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On the interaction of elementary particles. IV.

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§ 1. Introduction and Summary

In the previous papers¹⁾, it was shown that the introduction of the neutral Bose particles, which were named N-particles by Bhabha²⁾, in addition to the charged Bose particles, i.e. U-particles, was necessary in order to account for approximate equality of unlike particle and like particle forces.³⁾ In this paper, we want to deal with this problem by assuming the wave equations for N-particles, which have the same form as those for U-particles. We can assume, further, that the interaction between the heavy particle with the former has the same form as that with the latter, except that the factor $a + b \tau_3$ of the form $a + b \tau_3$ appears in place of $\tau_1 + i \tau_2$ or $\tau_1 - i \tau_2$, where a, b or $a + b \tau_3$ are arbitrary complex numbers.⁴⁾ Thus, the forces between like particles as well as the ordinary forces between unlike particles⁵⁾ can be obtained in a manner the ~~similar~~ usual way.⁶⁾ The arbitrary constants a, b can be determined by comparing the forces thus obtained with those employed in ~~the~~ choosing the constants a, b suitably, we arrive at the forces thus obtained can take the ~~same~~ ^{exactly same} form ~~as~~ ^{as obtained} ~~the~~ ^{as} required from the forces assumed in the current theory. Very recently, Kemmer ~~considered~~ ^{has} ~~investigated~~ ^{investigated} the problem in detail as ~~investigated~~ ^{shown} by Kemmer recently.³⁾

As already ~~shown~~ ^{mentioned} in ~~§ 7~~ ^{§ 7}, the forces between ~~thus obtained~~ ^{thus obtained} are not strictly central, so that we have to consider

- a) Yukawa, Proc. Phys.-Math. Soc., **17**, 48, 1935; Yukawa and Sakata, *ibid.* **19**, 1084, 1937; Yukawa, Sakata and Taketani, *ibid.* **20**, 319, 1938. They will be referred to as I, II and III.
- b) ~~Kemmer~~ The authors wish to express their hearty thanks to Dr. Kemmer for sending the manuscript of his paper to be published in Proc. Camb. Phil. Soc.

the coupling between the stationary states and energy body levels of the system become rather complicated. Already, the solution of the deuteron problem can not easily be solved owing to the coupling of, for example, of 2S_1 state with 3D_1 . Moreover the singular on the one hand and the strong singularity of the force of the proportional to $1/r$ small the appearance of at small distance on the other, F_1 and F_2 in the expression of

The theory of the δ -disintegration can be explained as already proposed in I § 4, the δ -disintegration is emitted initially shown that our the we can arrived at the modification of the theory leads to the result, which is essentially the same as that of Konopinski and Uhlenbeck, instead of a combination of Fermi-Thomas of Fermi and δ . But in this case, the assumptions for the interaction between the light particle and the δ -particle in this case takes rather complicated form. It is a pity, however, more elegant simple more satisfactory solution of this problem will be obtained is this the final solution of the problem.

Next, the creation problem of the creation of the δ -particle is important in case of γ by the following process is discussed in detail. A γ -ray of energy larger than $2m_p c^2$ colliding with a neutron (or a proton) in the nucleus can change into a δ -particle with the negative charge (or with the positive charge) and is emitted at the same time, leaving a neutron (or a proton) in the nucleus being transformed into a proton (or a neutron). The cross section of this process is far larger than that of the pair-creation of pair of δ -particles by γ -rays, so that, the hard component of the

3) Fermi

it is possible that

observed on the
cosmic ray consists mainly of the U-particles produced in
the atmosphere ~~for~~ by the soft primary ~~above~~ process from
the soft primary. The absorption of the \pm U-particle
by the inverse process, ~~etc.~~ ^{probability} ~~by emitting~~ is also discussed. (§6,
§7)

Finally, the formal μ In addendum, the equations for the U-field
is written in spinor form in order to make clear that they are
~~the~~ a special case ~~of~~ of the generalized wave equations of
Dirac

§2. Linear Equations for the N-Field.

We introduce the six vector \vec{F}, \vec{G} and the four vector \vec{U}, \vec{U}_0 for the N-field corresponding to \vec{F}, \vec{G} and \vec{U}, \vec{U}_0 for the U-field. The equations for the N-field in the presence of the heavy particle to have the form

$$\frac{1}{c} \frac{\partial \vec{F}}{\partial t} - \text{curl } \vec{G} - \kappa \vec{U} = -4\pi g_1 \vec{M} \quad \text{div } \vec{F} + \kappa \vec{U}_0 = 4\pi g_2 \vec{M}_0^{(1)}$$

$$\frac{1}{c} \frac{\partial \vec{U}}{\partial t} + \text{grad } \vec{U}_0 + \kappa \vec{F} = 4\pi g_2 \vec{I} \quad \text{and } \vec{U} - \kappa \vec{G} = 4\pi g_2 \vec{I}^{(2)}$$

Corresponding to the equations (36), (37) in III for the U-field. The constants κ, g_1, g_2 are ~~not necessarily assumed to be same as in the case previous case, although it is not more general~~ ^{considered for simplicity} ~~assumptions are not excluded, though they are not more general~~ ^{each of} ~~assumptions are not excluded.~~ ^{each of} The four vector \vec{M}, \vec{M}_0 and the six vector \vec{F}, \vec{G} on the right hand sides of (1) and (2) involve each the operator which causes the transition of the heavy particle between two neutron states or between two proton states, the charge being conserved throughout.

The simplest possible ^{expressions} forms for them are

$$\vec{M} = \vec{\Psi} \vec{\alpha} (a + b \tau_3) \Psi, \quad \vec{M}_0 = \vec{\Psi} (a + b \tau_3) \Psi \quad (3)$$

$$\text{and } \vec{I} = -\vec{\Psi} \rho_2 \vec{\beta} (a + b \tau_3) \Psi, \quad \vec{I}_0 = \vec{\Psi} \rho_3 \vec{\beta} (a + b \tau_3) \Psi, \quad (4)$$

corresponding to (38) and (39) in III, where a, b are complex constants of the order of 1.

From these the equations (1) and (2) the quantization of this neutral field can be performed in a manner similar to that of the U-field. ~~Also the interaction of two heavy particles 1 and 2 by virtual emission of the N-particles can be deduced determined by absorption accordingly, and takes the following form.~~ ^{by virtual}

$$\mathcal{H}_{12} = [2(|a|^2 + |b|^2) \tau_3^{(1)} \tau_3^{(2)} + (\bar{a}b + a\bar{b})(\tau_3^{(1)} + \tau_3^{(2)})] \times [g_1^2 + g_2^2 \tau_3^{(1)} \tau_3^{(2)} - \frac{(\partial^{(1)} \text{grad})(\partial^{(2)} \text{grad})}{x^2}] \frac{e^{-\kappa r}}{r} \quad (5)$$

In order that the forces \mathcal{H}_{12} involves the forces between two neutrons and two protons as well as those of Wigner and Bartlett types between the neutron and the proton. In order that the forces are symmetric with respect to the neutron and the proton, the condition

should be fulfilled, $\vec{a} \cdot \vec{b} + a \vec{b} = 0$ which implies, $a = 0$ or $\vec{b} = 0$ or $\vec{a} = -\frac{b}{a} \vec{a}$ which implies $\frac{b}{a} = c i$

where c is a real constant. According to recent results, Kemmer suggested that the case adopted the case $a = 0$, so that we obtained the force the resultant

$$\mathcal{H}_{12} = \frac{1}{2} (\tau_3^{(1)} \tau_3^{(2)}) \{ g_1^2 + g_2^2 \} \frac{z_{11} z_{12}}{r^2} (0^{(1)} \text{ grad}) (0^{(2)} \text{ grad}) \frac{e^{-\kappa r}}{r}$$

$\kappa = 1$ seems to be the correct one and we can adopt the case $a = 0$ $b = \frac{1}{2}$, so that we obtain

$$\mathcal{H}_{12} = \frac{(\tau_3^{(1)} \tau_3^{(2)})}{2} \left[\frac{e^{-\kappa r}}{r} \right] \quad (6)$$

Thus, the resultant force between two heavy particles due to the Λ -particles and that of the N -particles becomes $\frac{e^{-\kappa r}}{r}$ (7)

$$V_{12} + \mathcal{H}_{12} = \frac{(\tau_3^{(1)} \tau_3^{(2)})}{2} \left[\frac{e^{-\kappa r}}{r} + \frac{g_1^2 + g_2^2}{2} \frac{z_{11} z_{12}}{r^2} (0^{(1)} \text{ grad}) (0^{(2)} \text{ grad}) \frac{e^{-\kappa r}}{r} \right]$$

In the S -state, the (7) takes the approximate

$$H_{12} + \mathcal{H}_{12} = \frac{(\tau_3^{(1)} \tau_3^{(2)})}{2} \left\{ g_1^2 + \frac{2g_2^2}{3} (z_{11} z_{12}) \right\} \frac{e^{-\kappa r}}{r} \quad (8)$$

in the first approx. (8) takes the form $\frac{e^{-\kappa r}}{r}$. This result is in fair agreement with that obtained by Kemmer phenomenologically.

(6) The authors wish to express their hearty thanks to Dr. Kemmer for sending the manuscript of their paper before publication.

①

If we take

$$\left. \begin{aligned} g_1^2 &= \frac{5}{24} g^2 \frac{g^2}{8} \\ g_2^2 &= \frac{5}{16} g^2 \end{aligned} \right\} (6)$$

18) becomes

$$H_{12} + \mathcal{K}_{12} = \frac{\tau_1^{(1)} \tau_1^{(2)} + \tau_2^{(1)} \tau_2^{(2)} + \tau_3^{(1)} \tau_3^{(2)}}{2} \left\{ \frac{1}{8} + \frac{5}{24} (\tau_1^{(1)} \tau_1^{(2)}) \right\} g^2 \frac{e^{-x r}}{r} \quad (7)$$

which has the same form as the exact expression obtained by Kemmer phenomenologically⁷⁾. This seems very satisfactory. It should be noticed, however, that the expression thus, whether such an agreement between the results obtained reached by two different methods, is obtained, agrees well, to be very satisfactory, but, owing to various approximations made in both calculations we can say that not take it too seriously. ^{each pair of the states with the same j_z combination with each other}

In the former case, in which both sets of states have the same j_z component of the inner quantum number $j=1$. ~~It is~~ corresponding to the inner quantum number $j=1$ is 1 for both both sets of states, which and the ~~z-component~~ ^{magnetic quantum number} j_z of the total. The ~~z-component~~ ^{magnetic quantum number} j_z of the total angular momentum (13) can take the one of the values $-1, 0$ and $+1$ in units of $\frac{\hbar}{2}$. The eigenfunctions for $j_z = -1$ to the 3S_1 states with $j_z = -1$ can be written in the forms ~~with 3D_1~~ ^(for example)

$$u(r) \frac{X_{-1}}{r} X_{+1}(s_1, s_2),$$

where $X_{-1} = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$ ~~state with $j_z = -1$ and~~ $X_{+1} = \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}$ while that for 3D_1

7) Kemmer, Nature 140, 192, 1937. See further Heisenberg, Naturwiss. 25, 749, 1937; Flügel, Zeits. f. Phys. 108, 545, 1938.

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of Fermi type
which lead to energy distribution for β rays
as shown in the previous section.

spontaneous

§6. Annihilation of Heavy Quanta in Free Space.

In §8, III, the probability of annihilation of a heavy quantum in vacuum with the positive (or negative) charge by emitting a positive (or negative) electron and a neutrino (or an anti neutrino) was calculated on the assumption that the interaction between the heavy quantum and the light particle could be given by the simplest expressions such as (25) and (26) § in §4 of this paper. Thus, we found that the mean life time of the heavy quantum at rest was about 0.5×10^{-6} sec, which agrees qualitatively with the recent result of Euler¹⁸⁾ determined from the analysis of the cosmic ray¹⁸⁾ obtained by Euler¹⁸⁾ as the result of

It is a pity, however, that there was an error of a factor 2 in our calculation and the correct value was to be 2×10^{-6} sec. 0.25×10^{-6} sec instead of 0.5×10^{-6} sec.¹⁹⁾ This makes the agreement between theory and experiment a little worse,

~~although~~ In the previous section, it was shown that the

~~Now we want to~~ It is important, however, to calculate
~~Sakata and Tanikawa~~
~~We can easily extend the above calculation to more general~~

18) Euler, Naturwiss.

~~19) Prof. Heisenberg called informed us of the result of Euler.~~

The present authors wish to express their thanks to Prof. W. Heisenberg, who was kind informed them of the result of Euler.

19) The energy E in (62), §8, III, should be $m_0 c^2$ instead of $m_0 c$, so that (63), §8, III becomes $-dW_0 = \frac{m_0 c}{16 \pi^2 \epsilon_0^2} \sum |M|^2 d\Omega$. Thus, w_0 and w_0^2

~~In III should be multiplied by 2, so that~~
 ~~E in III should be divided by 2.~~

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as given in §4 of this paper,
 by assuming more general ~~or~~ form for the interaction
 between the heavy quantum and the light particle as made
 in §4 of this paper. By taking $\lambda_3 = \mu_3 = 0$, they obtained the
 probability for the annihilation of the heavy quantum with
 the ~~$w = w_0 \sqrt{1 - \beta^2}$ velocity~~ energy E

$$W = \frac{g'^2}{hc} \frac{m_0 c^2}{h} \frac{m_0 c^2}{E} \left\{ \frac{2}{3} |\lambda_1 + i \frac{1-\mu}{2} \lambda_2|^2 + \frac{1}{3} |\mu_1 + i \frac{1-\mu}{2} \mu_2|^2 \right\}$$

~~the value~~ g' is the same order of m (7D)
~~this value~~ Since $\lambda_1, \lambda_2, \mu_1, \mu_2$ are quantities of the order
 of 1 and g' is the same order of magnitude of with g_1, g_2
 in §8, III, the ~~probabilities~~ w () does not
 differ essentially from that the previous one. Thus,
 the qualitative agreement. The mean life time $\tau = \frac{1}{w}$ ~~for this~~
~~case~~ and is still in qualitative agreement with the
 experiment. ~~Thus~~

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