X'' - X'))$$

$$\bar{\varphi}_{n'n''n'''} \propto \varphi_n^{(in)}(2(X' - X'')) \delta(2(X'' - X'))$$

$$L_1 = g \int \psi^\alpha(x) \varphi_{n''}(x) \varphi_{n''}(x) dx$$

(ea)

(P n'' n'')

One cannot go over to the limit $\varphi_{n''} \rightarrow \delta$ or $\varphi, \psi \rightarrow \delta$ because then the h will be infinite (one has too many δ -funct.)

$$L_0(\{p_\mu, \varphi\}, \frac{1}{2}\{p_\mu, \varphi\}_+, \{x_\mu, \varphi\})$$

$$X_\mu = \frac{1}{2}(x'_\mu + x''_\mu)$$

$$Y_\mu = x'_\mu - x''_\mu$$

$$x'_\mu = X_\mu + \frac{1}{2} Y_\mu$$

$$x''_\mu = X_\mu - \frac{1}{2} Y_\mu$$

$$\frac{\partial}{\partial x_\mu} = \frac{1}{2} \left(\frac{\partial}{\partial x'_\mu} - \frac{\partial}{\partial x''_\mu} \right)$$

$$\frac{\partial}{\partial x_\mu} = \frac{\partial}{\partial X_\mu} + \frac{\partial}{\partial Y_\mu}$$

$$L_0 \circledast = h^{(ea)}(\{p_\mu, \varphi\}) + h^{(in)}\left(\frac{1}{2}\{p_\mu, \varphi\}_+, \{x_\mu, \varphi\}\right)$$

$$L_1(\varphi, \psi, \tilde{\psi}) \circledast$$

the form for

$$\Phi_{n''n'n''} = \frac{\psi_{n'}^{(in)}(2(X''-X')) \varphi_{n''}^{(in)}(2(X'-X''))}{x \varphi_{n''}^{(in)}(2(X''-X'))}$$

is not invariant by itself.

$$\Phi_{n''n'n''}$$

In order to have only one particle with real masses, we have to take, for instance:

$$h^{(in)} = \left\{ \left(\frac{1}{2} [p_\mu, \varphi] \right)^2 + \frac{1}{\lambda^2} [x_\mu, \varphi]^2 \right\}^2$$

together with supplementary conditions,

eigenvalues:

$$m^2 n = \mu^2 (n_1 + n_2 + n_3 - n_0)^2$$

~~(in)~~ degenerated

$n = 0; \quad n_1 = n_2 = n_3 = n_0 = 0.$

Rotational
 Supplementary condition
 or Rotational energy

$$f(r_\mu \frac{\partial}{\partial x_\mu}) \varphi = 0$$

$$\varphi \propto e^{i k_\mu x_\mu} \varphi_{k_\mu}(r_\mu)$$

$$k_\mu r_\mu \cdot \varphi_{k_\mu} = 0$$

$$\varphi_{k_\mu}^{(in)} \delta(k_\mu r_\mu)$$

$$= \varphi_{k_\mu}^{(in)}(r'_1, r'_2, r'_3) \delta(r'_4)$$

$r'_4 \rightarrow \frac{r'_1 + r'_2 + r'_3}{\sqrt{3}}$

097 (4)

$$\left(\frac{\partial^2}{\partial x_4'^2} + \frac{1}{\lambda^4} \nabla_4'^2 \right) \varphi_{k\mu}^{(in)}(x'_\mu) = 0 \quad \varphi_{k\mu}^{(in)}(x'_\mu)$$

$$n_0 = 0:$$

$$\mu^2 n^2 = \mu^2 (n_1 + n_2 + n_3)$$

$$n = 0: \quad n_1 = n_2 = n_3 = 0$$

$$\gamma_4' = \gamma_\mu' \frac{\partial}{\partial x_\mu}$$

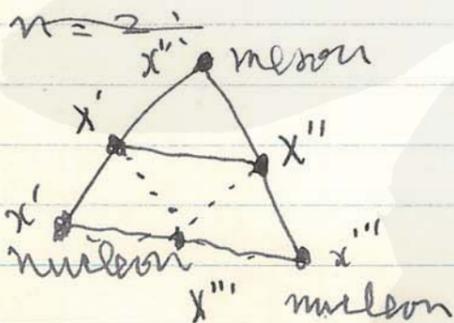
$$\frac{\partial}{\partial x_4'} = \frac{1}{\mu^2} \frac{\partial}{\partial x_\mu} \gamma_\mu'$$

$$\frac{1}{\mu^2} \left\{ \left(\frac{\partial}{\partial x_\mu} \right)^2 + \frac{1}{\lambda^4} \left(\gamma_\mu' \frac{\partial}{\partial x_\mu} \right)^2 \right\} \varphi(x, \gamma) = 0$$

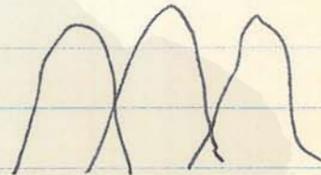
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degenerate

$$n=1: \quad \begin{cases} n_1=1, n_2=n_3=0 \\ n_1=0, n_2=1, n_3=0 \\ n_1=0, n_2=0, n_3=1 \end{cases}$$



$$\varphi_{1,0,0}^{(in)}(x'-x''') \propto (x'-x''')$$



Selection rule

$$n' + n'' + n''' : \text{even}$$

①

$$\left(\frac{\partial^2}{\partial r_4'^2} + \frac{1}{x_4} \frac{\partial}{\partial r_4'} \right) \varphi_{k\mu}^{(in)}(r_4') = 0$$

~~$\left(\frac{\partial^2}{\partial r_4'^2} + \frac{1}{x_4} \frac{\partial}{\partial r_4'} \right) \varphi_{k\mu}^{(in)}(r_4') = 0$~~

[Faint handwritten notes and diagrams, including a sine wave and a network diagram]

structure

Space-Time structure
 of the universe
 Modified Minkowski
 world

Dec. 17
1952

098 (1)

$$\vec{z}_\mu = x_\mu \left(1 - \frac{\lambda^2}{x_1^2 + x_2^2 + x_3^2 + x_4^2} \right)$$

$\mu = 1, 2, 3, 0$ $x_0 = ct$

$$\rho^2 = z_1^2 + z_2^2 + z_3^2 + z_0^2$$

$$= R^2 \left(1 - \frac{\lambda^2}{R^2} \right)^2$$

$$\frac{\rho^2}{R^2} = \left(1 - \frac{\lambda^2}{R^2} \right)^2$$

$$1 - \frac{\lambda^2}{R^2} = \pm \frac{\rho}{R}$$

$$\frac{\lambda^2}{R^2} \pm \frac{\rho}{R} - 1 = 0$$

$$\textcircled{+} \quad \frac{\lambda}{R} = \frac{1}{2} \left(\frac{\rho}{\lambda} \pm \sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} \right)$$

$$\textcircled{+} \quad \frac{\lambda}{R} = \frac{1}{2} \left(\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} - \frac{\rho}{\lambda} \right)$$

$$z_\mu = z_\mu \left(1 - \frac{1}{4} \left(\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} - \frac{\rho}{\lambda} \right)^2 \right)$$

$$\frac{R}{\lambda} = \frac{x_\mu}{\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} - \frac{\rho}{\lambda}}$$

$$\rho \rightarrow 0 \quad R \rightarrow \lambda$$

To the point $x_\mu, z_\mu = 0$, the sphere

$$x_1^2 + x_2^2 + x_3^2 + x_0^2 = \lambda^2$$

corresponds.

Minkowski world

$$ds^2 = \epsilon_{\mu\nu} dz_\mu dz_\nu$$

$$dz_\mu = dx_\mu \left(1 - \frac{\lambda^2}{x_\mu x_\mu} \right) + \frac{\lambda^2}{(x_\mu x_\mu)^2} 2x_\mu x_\nu dx_\nu$$

$$= \left[\left(1 - \frac{\lambda^2}{x_\nu x_\nu} \right) \delta_{\mu\lambda} + \lambda^2 \frac{2x_\mu x_\lambda}{(x_\nu x_\nu)^2} \right] dx_\lambda$$

$$\begin{pmatrix} 1 - \frac{\lambda^2}{(x_\nu x_\nu)^2} (x_\nu x_\nu - 2x_\mu^2) & \frac{2\lambda^2 x_\mu x_\nu}{(x_\nu x_\nu)^2} & & \\ \frac{2\lambda^2 x_1 x_2}{(x_\nu x_\nu)^2} & 1 - \frac{\lambda^2}{(x_\nu x_\nu)^2} (x_\nu x_\nu - 2x_2^2) & & \\ \frac{2\lambda^2 x_1 x_3}{(x_\nu x_\nu)^2} & \frac{2\lambda^2 x_2 x_3}{(x_\nu x_\nu)^2} & & \\ \frac{2\lambda^2 x_1 x_0}{(x_\nu x_\nu)^2} & \frac{2\lambda^2 x_2 x_0}{(x_\nu x_\nu)^2} & & \end{pmatrix}$$

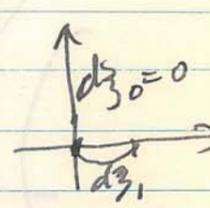
$$dz_\mu = \frac{\partial z_\mu}{\partial x_\lambda} dx_\lambda$$

$$\epsilon_{\mu\nu} dz_\mu dz_\nu = \epsilon_{\mu\nu} \frac{\partial z_\mu}{\partial x_\lambda} \frac{\partial z_\nu}{\partial x_\kappa} dx_\lambda dx_\kappa$$

$$g_{\mu\nu} \frac{\partial z^\mu}{\partial x^\lambda} \frac{\partial z^\nu}{\partial x^\kappa} = g_{\lambda\kappa}$$

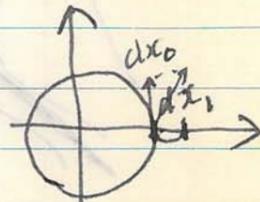
$\parallel \delta_{\mu\nu} - \eta_{\mu\nu} \parallel \leq 1$ $\parallel \eta_{\mu\nu} \parallel \leq 1$

$R \geq \lambda$ $\parallel \delta_{\mu\nu} - \eta_{\mu\nu} \parallel \leq 1$



$x_2 = x_3 = x_0 = 0$ $|x_1| \geq \lambda$

$$1 + \frac{\lambda^2}{x_1^2} x_1^2 = 1 + \frac{\lambda^2}{x_1^2}$$



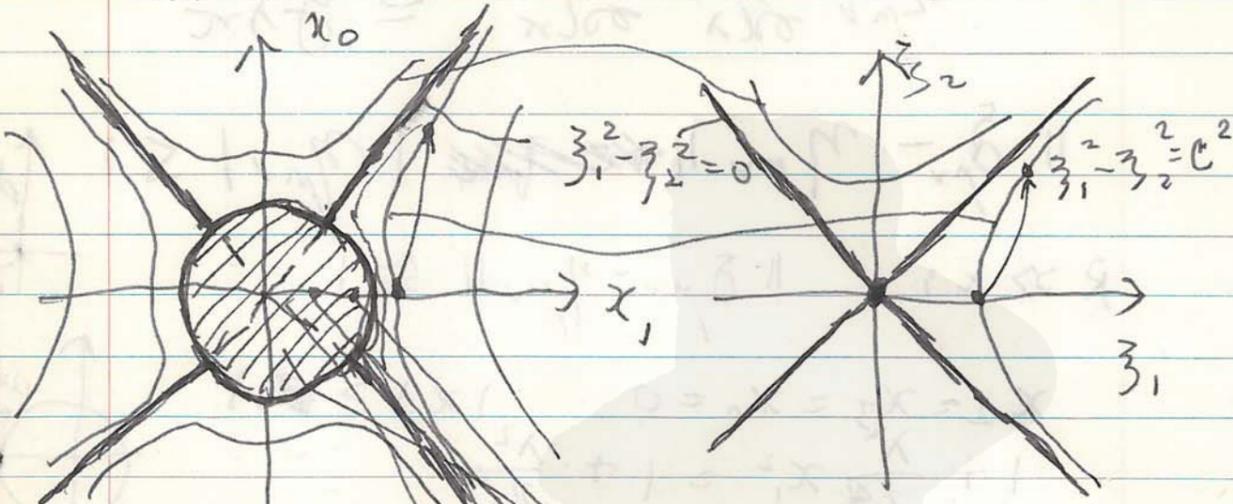
$$\begin{vmatrix} 1 + \frac{\lambda^2}{x_1^2} & 0 & 0 & 0 \\ 0 & 1 - \frac{\lambda^2}{x_1^2} & 0 & 0 \\ 0 & 0 & 1 - \frac{\lambda^2}{x_1^2} & 0 \\ 0 & 0 & 0 & 1 - \frac{\lambda^2}{x_1^2} \end{vmatrix} > 0$$

$dz_1 = (1 + \frac{\lambda^2}{x_1^2}) dx_1$
 $dz_2 = (1 - \frac{\lambda^2}{x_1^2}) dx_1$
 $dz_3 = (1 - \frac{\lambda^2}{x_1^2}) dx_0$

for $x_1^2 > \lambda^2$

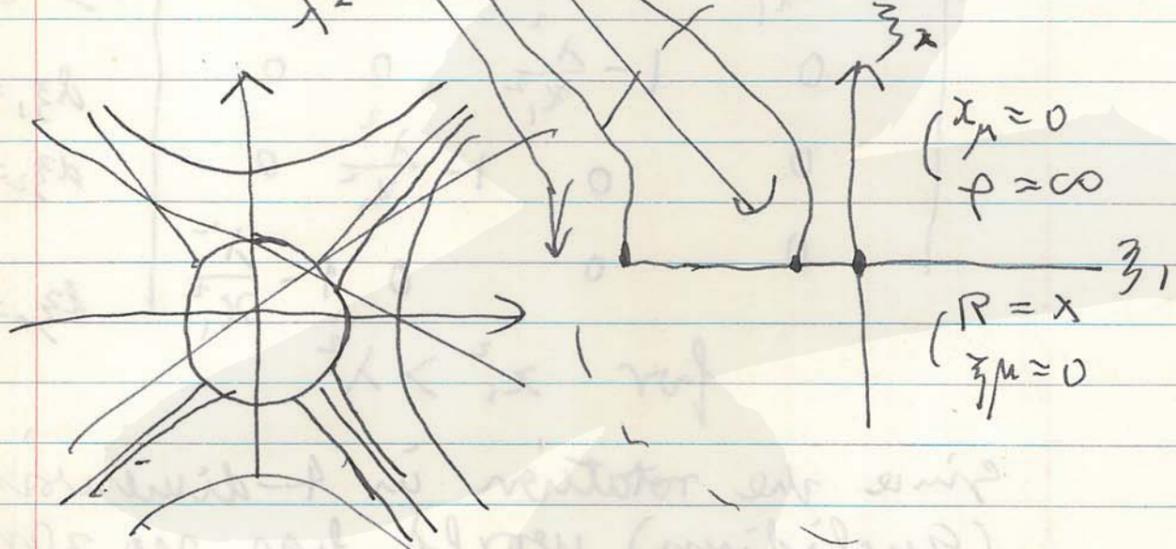
Since the rotation in 4-dimensional (Euclidian) world has no zero jacobian and transforms any point with $R \geq \lambda$ into another point with $R \geq \lambda$, the jacobian for the transformation $x_\mu \rightarrow z_\mu$ or its inverse is different from zero, provided that $R > \lambda$

Let us take two dimensional case:



$$\delta(x_1 + x_2^2 + x_3^2 + x_0 - \lambda^2)$$

$$\delta(z_1) \delta(z_2) \delta(z_3) \delta(z_0)$$



$$\square \varphi(x_\mu) = -\delta(x_\mu / \mu - \lambda^2) \quad \square \varphi = -\delta$$

$$\begin{aligned}
 & dz_1 dz_2 dz_3 dz_0 \\
 &= \sqrt{\left(1 + \frac{\lambda^2}{R^2}\right) \left(1 - \frac{\lambda^2}{R^2}\right)^3} \cdot dx_1 dx_2 dx_3 dx_0 \\
 &\delta(z_1) \delta(z_2) \delta(z_3) \delta(z_0) \int f(x_1, x_2, x_3, x_0) \frac{\delta(x_1^2 + x_2^2 + x_3^2 + x_0^2 - \lambda^2)}{\sqrt{\left(1 + \frac{\lambda^2}{R^2}\right) \left(1 - \frac{\lambda^2}{R^2}\right)^3} \cdot \lambda^2}
 \end{aligned}$$

Modified Lorentz transformation

$$z'_\mu = a_{\mu\nu} z_\nu$$

$$\begin{aligned} x'_\mu(z'_\mu) &= x'_\mu(a_{\mu\nu} z_\nu) \\ &= x'_\mu(a_{\mu\nu} z_\nu(\lambda^2)) \end{aligned}$$

$$x_\mu x_\mu \gg \lambda^2: \quad z_\nu \approx x_\nu$$

$$\begin{aligned} x'_\mu &= x'_\mu(a_{\mu\nu} x_\nu) \\ x'_\mu x'_\mu &> \lambda^2 \end{aligned}$$

$$\varphi(x', x'') = \varphi(X, r)$$

X : Minkowski space

r : modified space

$$x' = X + \frac{r}{2} \quad x'' = X - \frac{r}{2}$$

$$e^{ik_\mu x_\mu} f(r_\mu)$$

limit to the point particle

$$f(r_\mu) \approx 0 \quad \text{for} \quad r_\mu r_\mu > \lambda^2$$

Momentum space.

$$f(x_\mu) = \int d^4k e^{ik_\mu x_\mu} g(k_\mu) \cdot d^4k$$

Distance of two points in x -space is approximately

$x_\mu, y_\mu (x_\mu - y_\mu) (x_\mu^2 - y_\mu^2)$,
if $(x_\mu - y_\mu) (x_\mu - y_\mu)$ is large compared with λ^2 .

On the contrary, if two points x_μ, y_μ are close to each other, the distance is given by

$$(\xi_\mu - \eta_\mu) (\xi_\mu - \eta_\mu),$$

where

$$\xi_\mu = x_\mu \left(1 - \frac{\lambda^2}{(x_\mu - y_\mu)(x_\mu - y_\mu)}\right)$$
$$\eta_\mu = y_\mu \left(1 - \frac{\lambda^2}{(x_\mu - y_\mu)(x_\mu - y_\mu)}\right)$$

If $(x_\mu - y_\mu)(x_\mu - y_\mu) \ll \lambda^2$, the distance could be very large again compared with λ^2 .

The metric in x -space is very complicated, but for any two points

$$x_\mu, y_\mu: \begin{cases} x_\mu = x_\mu + \frac{1}{2} r_\mu \\ y_\mu = x_\mu - \frac{1}{2} r_\mu \end{cases}$$

x_μ -space: Minkowski
 y_μ -space: modified Minkowski space
 ~~$\tilde{F}(x_\mu) \tilde{f}(y_\mu)$~~

$$\Phi(x_\mu - y_\mu) \Psi(x_\mu) \varphi(y_\mu)$$

$$x_\mu = x_\mu + \frac{1}{2} r_\mu$$

$$y_\mu = x_\mu - \frac{1}{2} r_\mu$$

$$x'_\mu = a_\mu + a_{\mu\nu} x_\nu$$

$$r'_\mu = r'_\mu(r_\mu)$$

$$r_\mu r_\mu, r'_\mu r'_\mu \geq \lambda^2$$

$$x'_\mu = a_\mu + a_{\mu\nu} x_\nu + \frac{1}{2} r'_\mu(r_\mu)$$

$$y'_\mu = a_\mu + a_{\mu\nu} x_\nu - \frac{1}{2} r'_\mu(r_\mu)$$

$$x'_\mu = a_\mu + a_{\mu\nu} \frac{x_\nu + y_\nu}{2} + \frac{1}{2} r'_\mu(x_\nu - y_\nu)$$

$$y'_\mu = a_\mu + a_{\mu\nu} \frac{x_\nu + y_\nu}{2} - \frac{1}{2} r'_\mu(x_\nu - y_\nu)$$

~~dz_μ~~
 Ordinary space-time world

$$dx_\mu = dz_\mu \left(1 - \frac{1}{4} \left(\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} - \frac{\rho}{\lambda} \right)^2 \right)$$

$$+ \sum_{\mu} \frac{\frac{1}{4\lambda^2} \sum_{\nu} z_\nu dz_\nu}{\left(1 - \frac{1}{4} \left(\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} - \frac{\rho}{\lambda} \right)^2 \right)^2} dz_\nu$$

$$+ \sum_{\mu} \frac{\frac{1}{2} \left(\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} - \frac{\rho}{\lambda} \right)}{\left(1 - \frac{1}{4} \left(\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} - \frac{\rho}{\lambda} \right)^2 \right)^2} \left\{ \frac{\frac{1}{\lambda^2} \sum_{\nu} z_\nu dz_\nu}{\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4}} - \frac{\sum_{\nu} z_\nu dz_\nu}{\lambda \cdot \rho} \right\}$$

$$= \frac{dz_\mu}{\left(1 - \frac{1}{4} \left(\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} - \frac{\rho}{\lambda} \right)^2 \right)}$$

$$+ \frac{\frac{1}{2} \frac{1}{\lambda^2} \sum_{\mu} z_\mu dz_\mu}{\left(1 - \frac{1}{4} \left(\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} - \frac{\rho}{\lambda} \right)^2 \right)^2} \frac{- \left(\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} - \frac{\rho}{\lambda} \right)^2}{\sqrt{\left(\frac{\rho}{\lambda}\right)^2 + 4} \cdot \frac{\rho}{\lambda}}$$

$$g_{\mu\nu} dx_\mu dx_\nu = \Gamma_{\mu\nu} dz_\mu dz_\nu$$

if $\Gamma_{\mu\nu}$ is finite (non-zero)
 at $x_\mu x_\nu = \lambda^2$.

Connection Between Two Spaces

(i) There is one-to-one correspondence between the whole space of z_μ and the space of x_μ outside of the sphere $x_\mu x_\mu = \lambda^2$.

(ii) As for the metric properties of z_μ (or x_μ), one may either assume that $g_{\mu\nu}$ vanishes on $x_\mu x_\mu = \lambda^2$ in such a way that $\Gamma_{\mu\nu}$ is regular at $z_\mu = 0$ or $g_{\mu\nu}$ is regular on $x_\mu x_\mu = \lambda^2$ but does not behave so as to give regular $\Gamma_{\mu\nu}$ at $z_\mu = 0$.

If x_μ is our actual space, the quantity which corresponds to the relative coordinates of two particles, one can say that it is ~~not~~ impossible that two particles come closer than λ four dimensionally (in any coordinate system and irrespective of the metric). This is clear, if z_μ is a Minkowski space. If, on the contrary, x_μ is a space which is not too singular on $x_\mu x_\mu = \lambda^2$, then z_μ -space may be singular at $z_\mu = 0$.

construct a field theory in x -space
 and then
 then, if we go over to the limit $\lambda=0$,
 we may ~~not~~ have a ^{field} theory which
 is quite different from the
 ordinary ^{field} theory.

$$z_\mu = x_\mu \sqrt{1 - \frac{\lambda^2}{x_1^2 + x_2^2 + x_3^2 + x_0^2}}$$

$$dz_\mu = \frac{dx_\mu \sqrt{1 - \dots}}{\sqrt{1 - \frac{\lambda^2}{x_1^2 + \dots}}} + \frac{\lambda^2 x_\nu dx_\nu}{(x_1^2 + \dots)^2} \frac{x_\mu}{\sqrt{1 - \frac{\lambda^2}{x_1^2 + \dots}}}$$

$$= \frac{\left(1 - \frac{\lambda^2}{x_\mu x_\mu}\right) dx_\mu + \frac{\lambda^2 x_\mu x_\nu dx_\nu}{(x_\mu x_\mu)^2}}{\sqrt{1 - \frac{\lambda^2}{x_\mu x_\mu}}}$$

$$\begin{pmatrix} \frac{1}{\sqrt{1 - \frac{\lambda^2}{x_\nu x_\nu}}} \left\{ 1 - \frac{\lambda^2}{(x_\nu x_\nu)^2} (x_\nu x_\nu - x_1^2) \right\} & \frac{1}{\sqrt{1 - \frac{\lambda^2}{x_\nu x_\nu}}} \frac{x_1 x_2}{(x_\nu x_\nu)^2} \dots \\ \frac{1}{\sqrt{1 - \frac{\lambda^2}{x_\nu x_\nu}}} \frac{x_1 x_2}{(x_\nu x_\nu)^2} & \frac{1}{\sqrt{1 - \frac{\lambda^2}{x_\nu x_\nu}}} \left\{ 1 - \frac{\lambda^2}{(x_\nu x_\nu)^2} (x_\nu x_\nu - x_2^2) \right\} \dots \end{pmatrix}$$

$x_2 = x_3 = x_0 = 0$

$$\begin{pmatrix} \frac{1}{\sqrt{1 - \frac{\lambda^2}{x_1^2}}} & 0 \\ 0 & \frac{1}{\sqrt{1 - \frac{\lambda^2}{x_1^2}}} \left(1 - \frac{\lambda^2}{x_1^2}\right) \end{pmatrix}$$

$$= \frac{(1 - \frac{\lambda^2}{x_1^2})^3}{(\sqrt{1 - \frac{\lambda^2}{x_1^2}})^4} = 1 - \frac{\lambda^2}{x_1^2} > 0$$

for $x_1^2 > \lambda^2$.

$$\rho^2 = \xi_\mu \xi_\mu$$

$$= x_\mu x_\mu \left(1 - \frac{\lambda^2}{x_\mu x_\mu}\right)$$

$$\frac{\rho^2}{x_\mu x_\mu} = 1 - \frac{\lambda^2}{x_\mu x_\mu} \rightarrow \approx \frac{\rho^2}{\rho^2 + \lambda^2} \quad \left(= \frac{1}{1 + \frac{\lambda^2}{\rho^2}} \right)$$

$$\frac{\rho^2 + \lambda^2}{x_\mu x_\mu} = 1 \quad x_\mu x_\mu = \rho^2 + \lambda^2$$

$$\xi_\mu = x_\mu \sqrt{1 - \frac{\lambda^2}{\rho^2 + \lambda^2}}$$

$$x_\mu = \frac{\xi_\mu}{\sqrt{1 - \frac{\lambda^2}{\rho^2 + \lambda^2}}} = \xi_\mu \frac{\sqrt{\rho^2 + \lambda^2}}{\rho}$$

$$= \xi_\mu \sqrt{1 + \frac{\lambda^2}{\rho^2}}$$

$\rho \rightarrow 0$: $x_\mu x_\mu \rightarrow \lambda^2$

$$\xi_{\mu\nu} \xi_\mu \xi_\nu = \xi_{\mu\nu} x_\mu x_\nu \left(1 - \frac{\lambda^2}{x_\mu x_\mu}\right)$$

$$\delta d\zeta_1 d\zeta_2 d\zeta_3 d\zeta_0 = \left(1 - \frac{\lambda^2}{x_\mu x_\mu}\right) dx_1 dx_2 dx_3 dx_0$$

$$\delta(\zeta_1) \delta(\zeta_2) \delta(\zeta_3) d(\zeta_0)$$

$$= \text{const.} \frac{\delta(x_\mu x_\mu - \lambda^2)}{1 - \frac{\lambda^2}{x_\mu x_\mu}} \quad \mathbb{R}^4$$

$$\varphi(\zeta_1, \zeta_2, \zeta_3, \zeta_0) = \varphi(x_1, x_2, x_3, x_0)$$

$\varphi(x_\mu)$ may depend

$$\varphi(\zeta_\mu, \zeta_\nu) = \varphi\left(\frac{x_\mu x_\nu}{x_\mu x_\mu}\left(1 - \frac{\lambda^2}{x_\mu x_\mu}\right)\right)$$

$$\zeta_\mu \rightarrow 0, \quad x_\mu x_\mu \rightarrow \lambda^2$$

Any function of the form

$$f(x_1, x_2, x_3, x_4) \delta(x_\mu x_\mu - \lambda^2)$$

can be an invariant. However,

$$\int f(x_1, x_2, x_3, x_4) \delta(x_\mu x_\mu - \lambda^2) dx_1 dx_2 dx_3 dx_4$$

depends on the distribution of f on the sphere $x_\mu x_\mu = \lambda^2$.

Jan. 1953

Carl Hevinson, S-Matrix in
 local and Non-local Field Theories

$$A_H(x) = U(-\infty, t) A_I(x) U^{-1}(-\infty, t)$$

$$A_M(x) \equiv S A_H(x) = U(\infty, t) A_I(x) U(t, -\infty)$$

$$A_M(x) = \sum_{n=0}^{\infty} \frac{(-i)^n}{n!} \int_{-\infty}^{\infty} T(H_1 \dots H_n, A_I(x))$$

$$\int_{\sigma(t)} H_1(x_1) \theta(x_1, x_2) H_2(x_2) \dots \theta(x_j, x_{j+1}) A(x) \theta(x_{j+1}, x_n)$$

where $\theta(x_j, x_{j+1})$ can be 1, 2, ..., n, depending
 according as if we take together
 various term in the product of
 two ψ 's which are proportional

to g^n } The net result is

$$\sum_j \int_{-\infty}^{\infty} H_1(x_1) \theta(x_1, x_2) H_2(x_2) \dots \theta(x_j, x) A(x) \theta(x, x_{j+1}) H_{j+1}(x_{j+1})$$

$$= \frac{1}{n!} P \sum_j \int_{-\infty}^{\infty} H_1(x_1) \theta(x_1, x_2) H_2(x_2) \dots$$

S Matrix Local Theory - Call Levinson

First Proof - Valid only for Scalar Coupling

For any field operator (time derivatives of field operators excluded) we have:

$$A_H(x) = U(-\infty, \pm) A_I(x) U^{-1}(-\infty, \pm)$$

$A_H(x)$ = Heisenberg Representation

$A_I(x)$ = Interaction Representation

$U(a,b)$ = Dyson's "U" Matrix

Also:

$$(1) A_M(x) \equiv S A_H(x) = U(\infty, \pm) A_I(x) U(\pm, -\infty) \left\{ \begin{array}{l} \text{Dyson's "Mixed"} \\ \text{Representation} \end{array} \right.$$

$$(2) A_M(x) = \sum_{n=0}^{\infty} \frac{(-i)^n}{n!} \int_{-\infty}^{\infty} T(H_1, \dots, H_n, A_I(x)) dx_1 \dots dx_n \left\{ \begin{array}{l} \text{Proven in} \\ \text{Dyson's First} \\ \text{Paper} \end{array} \right.$$

$H_1 = H_I^{\text{int}}(x_1)$ the interaction representation Hamiltonian.

Define $H_M(x) \equiv \sum_{n=1}^{\infty} H_M^{(n)}(x)$ where $H_M^{(n)}(x)$ is

the term that goes as g^n in an expansion of $H_M(x)$ in a power series in g .

$$\therefore H_M^{(n)}(x) = \frac{(-i)^{n-1}}{(n-1)!} \int T(H_1, \dots, H_{n-1}, H(x)) dx_1 \dots dx_{n-1} \quad n=1, 2, \dots$$

Now Dyson's Expansion for the S Matrix is:

$$\begin{aligned}
 (3) \quad S &= \sum_{n=0}^{\infty} \frac{(-i)^n}{n!} \int T(H_1, \dots, H_n) dx_1 \dots dx_n = \\
 &= 1 + \sum_{n=1}^{\infty} \frac{(-i)^{n-1}}{(n-1)!} \left(-\frac{i}{n}\right) \int T(H_1, \dots, H_{n-1}, H_n) dx_1, \dots, dx_{n-1}, dx_n \\
 &= 1 + \sum_{n=1}^{\infty} \left(-\frac{i}{n}\right) \int H_M^{(n)}(x) dx
 \end{aligned}$$

but $H_M(x) = S H_H(x) \quad \therefore H_M^{(n)}(x) = (S H_H(x))^{(n)}$

where H_H = interaction part of the Hamiltonian in the Heisenberg representation.

$$\begin{aligned}
 \therefore S &= 1 + \sum_{n=1}^{\infty} \left(-\frac{i}{n}\right) \int (S H_H(x))^{(n)} d^4x \\
 &= 1 + \sum_{n=1}^{\infty} \left(-\frac{i}{n}\right) \left(S \int H_H(x) dx\right)^{(n)}
 \end{aligned}$$

set $\int H_H(x) dx = -\int \mathcal{L}_H^{int}(x) dx = -I$

$\mathcal{L}_H^{int}(x)$ = interaction part of Lagrangian in the Heisenberg representation:

$$\therefore S = 1 + \sum_{n=1}^{\infty} \left(\frac{+i}{n}\right) (SI)^{(n)}$$

or $S^{(n)} = \frac{i}{n} (SI)^{(n)} \quad n=1, 2, \dots$

(4) or $n S^{(n)} = i(SI)^{(n)} \quad n=0, 1, 2, \dots$

$$\text{Now } \left(g \frac{\partial S}{\partial g} \right)^{(n)} = N S^{(n)}$$

$$\therefore \left(g \frac{\partial S}{\partial g} \right)^{(n)} = i (S I)^{(n)}$$

Since the expressions are equal term by term they are equal in general.

$$\therefore \boxed{g \frac{\partial S}{\partial g} = i S I}$$

This follows also from Schwinger's

$$\boxed{\delta S = i S \delta W}$$

P.R. 82, 916 Equation 2.7

where $U_{0,-\infty}^{-1} = S$

where $\delta W = \frac{\delta g}{g} \int \mathcal{L}_+^{\text{int}}(x) dx$ is the change in the Lagrangian due to replacing g by $g + \delta g$.

Lemma $[I, S]_- = 0$

We now ~~red~~ repeat the above proof except that we transform everything by $S^{-1} \dots S$.

$$(1)' \quad A'_m(x) = A_m(x) S \quad \text{where } S^{-1} A_m S = A_{\text{out}}$$

$$(2)' \quad A'_m(x) = \sum_{n=0}^{\infty} \frac{(-i)^n}{n!} \int_{-\infty}^{\infty} T(N_1^{\text{out}} \dots N_n^{\text{out}}, A^{\text{out}}(x)) dx_1 \dots dx_n$$

$$(3)' \quad S^{-1} S S = S = \sum_{n=0}^{\infty} \frac{(-i)^n}{n!} \int T(N_1^{\text{out}} \dots N_n^{\text{out}}) dx_1 \dots dx_n$$

$$= \mathbb{1} + \sum_{n=1}^{\infty} \left(\frac{-i}{n} \right) \int H_n^{(n)}(x) dx$$

$$= \mathbb{1} + \sum_{n=1}^{\infty} \left(\frac{-i}{n} \right) \left(\int H_n(x) dx S \right)^{(n)}$$

$$(4)' \therefore n S^{(n)} = i(I S)^{(n)}$$

Comparing (4) and (4)' we see

$$\text{that } (SI)^{(n)} = (IS)^{(n)} \quad \text{or } [I, S]_- = 0 \quad \text{Q.E.D.}$$

Second Type of Proof Applicable to
 Derivative Coupling Theories.

We shall prove the above theorem for the
 case of Pseudo-scalar mesons with Pseudo-vector
 coupling. First we must verify Schwinger's
 variation formula.

$$\delta \Psi_H(x, g) = -\frac{i \delta g}{g} \left[\int_{-\infty}^t \mathcal{L}_H^{\text{int}}(x) d^4x, \Psi_H(x, g) \right].$$

i.e. $\delta \Psi \equiv$ the change in $\Psi_H(x)$ due to a change
 in the equations of motion when g is increased
 to $g + \delta g$. (As a boundary condition we have
 taken $\delta \Psi = 0$ for $t = -\infty$.)

$$\mathcal{L}_H^{\text{int}}(x) = ig \bar{\Psi}_H(x) \gamma_5 \gamma_\mu \Psi_H(x) \frac{\partial}{\partial x_\mu} \phi(x)$$

$$(\gamma_\mu \partial_\mu + m) \Psi_H(x) = ig \gamma_5 \gamma_\mu \Psi_H(x) \frac{\partial}{\partial x_\mu} \phi(x)$$

$$(\square - \mu^2) \phi(x) = ig \frac{\partial}{\partial x_\mu} (\bar{\Psi}_H(x) \gamma_5 \gamma_\mu \Psi_H(x))$$

$$(\gamma_\mu^T \partial_\mu - m) \bar{\Psi}_H(x) = -ig \bar{\Psi}_H(x) \gamma_5 \gamma_\mu \frac{\partial}{\partial x_\mu} \phi(x)$$

} Equations
 of
 motion

(5)

Canonical Commutation Relations -

$$\pi_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \dot{\phi} + g \bar{\psi} \gamma_5 \gamma_4 \psi$$

$$[\phi(x, t), \pi_\phi(x', t)]_- = i \delta(x - x')$$

$$[\phi(x, t), \phi(x', t)]_- = 0$$

$$[\psi_\alpha(x, t), \psi_\beta(x', t)]_- = \delta_{\alpha\beta} \delta(x - x')$$

$$[\psi_\alpha(x, t), \pi_\phi(x', t)]_- = 0$$

$$[\psi_\alpha(x, t), \phi(x', t)]_- = 0$$

Hamiltonian: $H^{\text{int}}(x) = -\mathcal{L}^{\text{int}}(x) + \frac{1}{2} g^2 (\bar{\psi} \gamma_5 \gamma_4 \psi)^2$

We must verify that if:

$$\psi' = \psi + \delta\psi \quad \bar{\psi}' = \bar{\psi} + \delta\bar{\psi} \quad \phi' = \phi + \delta\phi$$

then $\psi', \bar{\psi}', \phi'$ satisfy the Heisenberg equations of motion with g replaced by $g + \delta g$ (up to first order in δg of course) This is a straight-forward calculation and will not be given here.

Therefore, given that $\delta\psi_H(x, g) = -\frac{i\delta g}{g} \left[\int_{-\infty}^{\infty} \mathcal{L}^{\text{int}}(x) dx, \psi_H(x, g) \right]_-$
 we find set $\int_a^b \mathcal{L}^{\text{int}}(x) dx = I_{ab}^{\text{int}}$

then: $\delta\psi_H(x, \infty; g) = -\frac{i\delta g}{g} \left[I_{-\infty}^{\text{int}}, \psi_H(x, \infty; g) \right]_-$

or $\delta\psi^{\text{out}}(x, \infty; g) = -\frac{i\delta g}{g} \left[I_{-\infty}^{\text{int}}, \psi^{\text{out}}(x, \infty; g) \right]_-$

(6)
 The last equation implies:

$$\delta \Psi^{\text{out}}(x, g) = -i \frac{\delta g}{g} [\mathbb{I}_{-\infty}^{\infty}, \Psi^{\text{out}}(x, g)]_-$$

Proof. Both sides are equal for $t = \infty$ + both sides satisfy $(\gamma_{\mu} \partial_{\mu} + m) \Psi(x) = 0$ Q.E.D.

Now $S^{-1} \Psi_{\text{in}}(x) S = \Psi^{\text{out}}(x, g)$

$$\therefore \frac{\partial}{\partial g} (S \Psi^{\text{out}}(x, g) S^{-1}) = 0$$

$$\text{or } \frac{\partial}{\partial g} \Psi^{\text{out}}(x, g) = [\Psi^{\text{out}}(x, g), S^{-1} \frac{\partial S}{\partial g}] \quad (\text{A})$$

$$\text{but } \frac{\partial}{\partial g} \Psi^{\text{out}}(x, g) = [\Psi^{\text{out}}(x, g), \frac{i}{g} \mathbb{I}]_- \quad \text{proven above. (B)}$$

We now prove that (A) and (B) imply $S^{-1} \frac{\partial S}{\partial g} = \frac{i}{g} \mathbb{I}$

Proof. Define $Q^{-1} \frac{\partial Q}{\partial g} \equiv \frac{i}{g} \mathbb{I} \quad Q = 1 \text{ for } g = 0$

then $Q^{-1} \Psi_{\text{in}}(x) Q \equiv \Psi_Q(x, g)$

$$\text{or } \frac{\partial \Psi_{\text{in}}}{\partial g} = 0 = \frac{\partial}{\partial g} (Q \Psi_Q(x, g) Q^{-1})$$

$$\text{or } \frac{\partial \Psi_Q(x, g)}{\partial g} = [\Psi_Q(x, g), Q^{-1} \frac{\partial Q}{\partial g}] = [\Psi_Q(x, g), \frac{i}{g} \mathbb{I}]$$

hence $\Psi_Q(x, g) = \Psi_{\text{out}}(x, g)$ for $g = 0$

+ both satisfy the same linear differential equation

$\therefore \Psi_Q = \Psi_{\text{out}}$ + hence $Q = S$ since $Q^{-1} \Psi_{\text{in}} Q = \Psi_{\text{out}}$ defines S .

(7)
 we must now prove that $(S, \mathcal{I})_- = 0$

Proof:

We have verified (elsewhere) that:

$$\delta \Psi_{\pm}(x, g) = [\Psi_{\pm}(x, g), \frac{i\delta g}{g} \mathcal{I}_a^{\pm}]_- \quad \text{where } a \text{ is}$$

any given time. We now choose $a = +\infty$
 and set $t = -\infty$ then we get:

$$\delta \Psi_{\pm}(x, -\infty, g) = -[\Psi_{\pm}(x, g), \frac{i\delta g}{g} \mathcal{I}]_- \quad (A')$$

with the boundary condition $\delta \Psi_{\pm} = 0$ for $t = \infty$
 under these kinds of boundary conditions we are
 holding $\Psi_{out}(x)$ fixed & varying $\Psi_{in}(x)$ therefore

we write $S^{-1} \Psi_{in}(x, g) S = \Psi_{out}(x)$

or $\frac{\partial}{\partial g} (S^{-1} \Psi_{in}(x, g) S) = 0$

or $\frac{\partial}{\partial g} \Psi_{in}(x, g) = [\frac{\partial S}{\partial g} S^{-1}, \Psi_{in}(x, g)]_-$

also $\frac{\partial}{\partial g} \Psi_{in}(x, g) = [\frac{i\delta g}{g} \mathcal{I}, \Psi_{in}(x, g)]_-$ from (A')

Again we can identify

$$\frac{i}{g} \mathcal{I} = \frac{\partial S}{\partial g} S^{-1}$$

or $\frac{i}{g} \mathcal{I} S = \frac{\partial S}{\partial g}$

O.E.D.

Relativistic Two-Body Problem
京都大学基礎物理学研究所 湯川記念館史料室

M. Gell-Mann and F. Low, P.R. 84 (1951),
350.

$$\frac{\partial F}{\partial x_\mu} = i[F, P_\mu]$$

$$P_\mu \Psi_n = p_\mu^n \Psi_n$$

$$P_0 = H = H_0(t) + H_I(t)$$

$t=0$: interaction repr. $O(t)$ and Heisenberg representation $O(t)$ coincide, where

$$O(t) = \exp(iH_0(t)t) O(0) \exp(-iH_0(t)t)$$

$$O(t) = \exp(iHt) O(0) \exp(-iHt)$$

or

$$O(t) = U(t, 0) O(0) U^{-1}(t, 0)$$

$$U(t, 0) = \exp(iH_0(t)t) \exp(-iHt)$$

$$i \frac{dU(t, 0)}{dt} = H_I(t) U(t, 0)$$

$$\text{with } U(0, 0) = 1$$

True vacuum state: Ψ_0 , (eigenstate of H with the lowest energy, $p_0^0 = E_0$)
 vacuum of free particles: Φ_0 , (eigenstate of $H_0 = H_0(0)$ with the lowest value ϵ_0)

First Method:

1. Assume that at $t = -cs$, coupling vanishes and
 $H(-cs) = H_0(-cs)$

$$\text{and that } \Psi_0 = \Phi_0$$

2. Second method: interact. repr. is defined at $t=0$ and

$$c\Psi_0 = [1 + (H_0 - \epsilon_0)^{-1} (1 - V) (H_I(0) - E_0 + \epsilon_0)]^{-1} \Phi_0$$

where V is the projection operator on Φ_0 .

$$\begin{aligned} \because (H_0 + H_I) \Psi_0 &= E_0 \Psi_0 \\ \Psi_0 &= a \Phi_0 + \Phi' & V \Psi_0 &= a \Phi_0 \\ & & (1-V) \Psi_0 &= \Phi' \\ H_0 \Phi_0 &= E_0 \Phi_0 \\ a(E_0 + H_I - E_0) \Phi_0 + (H_0 + H_I - E_0) \underbrace{(\Psi_0 - a \Phi_0)}_{(1-V) \Psi_0} &= 0 \end{aligned}$$

or

$$\Phi_0 = a \Psi_0 +$$

$$c \Psi_0 = U^{-1}(\pm\infty, 0) \Phi_0 / (\Phi_0, U^{-1}(\pm\infty, 0) \Phi_0)$$

Define Feynmann's two-body kernel:

$$K(x_1, x_2; x_3, x_4) = \varepsilon \underbrace{(\Psi_0, P[\psi(x_1)\phi(x_2)\bar{\psi}(x_3)\bar{\phi}(x_4)]}_{\Psi_0})}$$

$$\varepsilon = \begin{cases} -1 & \text{perm. is even} \\ +1 & \text{odd} \end{cases}$$

$$t_1 \geq t_2 \geq t_3 \geq t_4:$$

$$\begin{aligned} K(x_1, x_2; x_3, x_4) &= - (\Phi_0, U(+\infty, 0) U^{-1}(t_1, 0) \\ &\quad \psi(x_1) U(t_1, 0) U^{-1}(t_2, 0) \phi(x_2) \dots \\ &\quad \times \bar{\phi}(x_4) U(t_4, 0) U^{-1}(-\infty, 0) \Phi_0) / \\ &\quad (\Phi_0, U(+\infty, 0) U^{-1}(-\infty, 0) \Phi_0) \end{aligned}$$

$$U(t,0)U^{-1}(t',0) = U(t,t')$$

$$= \sum_{n=0}^{\infty} \frac{(-i)^n}{n!} \int_{t'}^t d\tau_1 \dots d\tau_n P[H_I(\tau_1) \dots H_I(\tau_n)]$$

$$K = \epsilon (\Phi_0, P[U(\infty, -\infty) \psi(x_1) \phi(x_2) \bar{\psi}(x_3) \bar{\phi}(x_4)] \Phi_0)$$

$$(\Phi_0, U(\infty, -\infty) \Phi_0)$$

where

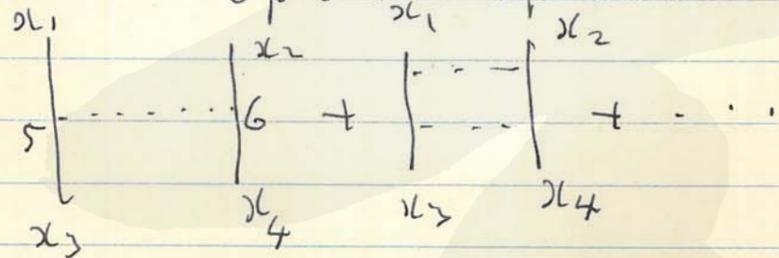
$$P[U(\infty, -\infty) \psi \phi \bar{\psi} \bar{\phi}]$$

$$= \sum_{n=0}^{\infty} \frac{(-i)^n}{n!} \int_{-\infty}^{\infty} d\tau_1 \dots d\tau_n P[H_I(\tau_1) \dots H_I(\tau_n) \times \psi \phi \bar{\psi} \bar{\phi}]$$

$$K(1,2,3,4) = S_F^P(1,3) S_F^N(2,4)$$

$$-g^2 \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} d\omega_5 d\omega_6 S_F^P(1,5) S_F^N(2,6) \Delta_F(5,6)$$

$$\times S_F^P(5,3) S_F^N(6,4) + \dots$$



$$S_F^P(x-y) = \epsilon(x-y) (\Phi_0, P[\psi(x) \bar{\psi}(y)] \Phi_0)$$

$$S_F^N(x-y) = \epsilon(x-y) (\Phi_0, P[\phi(x) \bar{\phi}(y)] \Phi_0)$$

$$\Delta_F(x-y) = (\Phi_0, P[A(x)A(y)] \Phi_0)$$

$$K(1, 2; 3, 4) \cong S_F^P(1, 3) S_F^N(2, 4)$$

$$- g^2 \int_{-\infty}^{\infty} d\omega_5 d\omega_6 S_F^P(1, 5) S_F^N(2, 6) \Delta_F(5, 4)$$

$\times K(5, 6; 3, 4)$

(Bethe-Salpeter)

$$K(1, 2; 3, 4) = \sum_n (\Psi_0, \varepsilon(1, 2) P[\psi(x_1) \phi(x_2)] \Psi_n)$$

$$\times (\Psi_n P[\bar{\psi}(x_3) \bar{\phi}(x_4)] \varepsilon(4, 3) \Psi_0)$$

$$= \sum_n \chi_n(1, 2) \tilde{\chi}_n(3, 4)$$

$$\chi_n = V^{-\frac{1}{2}} \exp[i(p_\mu^n - p_\mu^0) X_\mu] f_n(x)$$

$$\chi_n(1, 2) = g^2 \int d\omega_5 d\omega_6 S$$

$$- \int d\omega_5 d\omega_6 d\omega_7 d\omega_8 S_F^P(1, 5)$$

$$S_F^N(2, 6) G(5, 6; 7, 8) \chi_n(7, 8)$$

C. Gregory, Phys. Rev. 89 (1953)
1199 (March 15)

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1953, Feb.

Fusion Theory of Bosons

GRUBBERG → Fred M. Johnson

new Non-local Field

by Satoru Watanabe

$$(D(x') + mE_5) \Psi(x', x'') (D(x'') - mE_5)$$

$$\approx g^{\alpha\beta} W(x' - x'') O'^{\alpha} \Psi(x', x'') O''^{\beta}$$

$$D(x) = -i \frac{\partial}{\partial x_{\mu}} \cdot E_{\mu}$$

$$O'' = J O' J^{-1}$$

$$\bar{\Psi} = -\Psi^{\dagger} \gamma_5 \quad F_a = i \gamma_5 \delta_a$$

$$E_5 = \gamma_5 \quad E_0 = -J = \gamma_4 \gamma_5$$

$$\Psi_{r5} \rightarrow \Psi_r(x') \Psi_s(x'')$$

$$(D(x') + D(x'') + \kappa E_5) \Psi(x', x'') = 0$$

(W - \Delta_F in ordinary theory P.T.P. 2 (1958), 6(4))

$$\Psi \approx (iU + U^{\dagger} E_{\mu} E_5) + (V E_5 + V^{\dagger} E_{\mu} + \frac{1}{2} i V^{\dagger} E_{\mu} E_{\nu} E_5)$$

$$O' = O'' = E_5$$

$$W = -\gamma_{\mu} \gamma^{\mu} \quad g^{\mu\nu} = 4/4$$

$$\begin{cases} \{ p_{\mu}, \{ p^{\mu}, U \} \} + \kappa^2 U = 0 \\ \{ \gamma_{\mu}, \{ p^{\mu}, U \} \} + (\kappa - 2m) U \end{cases}$$

$$+ \frac{16}{\lambda^4} [x_\mu [x^\mu U]] = 0$$

$$\{p_\mu p^\mu U\} - \{g_{\mu\nu} [p^\nu U]\} = 0$$

$$U^\mu = \kappa \frac{\partial U}{\partial x_\mu}$$

$\frac{\partial U}{\partial x_\mu}$: trial-like

$U \propto \exp(r_\mu r^\mu / \lambda^2)$

$r_\mu \propto k_\mu$

$$\kappa = 2m \pm 4\sqrt{2} / \lambda$$