

YHAL E18010P02

23

三  
數物報  
拾月

數物記子稿

校正 逸先

兵庫縣西宮市花樂園  
湯川秀樹

所安別刷

著名宛百五十部

阪大理子部予稿室宛  
八百部

内部交換用紙

陰陽電子發生の理論

3

① 物 報 告  
拾 頁

24

On the Theory of Internal Pair Production.<sup>1)</sup>

By Hideki Yukawa and Shoichi Sakata

Abstract

The probability of internal pair production by radiationless transition of the radioactive nucleus was calculated and its ratio to the probability of emission of electrons from K levels by the same nuclear transition was compared with the experiment in the case of RaC'.

---

1) This paper was read before the meeting of Osaka branch on July 6, 1935.

25  
數物報  
拾月

On the Theory of Internal Pair Production.<sup>1)</sup>

By Hideki Yukawa and Shoichi Sakata

§ 1. Introduction and Summary

The theory of the production of the pair of an electron and a positron by the internal conversion of the  $\gamma$ -ray with the energy larger than  $2mc^2$ , was developed recently by Jaeger and Hulme<sup>2)</sup>, as a natural extension of the theory of ordinary internal conversion of the  $\gamma$ -ray with the emission of the  $\beta$ -ray.

They compared their result with the experiment of Alichanow and Kosodaew<sup>3)</sup>, who measured the number of positrons emitted from a Ra(B+C) source. According to the theory the second and the third maxima of the experimental distribution curve,<sup>4)</sup> at about  $0.7 \times 10^6$  eV and  $10^6$  eV, were to be due to the  $\gamma$ -rays of energies  $1.7 \times 10^6$  eV and  $2.2 \times 10^6$  eV, <sup>respectively</sup> as the coefficient of internal pair production ~~was~~ was maximum at the upper limit of the energy of the positron.

The exceptional case, where a pair of an electron and a ~~posi~~ positron is emitted <sup>with the absorption of</sup> ~~by taking up~~ the energy liberated by a nuclear transition between two S levels, was not <sup>dealt with</sup> ~~treated~~ by the former authors, because of lack of the corresponding  $\gamma$ -ray.

1) This paper was read before the meeting of Osaka branch on July 6, 1935.

2) J.C.Jaeger & H.R.Hulme, Proc.Roy.Soc. **148**, 708, 1935.

3) A.I.Alichanow & M.S.Kosodaew, Zeits.f.Phys. **90**, 249, 1934.

4) See Fig. 1 below, which is the same with Fig. 15 of Alichanow & Kosodaew, loc.cit.

3  
26 数物報告  
拾

~~Now~~ <sup>1</sup> If the maximum of the distribution curve of the positron produced by such a process lies also at the upper limit of the energy, the first hump of the curve of Fig. 1, at about  $0.4 \times 10^6$  eV, can be attributed to the missing  $\gamma$ -ray of energy  $1.4 \times 10^6$  eV.

Hence it seems worth while to investigate this case in detail and to ~~determine~~ <sup>see</sup> whether such a process is sufficient to account for the large hump of the curve. ~~of~~

The nuclear transition between two levels of energy difference  $\Delta W$  can be considered in general as a perturbing field on the electron varying with the frequency  $\frac{\Delta W}{\hbar}$ . This perturbing field extends only over the region of nuclear dimension in the case of S-S transition, so that the probability of transition of the electron is large only when the eigenfunction of the initial and the final states are not small in the neighborhood of the nucleus. Accordingly, we have only to consider the case, where the total angular momenta of both of the electron and the positron take the smallest values.

The distribution function of the positron calculated in this manner has the maximum in the neighborhood of the upper limit of the energy, which agrees with the first maximum of the curve of Fig. 1.

To find the absolute value of the probability of the pair production, we must have the detailed information of ~~which is too difficult as yet.~~

~~we have to know the nuclear structure in detail, to avoid this~~  
<sup>For this reason</sup> we considered only the ratio of the probability of the pair production to that of <sup>the</sup> emission of  $K$  electrons caused by the same nuclear transition, as they have the common factors <sup>concerning</sup> relating to the nuclear transition, if the vector potential of the perturbing field

随核の

(4)

27 數物報管  
拾頁

is negligible compared with the scalar potential. The result obtained shows that the ratio is so small that only a small part of the hump in the experimental curve can be attributed to the positrons emitted by the above process.

If the vector potential is also taken into account, the ratio depends on the special forms of scalar and vector potentials and in the favorable case it becomes large enough to be in accord with the experiment. *It should be noticed however, that we can not, however, have much confidence in our*

~~results anyhow,~~ the application of the present theory of the electron up to the immediate neighborhood of the nucleus may happen to be totally wrong. *It remains still uncertain, or not, therefore, hence we are not certain whether other processes such as the emission of the positron alone from the nucleus or the internal conversion of the  $\beta$ -ray may be considered* explanation of the experimental result, ~~or not.~~

A brief account of the calculation will be given in the following sections.

§ 2. Probability of Internal Pair Production

~~Now we~~ consider the general case, where an electron moving in the field of the nucleus with the charge  $Ze$  is perturbed by a nuclear transition between two levels of energy difference  $\Delta W$ . *no*

The perturbation can be expressed by scalar and vector potentials of the form  $A_0(\vec{r})\exp(-2\pi i\nu t)$  and  $\vec{A}(\vec{r})\exp(-2\pi i\nu t)$ , where  $\nu = \frac{\Delta W}{h}$ . If we assume ~~that~~ only one particle with the effective charge  $Qe$  and the reduced mass  $M$  in the nucleus ~~fall~~ <sup>to</sup> from a state  $\varphi(\vec{r})$  to  $\chi$  a



(4)  
(5)

28 數 物 報 告 拾 月

state  $\varphi_0(\vec{r})$ , the energy difference being  $\Delta W$ , the potentials can be expressed in general by

$$A_0(\vec{r}) = qe \int \frac{\exp(iq|\vec{r}-\vec{r}_1|)}{|\vec{r}-\vec{r}_1|} \varphi_0^*(\vec{r}_1) \varphi(\vec{r}_1) d\vec{r}_1, \quad (1)$$

$$\vec{A}(\vec{r}) = \frac{qeh}{4\pi mci} \int \frac{\exp(iq|\vec{r}-\vec{r}_1|)}{|\vec{r}-\vec{r}_1|} \{ \varphi_0^*(\vec{r}_1) \text{grad}_1 \varphi(\vec{r}_1) - \varphi(\vec{r}_1) \text{grad}_1 \varphi_0^*(\vec{r}_1) \} d\vec{r}_1, \quad (2)$$

where  $q = \frac{2\pi\nu}{c} = \frac{2\pi\Delta W}{hc}$ . The symbol  $\rightarrow$  means always a vector quantity and \* a conjugate complex quantity.

Now, according to the perturbation theory, the probability per unit time per unit energy range of the transition of the electron from the continuous state of negative energy  $-E_+ (E_+ > mc^2)$  to the continuous state of positive energy  $E_-( < mc^2)$ , satisfying the relation  $E_+ + E_- = \Delta W$ , is given by

$$p(E_+) = \frac{4\pi^2}{h} \sum_{j,u} \sum_{j,u} |\int \tilde{\Psi}_{E_-,j,u}(\vec{r}) \{ eA_0(\vec{r}) + e\vec{A}(\vec{r}) \} \times \Psi_{-E_+,j,u}(\vec{r}) d\vec{r}|^2, \quad (3)$$

where  $\Psi_{E,j,u}$  denotes the solution of Dirac's equation

$$(E + \frac{Ze^2}{r} + c\vec{\alpha}\vec{p} + \beta mc^2) \Psi = 0$$

with energy  $E$ , inner quantum number  $|j| - \frac{1}{2}$ , and its z-component  ~~$u$~~   $u + \frac{1}{2}$ .  $j$  is a positive or a negative integer excluding zero and  $u$  takes a value between  $-|j|$  and  $|j|-1$ .  $\vec{\alpha}, \beta$  are Dirac's matrices with four rows and four columns and  $\Psi$  can be considered as a matrix with one row and four columns, while  $\tilde{\Psi}$  as that with four rows and one column, the corresponding elements being complex conjugate to each other.

5) H.M. Taylor & N.F. Mott, Proc. Roy. Soc. **138**, 665, 1932.

(6)

The eigenfunctions are normalized with respect to energy, the definition of which is

$$\int \tilde{\Psi}_E \Psi_E d\vec{r} = \delta(E' - E).$$

This corresponds to the process, in which a positron of energy  $E$  and an electron of energy  $E$  are emitted simultaneously from the nucleus.

Hence, the total probability per unit time of the production of the pair is given by  $\frac{\Delta W}{mc^2}$

$$P_{\text{pair}} = \int_{mc^2}^{\infty} p(E_+) dE_+ \quad (4)$$

In the case of S-S transition,  $\varphi_0$  and  $\varphi$  are functions of  $r$  only, so that the scalar potential becomes a function of  $r$  and the vector potential  $\vec{A}$  reduced to the form  $\frac{\vec{r}}{r}$  multiplied by a function of  $r$ . Thus

the perturbing potential  $eA_0 + e\vec{r}\vec{A}$  can be written in a form

$$V_0(r) + \frac{\vec{\alpha}\vec{r}}{r} V(r)$$

which commutes with  $j$  and  $u$ .

Hence we obtain the selection rules

$$\Delta j = 0, \quad \Delta u = 0$$

for the transition of the electron.

In this case, moreover,  $A_0$  and  $\vec{A}$  can be reduced to zero outside of the nucleus by a suitable gauge transformation, which does not affect the transition probability in general. Consequently, the probability of the transition of the electron is appreciable only when the eigenfunctions of the initial and the final states are not small in the

①

neighborhood of  $r=0$ , i.e. only when  $j = \mp 1$ .

Thus, the expression (3) becomes approximately

$$p(E_+) = \frac{4\pi^2}{h} \sum_{j=\mp 1} \sum_{u=0, -1} |\tilde{\Psi}_{E_+, j, u}(\vec{r})| \int \tilde{\Psi}_{E_+, j, u}(\vec{r}) \{V_0(r) + \frac{\partial^2}{\partial r^2} V(r)\} \\ \times \Psi_{E_+, j, u}(\vec{r}) d\vec{r} \quad (5)$$

The eigenfunctions in this expression are well known.<sup>6)</sup> Since only

the values of the eigenfunctions in the neighborhood of  $r=0$  are

important, their radial parts are expanded in powers of  $r$  and the first

terms with the power  $r^{\delta-1}$  are taken as the first approximation, where

$\delta = \sqrt{1 - \alpha^2 Z^2}$ . Inserting these expressions for the eigenfunction in (5),  $\alpha$  being the fine structure constant, and performing the integrations with respect to  $\theta, \phi$ , we obtain at length

$$p(E_+) = \frac{4m}{\{\Gamma(2\delta+1)\}^2 h^3} \left(\frac{4\pi mc}{h}\right)^{4\delta} \{ (1 - \alpha^2 Z^2) R_p^2 + \alpha^2 Z^2 S_p^2 \} \\ \times F_p(E_+) \quad (6)$$

where

$$R_p = \int r^{2(\delta-1)} V_0(r) dr,$$

$$S_p = \int r^{2(\delta-1)} V(r) dr,$$

$F_p(E_+)$  is the only factor, which depends on the energy of the positron emitted, and has the form

$$F_p(E_+) = \eta_+ e^{-\pi b_+} |\Gamma(\delta + i b_+)|^2 \\ \times \eta_- e^{-\pi b_-} |\Gamma(\delta - i b_-)|^2 (\varepsilon_+ \varepsilon_- - \gamma^2), \quad (7)$$

where

$$\varepsilon_{\pm} = \frac{E_{\pm}}{mc^2}, \quad \eta_{\pm} = \sqrt{\varepsilon_{\pm}^2 - 1}$$

$$b_{\pm} = \alpha Z \frac{\varepsilon_{\pm}}{\eta_{\pm}}.$$

6) Jaeger & Hulme, loc.cit. We have to take  $k=0$  in (2A) for  $j=-1$

and  $k=1$  in (2B) for  $j=1$ ,  $u$  being the same as in our case.

The forms of the distribution function  $F_p(\epsilon_+)$  in special cases are shown in Fig. 2 and Fig. 3.

The total probability of pair production is now becomes

$$P_{\text{pair}} = \frac{4m}{\{\Gamma(2\delta+1)\}^2 h^3} \left( \frac{4\pi mc}{h} \right)^{4\delta} \int_{\Delta W-1}^1 F_p(\epsilon_+) d\epsilon_+ \quad (8)$$

where  $\Delta W = \frac{\Delta W}{mc^2}$ .  
 In these formulae,  $F_p$  and  $G_p$  depend on the detailed structure of the nucleus,  $R_p$  and  $S_p$  can not be estimated easily.

If we neglect the vector potential, the calculation ~~is~~ <sup>becomes</sup> formally similar to that of Beck and Sittler on  $\beta$ -disintegration. ~~The~~ physical interpretations, however, differ ~~in two cases~~. In the latter case, either the electron or the positron is reabsorbed by the original nucleus, whereas, in ~~the former~~ <sup>the former</sup> case, both of them escape <sup>from</sup> the nucleus.

§ 3. Probability of Electron Emission from K Levels

Next we consider the emission of  $K$  electrons by the same nuclear transitions.

In general, the probability per unit time of the transition of the electrons from  $K$  states ~~of~~ of energy  $mc^2 \sqrt{1-\alpha^2 Z^2}$  to continuous states of energy  $E' (> mc^2)$ , satisfying the condition  $E' = mc^2 \sqrt{1-\alpha^2 Z^2} + \Delta W$ , is given by

$$P_K = \frac{4\pi^2}{h} \sum_{j,u} \sum_{u=0,-1} | \int \tilde{\Psi}_{E',j,u}(\vec{r}) \{ e A_0(\vec{r}) + e \vec{\alpha} \vec{A}(\vec{r}) \} \times \Psi_{K,u}(\vec{r}) d\vec{r} |^2, \quad (9)$$

7) G. Beck und K. Sittler, Zeits.f.Phys. **86**, 105, 1933.

⑨

the meaning of  $\Delta W$ ,  $A_0(\vec{r})$ ,  $\vec{A}(\vec{r})$  and  $\Psi_{E',j,u}$  being the same as in the previous section.  $\Psi_{K,u}$  normalized eigenfunctions of two K states.  $\Psi_{K,u=0,-1}$  are

Now we consider the same S-S transition of the nucleus as in the previous section. The expression (9) becomes simply

$$P_K = \frac{4\pi^2}{R} \sum_{u=0,-1} |\int \tilde{\Psi}_{E',-1,u}(\vec{r}) \{ V_0(r) + \frac{\partial \vec{E}}{\partial r} V(r) \} \Psi_{K,u}(\vec{r}) d\vec{r}|^2 \quad (10)$$

By taking for  $\Psi_{K,u}$  and  $\tilde{\Psi}_{E',-1,u}$  the first terms of the expansion in powers of  $r$ , we ~~obtain~~ finally

$$P_K = \frac{4\pi m (\alpha Z)^{2\delta+1}}{\{\Gamma(2\delta+1)\}^3 R^3} \left( \frac{4\pi mc}{h} \right)^{4\delta} R_p^2 \eta'^{2\delta-1} e^{\pi b'} |\Gamma(\delta + i b')|^2 \times (\epsilon' + \delta) \quad (11)$$

where  $\epsilon' = \frac{E'}{mc^2}$ ,  $\eta' = \sqrt{\epsilon'^2 - 1}$ ,  $b' = \alpha Z \frac{\epsilon'}{\eta'}$ .

In this case, the probability does not depend on the magnitude of the vector potential. ~~By comparing (11) with (8),~~ the ratio of the probability of pair production by the S-S transition of the nucleus to that of the emission of electrons from K levels by the same transition is given by

$$\rho = \frac{P_{pair}}{P_K} = \frac{(1-\alpha^2 Z^2) \int F_p(\epsilon_+) d\epsilon_+ \cdot \left\{ 1 + \frac{\alpha^2 Z^2}{1-\alpha^2 Z^2} \left( \frac{S_p}{R_p} \right)^2 \right\}}{\pi (\alpha Z)^{2\delta+1} \Gamma(2\delta+1) \eta'^{2\delta-1} |\Gamma(\delta + i b')|^2 e^{\pi b'} (\epsilon' + \delta)} \quad (12)$$

8) See, for example, H.R.Hulme, Proc.Roy.Soc. **138**, 643, 1932.

§ 4. General Discussions

Before entering into numerical calculations in special cases, it should be noticed that the above formulation is incomplete in two points.

The first point is that the perturbing potentials  $V'$  were assumed to have the special forms (1) and (2). Now, we can show more generally that the perturbing potentials are expressed in the forms

$$F(r)\exp(-2\pi i\nu t),$$

$$\frac{\partial V}{\partial t} G(r)\exp(-2\pi i\nu t),$$

where  $F$  and  $G$  are  $r$  being certain functions of  $r$  only, whenever both the initial and the final states of the nucleus have zero spin or, more generally, <sup>still</sup> have the same spin in the same direction.

To prove this, we consider the fact that the total angular momentum of the system consisting of the nucleus and the electron should be constant throughout the above processes. So, if the spin of the nucleus does not change, the total angular momentum of the electron

$$\vec{M} = \vec{m} + \frac{\hbar}{4\pi} \vec{\sigma}$$

should be constant <sup>become</sup> ~~also~~, where  $\vec{m} = \vec{r} \times \vec{p}$  and  $\vec{\sigma}$  is the spin vector. Hence, the perturbed Hamiltonian of the electron should commute with  $\vec{M}$ , so that we have

$$\vec{M}(eA_0 + e\vec{\sigma} \cdot \vec{A}) - (eA_0 + e\vec{\sigma} \cdot \vec{A})\vec{M} = 0,$$

where  $A_0(\vec{r})\exp(-2\pi i\nu t)$ ,  $\vec{A}(\vec{r})\exp(-2\pi i\nu t)$  are perturbing potentials due to the nuclear transition. As the coefficients of 1 (the unit matrix),  $\alpha_x = \rho_1 \sigma_x$ ,  $\alpha_y = \rho_1 \sigma_y$ ,  $\alpha_z = \rho_1 \sigma_z$  in the left hand sides of these equations should vanish separately, we have



$$\vec{m} A_0 - A_0 \vec{m} = 0 \quad (13)$$

$$\left. \begin{aligned} m_x A_x - A_x m_x &= 0 \\ m_y A_x - A_x m_y &= -\frac{i\hbar}{2\pi} A_z \\ m_z A_x - A_x m_z &= \frac{i\hbar}{2\pi} A_y \end{aligned} \right\} \quad (14)$$

and similar equations for  $A_y, A_z$ .

The equations (13) show that  $A_0$  is a function of  $r$  only, while the equations show that the relations of commutation between  $\vec{m}$  and  $\vec{A}$  are the same as those between  $\vec{m}$  and  $\vec{r}$ .

Now, by using (14) and the relations of commutation between  $\vec{m}$  and  $\vec{r}$ , we have

$$m_y (x A_y - y A_x) - (x A_y - y A_x) m_y = \frac{i\hbar}{2\pi} (y A_z - z A_y) \quad (15)$$

On the other hand, by using (14) and the relations of commutation between  $\vec{p}$  and  $\vec{r}$ , we obtain

$$p_x (x A_y - y A_x) - (x A_y - y A_x) p_x = x (p_x A_y - A_y p_x) - \frac{i\hbar}{2\pi} A_y - (y p_x A_x - A_x y p_x). \quad (16)$$

From the third equation

$$(x p_y - y p_x) A_x - A_x (x p_y - y p_x) = \frac{i\hbar}{2\pi} A_y$$

of (14), the last two terms become  $-x p_y A_x + A_x x p_y$ , so that we have

$$p_x (x A_y - y A_x) - (x A_y - y A_x) p_x = x (p_x A_y - A_y p_x - p_y A_x + A_x p_y). \quad (17)$$

Further, by using the relations

$$(z p_x - x p_z) A_y - A_y (z p_x - x p_z) = 0, \text{ and } (y p_z - z p_y) A_x - A_x (y p_z - z p_y) = 0,$$

we can deduce the equation as follows:

$$p_z (x A_y - y A_x) - (x A_y - y A_x) p_z = z (p_x A_y - A_y p_x - p_y A_x + A_x p_y). \quad (18)$$

(12)

Combining (17) and (18), we have

$$(z p_x - x p_z)(x A_y - y A_x) - (x A_y - y A_x)(z p_x - x p_z) = 0.$$

Comparing this with (15), we ~~have~~ <sup>obtain</sup> finally

$$y A_x - z A_y = 0.$$

Similarly, we can further deduce the relations

$$z A_x - x A_z = 0, \quad x A_y - y A_x = 0,$$

and also

$$\vec{\nabla}(\vec{r} \cdot \vec{A}) - (\vec{\nabla} \cdot \vec{A})\vec{r} = 0.$$

These equations together show <sup>clearly</sup> that  $\vec{A}$  has the form  $\vec{\nabla} G(\lambda)$ .

Thus the proof is ended.

We can show, moreover, that the perturbing field is zero outside of the nucleus in this general case. Firstly, the magnetic field  $\vec{H}$  vanishes everywhere, since  $\vec{H} = \text{curl } \vec{A}$

$$\vec{H} = \text{curl } \vec{A} = \text{curl} \exp(-2\pi i \vec{r} \cdot \vec{t}), = 0.$$

Secondly, from the field equations

$$\frac{1}{c} \frac{\partial \vec{E}}{\partial t} = \text{curl } \vec{H} - \frac{4\pi}{c} \vec{I},$$

in which  $\vec{H} = 0$  and both the electric field  $\vec{E}$  and the current density  $\vec{I}$  have the time factor  $\exp(-2\pi i \vec{r} \cdot \vec{t})$ , we ~~have~~ <sup>obtain</sup> the relation

$$\vec{E} = \left(\frac{2}{c}\right) \vec{I},$$

~~we~~ <sup>ing</sup> show that the electric field vanishes outside of the nucleus, where the current density due to the nuclear transition vanishes.

Hence, in this case, we can reduce the potentials themselves to zero outside of the nucleus by a suitable gauge transformation.



(13)

The second point is that the solutions of Dirac's equation in Coulomb field of the nucleus were taken as the eigenfunctions of the unperturbed states of the electron. <sup>since</sup> the field is ~~no more~~ <sup>any more</sup> of Coulomb type in the neighborhood of the nucleus, we should modify the forms of the eigenfunctions in this region. Thus, each eigenfunction in the previous sections should be multiplied by a certain function, which deviates appreciably from 1 for ~~the~~ value of  $r$  comparable with the nuclear radius  $a_N$ .

As the exact determination of such functions is very complicated, ~~and~~ <sup>ing</sup> necessitate special assumptions on the nuclear structure, we should be contented ~~satisfied~~ with taking for the values of the eigenfunctions in the nucleus simply the values of the solutions in Coulomb field at  $r = a_N$ , while their values outside of the nucleus do not ~~concern us here~~ <sup>matter, as far as the results of this problem.</sup>

Then, in the formulae of the preceding sections, we have only to change the definitions of  $R_p, S_p$ , other points being unaltered. Namely,

$$R_p = a_N^{2(\delta-1)} \int V_0 d\vec{r},$$
$$S_p = a_N^{2(\delta-1)} \int V d\vec{r},$$

(our calculations are concerned.)

instead of the corresponding expressions in § 2.

(14)

§ 5. Numerical Results ~~and discussions~~

We want now to obtain the numerical values of  $F_p(\epsilon_+)$  and  $\rho$  in the case of RaC. <sup>By taking</sup> ~~Inserting~~ the values

$Z=84, \alpha = \frac{1}{13\eta}, \Delta W = 1.4 \times 10^6 \text{ eV} = 2.8mc^2,$

we ~~have~~ obtain

$\alpha Z = 0.6, \gamma = 0.8, \frac{\gamma^2}{\pi(\alpha Z)^{2\delta+1}\Gamma(2\delta+1)} = 0.5, \epsilon_+ + \epsilon_- = 2.8$

$\epsilon' = 3.6, \eta' = 3.5,$

$|\Gamma(\gamma+ib')|^2 e^{\pi b'} (\epsilon'+\gamma) \cong (4.5+1.6\eta') e^{-1.2\pi\frac{\epsilon'}{\eta'}} (\epsilon'+\gamma) = 444,$

so that

$F_p(\epsilon_+) \cong (4.5+1.6\eta_-)(4.5+1.6\eta_+) (\epsilon_-\epsilon_+-0.6) e^{-1.2\pi\frac{\epsilon_+}{\eta_+}}$

and

$\rho = 1.1 \times 10^{-2} \cdot \int_1^{1.8} F_p(\epsilon_+) d\epsilon_+ \cdot \left\{ 1 + 0.56 \left( \frac{S_p}{R_p} \right)^2 \right\}.$

The distribution function  $F_p(\epsilon_+)$  in this case is given in Fig. 2. Its maximum ~~is~~ lies near the upper limit of the energy of the positron, ~~and~~ coincides <sup>very</sup> approximately with the first maximum of Alichanow's curve, as shown in Fig. 1, which represents the energy distribution of positrons emitted from a ~~Ra~~ Ra(B+C) source.<sup>10)</sup>

From Fig. 2, we ~~have~~ obtain

$\int_1^{1.8} F_p(\epsilon_+) d\epsilon_+ \cong 0.18,$   
 $\rho \cong 2 \times 10^{-3} \left\{ 1 + 0.56 \left( \frac{S_p}{R_p} \right)^2 \right\}.$  (19)

so that

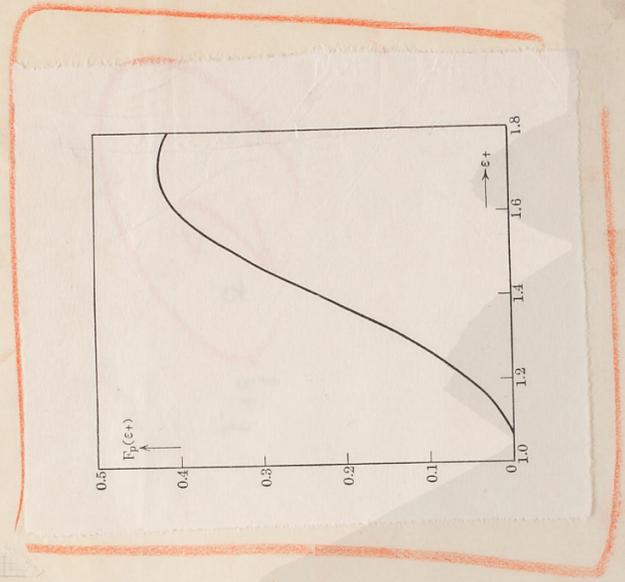
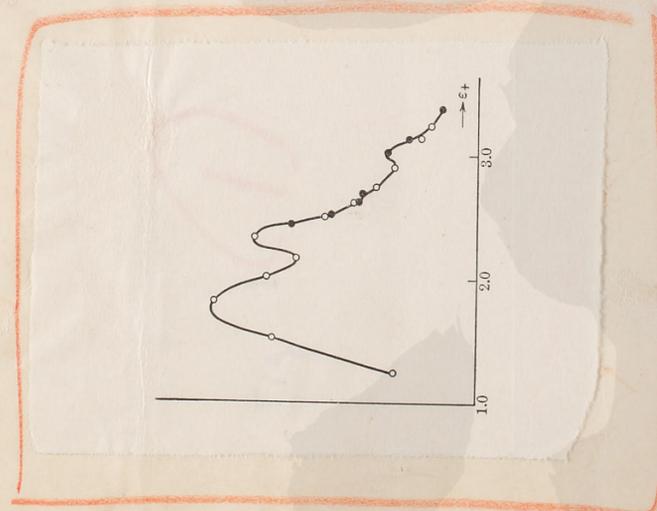
10) The energy is written in unit of  $mc^2 = 0.5 \times 10^6 \text{ eV}$ , the proper energy 1 being inclusive.

94) E. Fermi, Zeits. f. Phys. 88, 161, 1934.



38 拾物

(15) (10)



~~Fig. 1~~ The number of <sup>electrons</sup> the electron emitted from the K levels is known to be about 0.003 per disintegration in this case, ~~while~~ <sup>while</sup> on the other hand, according to Alichanow, total number of ~~the~~ positrons is also about 0.003 per disintegration.

Hence, if we assume <sup>all of</sup> the first hump in Fig. 1 to be due to the process (above considered),  $\rho$  should be at least about 0.2, ~~from~~ from which we ~~can~~ obtain

$$1 + 0.56 \left( \frac{S_p}{R_p} \right)^2 \geq 100$$

$$\text{or } \int r^{2\delta-1} V_0(r) d\vec{r} \ll \int r^{2\delta-1} V(r) d\vec{r}. \quad (20)$$

Otherwise,  $\rho$  will be of order  $10^{-2}$  <sup>as to satisfy</sup> ~~satisfying~~ the condition ~~(19)~~ (20). Since ~~the~~ the occurrence of such an extreme case seems very improbable, we ~~draw a conclusion~~ <sup>draw a conclusion</sup> that the calculated value of the pair production is probably too small to account for the experiment



Fig. 2

Fig. 1

~~The~~ The number of <sup>electrons</sup> the electron emitted from the K levels is known to be about 0.003 per disintegration in this case, ~~while~~ <sup>while</sup> on the other hand, according to Alichanow, total number of ~~the~~ positrons is also about 0.003 per disintegration.

Hence, if we assume the first hump in Fig. 1 to be due to the process above considered,  $\rho$  should be at least about 0.2, ~~from~~ from which we ~~can~~ <sup>obtain</sup>

$$1 + 0.56 \left( \frac{54}{R_p} \right)^2 \geq 100$$

or ~~or~~  $\int r^{28-1} V_0(r) d\vec{r} \ll \int r^{28-1} V(r) d\vec{r}$ . ~~(20)~~ (20)

Otherwise,  $\rho$  will be of order  $10^{-2}$  <sup>as to satisfy</sup> ~~satisfying~~ <sup>the condition</sup> ~~(20)~~ (20). Since ~~the~~ <sup>draw a conclusion</sup> the occurrence of such an extreme case ~~seems~~ <sup>seems</sup> very improbable, we ~~should conclude~~ <sup>conclude</sup> that the the calculated value of the pair production is probably too small to account for the experiment

exhaustively.

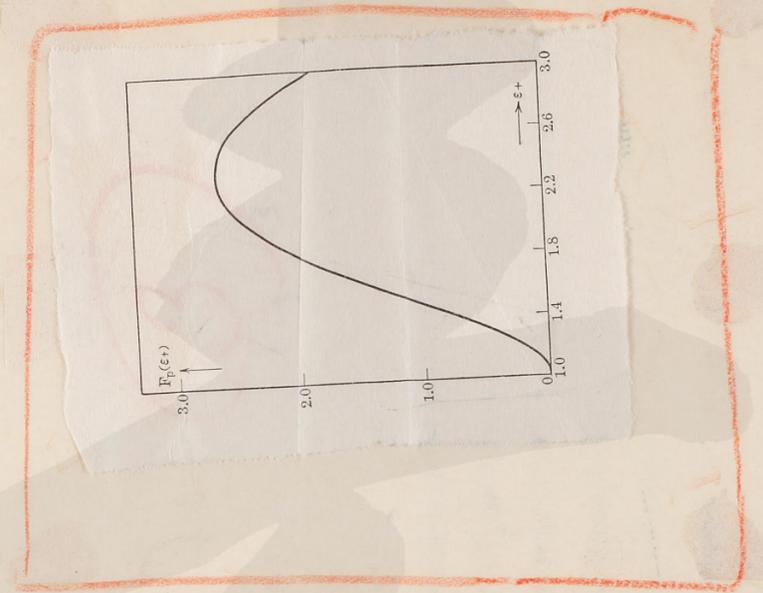
As shown by the formula (12),  $\rho$  depends, in general, on  $\Delta W$  and  $Z$ . It increases rapidly <sup>(with the increase of  $\Delta W$ )</sup> as  $\Delta W$  increases and the maximum of  $F_p(\xi_+)$  displaces towards  $\xi_+ = \frac{\Delta W}{2} = \frac{\Delta W}{2mc^2}$  at the same time.

For example, assuming  $Z=84$  and  $\Delta W=4mc^2$ , we have for  $F_p(\xi_+)$  a curve as shown in Fig. 3 and for  $\rho$  a value

$$4.0 \times 10^{-2} \left\{ 1 + 0.5 \left( \frac{\rho}{R_p} \right)^2 \right\}.$$

$\rho$  decreases rapidly as  $Z$  increases, owing to the factor  $(\alpha Z)^{2\delta+1}$  in the denominator.

In conclusion the authors wish to express their cordial thanks to Prof. G. Beck for valuable discussions on <sup>the occasion of</sup> his visit to Osaka.



Department of Physics,  
Osaka Imperial University.

(Received <sup>from</sup> Prof. G. Beck  
August 22, 1935)

exhaustively.

As shown by the formula (12),  $\rho$  depends on  $\Delta W$  and  $Z$ . It increases rapidly with the increase of  $\Delta W$  and displaces towards  $Z_+$  as  $\Delta W$  increases.

$Z_+ = \frac{\Delta W}{2} = \frac{\Delta W}{2mc}$  at the same time.

For example, assuming  $Z=84$  and  $\Delta W=4 \times 10^{-2}$  and for  $\rho$  a value

$$4.0 \times 10^{-2} \left\{ 1 + 0.56 \left( \frac{Z}{K_p} \right)^2 \right\}.$$

$\rho$  decreases rapidly as  $Z$  increases, owing to the factor  $(\alpha Z)^{2\delta+1}$  in the denominator.

In conclusion the authors wish to express their cordial thanks to Prof. G. Beck for valuable discussions on his visit to Osaka.

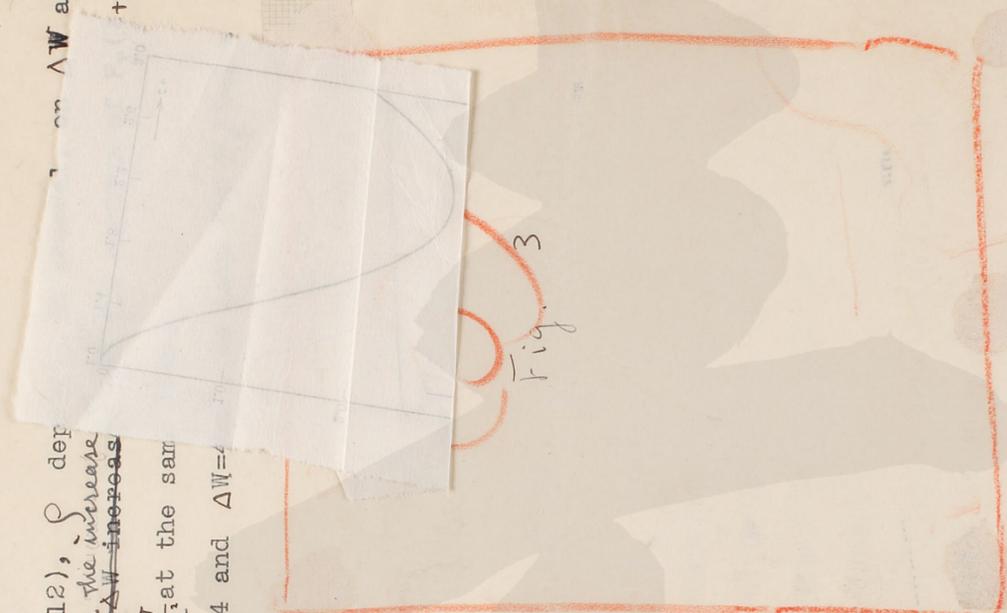


Fig. 3

Department of Physics,  
Osaka Imperial University.

(Received *fresh*  
August 22, 1935)

It will easily be seen that  
the magnitude of the  
Note added in Proof. — The ratio  $\rho$  depends much on  
the electrostatic field in the nuclear region. Especially,  
static potential acting on the electron  
if we assume rather large resolution ~~for~~ remaining  
~~that of the electron,~~  
potential, ~~to not~~ <sup>that</sup> near the nucleus  
exists in this region, the

BOND